

Multipoint Thermospheric Wind Observations: Networks of Fabry-Perot Interferometers and Plans for ICON

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- Daniel Fisher
- Rafael Mesquita
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- Kenneth Marr
- ICON Team
- NATION Team

ECE ILLINOIS

 ILLINOIS

Outline

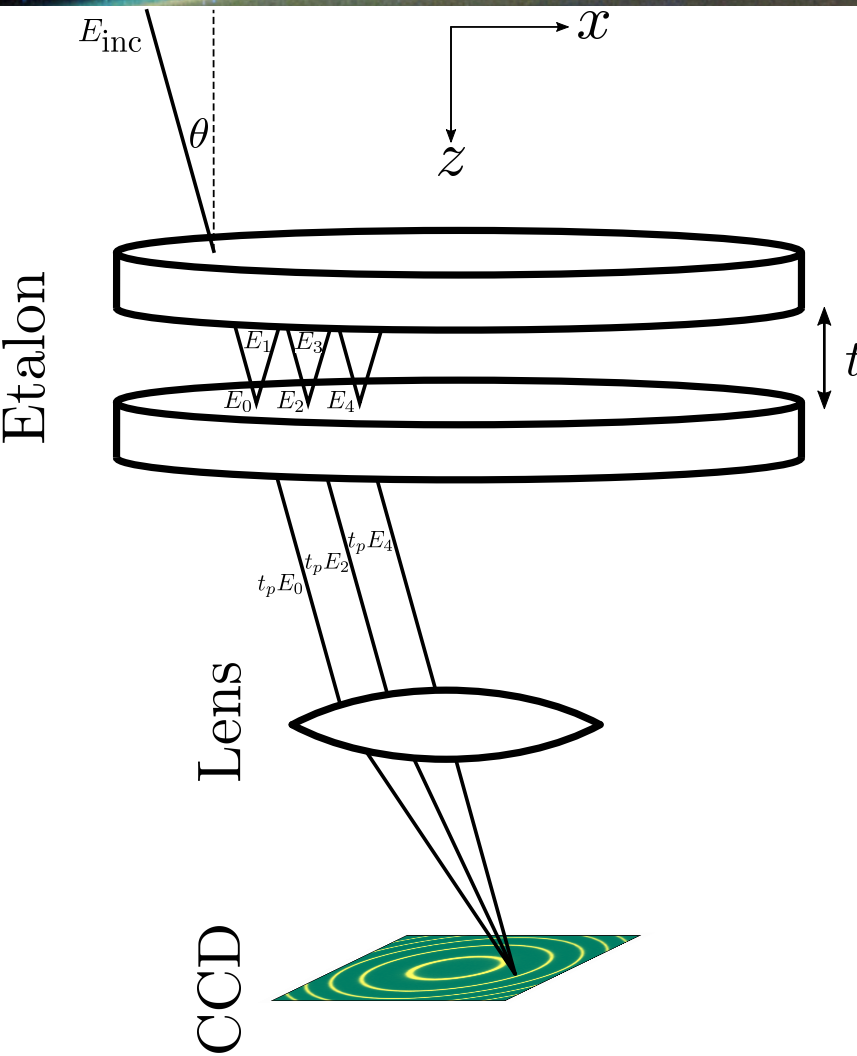
- Networks of ground-based wind instruments
- Midnight temperature maximum
 - Direct evidence of converging wind field
 - Propagation to midlatitudes
- ICON wind measurements

Thermosphere

630.0 nm

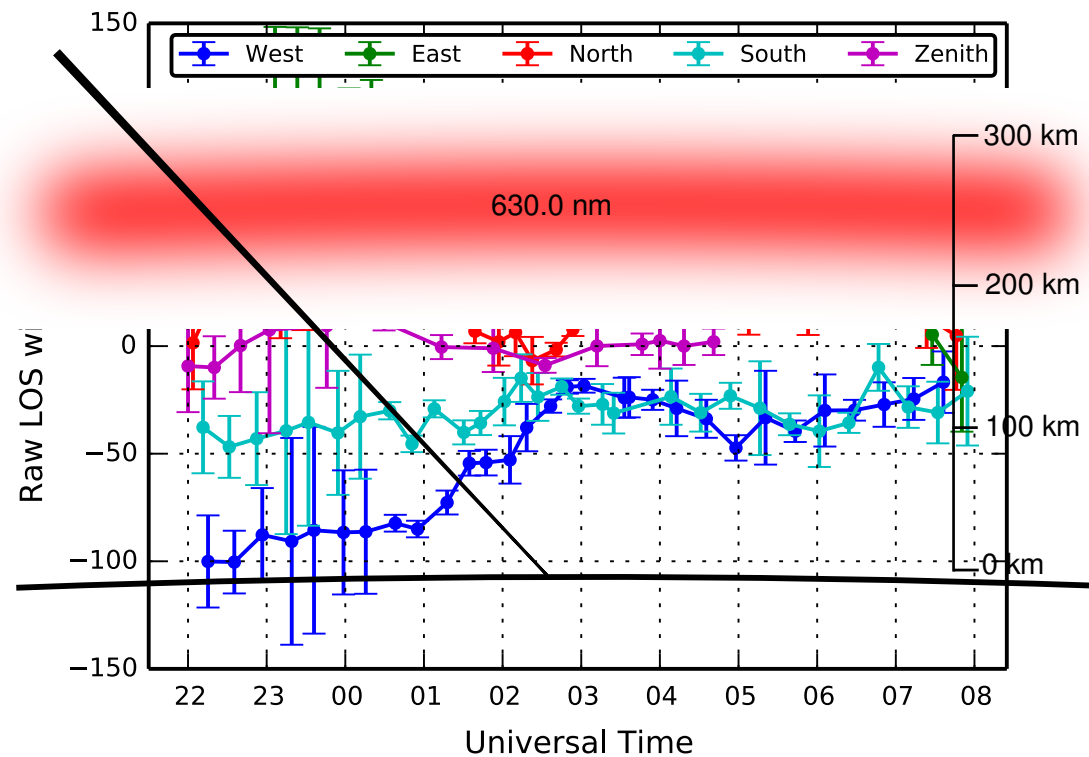
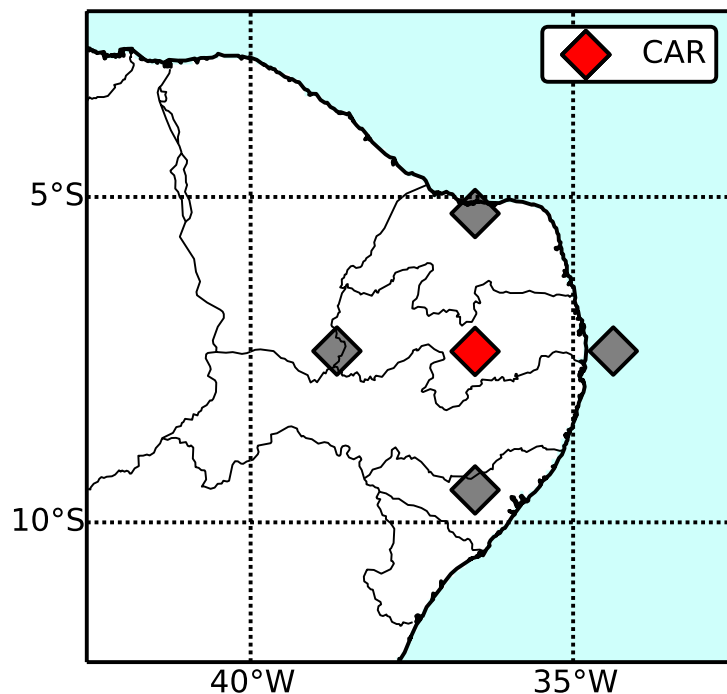
Mesosphere

557.7 nm



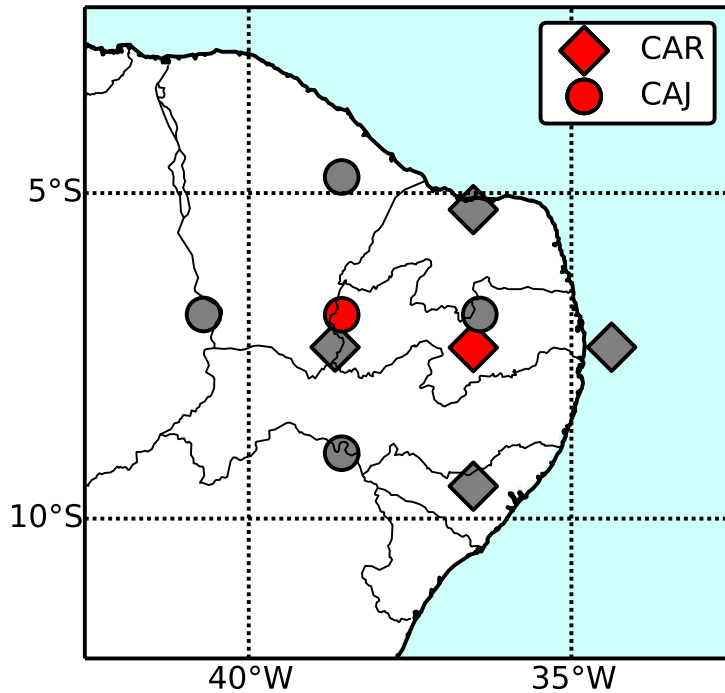
Doppler shift \rightarrow Wind
Doppler width \rightarrow Temperature

A single Fabry-Perot interferometer (FPI)

