

Stable Auroral Red (SAR) Arcs:

Complex Modes of Inner-Magnetosphere-Ionosphere Coupling

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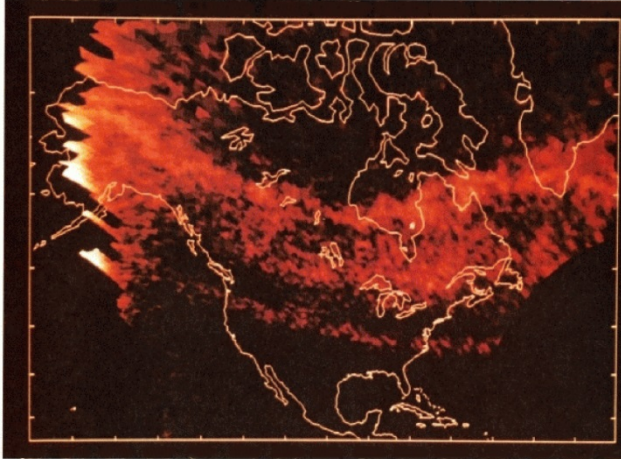
CEDAR/GEM Workshop

June 2016

Santa Fe

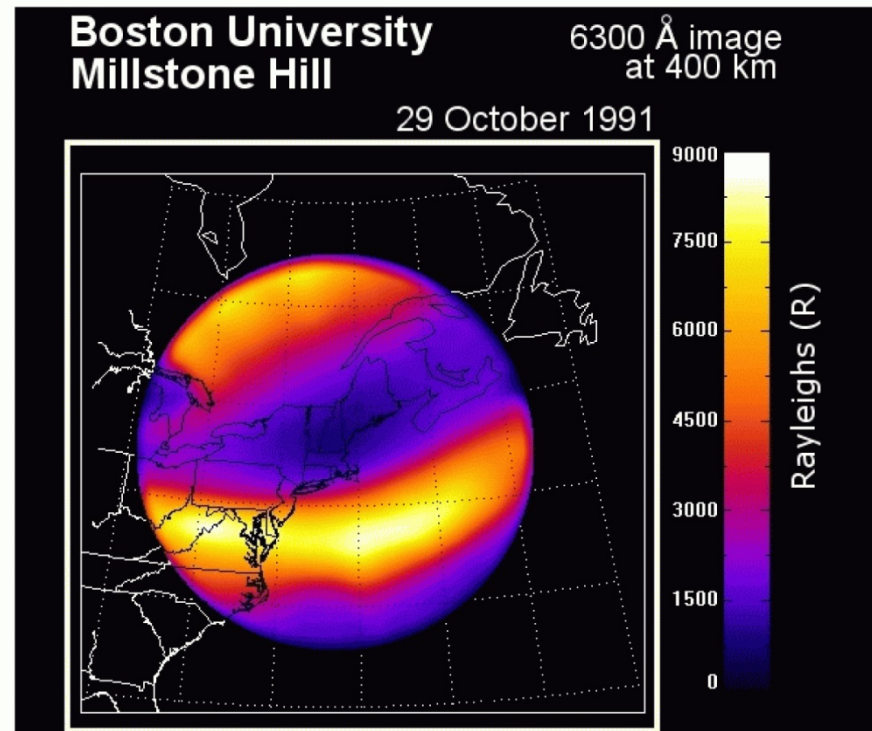
Sub-Auroral Inner Magnetosphere Science

SAR Arc from space



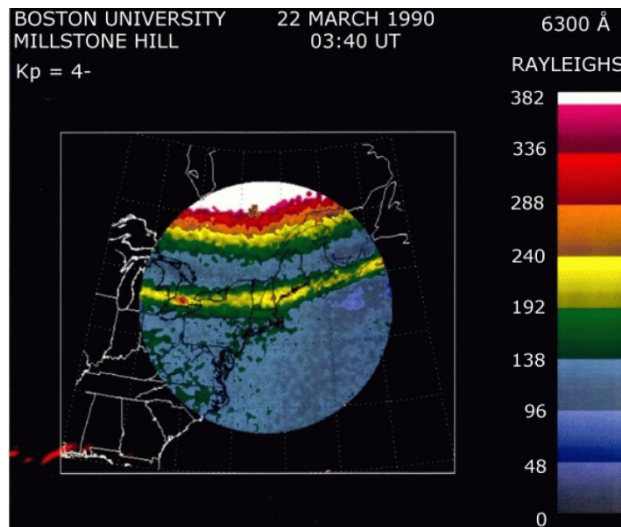
[Craven et al., 1980]

