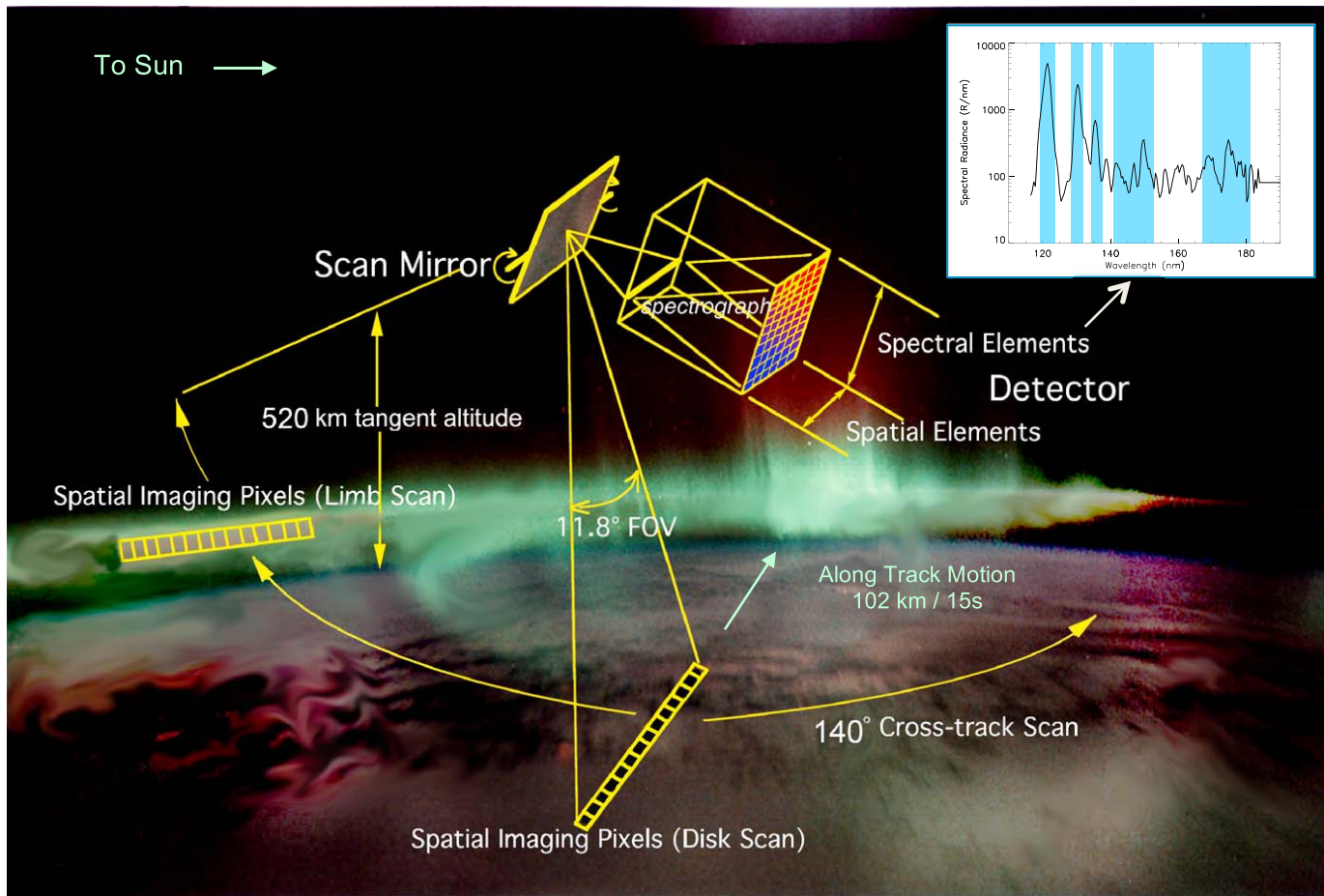




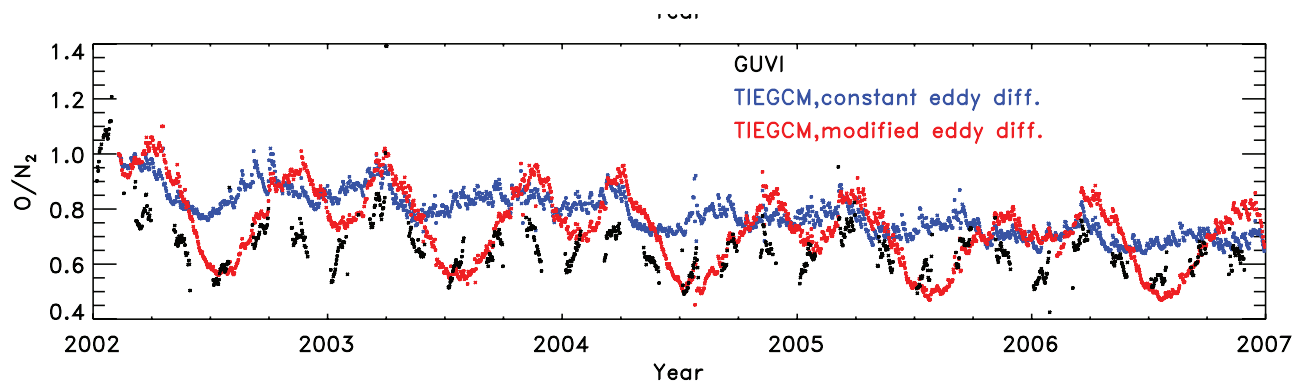
# ***Variations of Thermospheric Composition in TIMED/GUVI Limb Measurements***

Jia Yue, Yongxiao Jian, Wenbin Wang, Bob Meier, Alan Burns, Liying Qian, Mack Jones, Dong Wu, Marty Mlynczak



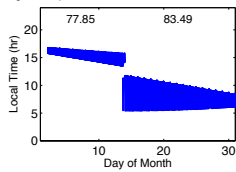
GUVI nadir and limb view schematics. The 14 spatial imaging pixels across the entrance slit each generate a spectrum on the detector focal plane. The scan mirror sweeps the slit across the limb and disc. There are 32 slit positions in a limb scan between approximately 510 and 110 km, and 159 samples on the disk, for a total of 191 angular slit positions in a complete cross-track scan [Meier et al., 2015].