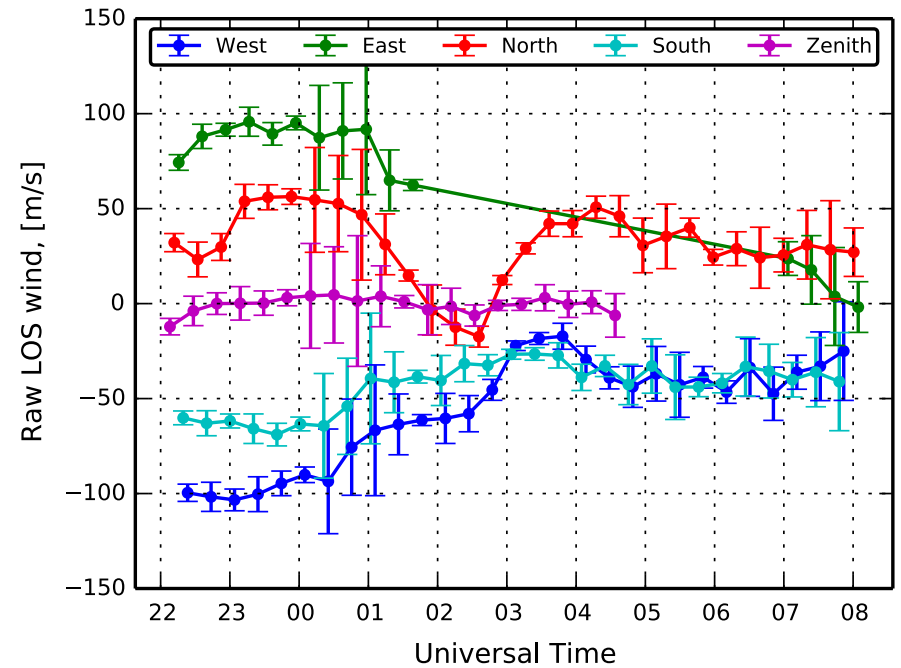
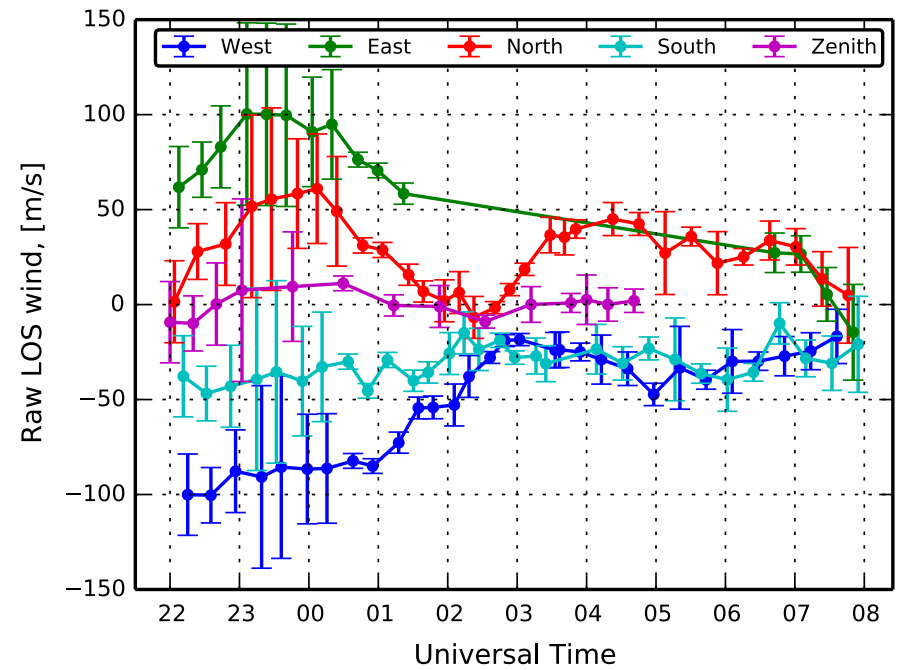


Estimate temporal evolution of the wind over a small (~500km) region.

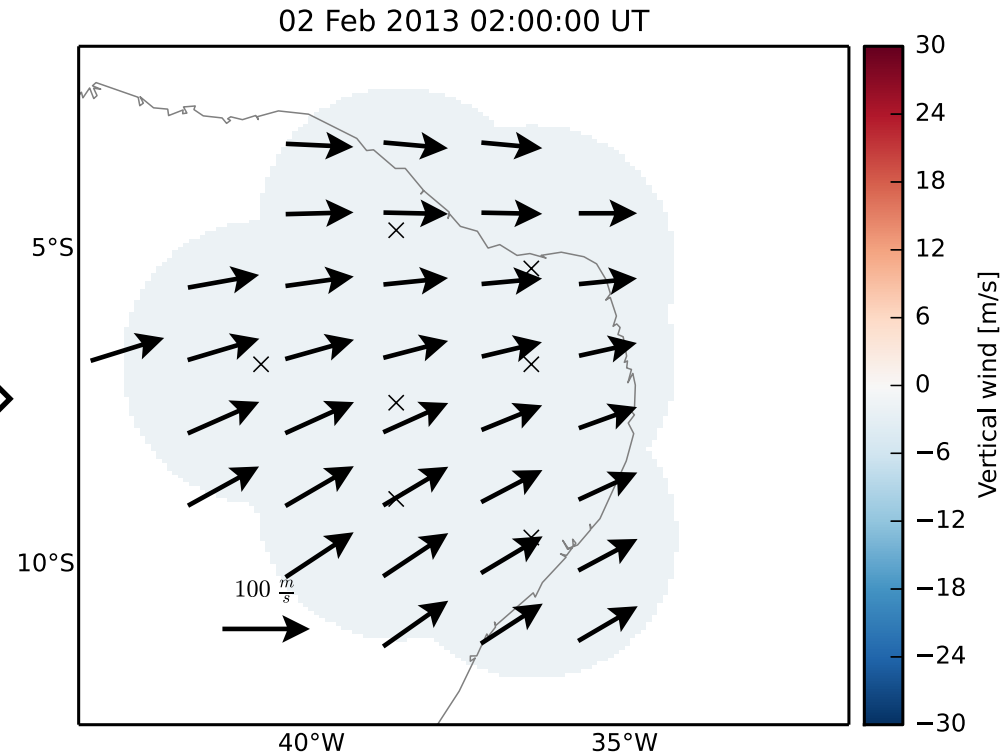
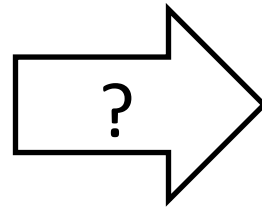
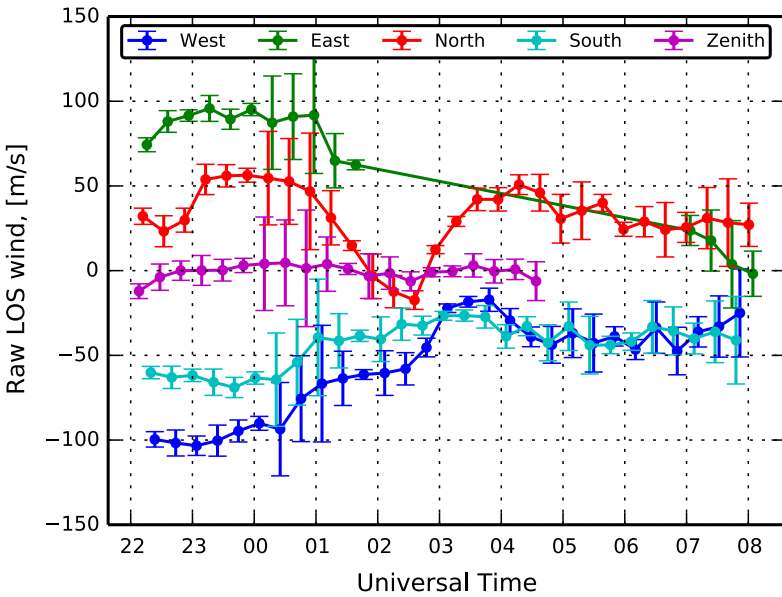
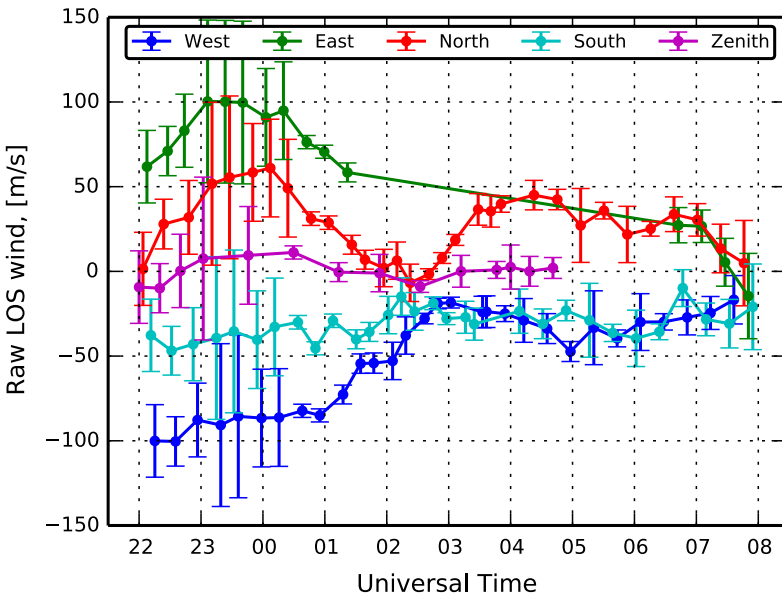
A network of FPIs



Observe temporal
and spatial
distributions



Wind Field Estimation



Wind Field Estimation

Unknowns:

$$m_i = \overset{\text{Zonal wind}}{u_i} \sin \theta \cos \phi + \overset{\text{Meridional wind}}{v_i} \sin \theta \sin \phi + \overset{\text{Vertical wind}}{w_i} \cos \theta \quad \bar{m} = A\bar{u}$$

Knowns:

$$\underbrace{m_i}_{\text{Measured LoS velocity}} \quad \underbrace{\phi}_{\text{Azimuth angle}} \quad \underbrace{\theta}_{\text{Zenith angle}}$$

minimize $\|D\bar{u}\|_2$ — Matrix operator measures “roughness” via gradient and curvature

