

Python Science Analysis Toolkit and Data Interpolating Empirical Orthogonal Functions

Pysat and DINEOFs:
A System for System Science

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Instrument Object

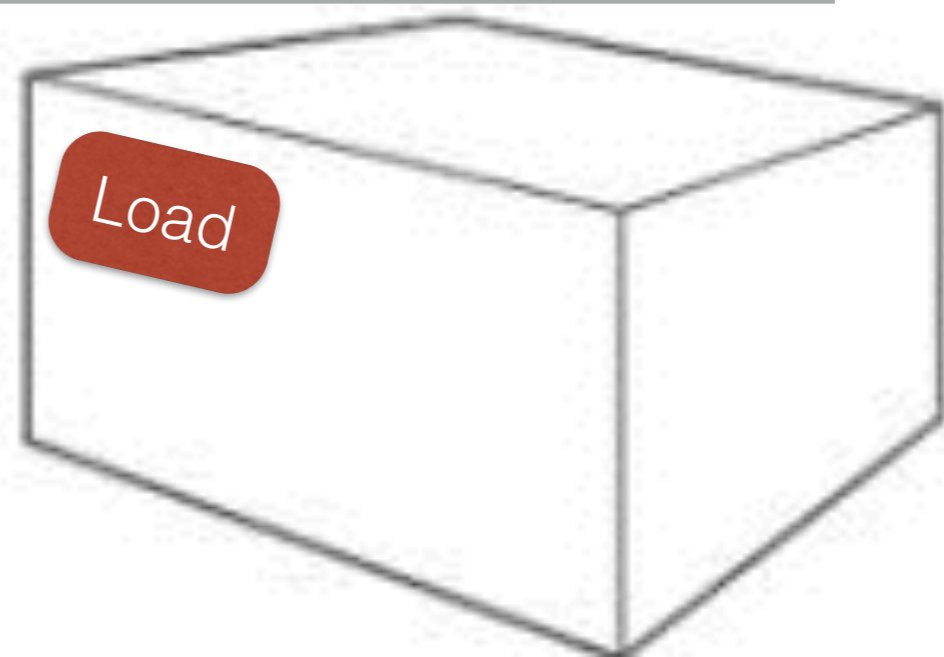
```
ivm = pysat.Instrument(platform='cnofs',  
                        name='ivm')
```

```
vefi = pysat.Instrument(platform='cnofs',  
                        name='vefi',  
                        tag='dc_b')
```

```
cosmic = pysat.Instrument(platform='cosmic2013',  
                          name='gps',  
                          tag='ionprf',  
                          clean_level='clean')
```

