

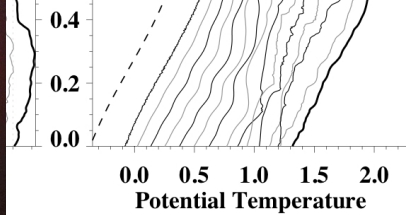
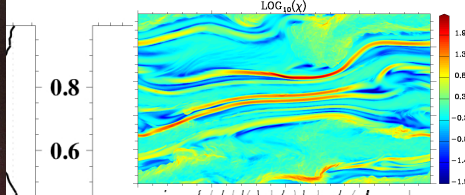
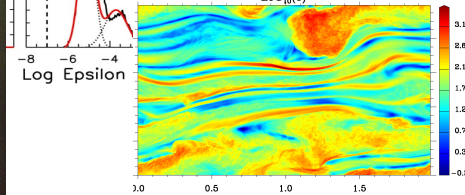
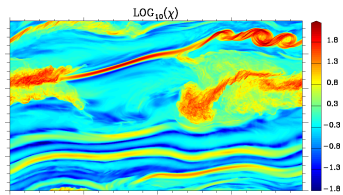
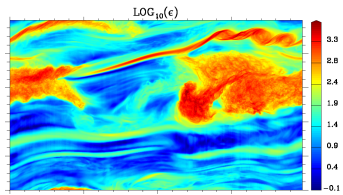
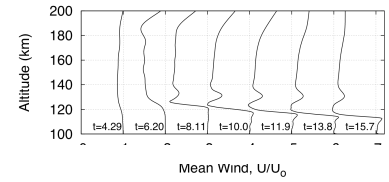
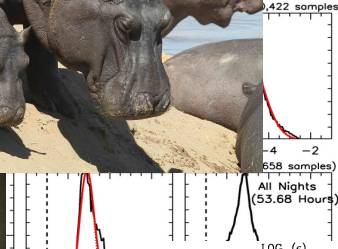
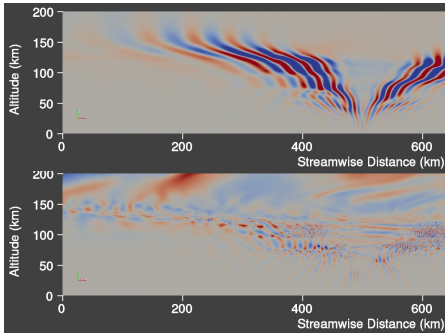
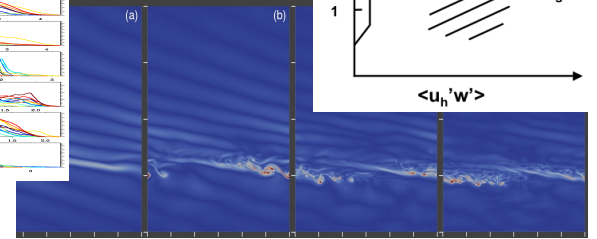
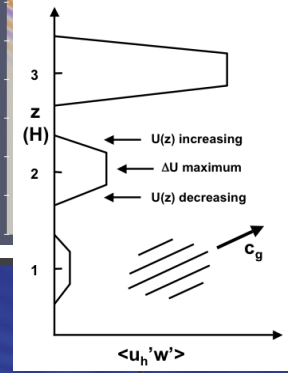
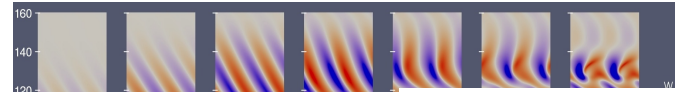
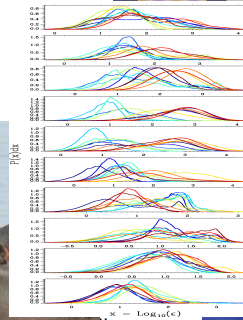
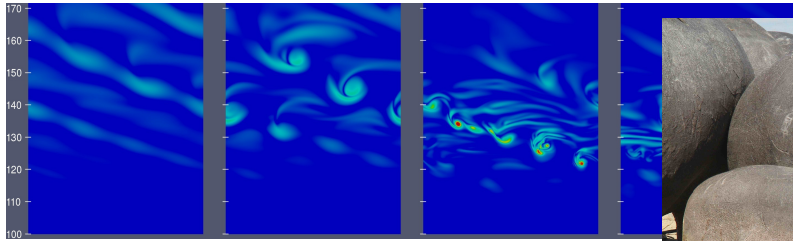