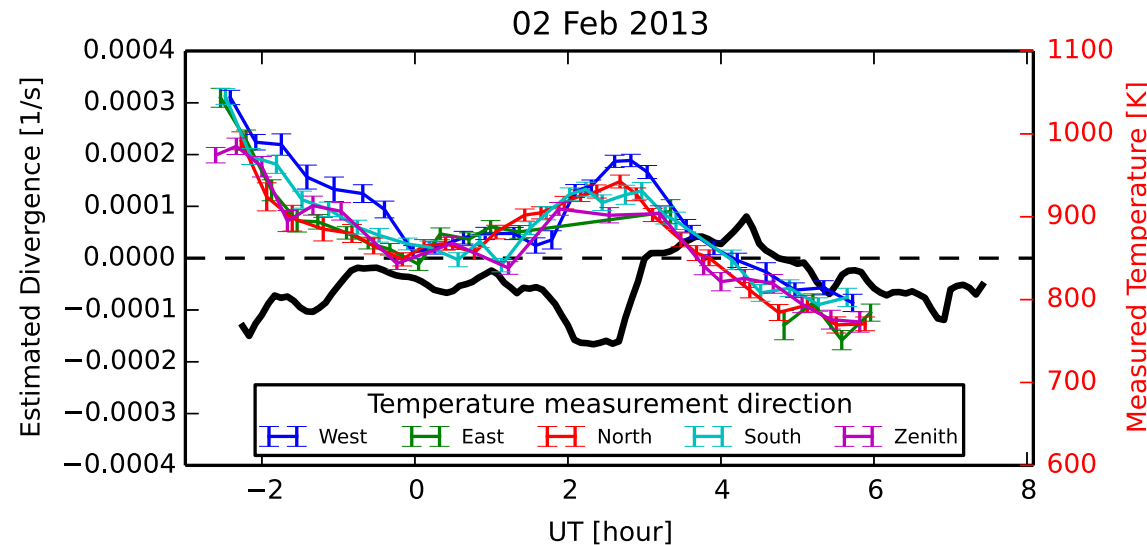
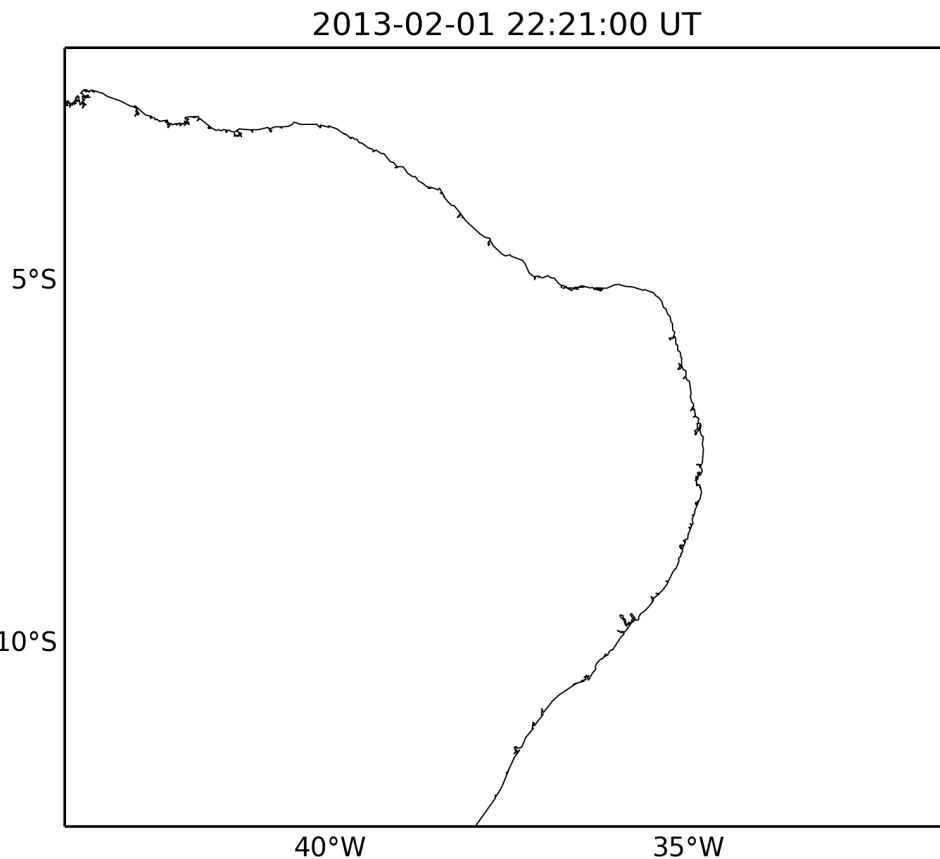
such that $\|\Sigma^{-\frac{1}{2}}(A\bar{u} - \bar{m})\|_2 \leq M$

Weight matrix incorporates measurement uncertainty

Number of measurements

Of all wind fields that match the data,
choose the smoothest.

Midnight Temperature Maximum

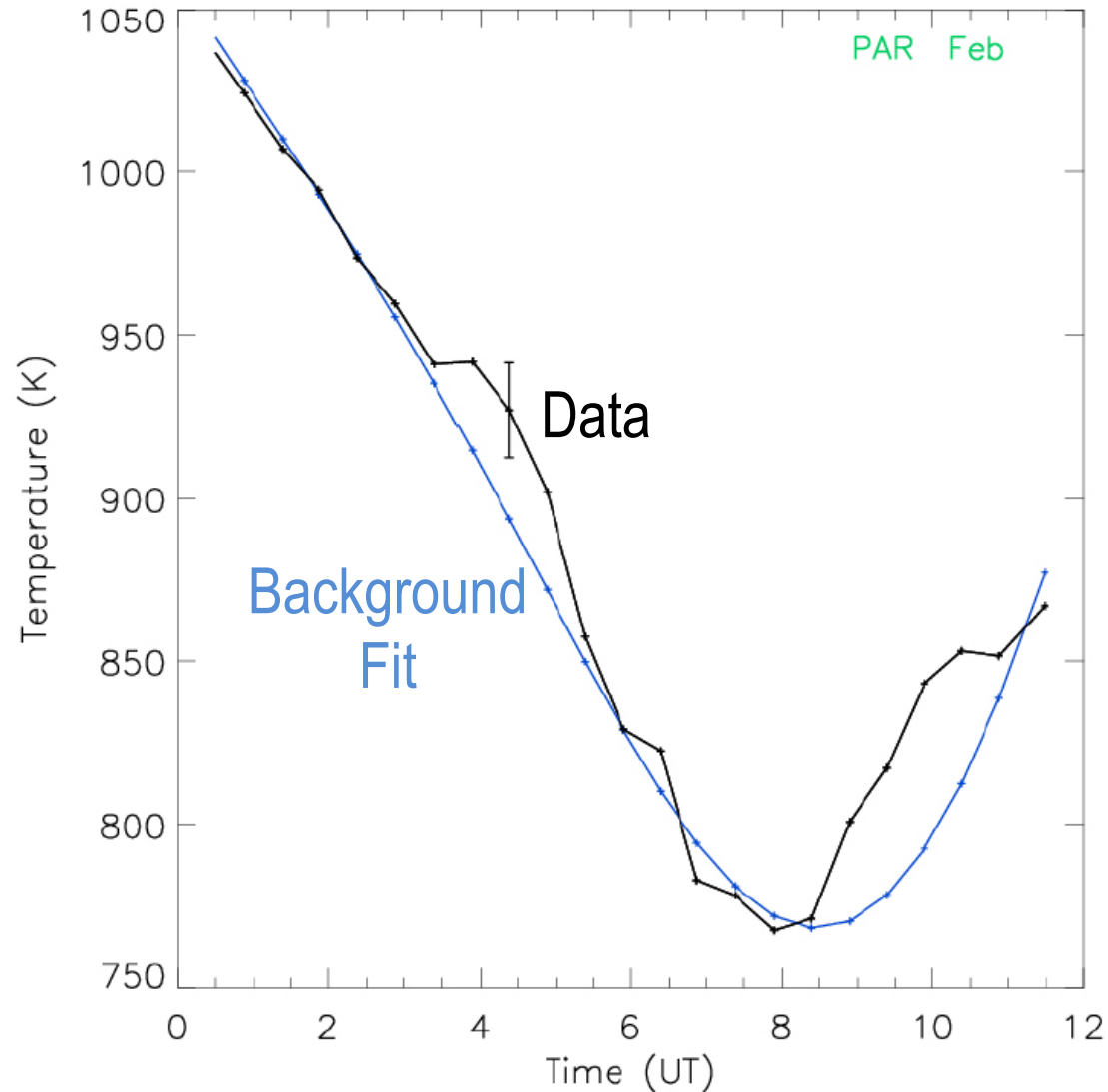
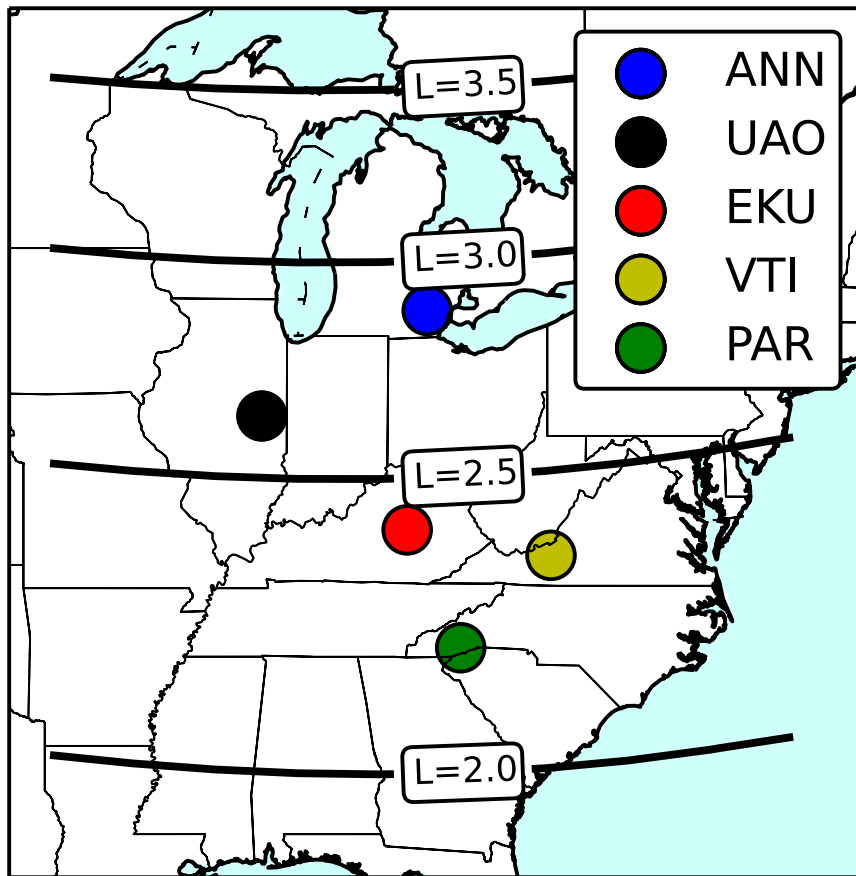


3 UT = 0 LT

Direct evidence of convergent wind field during MTM

Midnight temperature maximum... at midlatitudes?

NATION: North American Thermosphere Ionosphere Observing Network



[R. Mesquita]

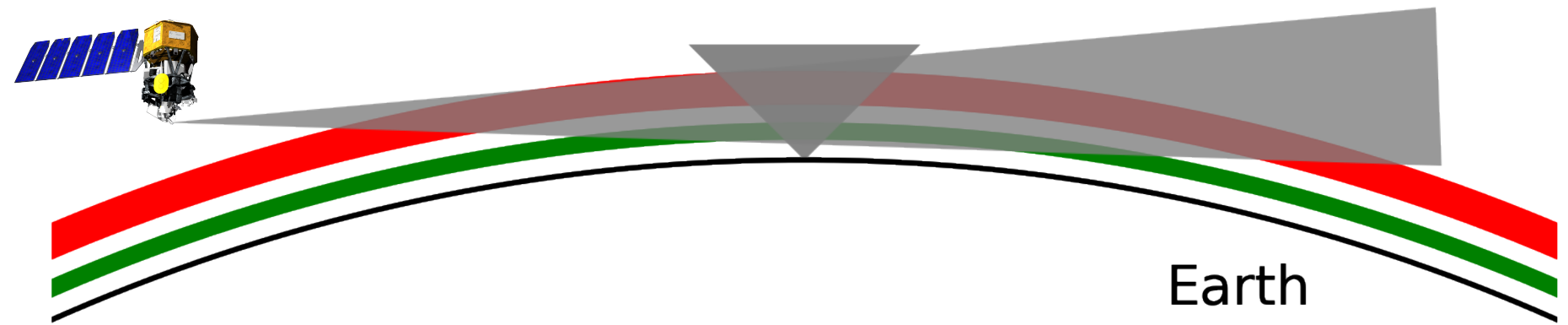
Midnight temperature maximum... at midlatitudes?



