Coordinated Investigation of Antarctic Total Solar Eclipse (TSE) using SuperDARN HF Radars

<u>S. Chakraborty</u>¹, X. Shi^{1,2}, G. Chisham³, J. M. Ruohoniemi¹, J. B. H. Baker¹, K. Stern¹, and E. G. Thomas⁴

¹Department of Electrical and Computer Engineering, Virginia Tech, Blacksburg, USA ²High Altitude Observatory, NCAR, Boulder, CO, USA ³British Antarctic Survey, Cambridge, UK ⁴Thayer School of Engineering, Dartmouth College, Hanover, NH, USA



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Outline

1. Introduction

- Super Dual Auroral Radar Network (SuperDARN)
- **Geometry of the 4th December TSE and coordination of SuperDARN radar.**
- Science goals of the study.

2. Experiments

- Normal Sounding Mode.
- Special Sounding Mode in 2nd Channel.

3. Observations and Current Findings

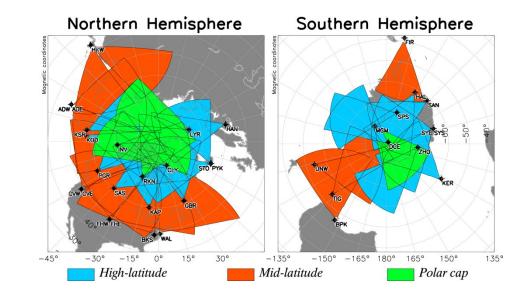
Instruments: SuperDARN HF Radar

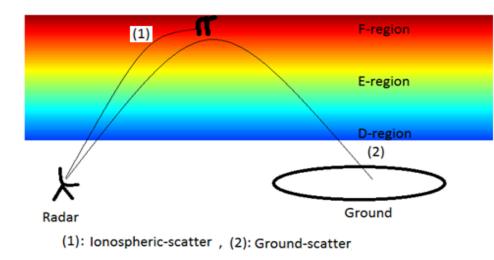
-3300

-50

150

- <u>SuperDARN</u>: An international HF radar network that probes the Earth's upper atmosphere and ionosphere.
- \circ $\,$ There are two primary backscatter targets:
 - 1. <u>Ionospheric scatter</u>: Backscatter from ionospheric irregularity structures.
 - 2. <u>Ground Scatter:</u> Backscatter from the Earth's surface.
- Ground scatter replicates a ground-to-ground communication link just like a 1-hop radio link does.
- We acknowledge the use of SuperDARN data. SuperDARN is a collection of radars funded by various national scientific funding agencies across the globe.



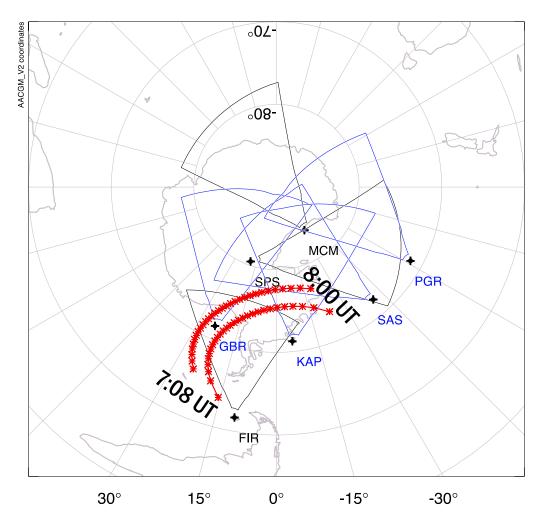


Radar Scan

:20:00 UT

-105°

Introduction: TSE Geometry and Science Goals

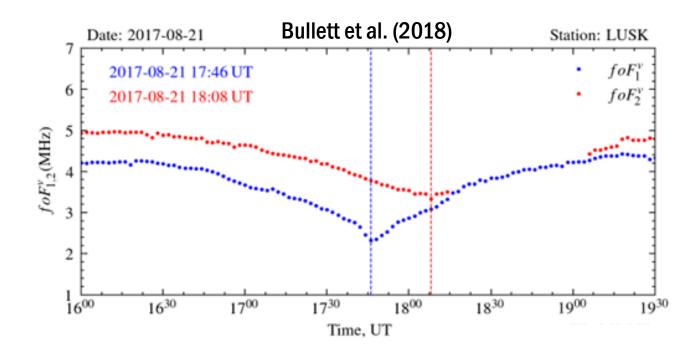


- 1. We are interested in understanding the HF wave propagation conditions following TSE. Specifically, ionospheric G-condition.
- 2. We are also interested to investigate ionospheric disturbances in conjugate hemispheres during solar eclipse.