Nonlinear Gravity Wave Dynamics in the Thermosphere

Dave Fritts, Brian Laughman, and Tom Lund

Outline

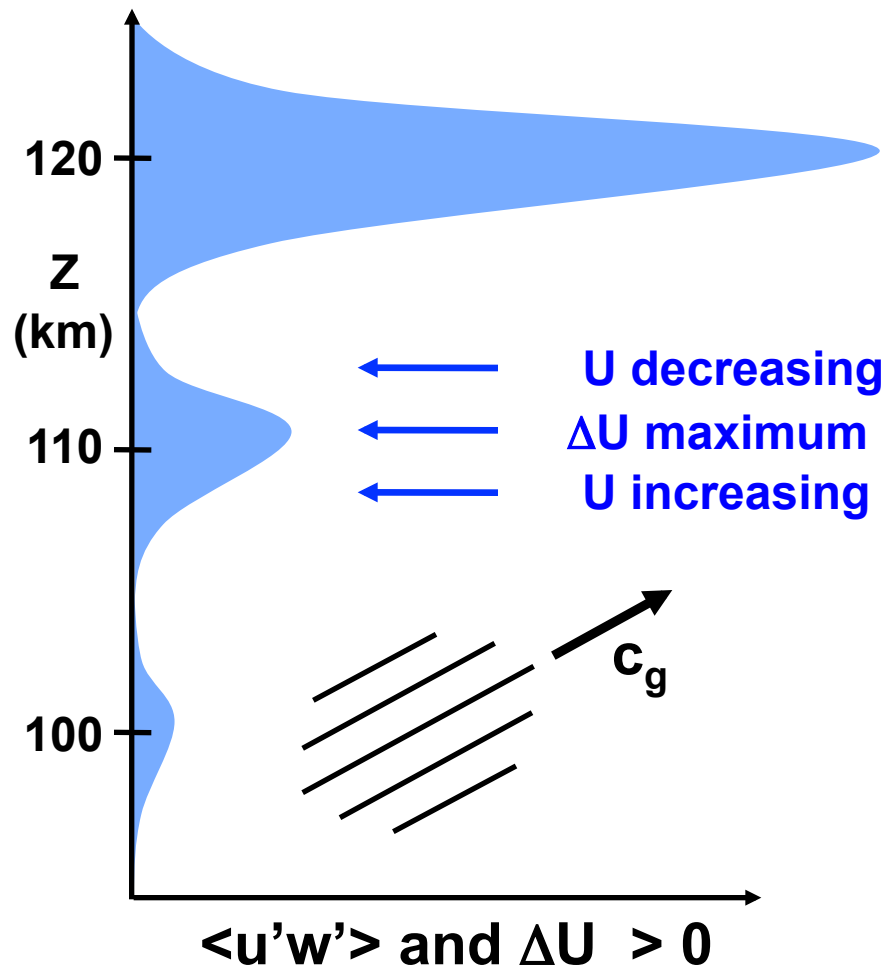
- Gravity wave momentum transport and self acceleration
- Gravity wave breaking in the thermosphere



GW Self Acceleration and Effects

Decreasing mean density

- =>
- GW amplitudes increase as $a \sim 1/\rho^{1/2}(z)$
 - $\Delta U \sim 1/\rho(z)$
 - large flow accelerations, altered GW phase structures