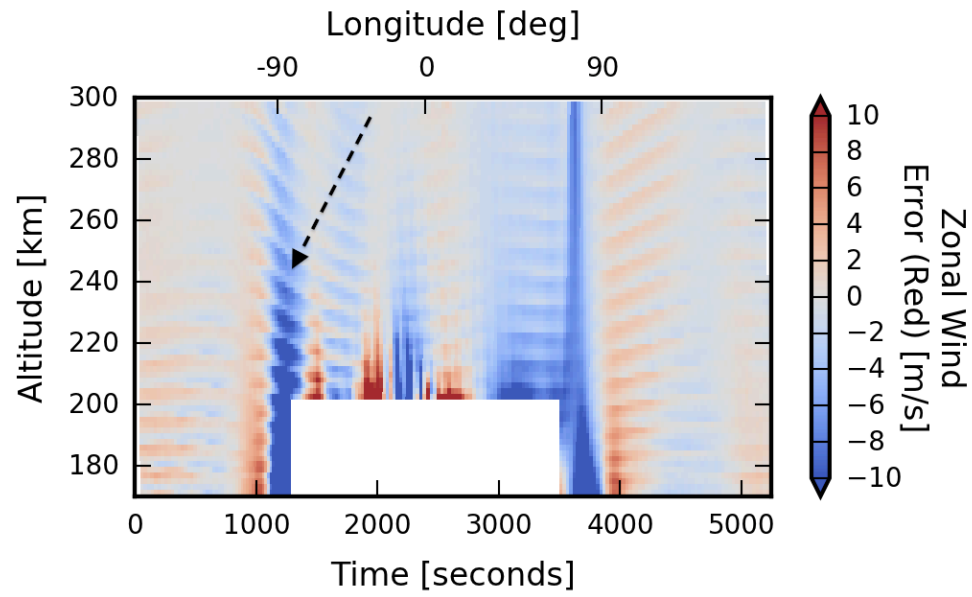
- Direct observation of winds, propagation
- Consistent with timing reported by *Hickey et al.* [2014] (ion and neutral temperatures)
- Propagation to 45°N

ICON

Ionospheric Connection Explorer



- MIGHTI is ICON's wind instrument
- System simulations suggest:
 - Expected Precision: **1-4 m/s** (1σ)
 - At 5x500km (Vert x Horz) res (Red 30x500)
 - Systematic Error: **3-6 m/s** (80th prctile)



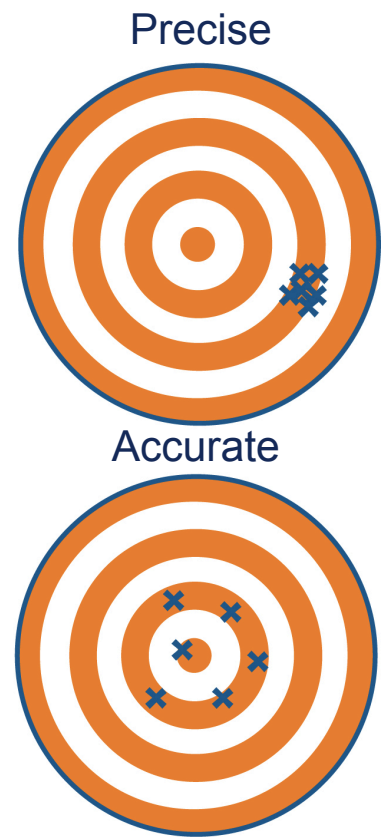
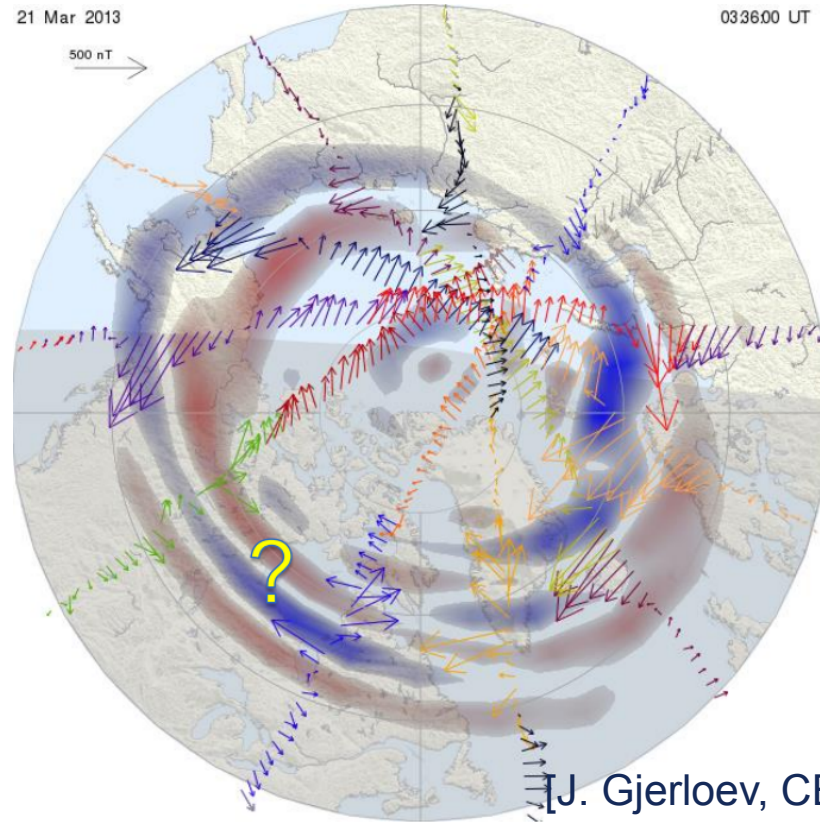
Bias

- Error
 - 2nd moment: precision, statistical uncertainty
 - 1st moment: accuracy, bias
- Data providers rarely report 1st moment
 - Important for assimilation and data fusion
- Geophysical data often have a bias towards “smoothness”
- Can quantify with resolution matrix:

$$\text{if } y = Ax + n$$

$$\text{and } x^* = A^p y$$

$$\text{then } R = A^p A$$



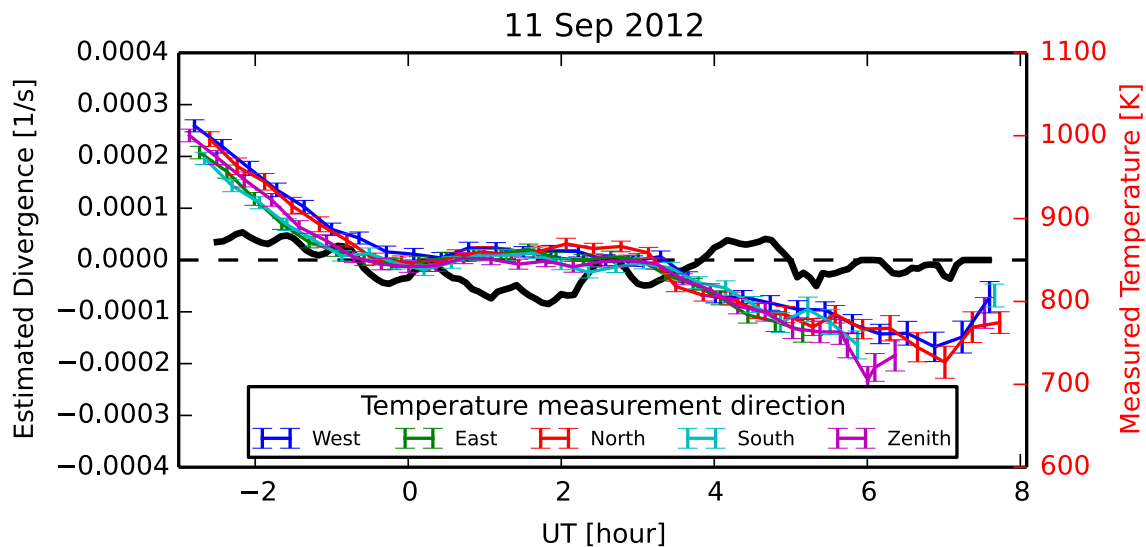
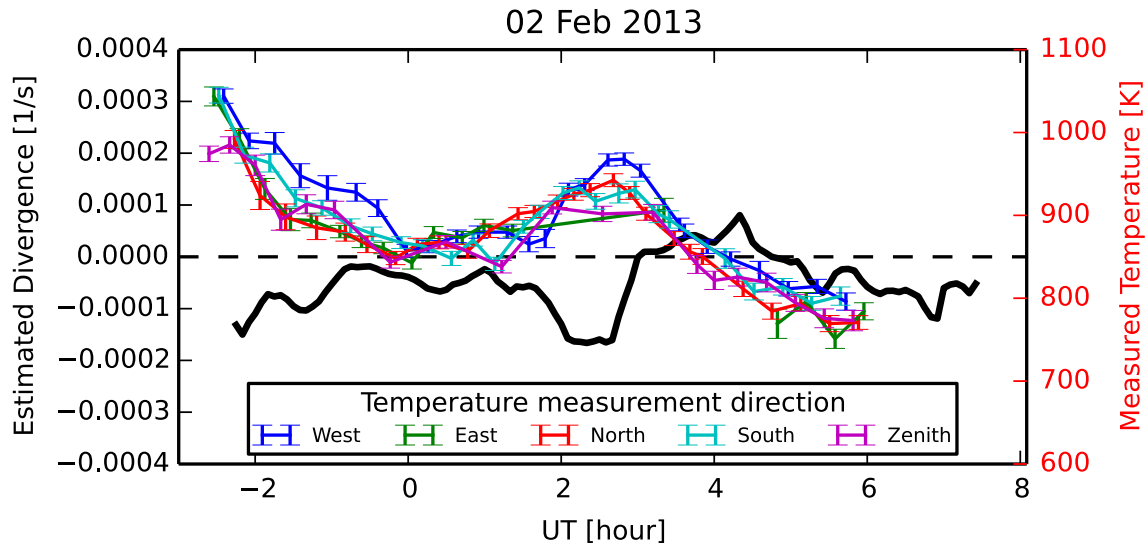
[J. Gjerloev, CEDAR Prize Lecture 2016]