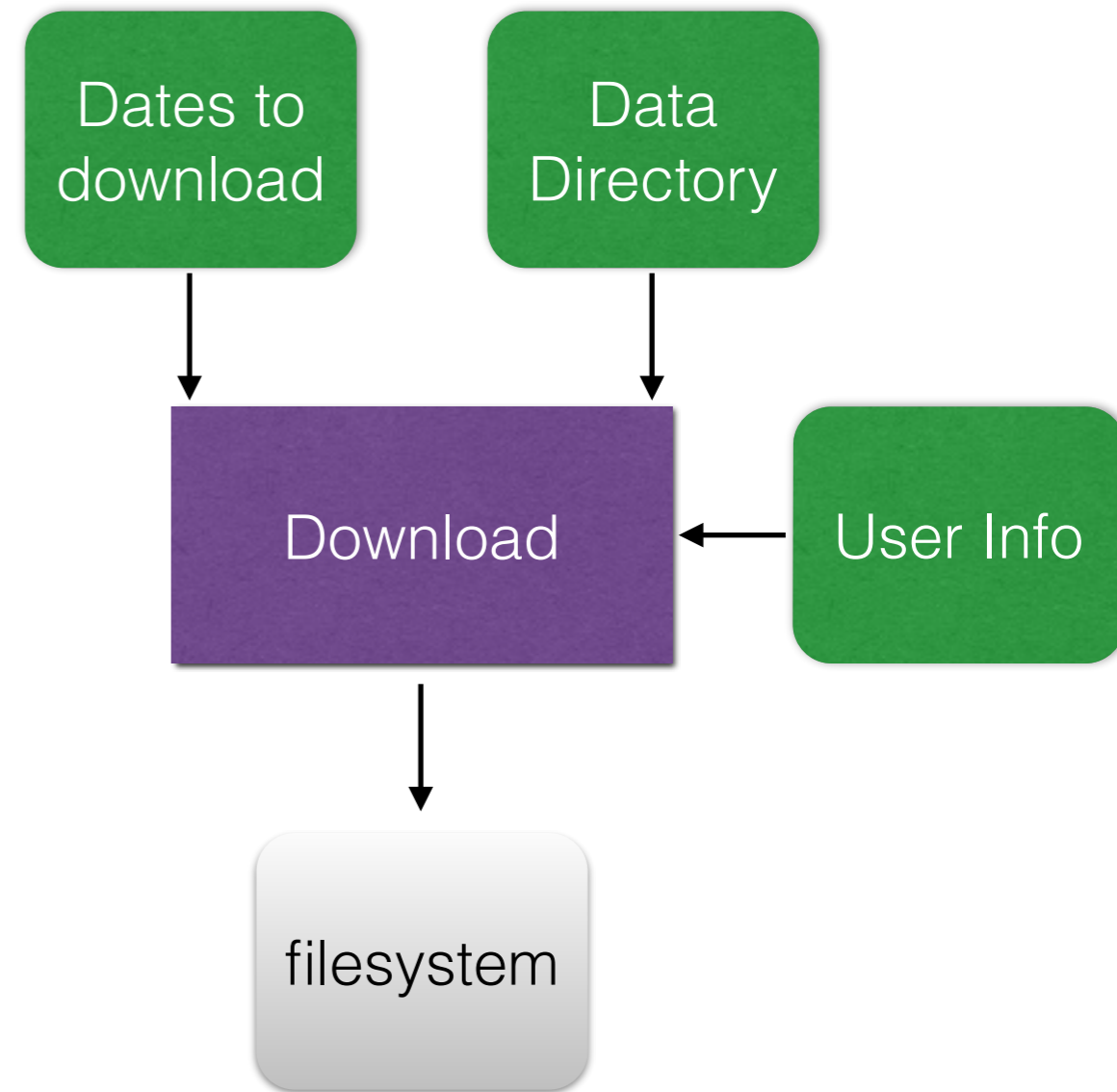
```
darn = pysat.Instrument(platform='superdarn',  
                        name='grdex',  
                        tag='north')
```

```
ivm.load(2012,1)  
vefi.load(2012,1)  
cosmic.load(2012,1)  
darn.load(2012,1)
```



Supports Data Downloads

```
In [53]: start = pysat.datetime(2010,1,1)
In [54]: stop = pysat.datetime(2010,1,3)
In [55]: ivm.download(start, stop)
Downloading file for 01/01/10
Downloading file for 01/02/10
Downloading file for 01/03/10
Updating pysat file list
In [56]: vefi.download(start, stop)
Downloading file for 01/01/10
Downloading file for 01/02/10
Downloading file for 01/03/10
Updating pysat file list
```