# GUVI DISC O/N<sub>2</sub> (Qian et al., 2009)

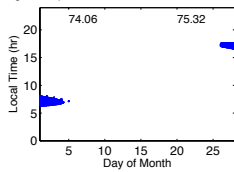


# GUVI limb data (2002-2007) is daytime only

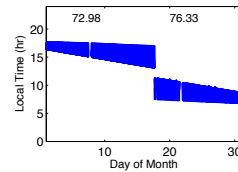
guvi footprint local time distribution in Jan., 2006



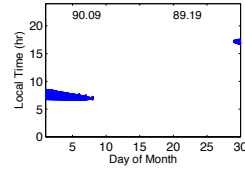
guvi footprint local time distribution in Feb., 2006



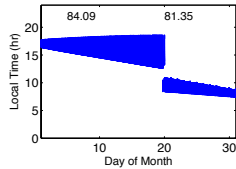
guvi footprint local time distribution in Mar., 2006



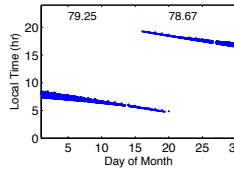
guvi footprint local time distribution in Apr., 2006



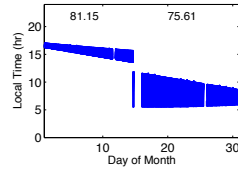
guvi footprint local time distribution in May., 2006



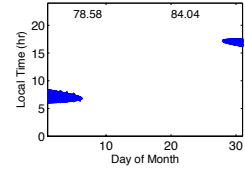
guvi footprint local time distribution in Jun., 2006



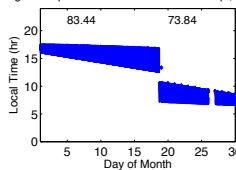
guvi footprint local time distribution in Jul., 2006



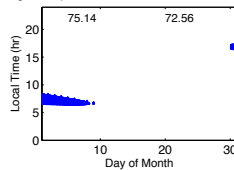
guvi footprint local time distribution in Aug., 2006



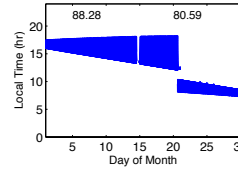
guvi footprint local time distribution in Sep., 2006



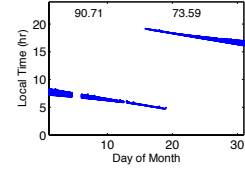
guvi footprint local time distribution in Oct., 2006



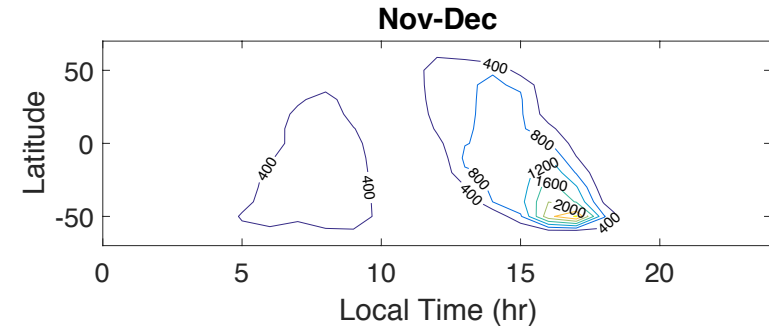
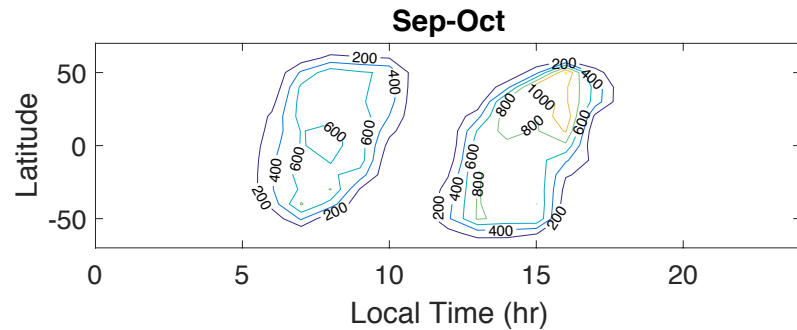
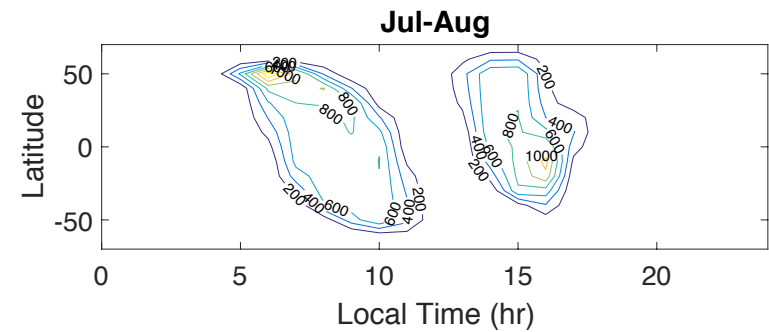
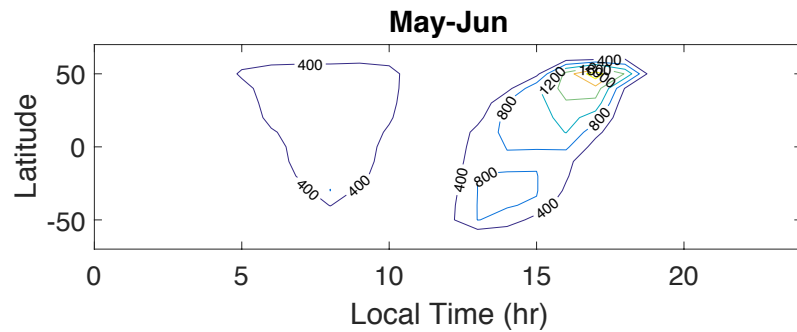
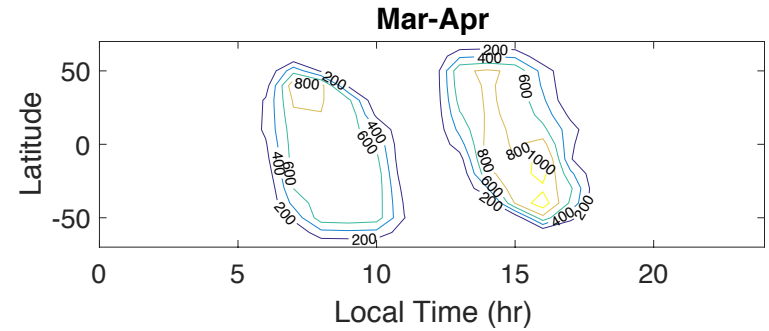
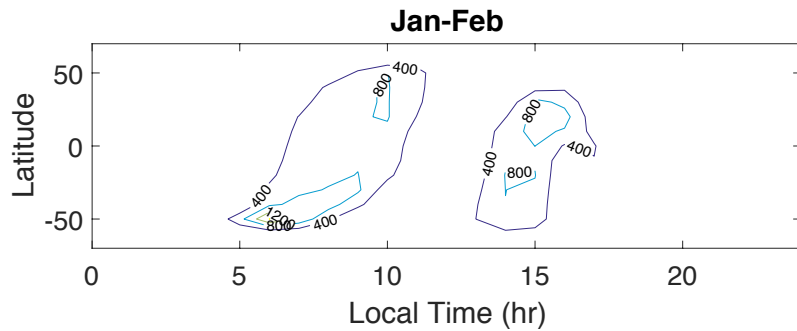
guvi footprint local time distribution in Nov., 2006