Conclusions

- Networks of FPIs allow us to see spatial structure, revealed by statistical estimation techniques
 - Direct evidence of converging wind field during MTM
 - Extension up to midlatitudes
- ICON/MIGHTI will provide global coverage
- Bias is just as important as uncertainty, especially for disparate data sets

Backup Slides

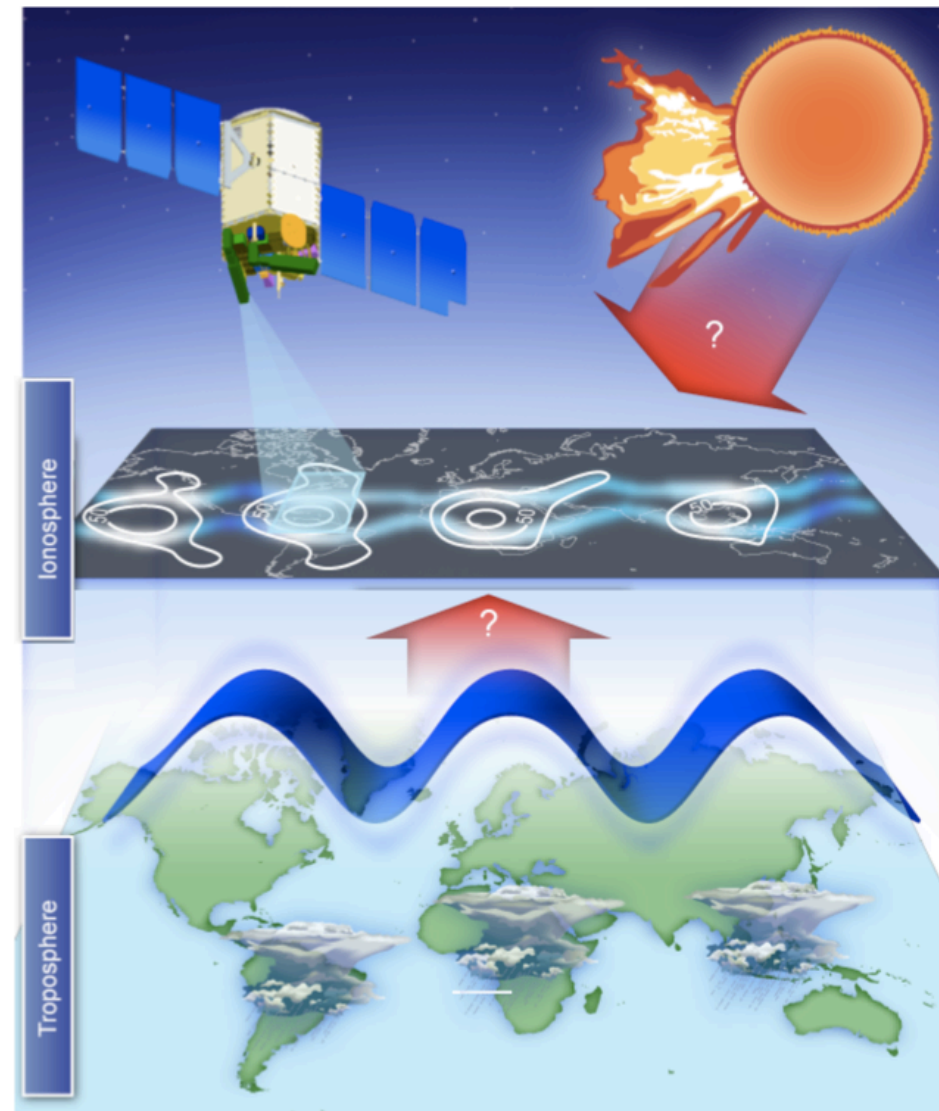
Application 1: Midnight Temperature Maximum



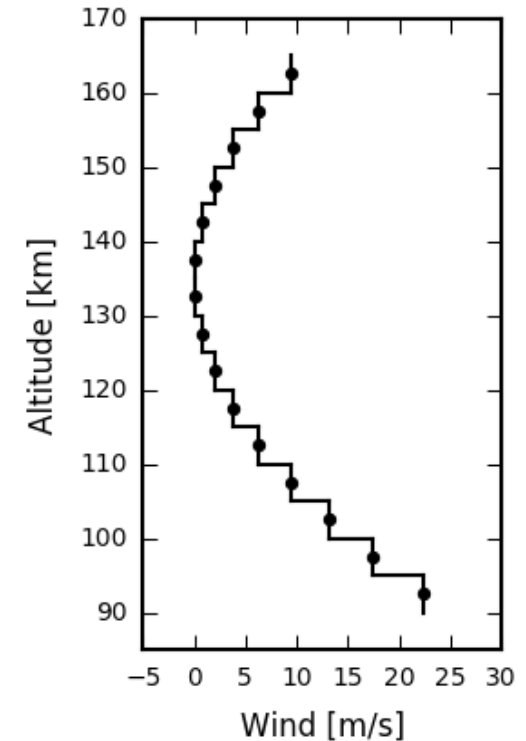
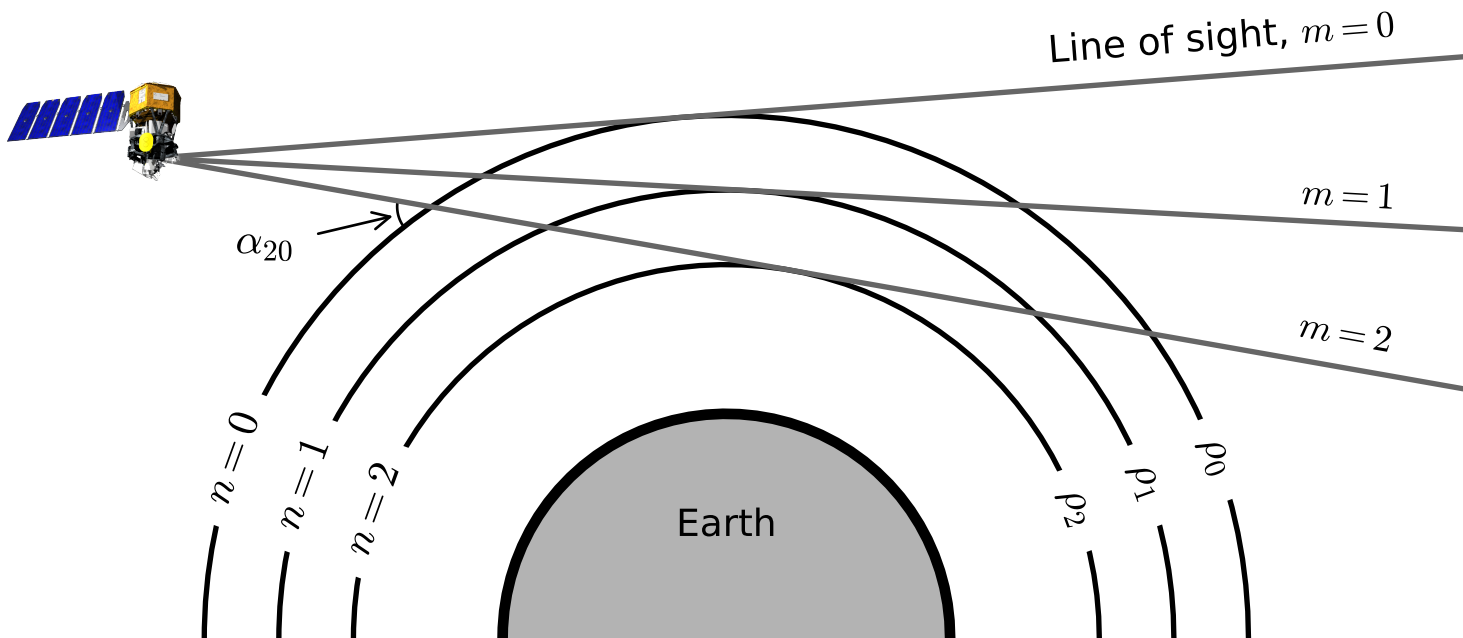
- Significant variability in amplitude of MTM
- Thought to be related to terdiurnal tide

ICON Mission Goal

- Identify drivers of extreme ionospheric variability
 - Cause of structure at low latitudes during quiet times
 - Coupling processes with large-scale atmospheric waves
 - Competition of these internal drivers with external drivers during geomagnetic storms