Data Organization

User Set Dir/platform/name/tag/

Legend

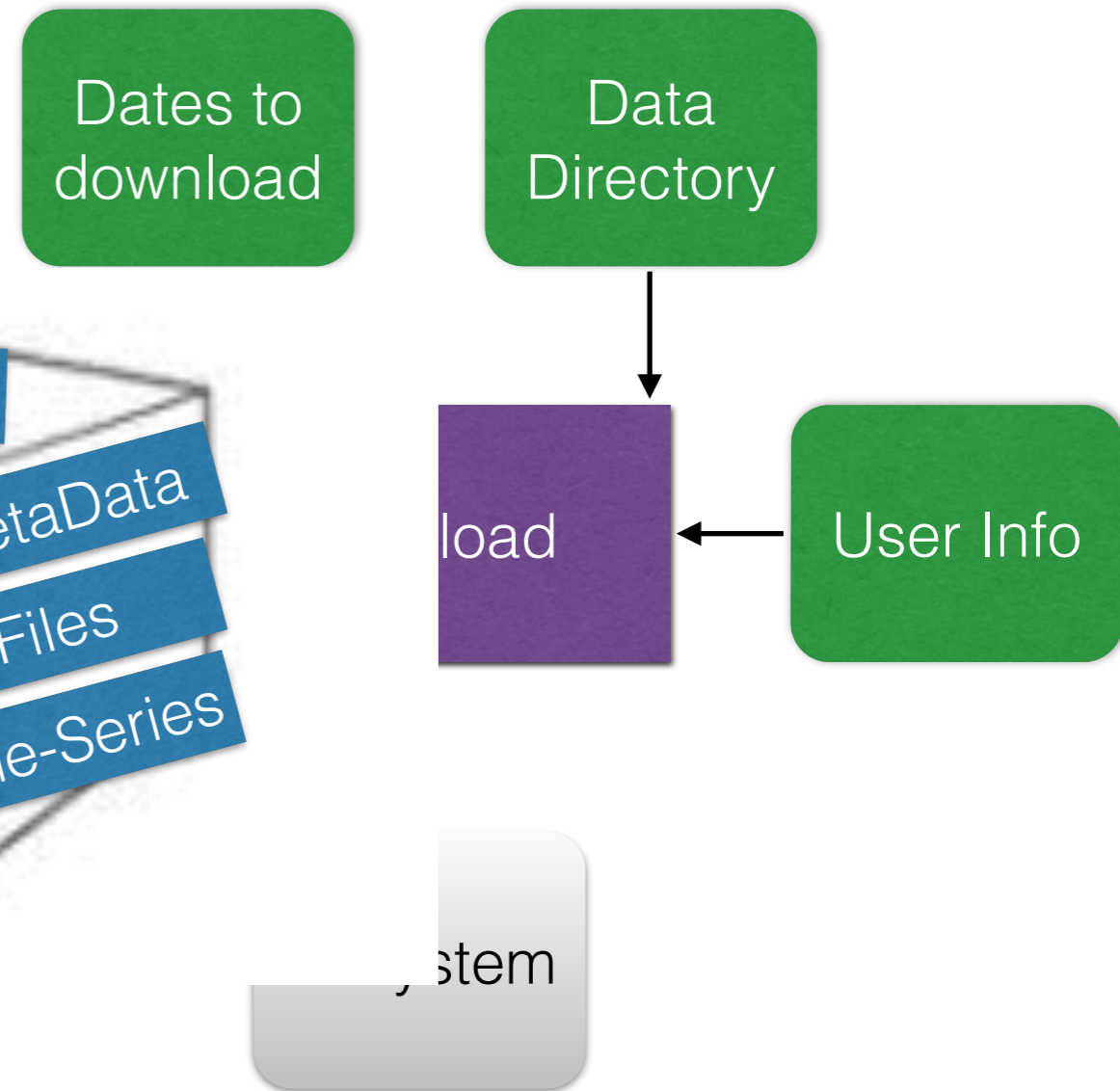
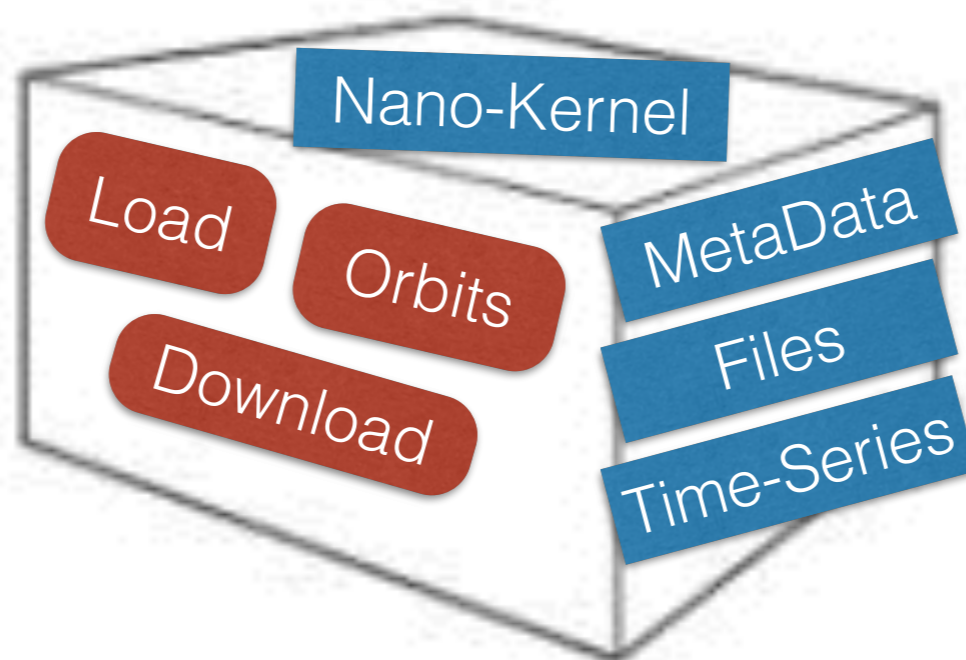
Instrument
Specific
Function