=> self-acceleration
& phase distortion

=> leading edge of GW packet
accelerates

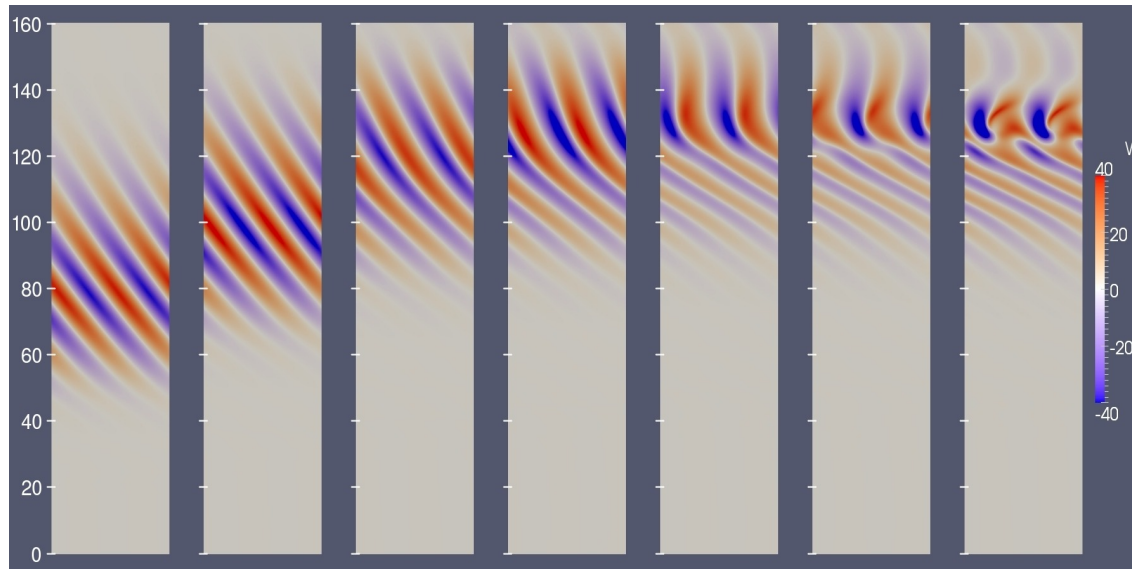
trailing edge decelerates

“Self-acceleration” of a localized GW packet

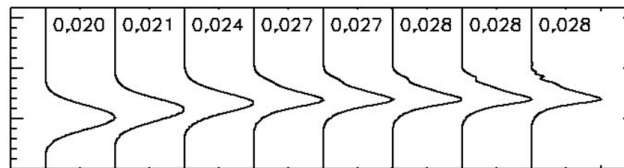
- =>
- steepening phase structures at leading edge
 - altered GW group velocities and GW instability

$\Delta t \sim 3$ GW periods

GW with initial
 $\lambda_x = \lambda_z = 20$ km,
 $N = 0.02$ s⁻¹,
 $\omega = N/1.4$,
 $c \sim 60$ m/s

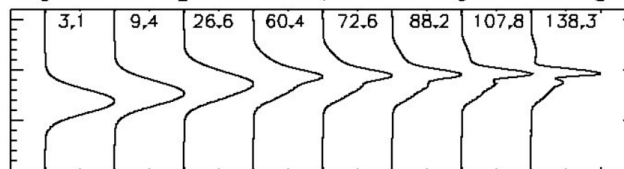


GW $\rho\langle u'w' \rangle$



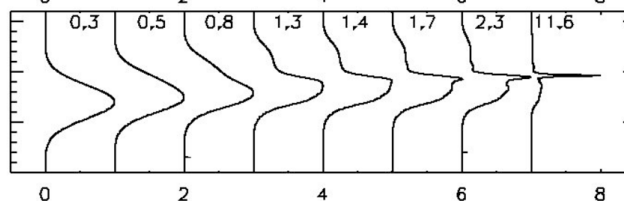
$\rho\langle u'w' \rangle$ nearly conserved