View from the south pole: black FOVs indicate radars in the southern hemisphere; blue FOVs are radars in the northern hemisphere projected along magnetic field lines into the southern hemisphere.

1. Introduction

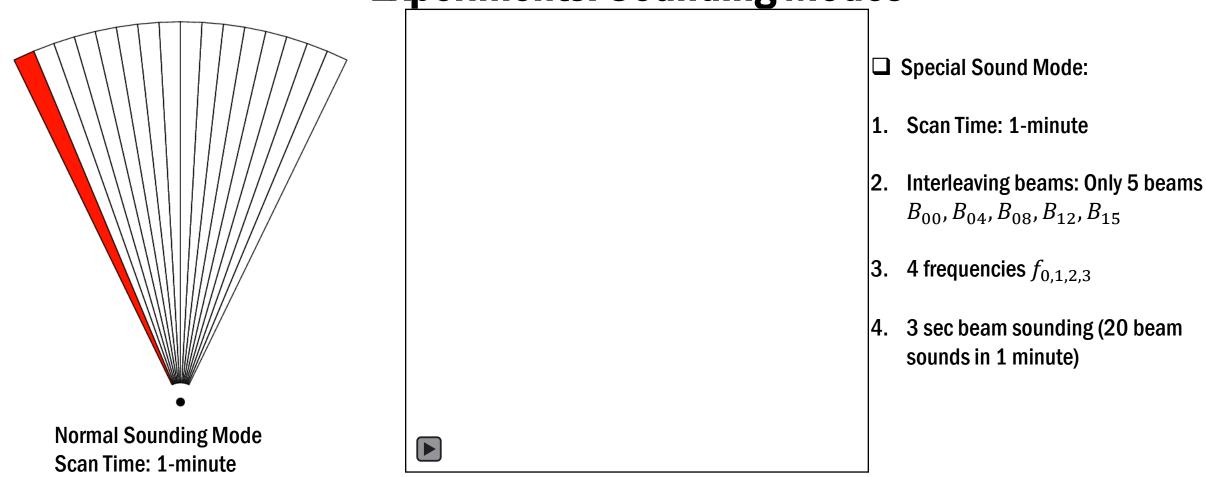
Introduction: Ionospheric G-Condition during 2017 (21st Aug) Great American Eclipse (GAE)



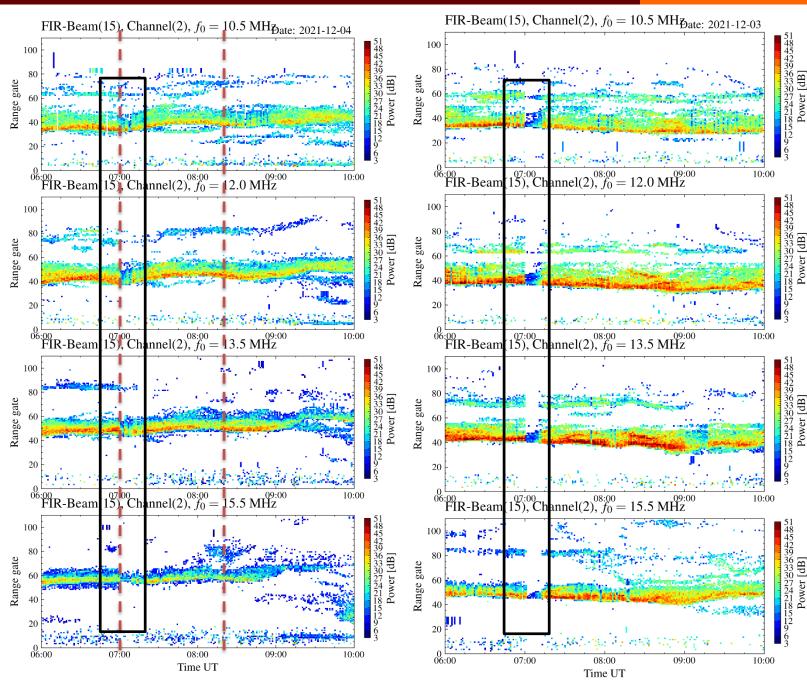
Ionospheric G-condition when $f \circ F_1 \ge f \circ F_2$