pysat
Function

pysat
Information

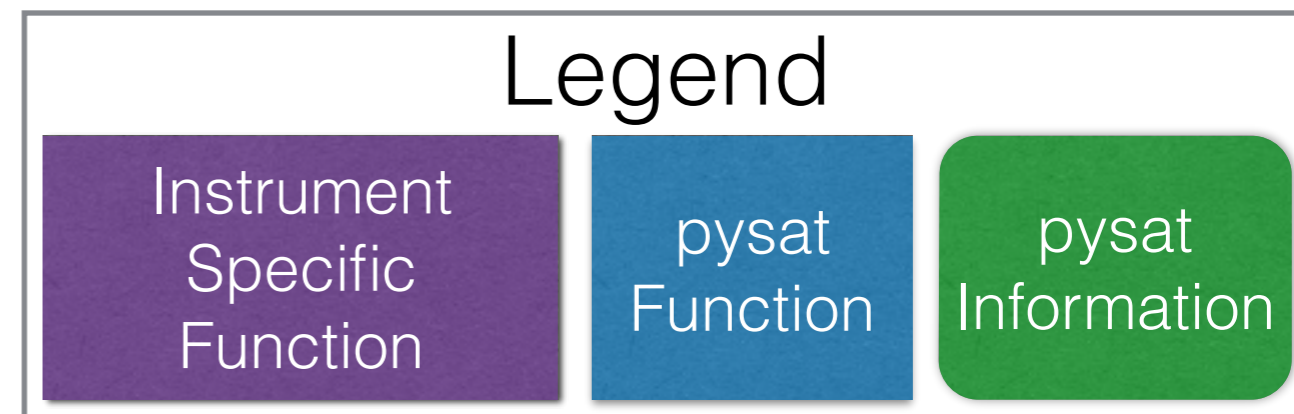
Supports Data Downloads

```
In [53]: start = pysat.datetime(2010,1,1)
In [54]: stop :
In [55]: ivm.d
Downloading fil
Downloading fil
Downloading fil
Updating pysat
In [56]: vefi.c
Downloading fil
Downloading fil
Downloading fil
Updating pysat FILE LIST
```