induced $U(z)$



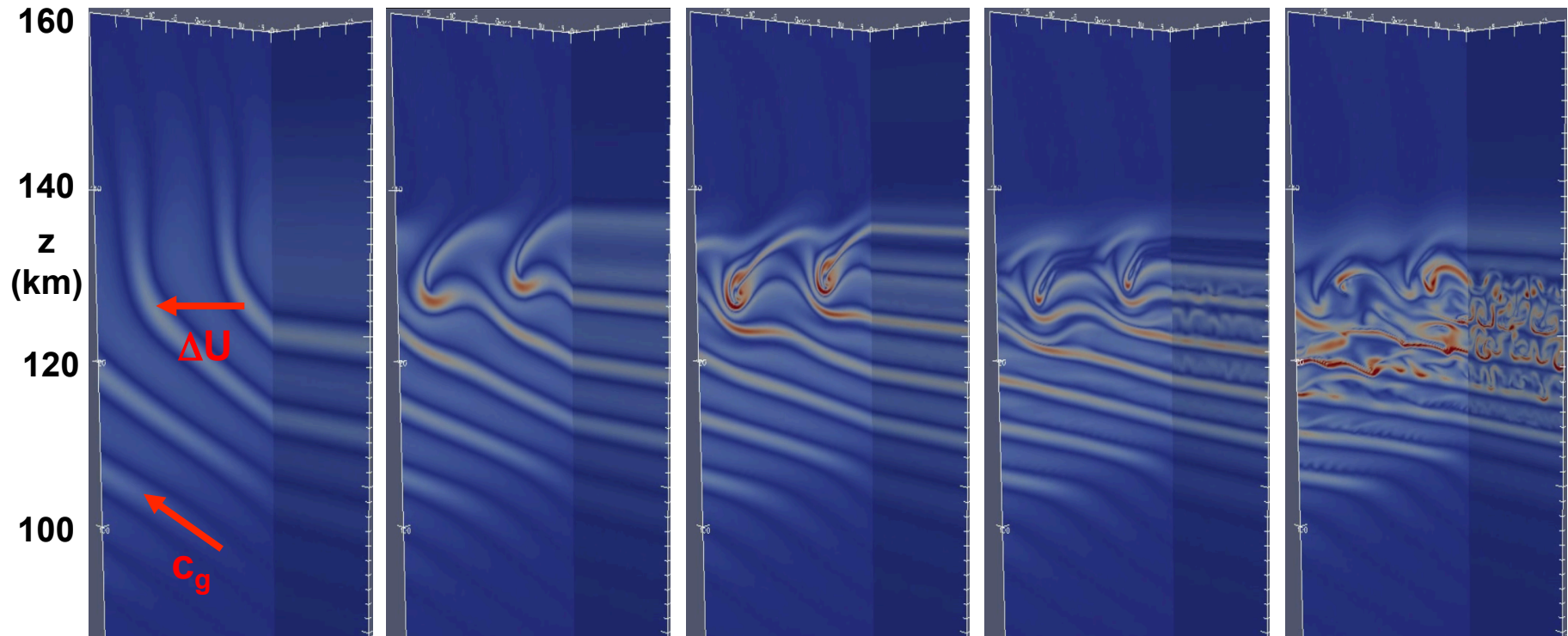
$U(z)$ increases beyond c

GW amplitude
 $(a = u'/(c-U))$



GW amplitude exceeds unstable value

Self Acceleration causes 2D and 3D instability