Problem: Scrambled spectra from different altitudes



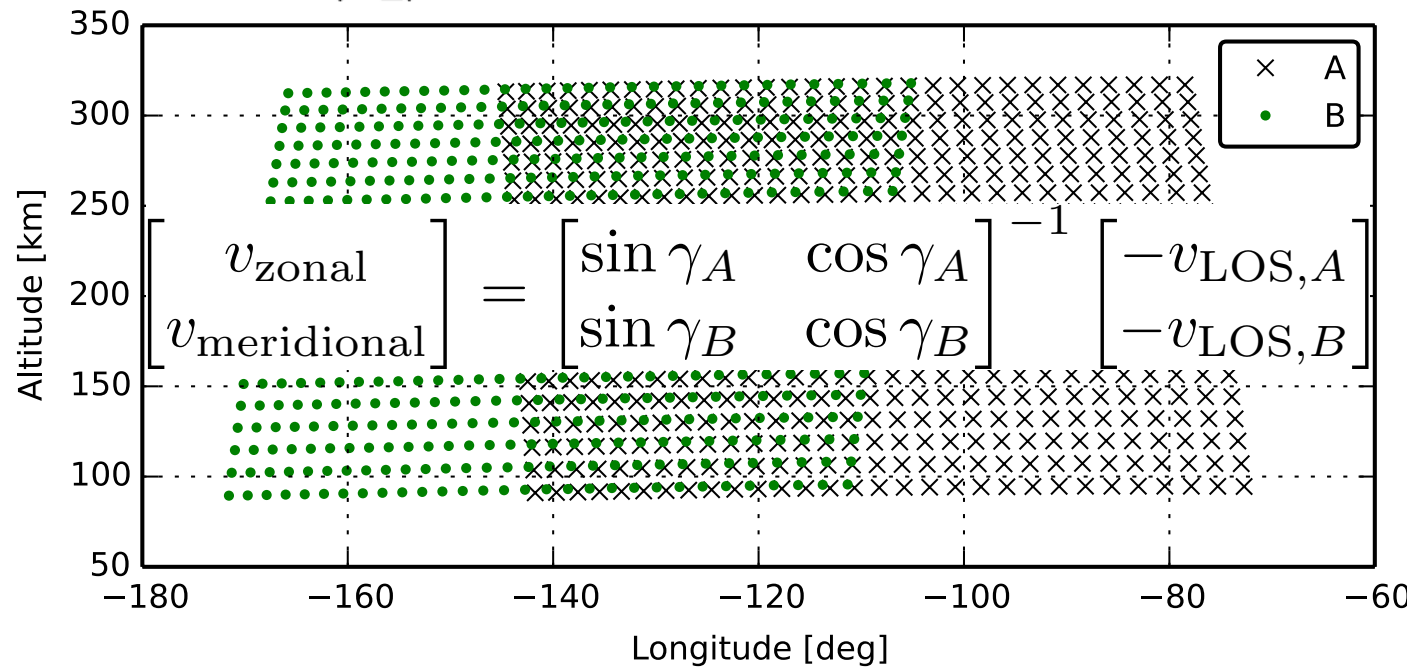
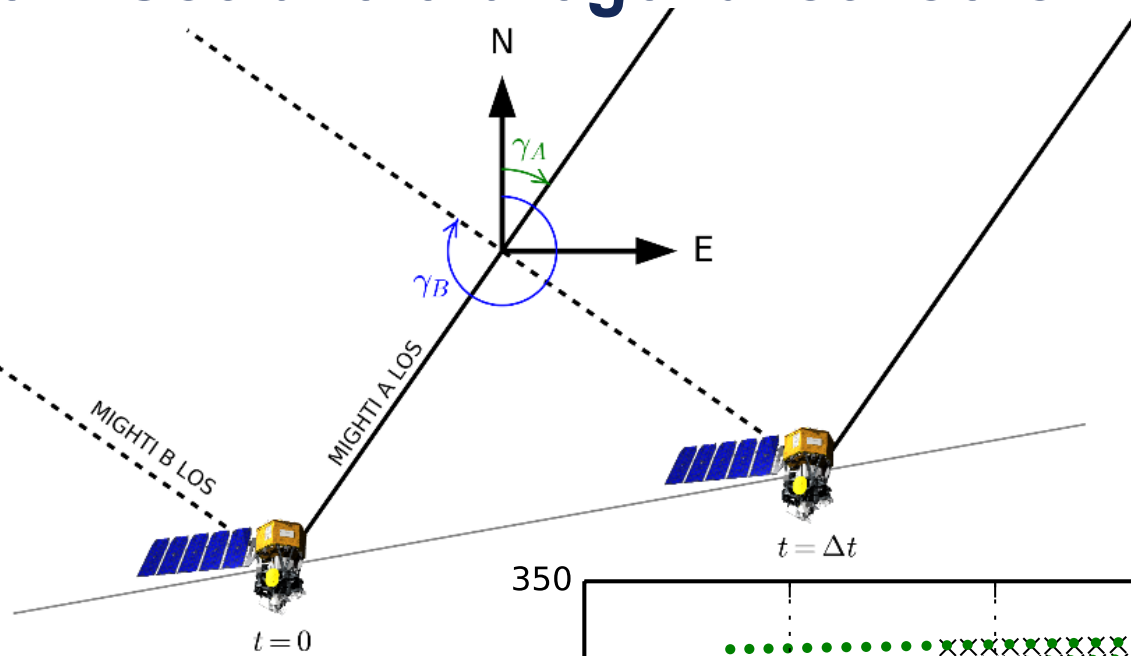
Solution: Abel-like inversion

Assume spherical symmetry and piecewise constant altitude dependence

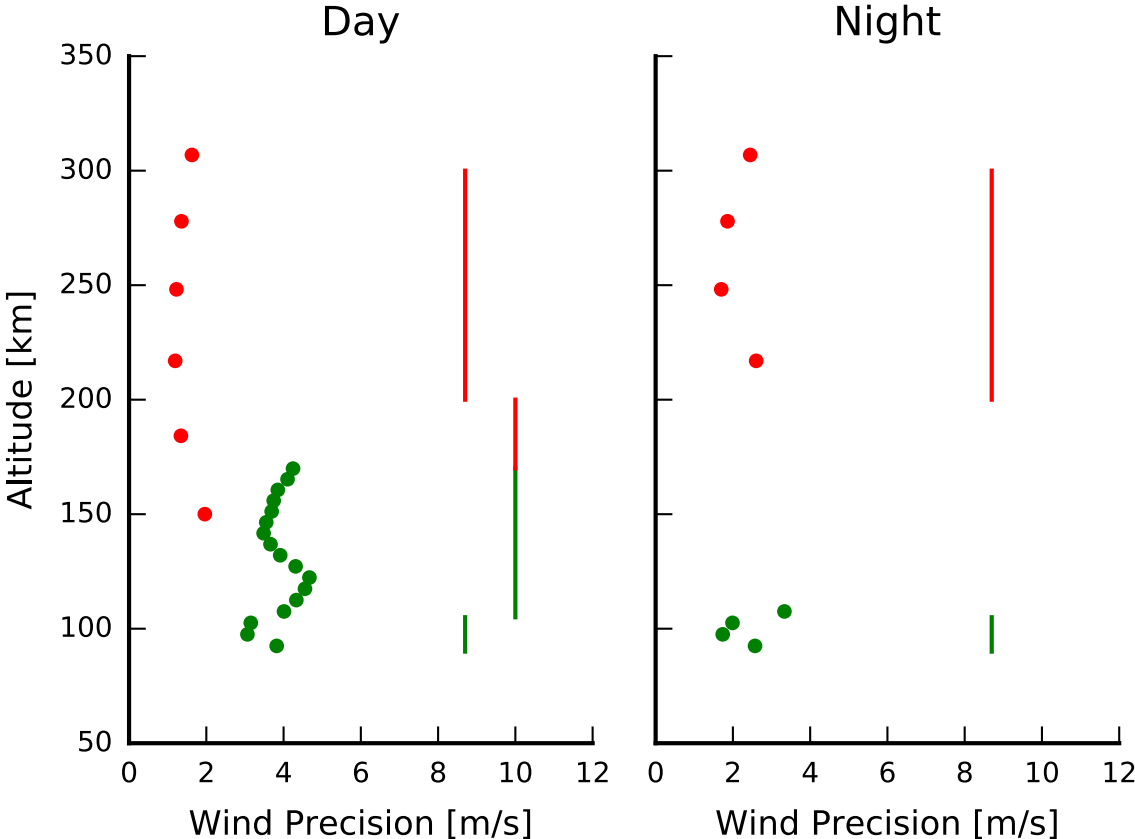
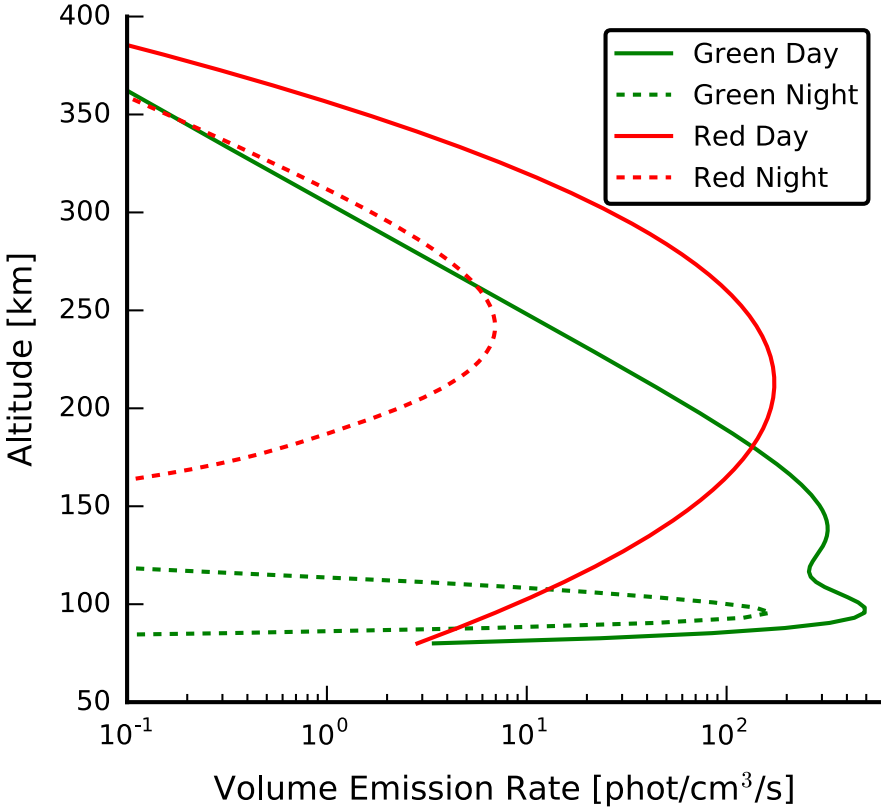
$$g = \int_0^{\infty} f(r) ds \longrightarrow g = \sum_{n=0}^{N-1} f_n \Delta s_n \longrightarrow \mathbf{g} = \mathbf{Sf}$$