guvi footprint local time distribution in Dec., 2006



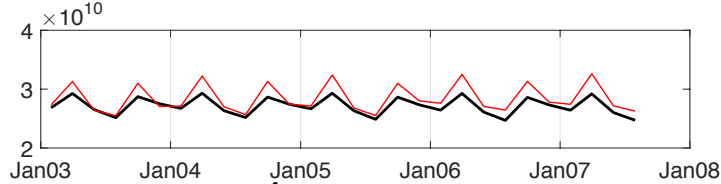
# Data distribution vary from month to month



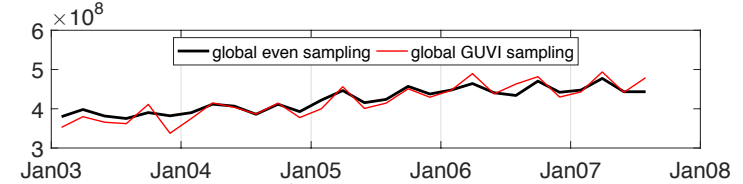
# *Impact of incomplete GUVI local time/latitude tested with MSIS global mean outputs*

## *The sampling does not change the main patterns*

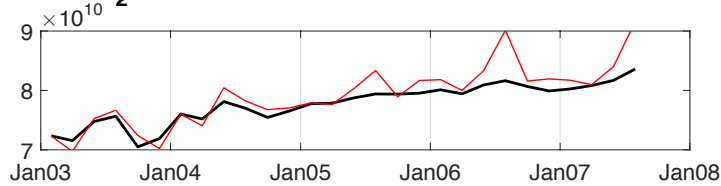
**MSIS Atomic O at  $8.4 \times 10^{-4}$  Pa (average pressure at 133km)**



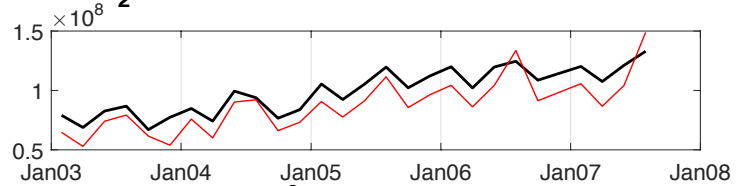
**MSIS Atomic O at  $6.35 \times 10^{-6}$  Pa (average pressure at 312km)**



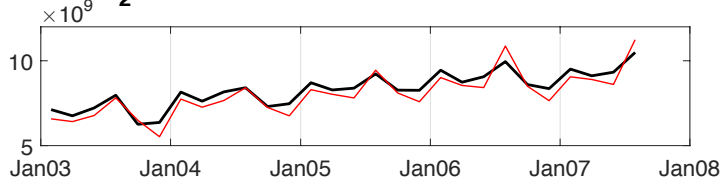
**MSIS N<sub>2</sub> at  $8.4 \times 10^{-4}$  Pa (average pressure at 133km)**



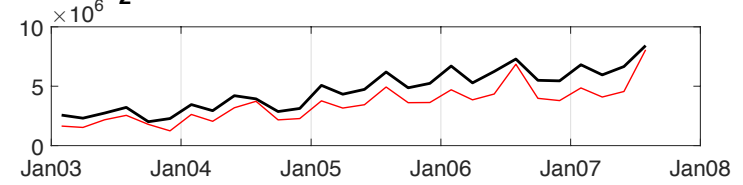
**MSIS N<sub>2</sub> at  $6.35 \times 10^{-6}$  Pa (average pressure at 312km)**



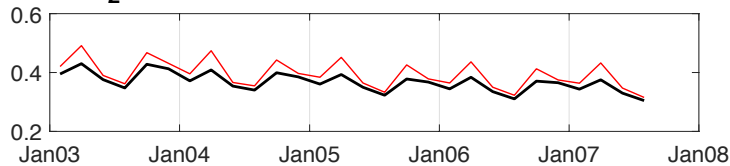
**MSIS O<sub>2</sub> at  $8.4 \times 10^{-4}$  Pa (average pressure at 133km)**



