The MST-ISR-EEJ mode at Jicamarca, fine structure and interferometry with mesospheric echoes

CLEMS

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JICAMARCA RADIO OBSERVATORY AT THE FOREFRONT OF EQUATORIAL GEOSPACE KNOWLEDGE

MST-ISR-EEJ Mode



- <u>Objective</u>: Gravity wave processes and coupling in D, E, and F region
- 4 fixed beams, East <u>IB</u>, West <u>IB</u>, South, Vertical
- Big transmitters, ~1.5 MW peak
- <u>MST</u>: 64-baud CC, 150 m, 0-200 km
- <u>ISR</u>: 3-baud Barker, 15 km, 0-1000 km
- <u>EEJ</u>: 1-baud, 150 m, 0-200 km
- 4-day campaigns, day and night, ~2 TB/day
- 8 campaigns between Jan-2014 and Jan-2016
- Data server at U of Illinois (for public access)

Outline



Nighttime E region echoes (MST mode)

2014-01-07



2014-01-08



E field anomaly?

2014-05-14

JRO MST-ISR-EEJ 2014133-2014134 Beam V



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morning ascent

2014-01-15



"elegant" descent

2014-10-01





lower "tidal" layer

//.

2015-01-06





LT

11.

meteor shower Quadrantids

Nighttime E region echoes - Summary

- Fundamental studies on nighttime electrojet, plasma waves and instabilities
- <u>Day-to-day variability</u>, especially during prereversal enhancement and spread-F development is less understood
- Lack of complementary measurements of
 - neutral metals and metal ions;
 - including sporadic and tidal metal layers
 - plasma gradient
 - neutral winds and temperatures (some by meteors)
 - electric fields

150-km echoes – bottom line



150-km echoes – bottom line



150-km echoes – layer fine structure

JR0 2014133 MST 150km



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150-km echoes – layer fine structure



JR0 2014274 MST 150km

150-km echoes – fast pulsations





10-15 minutes of data with <u>4 second</u> integrations high-resolution RTI reveal power modulations with periods of 20-30 seconds

150-km echoes – fast pulsations



Alfvén wave propagation along magnetic fieldlines?

Patra and Rao, 2007/6

150-km echoes - Summary

- <u>Day-to-day and seasonal variability</u> of "necklace"
 relation to local electron density profile
- <u>Fine structure in layers is very similar each day</u>
 - "forbidden" and enhanced layers
 - point to fundamental plasma resonances
- Very short period pulsations
 - possibly due to waves along the magnetic field lines
- Gravity wave modulation due to vertically propagating waves
 - similar to VIPER dynasonde results (P. Reyes; Negrea et al./Wallops Island)

Mesospheric echoes 2011130 - 10 May 2011 - RTI



Mesospheric echoes 2011131 - 11 May 2011 - RTI



Gravity wave – fine structure interactions

Direct Numerical Simulation (DNS), Fritts et al., CEDAR, 2012



Mesospheric echoes – Vertical velocity



Mesospheric echoes – Spectral width

(from generalized Gaussian fit)

FWHM Map: Channel 0 - 2011-05-10 Width (m/s)



Mesospheric echoes – interferometry



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Self-spectra, 16 coherent and 4 incoherent integrations, 256-point FFT, ~22 sec spectra

normalized self-spectra, -10 ... 10 m/s, Channel 0, SNR as color scale -10 ... 20 dB ²⁴



Cross-spectra, phase of normalized cross-spectra for coherence > 0.6



Cross-spectra; example for phase spectra in turbulent layer

complex variation of phase spectra, consistent horizontal bulk motion

Mesospheric echoes - interferometry

- Basic interferometry suggests evidence for individual, (partially) reflecting patches ("glints") in the radar volume
- Phase progression yields consistent horizontal trace velocity for all interferometric baselines



Mesospheric echoes - Summary

- First observations of <u>localized scatterers</u> in Jicamarca mesospheric echoes
- Example suggests a deep, convective and turbulent layer capped by a stable, non-turbulent layer
- Physical mechanism or plasma processes that cause glints is still unclear (similar to PMSE?)
- Interferometry and basic imaging may help understand details of gravity wave breaking and instabilities in the equatorial mesosphere

Thank you!



Example of KHI train on 1/8/2014 ~1 hour, ~77 km

Equatorial electron density profiles, 70 - 80 km 3-meter fine structure remains elusive



Goldberg et al., 1997; Lehmacher et al., 2006; Smith and Klaus, 1976; Das et al., 2009 ³⁰

Mesospheric instability layers (mid and high latitude)



Lehmacher and Lübken, 1995; Lehmacher et al., 2011; Collins et al., 2011