ISR Millstone Hill Data 16.30 UT 18.15 UT 400 400 hmf2=227 km hmf2=202 km 350 350 300 300 Height, km 250 250 200 200 150 150 $100 \\ 10^{10}$ 1011 1011 10^{12} 1012 19.15 UT 20.00 UT 400 400 hmf2=205 km hmf2=206 km 350 350 300 300 Height, km 250 250 200 200 150 150 100 L 10¹⁰ 100 L 10^{12} 1011 1011 10^{12} $[e^{-}], m^{-3}$ $[e^{-}], m^{-3}$

Experiments: Sounding Modes



- 1. Normal mode data should be sufficient for the current research interest on the interhemispheric electrodynamics.
- 2. To check propagation condition, with help of Dr. Chisham we requested a new mode to run on the channel 2 of the FIR radar. This sounding mode enables us to capture propagation mode in the ionosphere.

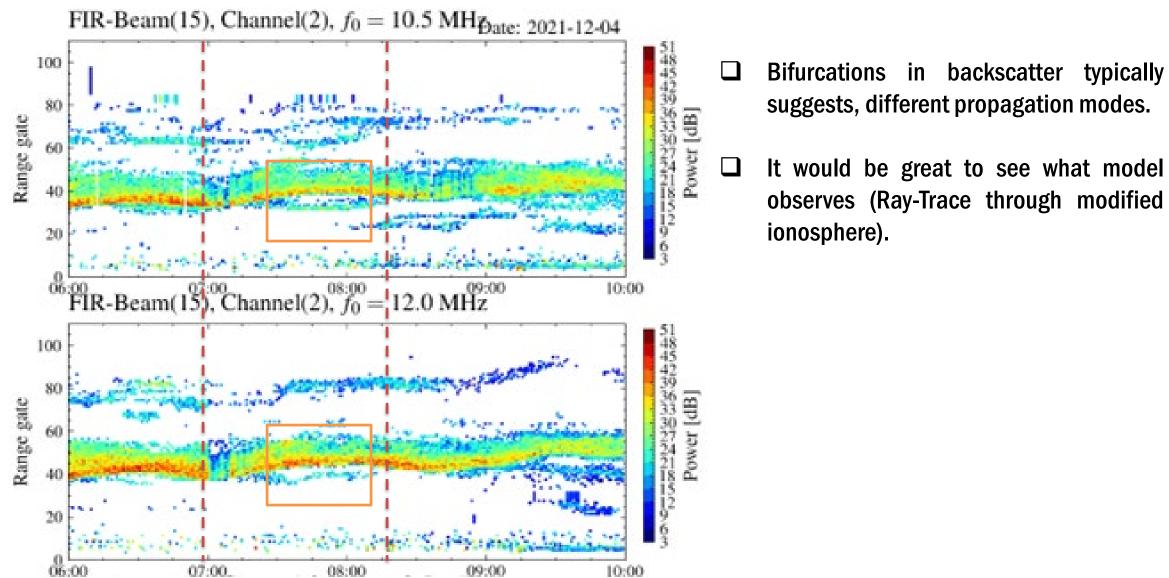


3. Observations

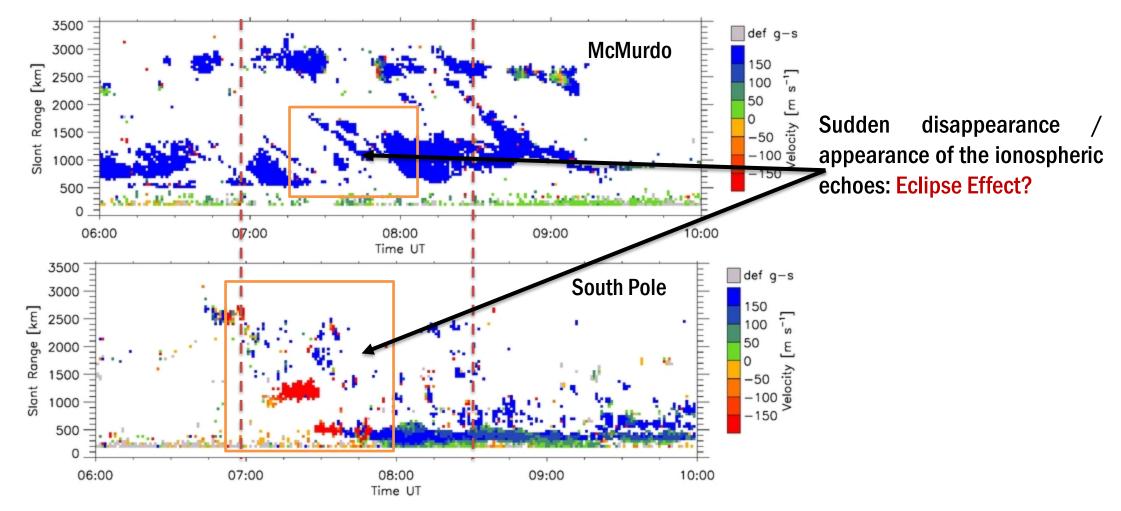
Observations: Antarctic Radars

- FIR radar observations, from beam 15 for 4 frequencies (10.5, 12, 13.5, 15.5) MHz.
- Left: Eclipse day, 4 December 2021; <u>Right:</u> Control day, 3 December 2021
