Modeling Energetic Particle Precipitation and Transport with WACCM

Laura Holt, Cora Randall, Anne Smith

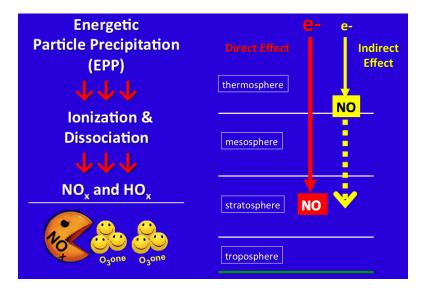
University of Colorado at Boulder Laboratory for Atmospheric and Space Physics National Center for Atmospheric Research

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Introduction

Results Conclusions Effects of energetic particle precipitation Science Questions

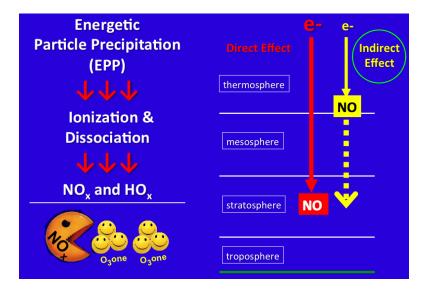


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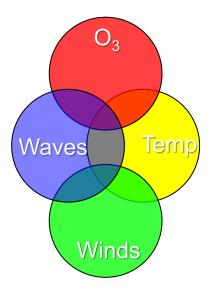


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Changes in polar ozone trigger a redistribution of solar and magnetospheric energy at Earth

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Effects of energetic particle precipitation Science Questions

Research Goal

Determine the coupling mechanisms that redistribute solar and magnetospheric energy in the atmosphere.

Science Questions

- How does transport affect the distribution of NO_x created by EPP?
- How are the different regions of the atmosphere coupled through EPP?

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Need to use a model \longrightarrow WACCM

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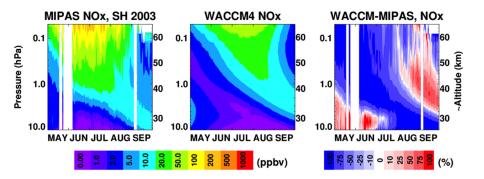
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WACCM compared to observations Effects of tuning taubgnd

How does WACCM compare to observations?



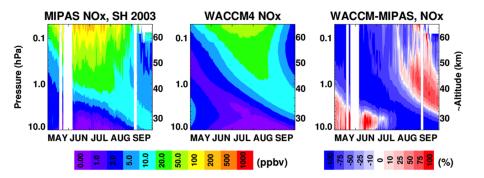
Too low at beginning of season and lasts too long — (1) residual circulation may be too low in WACCM (2) WACCM only has auroral and SPEs

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- Taubgnd controls amplitude of the GW source spectrum
- The amplitude of GWs determines where they break (lower amplitude=higher breaking)
- The GW forcing influences the residual circulation

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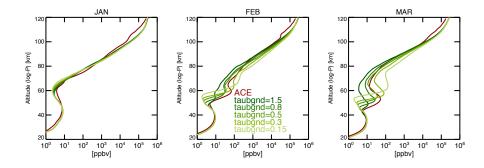
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WACCM compared to observations Effects of tuning taubgnd

WACCM NO_x sampled at ACE locations for the NH 2006 compared to ACE



Decreasing taubgnd increases altitude of peak NO_x enhancement, but moves peak too high compared to ACE

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Summary What do we need? End

Summary

- WACCM underestimates NO_x compared to observations
- Moving the residual circulation by tuning taubgnd doesn't appear to be enough to solve this problem
- A combination of tuning taubgnd and source strength will most likely produce better agreement between WACCM and observations

Other Results (Not Shown Here)

 However, WACCM shows evidence for coupling due to EPP even though the EPP Indirect Effect is underestimated in WACCM

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Introduction Summary Results What do we need? Conclusions End

What do we need?

Measurements

- NO_x throughout the polar winter, stratosphere to thermosphere
- NO_y and CI_y
- Polar Mesospheric Clouds
- Winds in MLT
- Global ozone and temperature profiles
- Precipitating particles: energy spectrum and spatial distribution

Models

- Improved treatment of gravity waves in CCMs
- Incorporation of higher energy electrons in CCMs

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Introduction Summary Results What do we need? Conclusions End

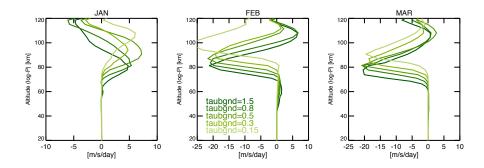
Thank You

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Summary What do we need? End

WACCM gravity wave drag poleward of 70 N



Decreasing taubgnd moves GW momentum spectrum up

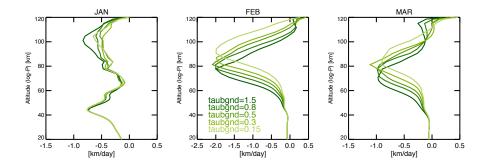
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Summary What do we need? End

WACCM vertical residual circulation poleward of 70 N



Decreasing taubgnd moves peak descent up

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Introduction	Summary
Results	What do we need?
Conclusions	End

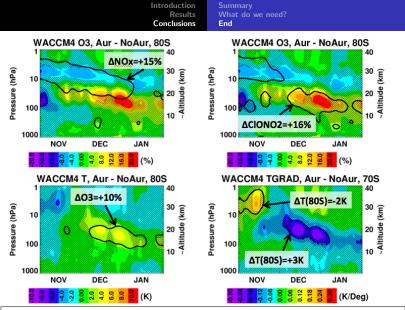
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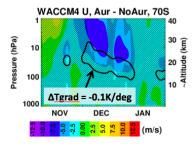
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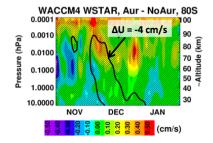
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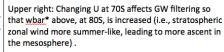
Upper left: Increased EPP-NOX causes O3 depletion and increased CIONO2. Upper right: Increased CIONO2 mitigates O3 loss. Lower left: Increased O3 causes temperature to increase. Lower right: Change in polar T leads to change in latitudinal T gradient at 70S.

Introduction Summary Results What do we need? Conclusions End





Upper left: Changing latitudinal gradient in T causes U at 70S above to become more negative (easterly/summer) .



Lower left: Increase in <u>wbar</u>* leads to mesospheric cooling at 80S.

