

The Effect of Atmospheric O and H on the Plasmasphere



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A challenge

The plasmasphere is "eroded" during a geomagnetic storm. Afterwards, it "refills."

The refilling rate is observed to decrease with increasing solar activity.

The challenge: Can SAMI3 simulations reproduce this result in agreement with observations?

Observed/simulated refilling: Krall et al., 2016, JGR SAMI3: Huba and Krall, 2013, GRL









Source of H^+ : $O^+ + H \iff O + H^+$



In order to obtain agreement with refilling rates (and pTEC) we modified the NRLMSISE-00 atmosphere by reducing atomic oxygen density

$$n_{O} = 0.7 n_{O, MSIS}$$

and by forcing the density to fall off faster with height above 600 km (in the exosphere)

$$T_{O,exo} = 0.8 T_{O,MSIS}$$
$$n_{O,exo} = n_{600} [n_{O,MSIS} / n_{600}]^{1.25}$$

(In Krall et al., 2016, JGR, $n_0 = 0.8 n_{O, MSIS}$ instead of 0.7)



Low solar activity: more O leads to more O^+ High activity: O acts as a barrier to O^+ (less O, more O^+)

Is our modified atmosphere realistic?



The overall temperature reduction isn't in the data.

This "temperature" effect can come from the Burnside Factor

$$\nu_{O-O^+} = F_B 2.4 \times 10^{-11} n_O T_r^{1/2}$$

SAMI3: $F_B = 1.3$
Salah [1993]: $F_B = 1.7$
Here: $F_B = 0.7$







Refilling rates are sensitive to the O density.

O (and probably H) densities vary on a time scale of days.

Deviations from empirical models can be significant.

We need a better model exosphere, where atoms go ballistic.

We need an experiment to determine the Burnside factor.

What about H? Preliminary results show that doubling the MSIS H density nearly doubles the refilling rate.





Emmert, J. T., S. E. McDonald, D. P. Drob, R. R. Meier, J. L. Lean, and J. M. Picone (2014), Attribution of interminima changes in the global thermosphere and ionosphere, JGR, doi:407 10.1002/2013JA019484

Huba, J. D., and J. Krall (2013), Modeling the plasmasphere with SAMI3, GRL, doi:10.1029/2012GL054300

Krall, J., J. D. Huba, V. K. Jordanova, R. E. Denton, T. Carranza, and M. B. Moldwin (2016), Measurement and modeling of the refilling plasmasphere during 2001, JGR, doi:10.1002/2015JA022126

Krall, J., J.T. Emmert, F. Sassi, S.E. McDonald, and J.D. Huba, Day-to-day variability in the thermosphere and its impact on plasmasphere refilling, JGR, in press

Krall, J. and J. D. Huba, The plasmasphere electron content paradox, submitted to JGR





NRL SAMI3 Ionosphere/Plasmasphere Model

- Magnetic field: HGRF-like Non-tilted dipole
- Interhemispheric
- Nonorthogonal, nonuniform fixed grid
- Seven (7) ion species (all ions are equal): H^+ , He^+ , N^+ , O^+ , N_2^+ , NO^+ , and O_2^+
 - Solve continuity and momentum for all 7 species
 - Solve temperature for H^+ , He^+ , O^+ , and e^-
- Plasma motion
 - $\mathbf{E} \times \mathbf{B}$ drift perpendicular to \mathbf{B}
 - Ion inertia included parallel to ${f B}$
- Neutral species: NRLMSISE00 and HWM93
- Chemistry: 21 reactions + recombination
- Photoionization: Daytime (EUVAC) and nighttime

SAMI3 is coupled to a magnetosphere potential model and a thermosphere model.

[Huba et al., JGR, 2000; Huba and Joyce, GRL, 2010; Huba and Krall, GRL, 2013]

HWM14





A challenge

The challenge: Can SAMI3 simulations reproduce...

- ... the observed decrease in the rate of post-storm plasmasphere refilling with increasing solar activity?
- 2. ... the observed increase in the plasmasphere contribution to "total electron content" (pTEC) in increasing solar activity?





pTEC increases with solar activity.



pTEC, is TEC contribution between JASON altitude (1340 km) and GPS altitude (20,200 km), increases with solar activity. Measured pTEC result based on data from 2002-2009.

Discussion





We need a better model exosphere, where atoms go ballistic.

We need an experiment to determine the Burnside factor.

What about H? Preliminary results show that doubling the MSIS H density nearly double the refilling rate.

Satellite Drag in the Exosphere