A SAR Arc of Visible Brightness



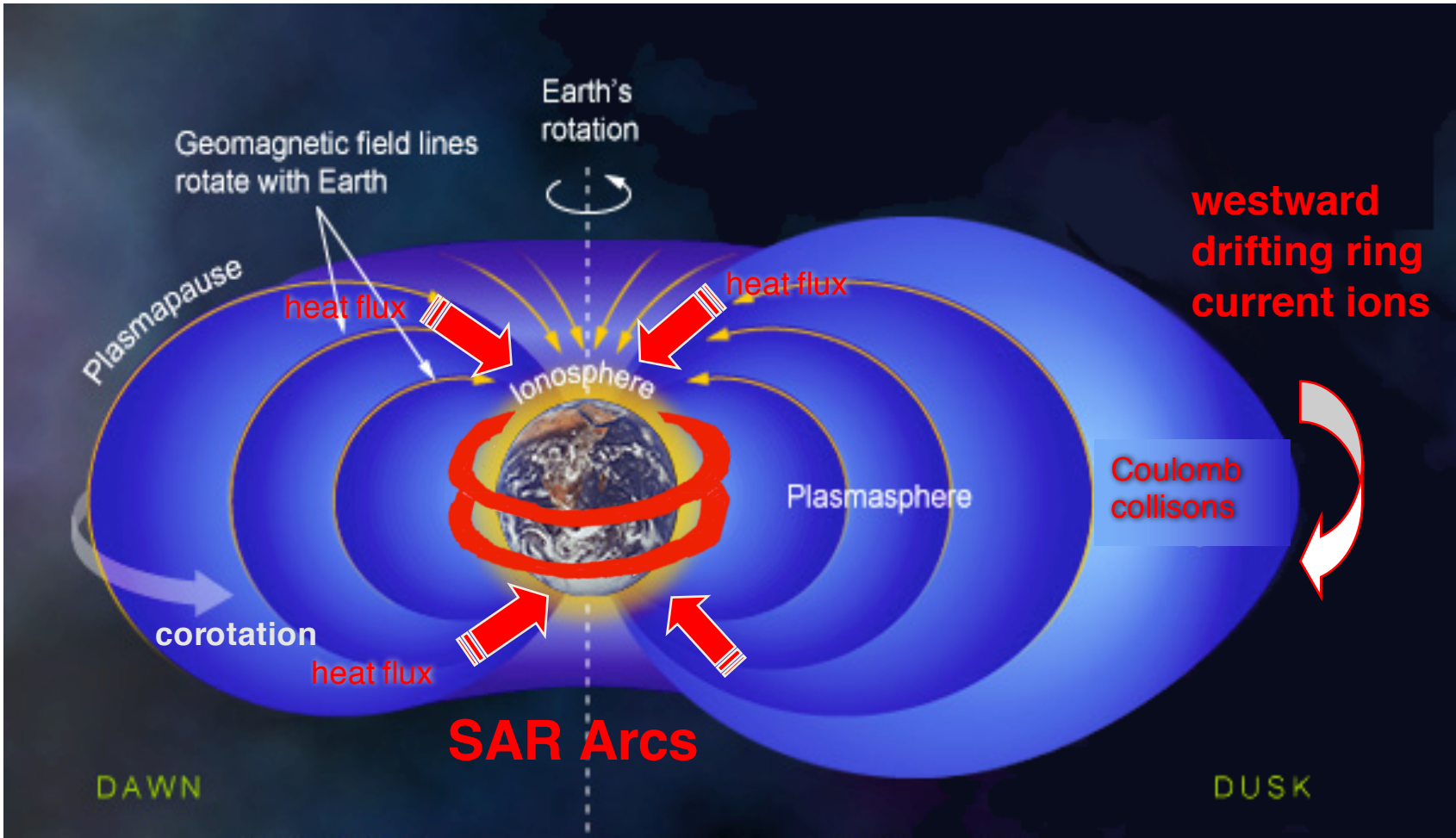
[Baumgardner et al., 2007]