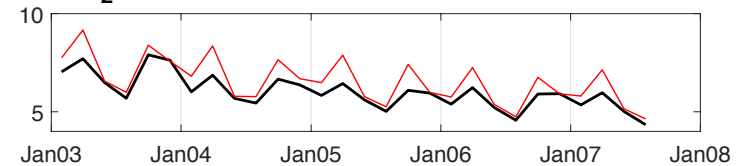
**MSIS O<sub>2</sub> at  $6.35 \times 10^{-6}$  Pa (average pressure at 312km)**



**MSIS O/N<sub>2</sub> ratio at  $8.4 \times 10^{-4}$  Pa (average pressure at 133km)**



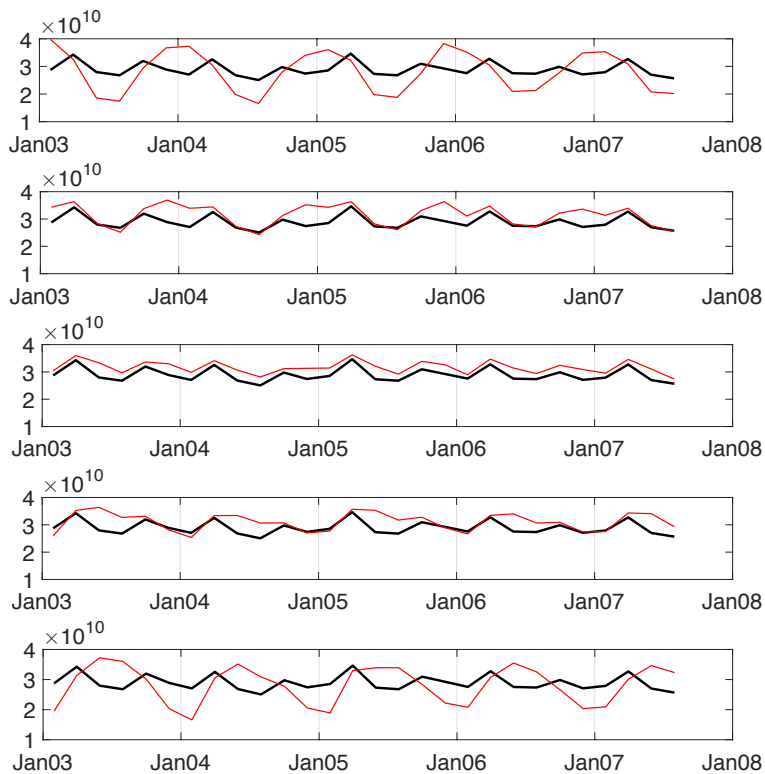
**MSIS O/N<sub>2</sub> ratio at  $6.35 \times 10^{-6}$  Pa (average pressure at 312km)**



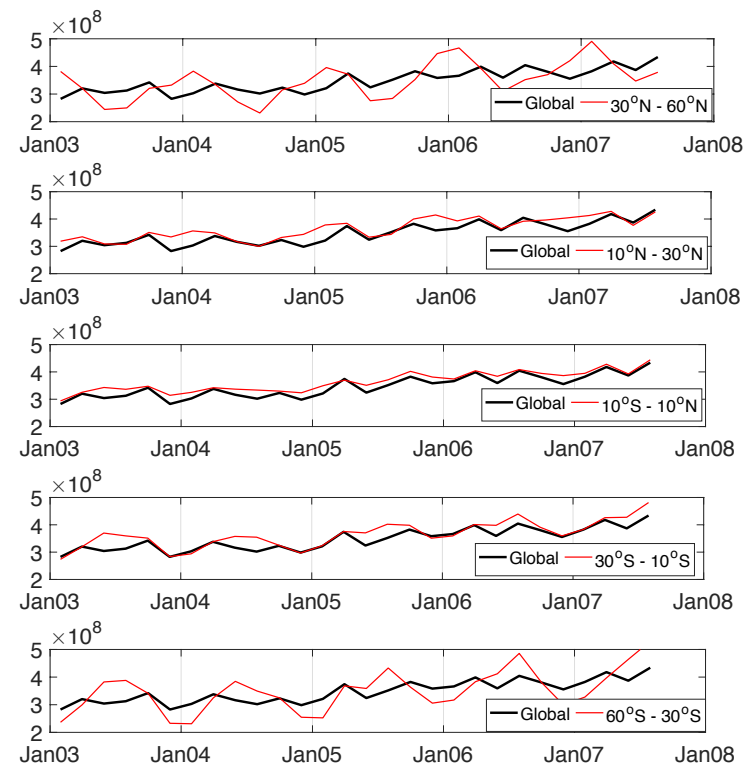
**Black: even sampling; red: GUVI limb sampling**

# GUVI O in the lower thermosphere and upper thermosphere

GUVI Atomic O at  $8.4 \times 10^{-4}$  Pa (average pressure at 133 km)



GUVI Atomic O at  $6.35 \times 10^{-6}$  Pa (average pressure at 312 km)