Data Organization

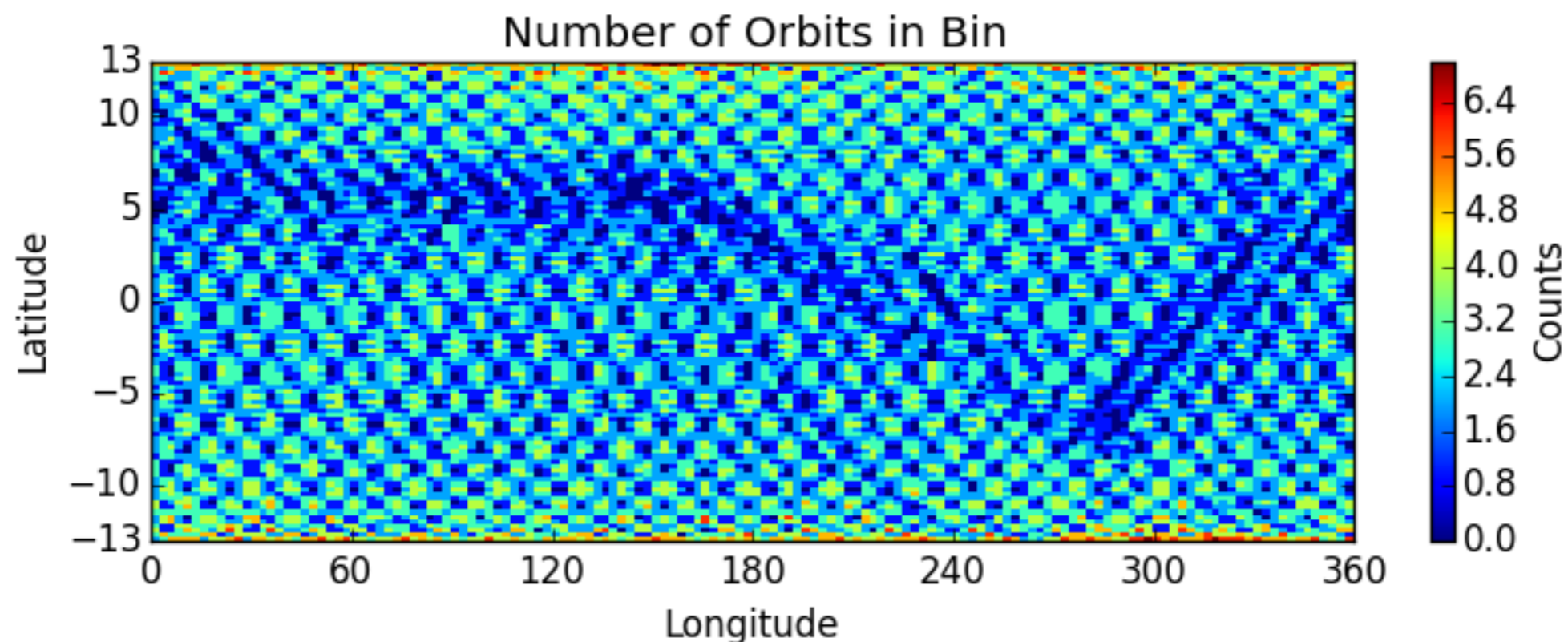
User Set Dir/platform/name/tag/



Instrument Independent Analysis

Attach a queue of functions to be applied whenever data is loaded, set and forget. (nano-kernel)

```
vefi = pysat.Instrument(platform='cnofs', name='vefi', tag='dc_b',
                        clean_level=None)
# define function to remove torque-rod firings
def filter_vefi(inst):
    idx, = np.where(vefi['B_flag']==0)
    vefi.data = vefi.data.iloc[idx]
    return
vefi.custom.add(filter_vefi, 'modify')
ans = pysat.ssnl.occure_prob.by_orbit2D(vefi, [0,360,144],
    'longitude', [-13,13,104], 'latitude', ['dB_mer'], [0.])
# plot commands go here
```