SAR Arc from Millstone Hill

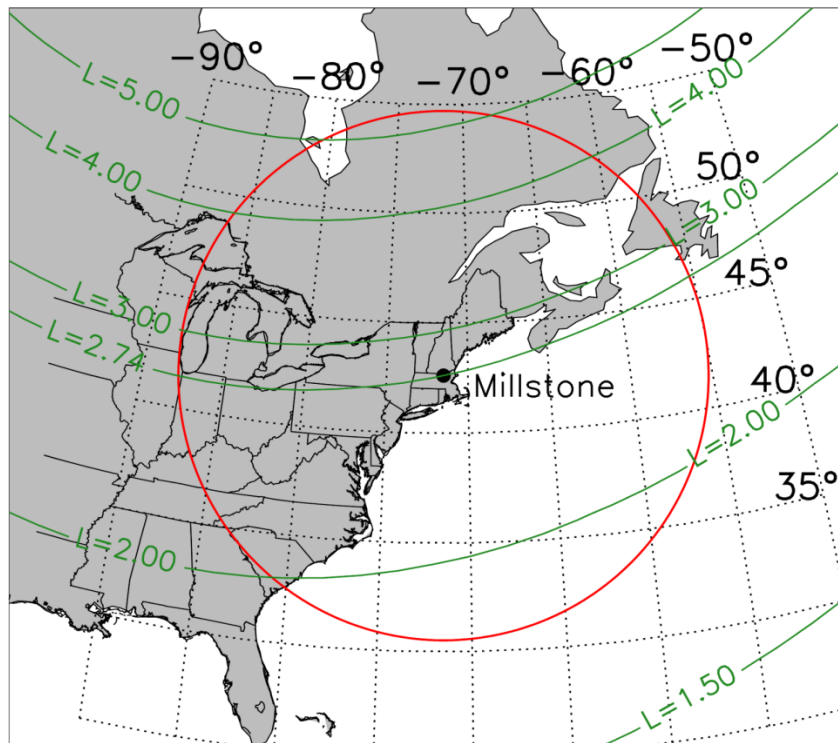


[Foster et al., 1994]