- **Eclipse duration** \sim 7:00 8:30 UT.
- ❑ We are still working on the effects observed by the radar.
- Eclipse effect before the totality OR interference?

Observations: Bifurcation in Backscatter



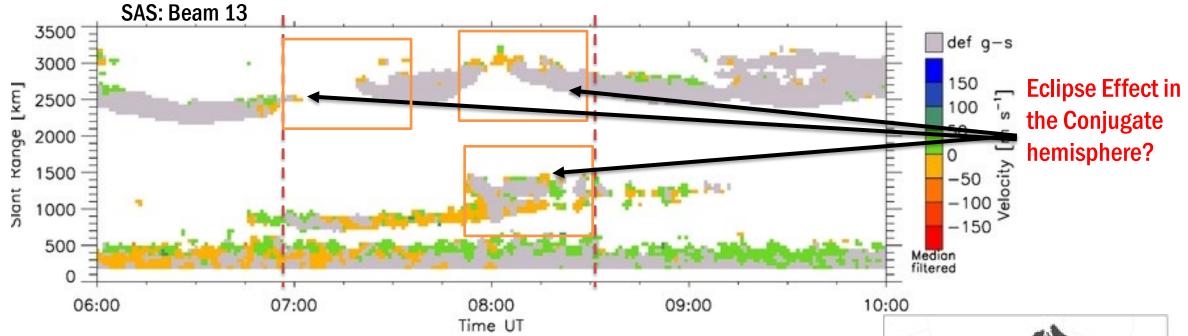
Observations: Antarctic Radars



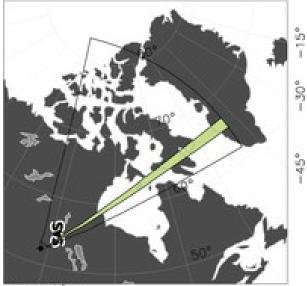
Observations from two other stations from the southern hemisphere.

□ We are not sure about any significant Eclipse effect here.

Observations: Northern Conjugate Radars



- Drop in back scatter counts at far ranges and increase in near ranges. Do we also see two different propagation modes?
 - Transport via magnetic field lines might cause this effect, this needs validation using other magnetospheric satellite instruments and modeling.



Questions?

shibaji7@vt.edu

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For more on G-Condition please visit poster: SOLA-9: Origination of Ionospheric G-condition following a Total Solar Eclipse Tuesday's Poster