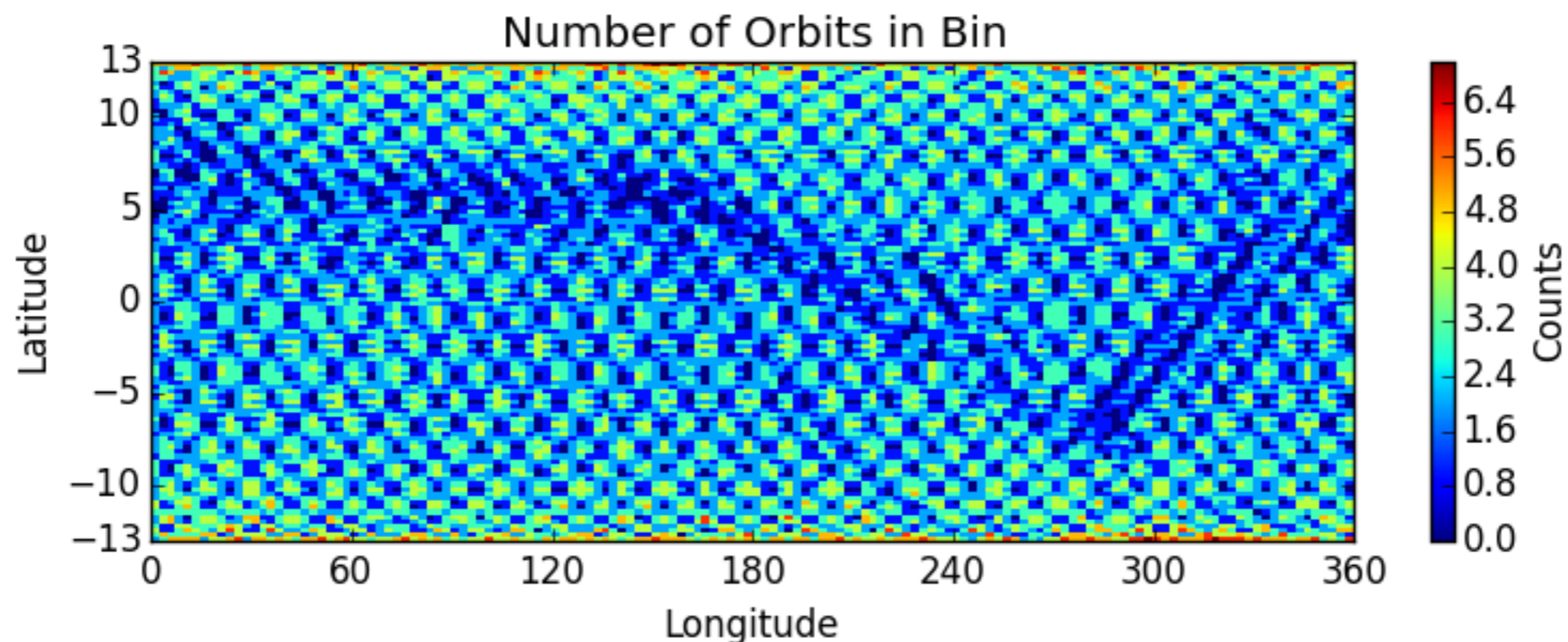


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Instrument Independent
Analysis Routine



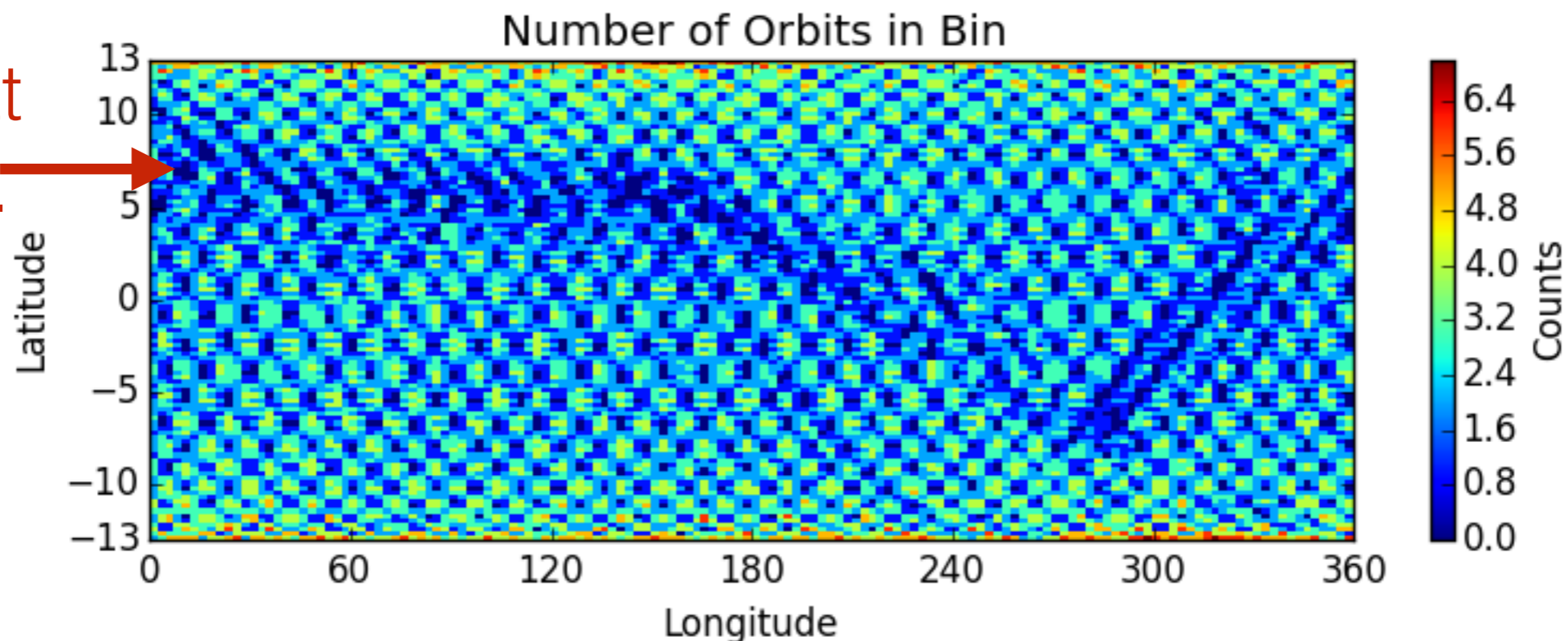
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```