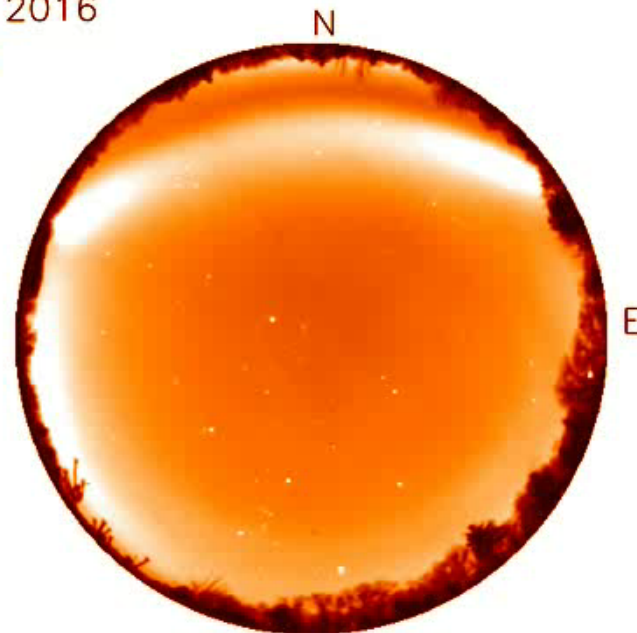
Formation of SAR Arcs



“Normal” SAR Arc

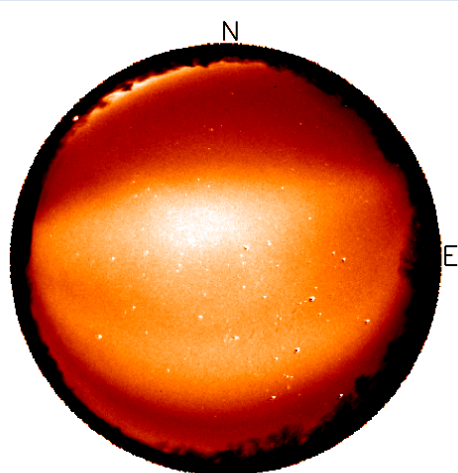


Boston University All-Sky Imager at Millstone Hill
Mar 07 2016
6300 Å

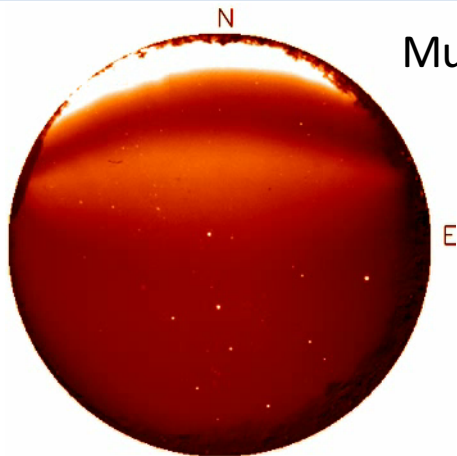


00:04:17 UT

Most SAR Arcs do not conform to “standard pattern”