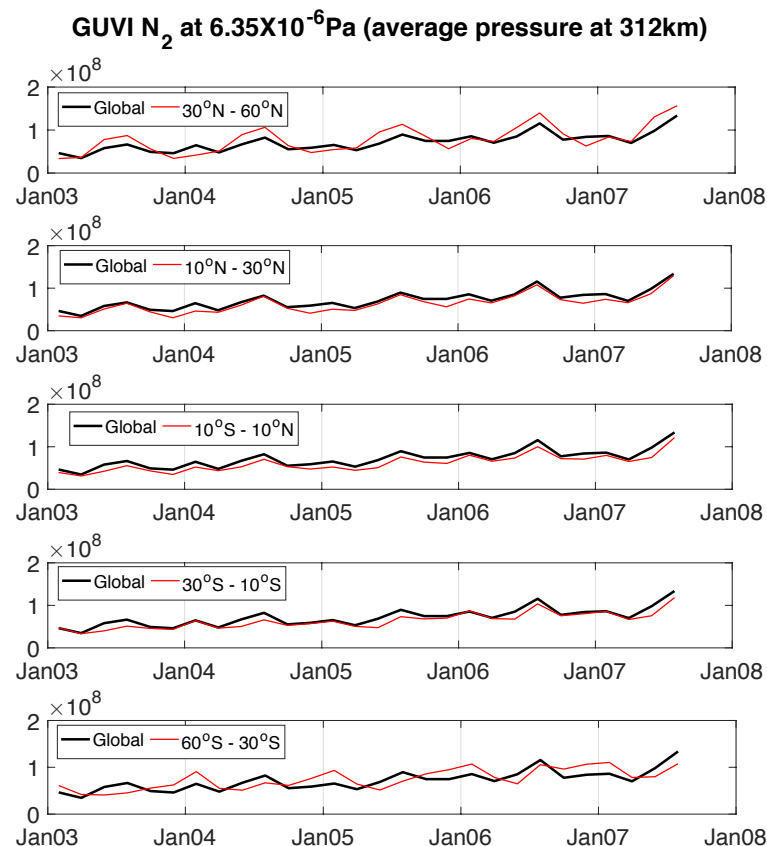
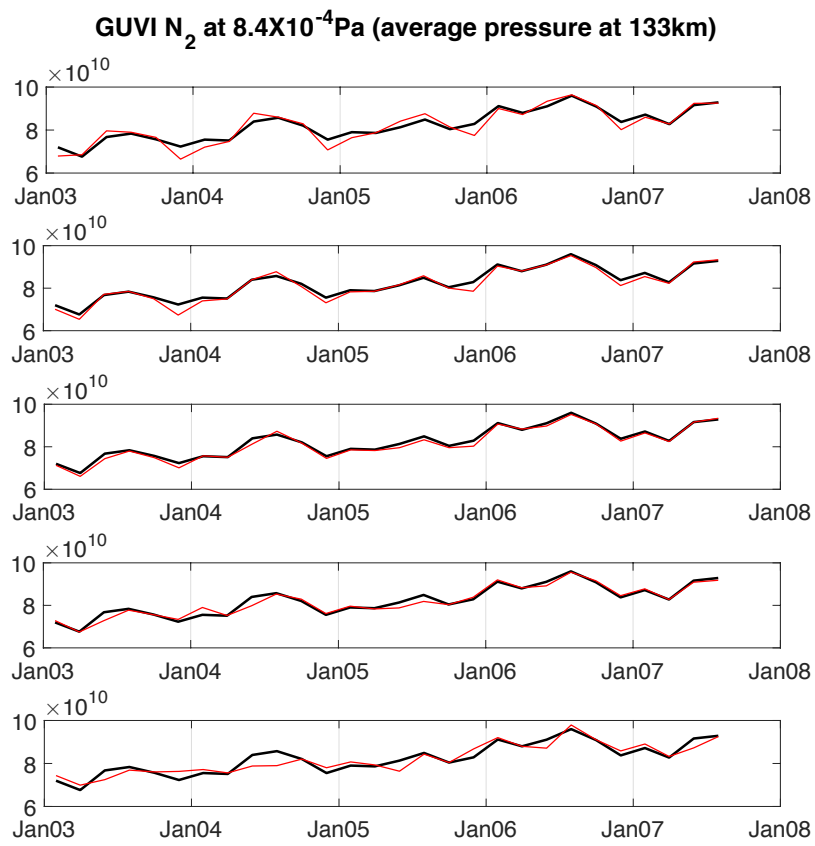


- At midlatitudes, the amplitude of annual variation decreases with solar cycle
- Semiannual variation is more important at low latitudes.
- No apparent O trend in the lower thermosphere
- Upper thermosphere O density increases with declining solar activity, following ideal gas law



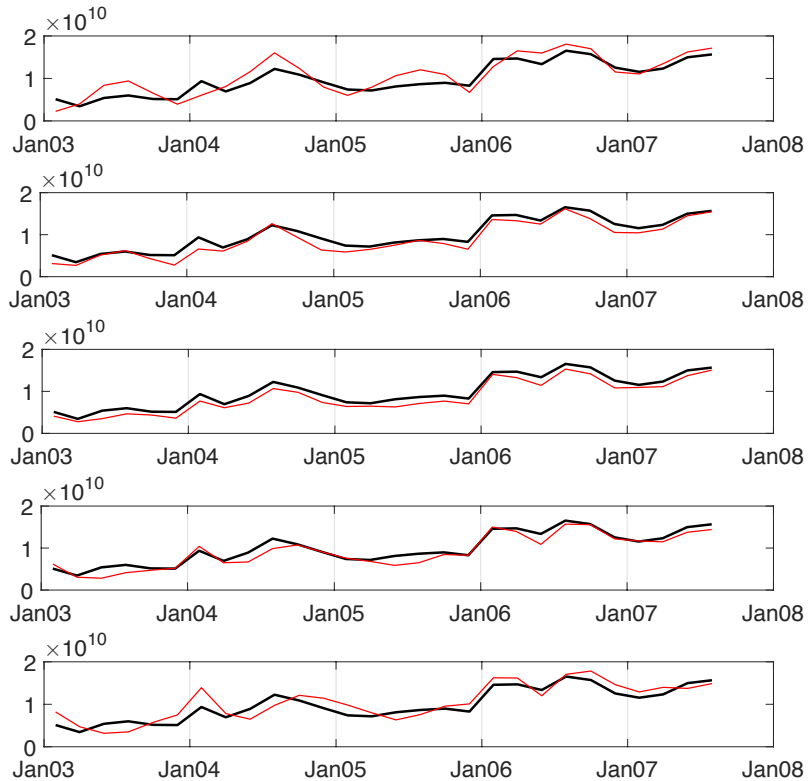
# GUVI N<sub>2</sub> in the lower thermosphere and upper thermosphere

Both lower and upper thermosphere N<sub>2</sub> density shows a negative correlation with solar cycle.