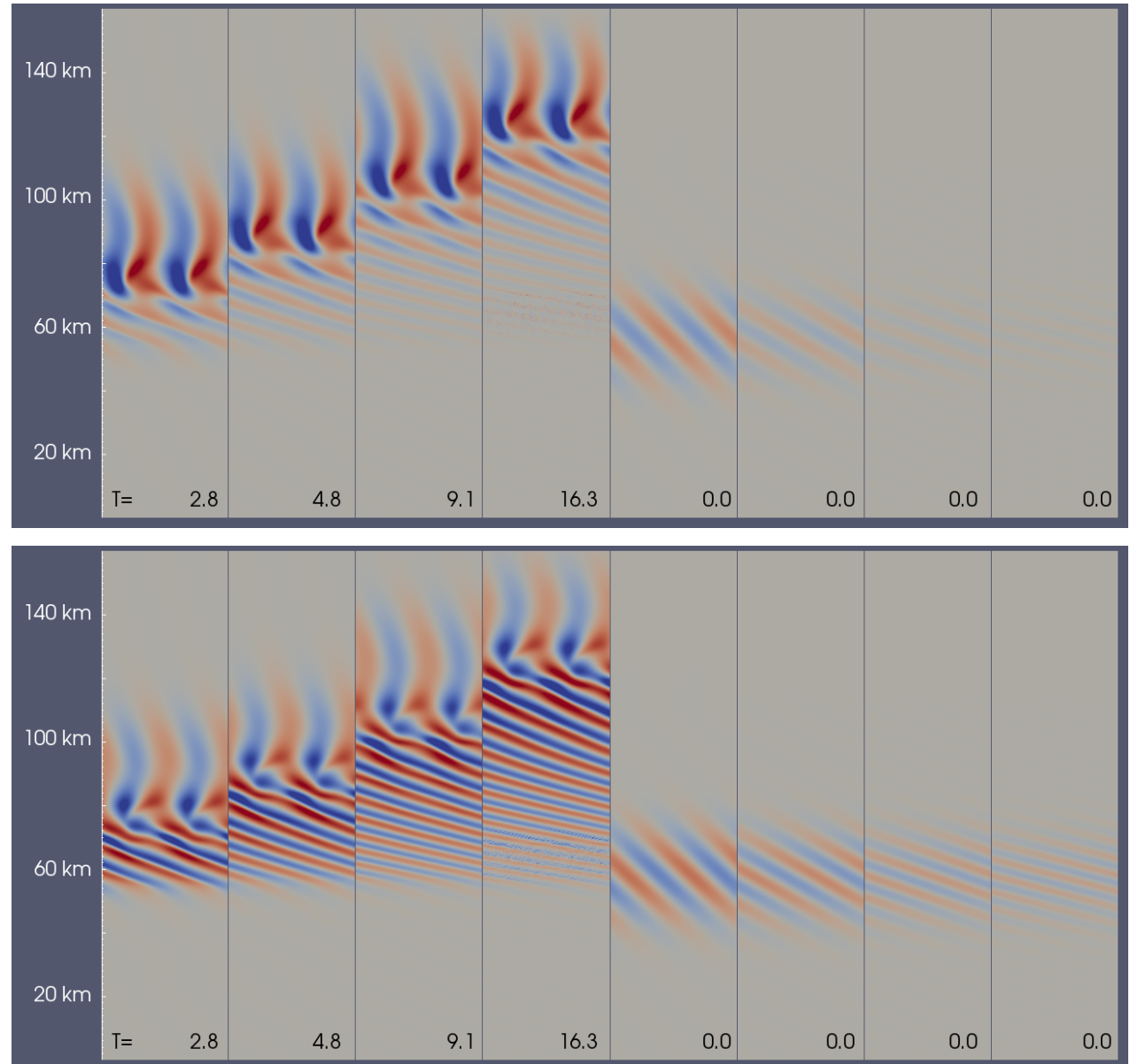
- 2D “self-acceleration” instability occurs first, drives 2D “wave-wave” interactions
- Secondary 3D instability accounts for most dissipation

