# Alfvén Dynamics in High-Latitude IT Heating



# **Alfvén Dynamics in High-Latitude IT Heating**





# Wave impedance measured in the *F*-region by DE-2



# Ionospheric profiles for Aflvén wave propagation



Alfvén wave propagation in a heightresolved ionosphere (*Lysak*, 1999)

- Alfvén wave is launched from 30000 km
- Driven for 1/2 wave period: 0.5 s
- Reflected power is removed by an absorbing layer at 10000 km





# **Normalized Alfvénic Heating Rates**



# **Reflected and Absorbed Alfvénic Power**



 $\approx$  1 km "cutoff" wavelength of CHAMP small-scale, field-aligned currents

### **Cumulative Heating from Impulsive Alfvénic Energy Deposition**



#### Stochastic heating in cusp



- Alfvén power in:  $\langle S_{\parallel} \rangle \approx 1 \text{ mW/m}^2$
- Neutral wind:  $u_n = 200 \text{ m/s}$
- Width of heating region: 200 km
- Duty cycle: 50% intermittency
- *F*-region heating: **2** × **10**<sup>-8</sup> W/m<sup>3</sup>



#### Neubert and Christiansen 2003

Comparable to Zhang et al. (2015) Brinkman et al. (2016)

Effects of soft precipitation on  $\sigma_P$  have not been included

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# Thermospheric density anomalies at 400 km altitude

CHAMP

Geomagnetic coordinates

Liu et al. 2005



#### Global simulation (LFM) model streamlines

 $\Rightarrow$  Anomalies straddle dayside (cusp) and nightside convection throats

 $\Rightarrow$  Anomalies are strongly controlled by magnetic geometry

# Effects of dynamic (Alfvénic) SSFACs during anomalies

### CHAMP | 400 km altitude

Alfvénic field-aligned currents are prevalent in the cusp region.

How does energy deposition by Alfvénic SSFACs differ from that of quasi-static SSFACs?

# Ionospheric reflection of small-scale Alfvén waves

Kilometer-scale (in  $\lambda_{\perp}$ ) waves are almost fully absorbed.

Larger scale waves are reflected and trapped.





Lessard and Knudsen 2001