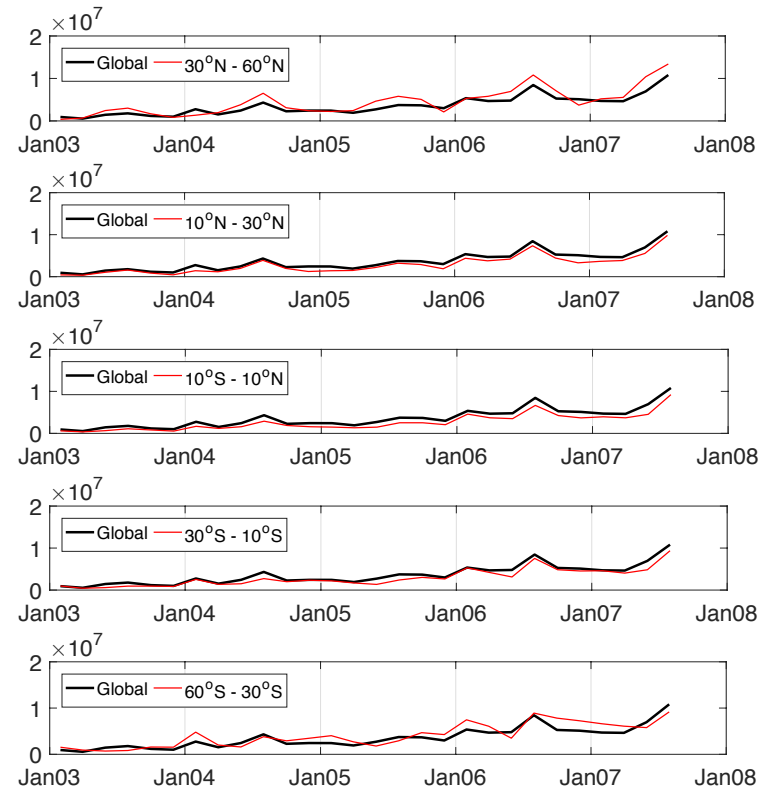


# GUVI O<sub>2</sub> in the lower thermosphere and upper thermosphere

GUVI O<sub>2</sub> at  $8.4 \times 10^{-4}$  Pa (average pressure at 133 km)

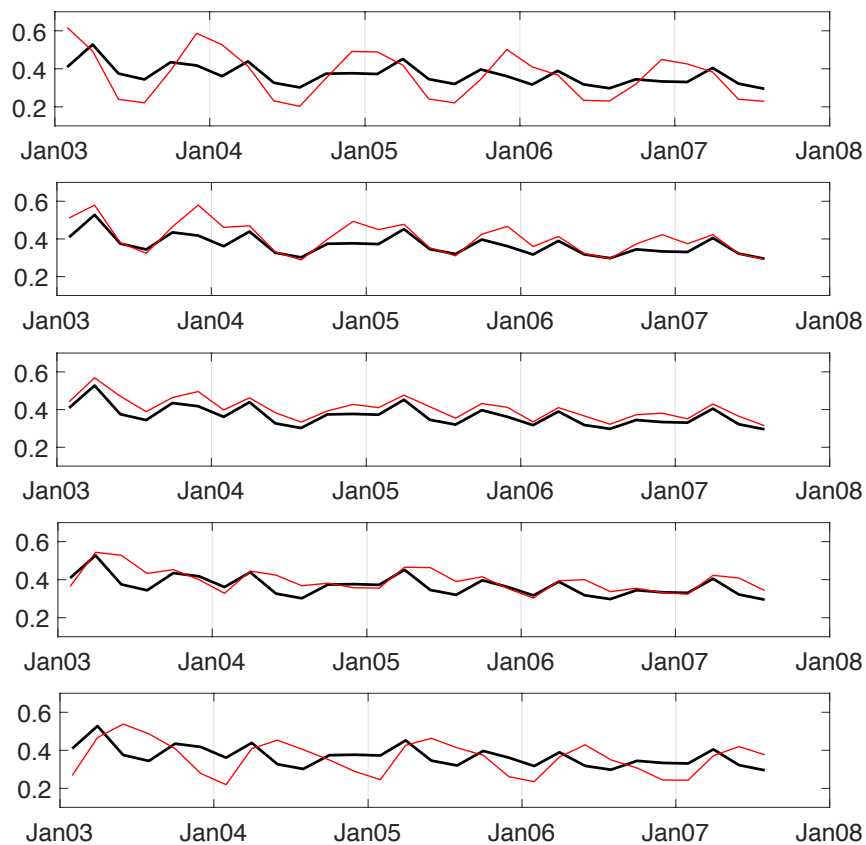


GUVI O<sub>2</sub> at  $6.35 \times 10^{-6}$  Pa (average pressure at 312 km)