Self Acceleration causes GW stalling and instability

GW w' (top),
 u' (bottom)

successive 4-panel
sets (right and left)
show forcing and
responses for

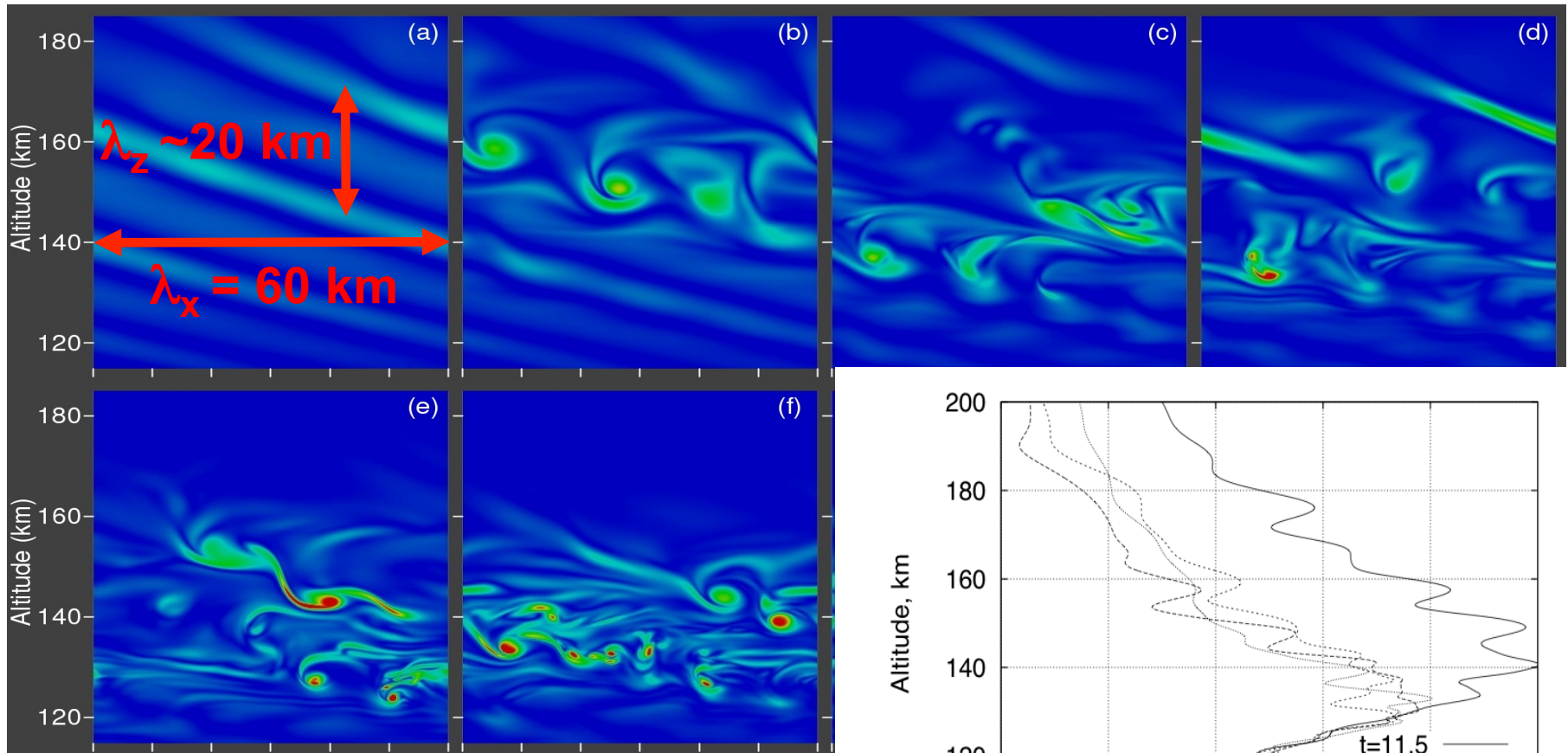
$\omega = N/1.4, N/2,$
 $N/3,$ and $N/5$
for the same u' .