Problem: Only one component measured

Solution: Use two orthogonal sensors

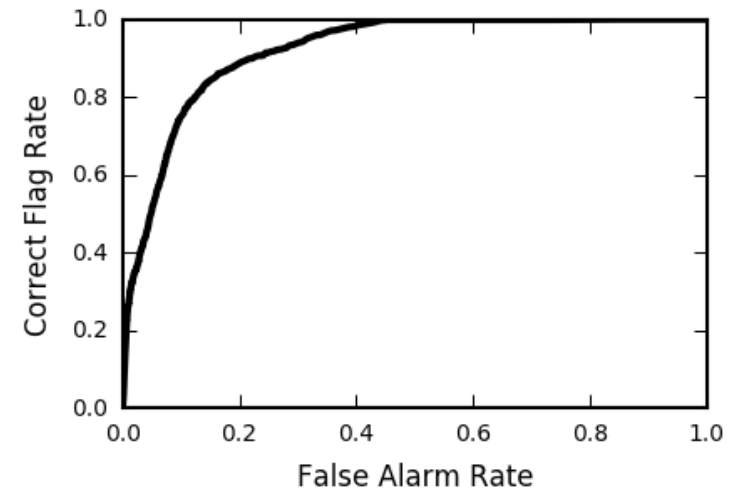
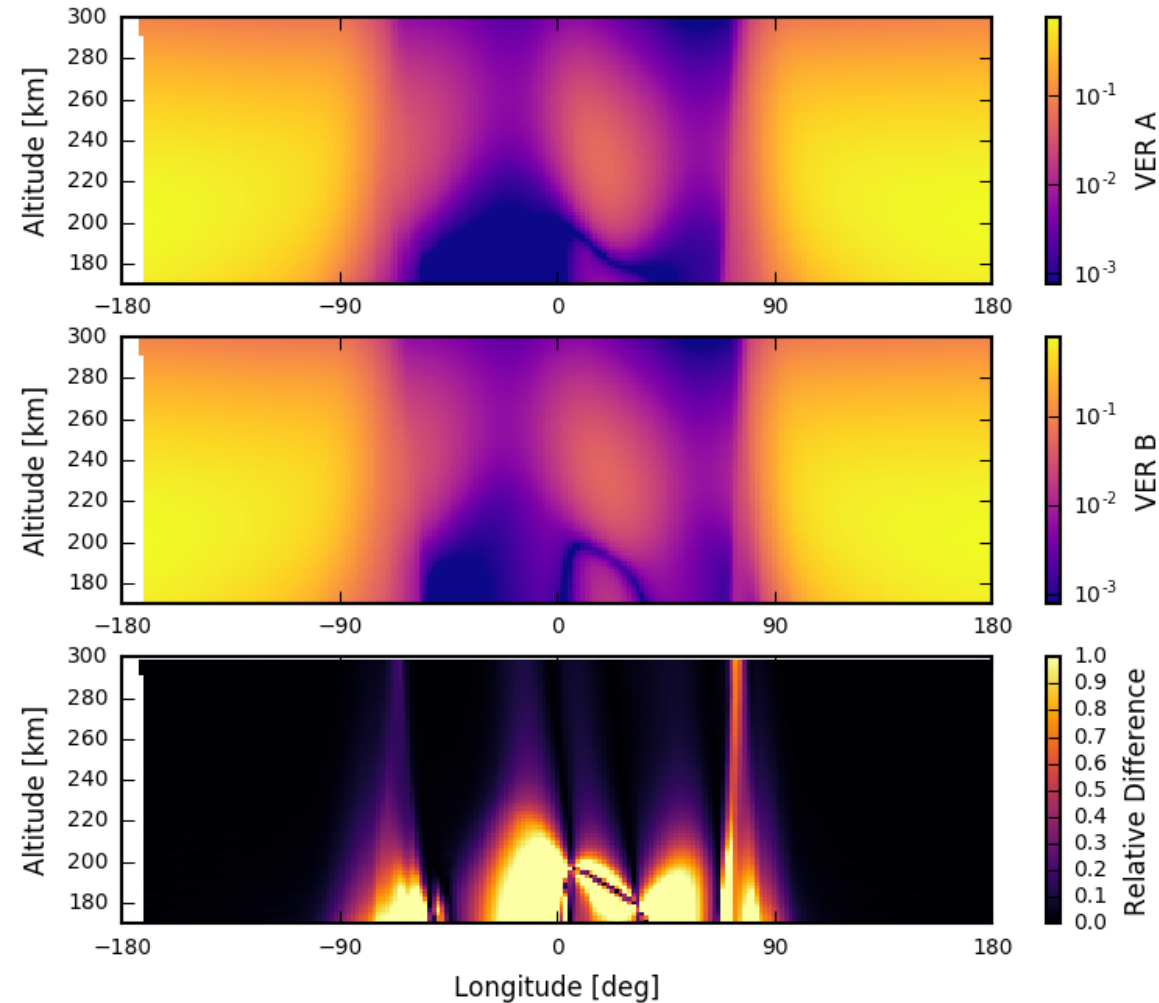


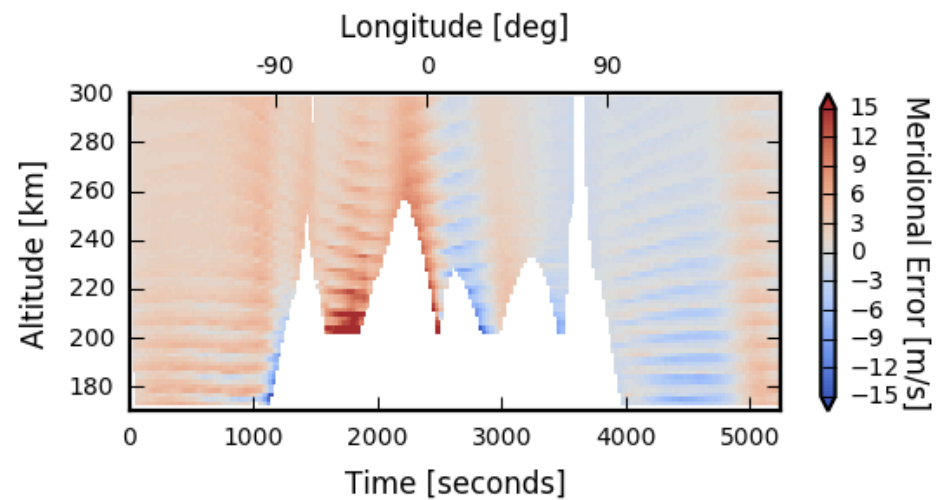
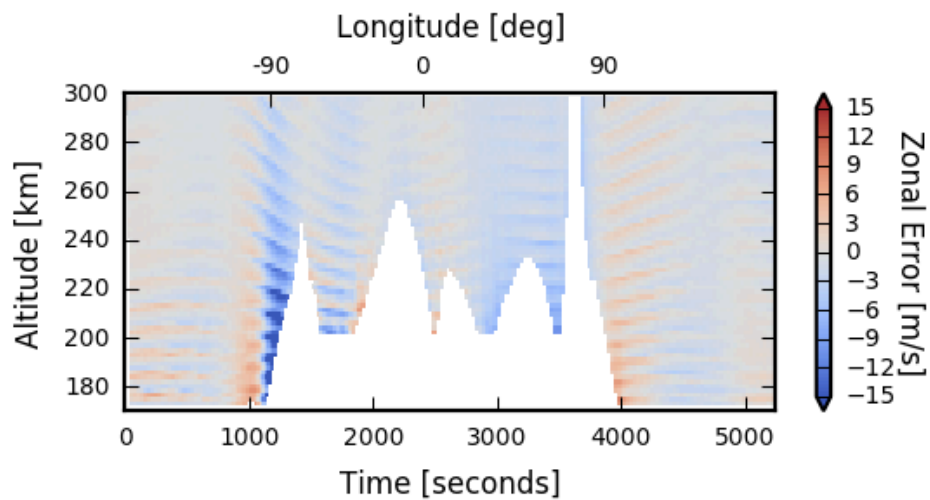
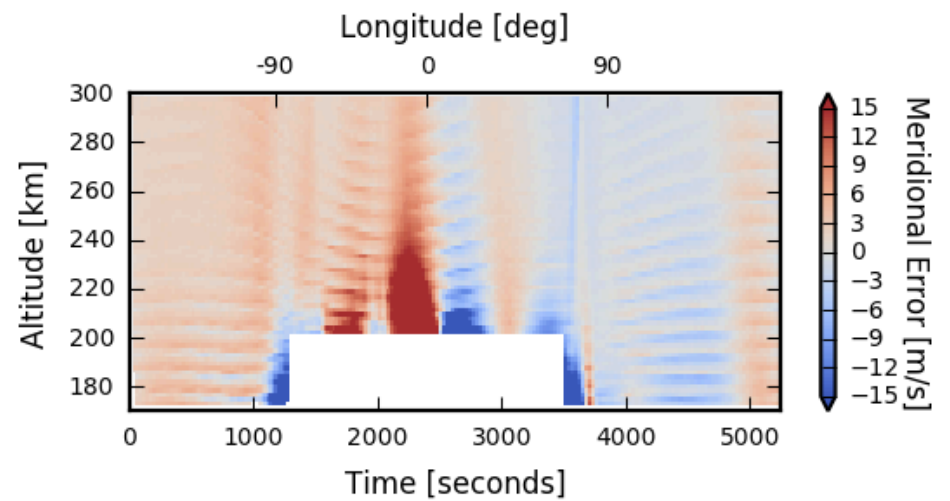
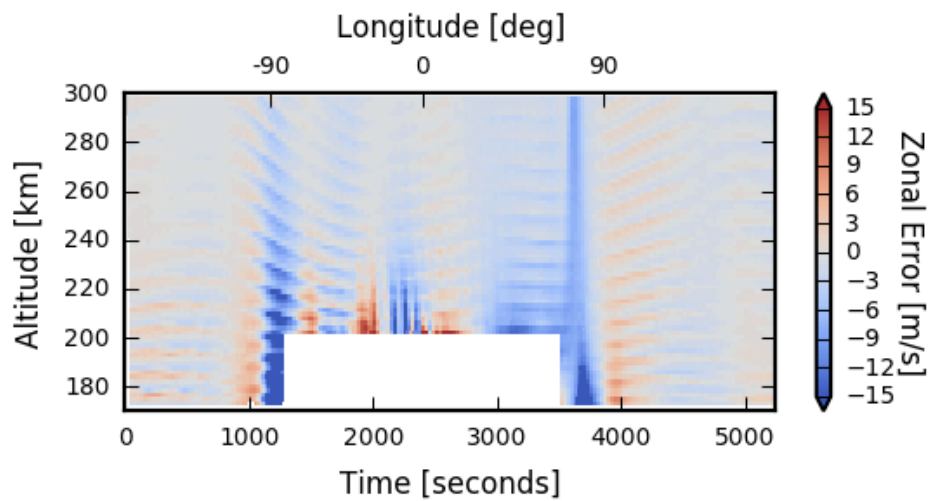
Expected Precision