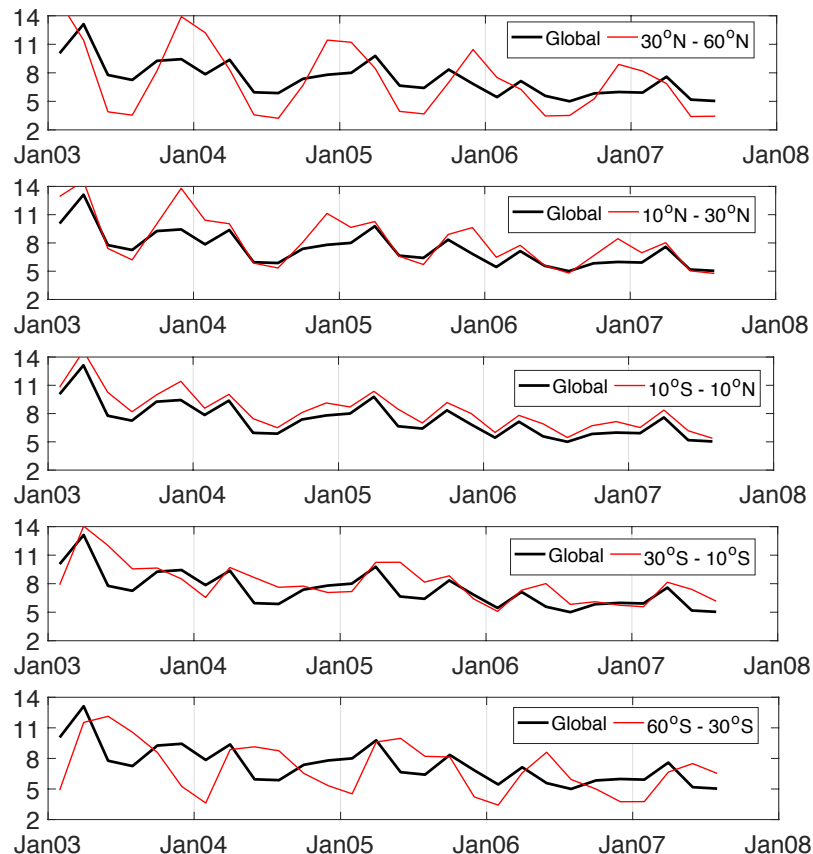


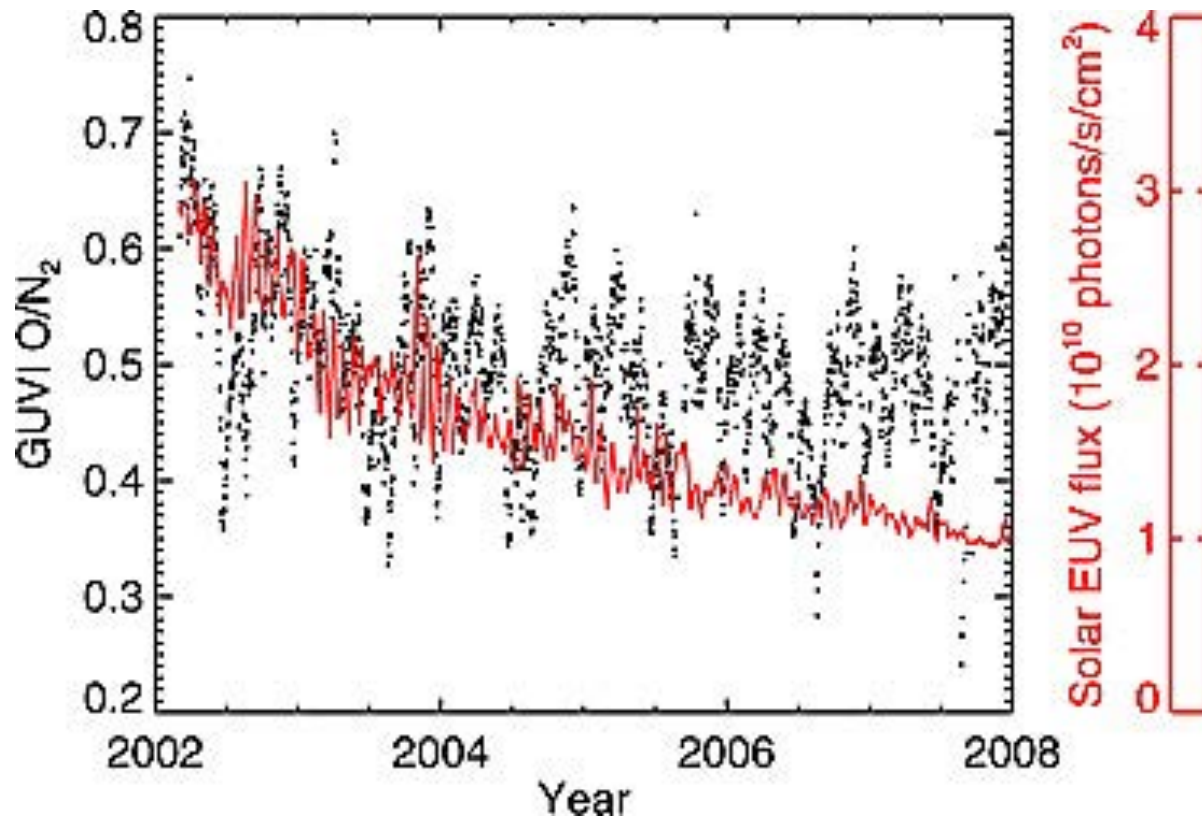
# GUVI O/N<sub>2</sub> in the lower thermosphere and upper thermosphere

GUVI O/N<sub>2</sub> ratio at  $8.4 \times 10^{-4}$  Pa (average pressure at 133 km)



GUVI O/N<sub>2</sub> ratio at  $6.35 \times 10^{-6}$  Pa (average pressure at 312 km)