GW breaking @ $z \sim 100\text{-}170$ km for constant $U(z)$

$$\lambda_x = 60 \text{ km}, \lambda_z \sim 20 \text{ km}, u' \sim c-U \sim 50 \text{ m/s}, \omega \sim N/3, \\ \text{Re} \sim 10^3 \text{ at } z \sim 160 \text{ km}$$

minimum scales of turbulence ~ 1 km



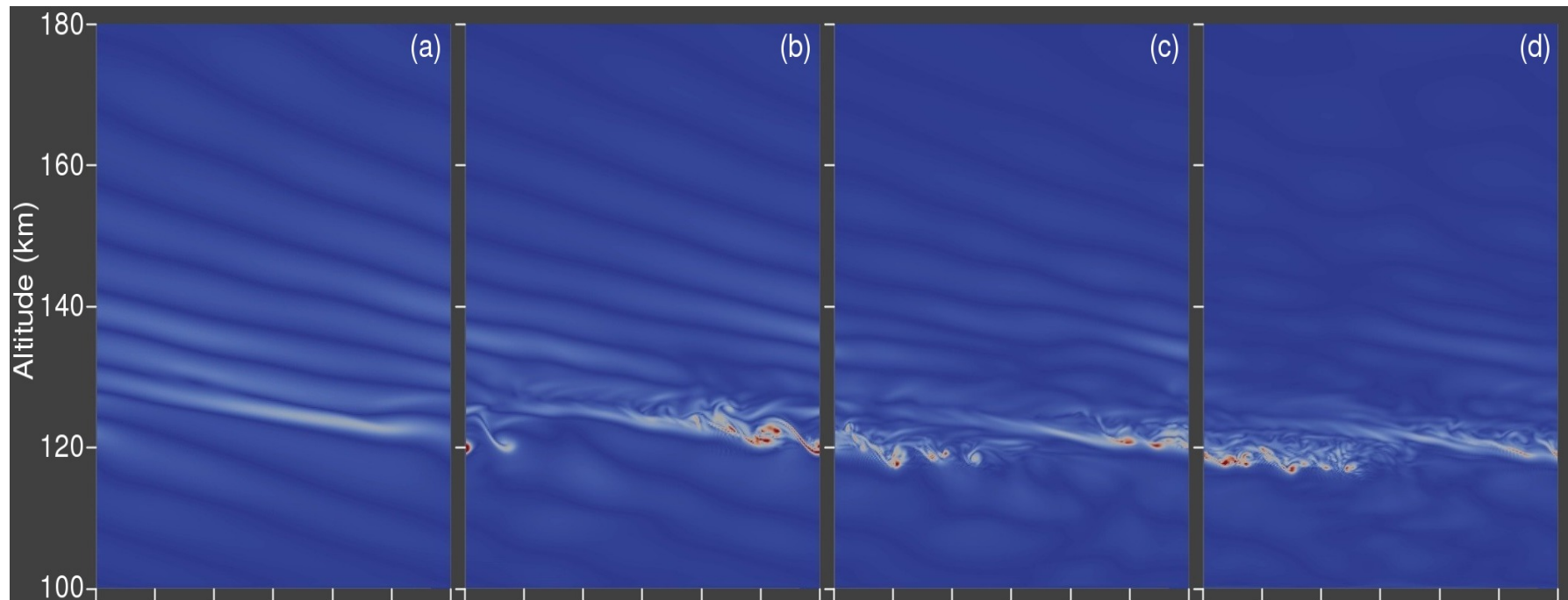
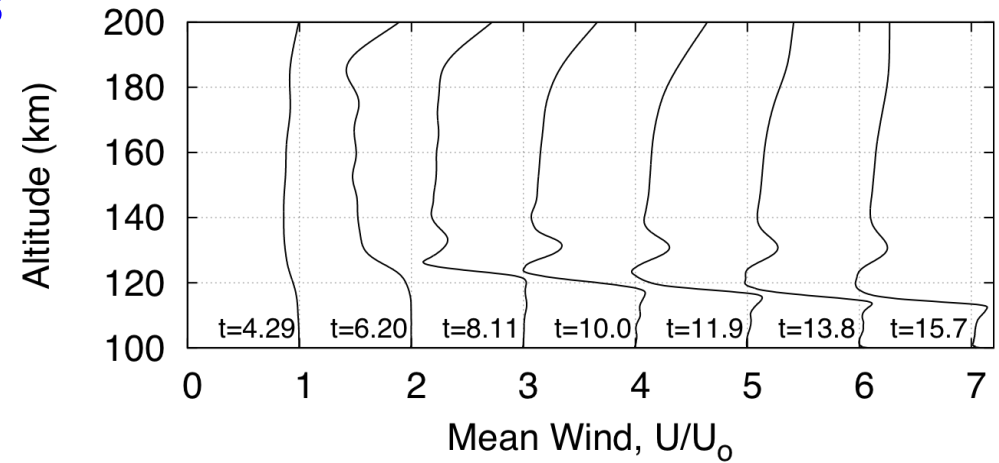
GW “saturates” at a ~ 0.7 . This would be a better value for parameterization purposes.