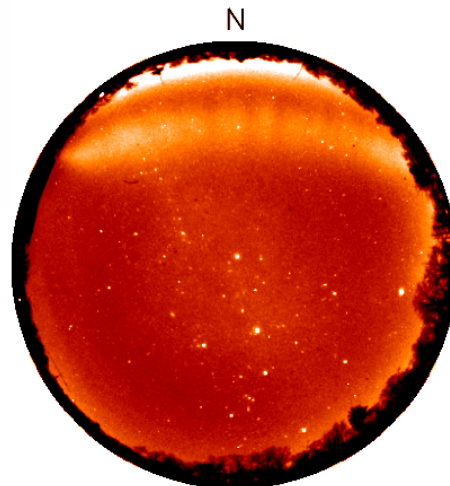


01:13:33 UT
23 Jan 2004

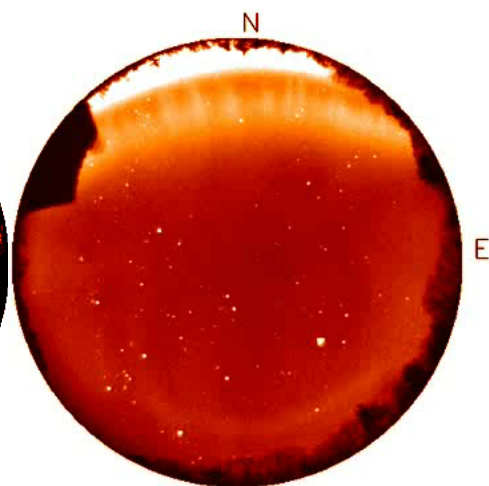


03:54:03 UT
27 Dec 2002

Multiple & Double SAR arcs

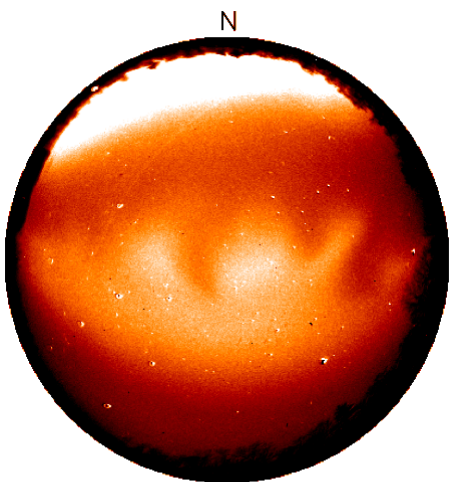


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02 Dec 2002

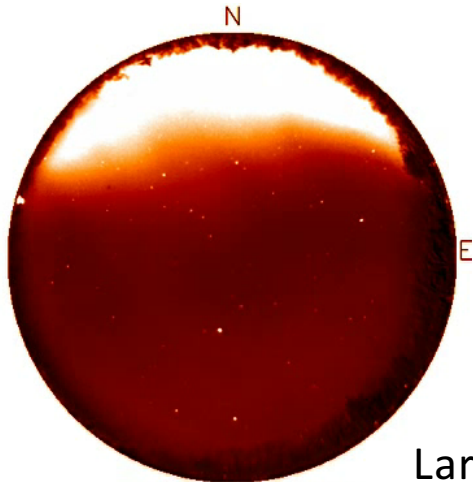


06:06:09 UT
30 Dec 2014

Small Scale Structures



05:41:29 UT
23 JAN 2004

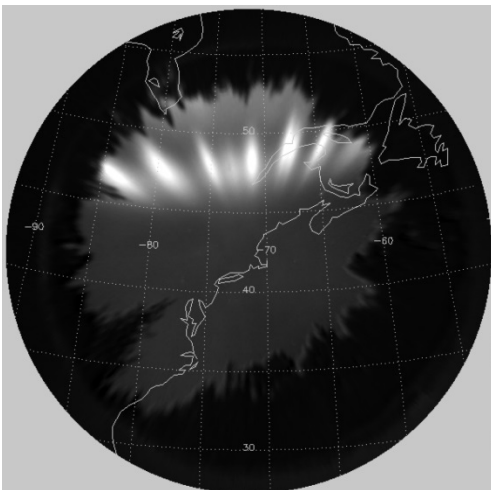


07:34:10 UT
17 Mar 2013

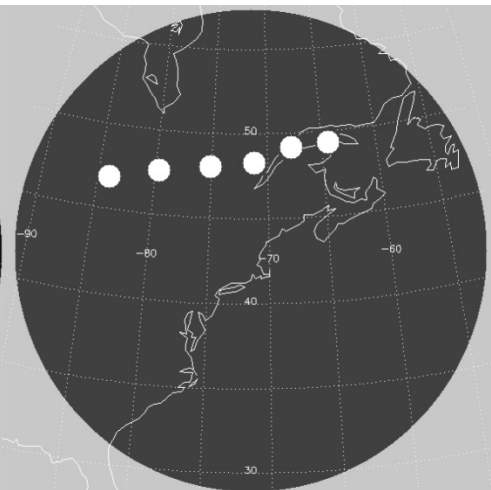
Large Scale Structures

Mapping 6300Å Brightness Patches to Geomagnetic Equatorial Plane

Streaks due to emission extended along B



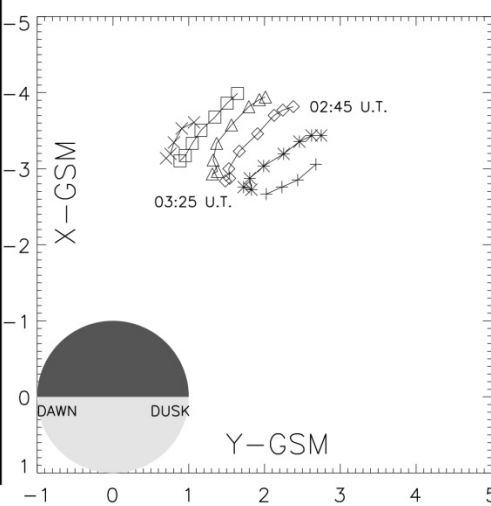
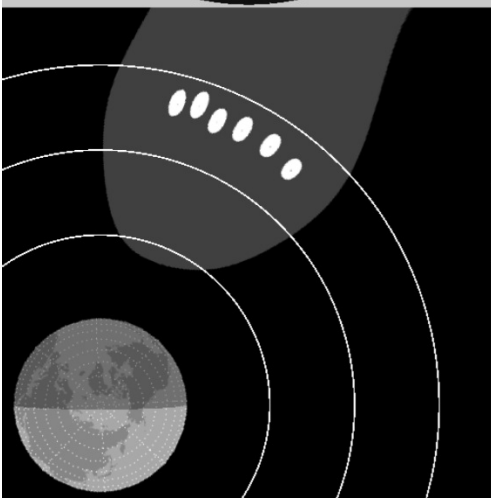
Estimate fluxtube
Circular footprints
of emission at
400km



Circular cross-sections
map to elliptical patterns
in equatorial plane:

Long/Zonal by $L^{3/2}$
Lat/Radial by $2L[L^{-3/4}]^{1/2}$

(Moser, 1970)



Radial distance vs.
time for each
emission patch

No Van Allen Probes or Themis
spacecraft in this sector on 2 Oct 2013

Summary

- Multiple Solar Cycle database of SAR arcs at Millstone Hill
[All images on BU website: <http://buimaging.com/data>]
- “Stability” term coined in 1950 prior to high performance CCD detectors
- At least 50% of all stable auroral red arcs are, in fact, variable
- Non-uniform brightness and dynamical features offer new insights into M-I Coupling in $L = 3-5$ domain
- SAR Arc Database during Van Allen Probes era at <http://buimaging.com/VanAllen/>