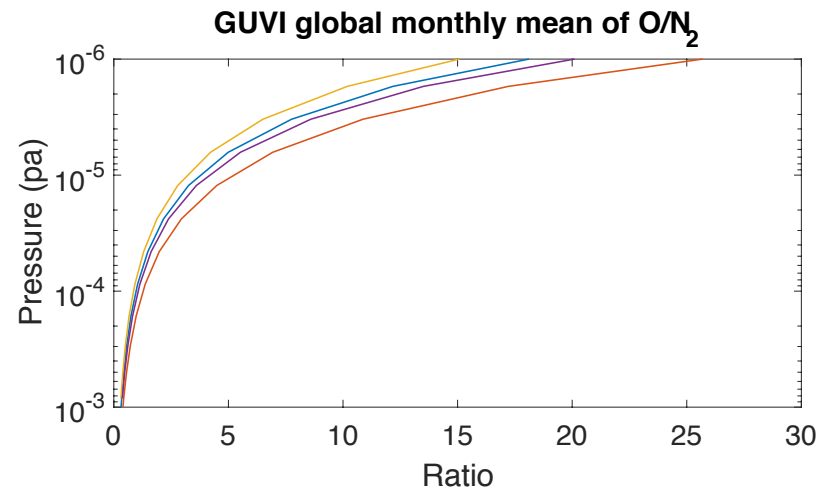
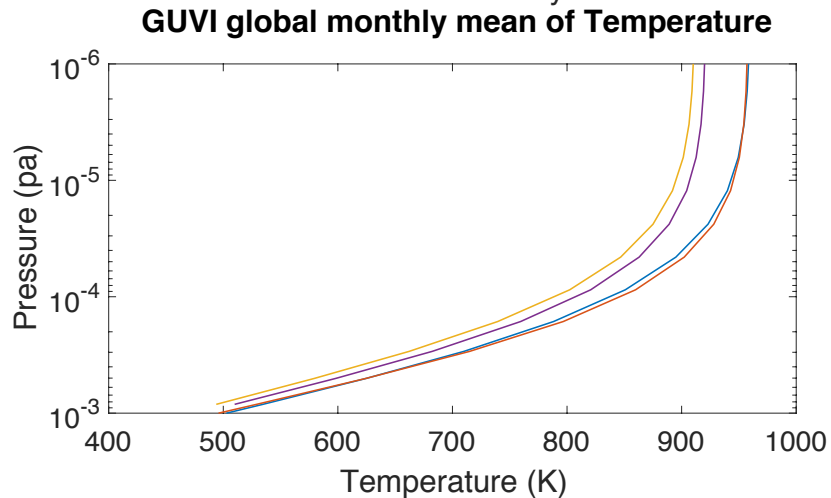
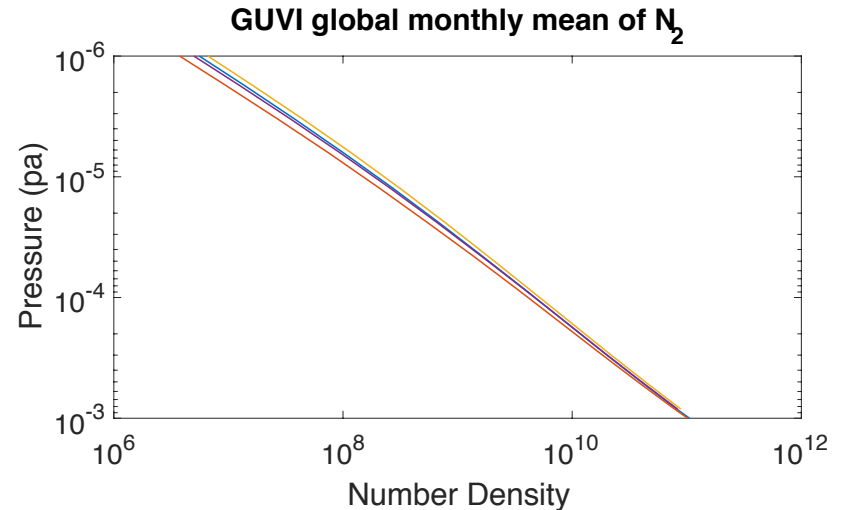
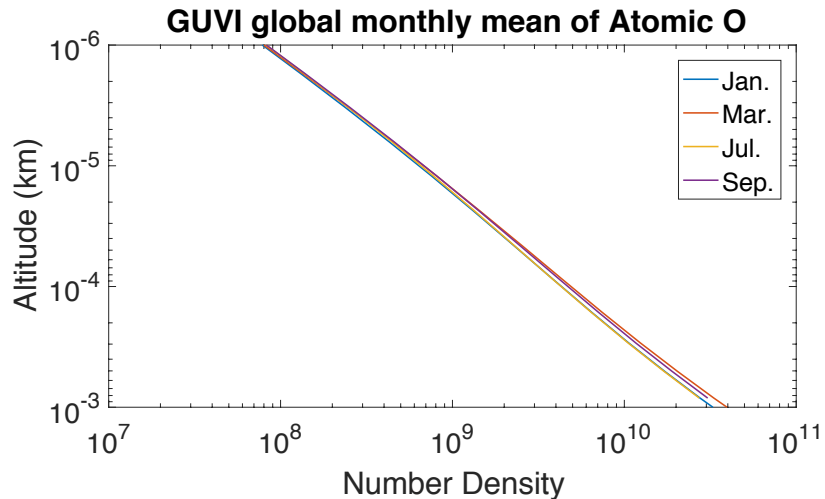




**GUVI  $\Sigma$ O/N<sub>2</sub> variations from 2002 to 2008  
(adapted from Zhang and Paxton [2011]).**

- $O/N_2$  in the upper thermosphere shows a strong annual variation with maximum in January and minimum in July, and the annual variation phases follow those in the lower thermosphere.
- The annual variation amplitude of  $O/N_2$  is stronger near solar maximum than solar minimum, following the annual variation amplitude of  $O$
- global  $O/N_2$  decreases with the declining solar phase from 2003 to 2007. This can be explained by the species dependent vertical density gradients: heavy species change more than light species [Burns et al., 2015]. The decrease of  $N_2$  density with solar flux at a constant pressure is larger than that in  $O$ .  $N_2$  is heavier than  $O$ , thus responding stronger to varying temperature.

# Vertical profiles of GUVI limb composition



# Conclusions

- GUVI limb column composition data is a useful dataset. Could be available from ICON.
- O and O/N<sub>2</sub> in the lower thermosphere show a strong annual variation at mid latitudes and a clear semiannual variation at lower latitudes.
- global mean GUVI O/N<sub>2</sub> shows the well-known annual variation, with slightly larger values in January than in July, and a semiannual variation with O/N<sub>2</sub> greater during equinoxes than at the solstices.
- O and N<sub>2</sub> densities on fixed pressure levels in the upper thermosphere are anti-correlated with solar flux, and O/N<sub>2</sub> follows the solar cycle.