Spherical Asymmetry Flag – Red

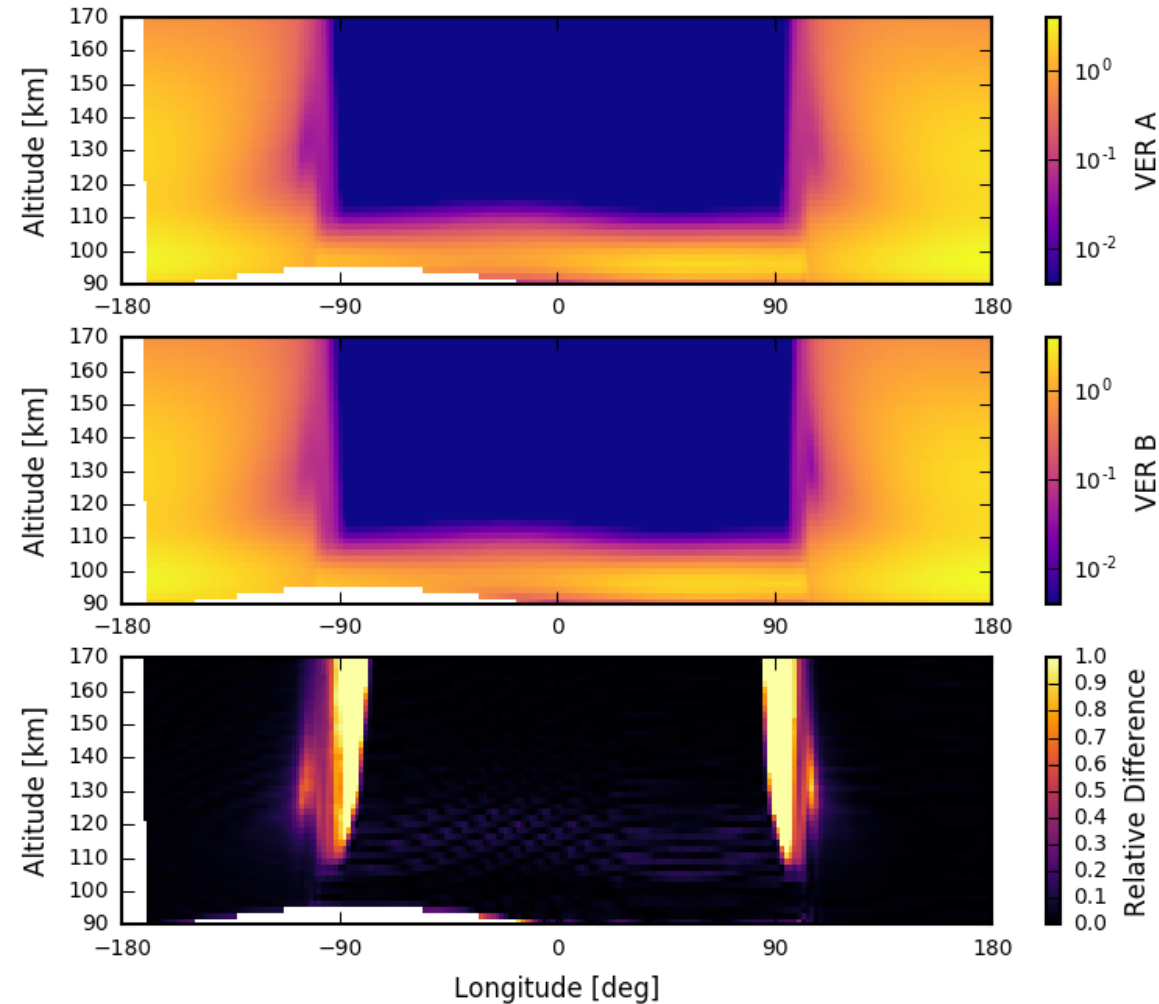
- ❑ Compare VER estimates from A and B
- ❑ Use threshold to flag expected errors in wind estimate



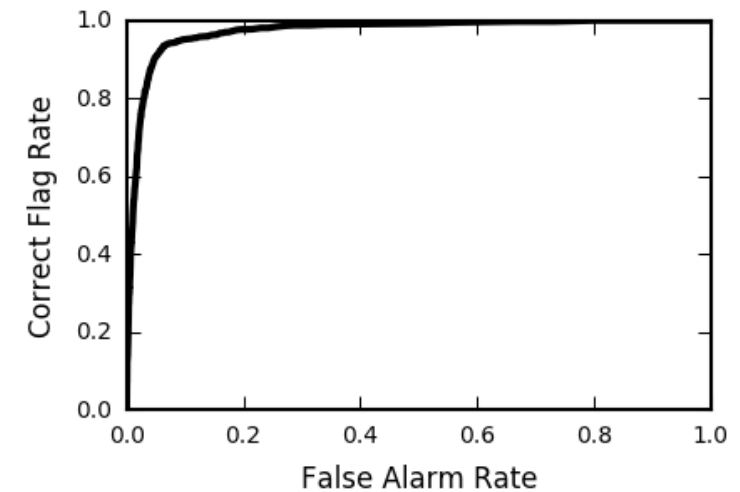


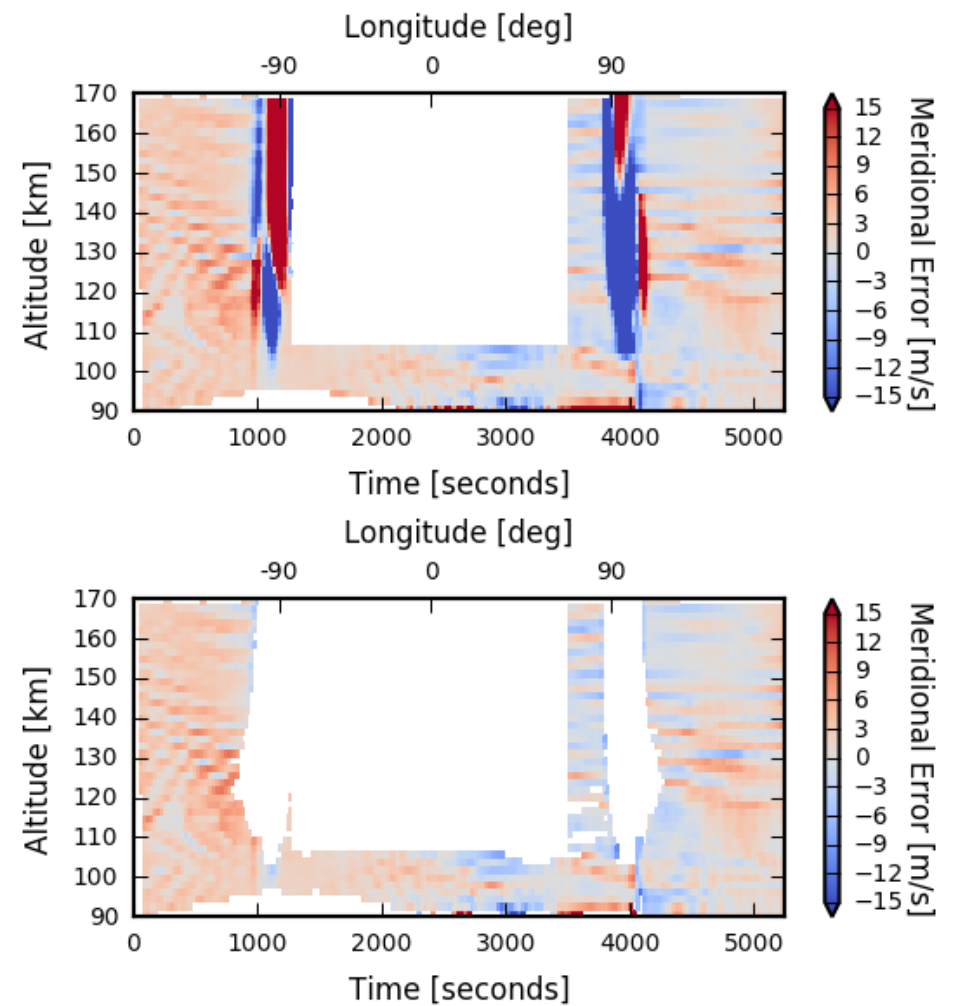
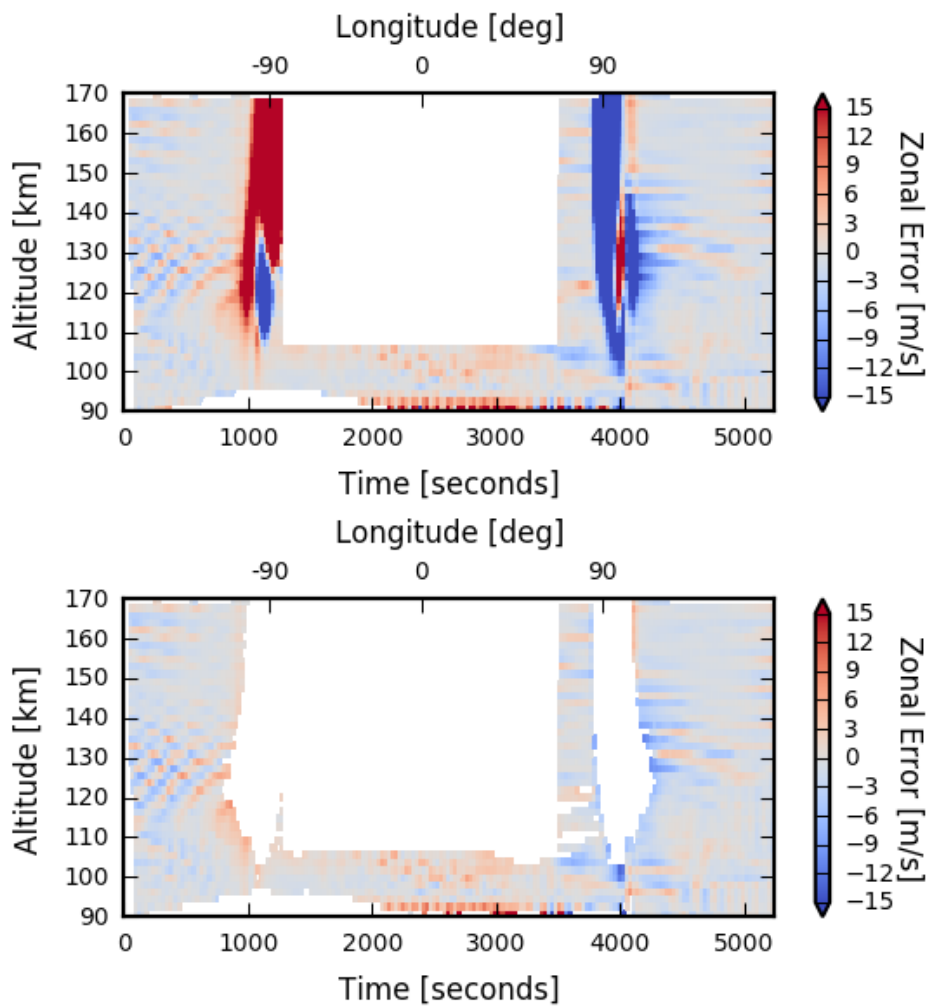
6.5 \rightarrow 3.9 m/s rms error

Spherical Asymmetry Flag – Green



- ❑ Compare VER estimates from A and B
- ❑ Use threshold to flag expected errors in wind estimate





26.0 \rightarrow 4.5 m/s rms error

Airglow above 300 km

- Assuming:
 - Constant wind & temp. above 300km
 - Exponential falloff in VER (H=26km)
- Effective path length in top layer:
- New algorithm works well for contrived cases
- Doesn't make a difference in our realistic end-to-end simulation

$$w_{m0}^{\text{above}} = \int_{\sqrt{r_0^2 - r_m^2}}^{\infty} e^{-\frac{1}{H}(\sqrt{x^2 + r_m^2} - r_0)} dx$$