**Same GW evolution
allowing mean flow changes**

**this case exhibits large
induced mean winds**

**=> confinement of shears and
instabilities to lower altitudes**

**evolving $U(z,t) \Rightarrow$ accelerations
towards GW phase speed**

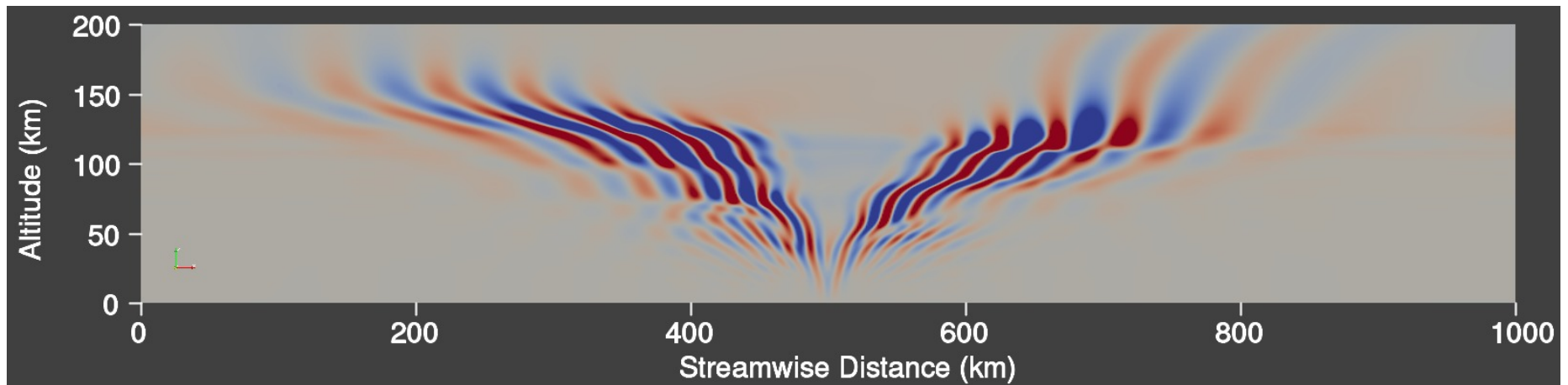


GWs are also strongly influenced by mean and tidal winds

- exhibit refraction, filtering, instabilities, and body forcing in more general environments

- example of GWs from a single convective plume propagating in TIME GCM tidal winds

t = 45 min



t = 90 min