This filter function modifies the vefi data available in the probability routine

Firings at Mag Eq.



Averaging Drifts and Profiles

```
# instantiate IVM Object
ivm = pysat.Instrument('cnofs', 'ivm', '', 'clean')
ivm.custom.add(restrictMLAT, 'modify', maxMLAT=25.)
ivm.bounds(startDate, stopDate)
ivmResults = pysat.ssn1.median2D(ivm, [0, 360, 24], 'apex_long',
                                  [0, 24, 24], 'mlt', ['iv_mer'])

# create COSMIC instrument object
cosmic = pysat.Instrument('cosmic2013', 'gps', 'ionprf',
                          'clean', altitude_bin=3)
# apply custom functions to all data that is loaded through cosmic
cosmic.custom.add(addApexLong, 'add')
cosmic.custom.add(filterMLAT, 'modify', mlatRange=(0., 10.) )
cosmic.custom.add(addlogNm, 'add')
cosmic.custom.add(addTopsideScaleHeight, 'add')

# do an average of multiple COSMIC data products from startDate
through stopDate
cosmic.bounds(startDate, stopDate)
cosmicResults = pysat.ssn1.median2D(cosmic, [0, 360, 24], 'apex_long',
                                     [0, 24, 24], 'edmaxlct', ['profiles', 'edmaxalt', 'lognm', 'thf2'])
# plot commands
```

COSMIC profiles are aligned by altitude before averaging

Averaging Drifts and Profiles

Instrument Independent

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# instantiate IVM Object
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Analysis Routine

Instrument Independent

Analysis Routine

COSMIC profiles are aligned by altitude before averaging

Averaging Drifts and Profiles

Instrument Independent

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```

Analysis Routine

Custom functions adds magnetic coordinates, filters data, and adds log density and topside scale height. All the processing needed to compare with IVM.

```
# apply custom functions to all data that is loaded through cosmic
```

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cosmic.custom.add(addApexLong, 'add')
```

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cosmic.custom.add(filterMLAT, 'modify', mlatRange=(0., 10.) )
```

```
cosmic.custom.add(addlogNm, 'add')
```

```
cosmic.custom.add(addTopsideScaleHeight, 'add')
```

```
# do an average of multiple COSMIC data products from startDate  
through stopDate
```

```
cosmic.bounds(startDate, stopDate)
```

```
cosmicResults = pysat.ssn1.median2D(cosmic, [0, 360, 24], 'apex_long',  
[0, 24, 24], 'edmaxlct', ['profiles', 'edmaxalt', 'lognm', 'thf2'])
```

```
# plot commands
```

Instrument Independent

Analysis Routine

COSMIC profiles are aligned by altitude before averaging

Averaging Drifts and Profiles

Instrument Independent

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ivm.bounds(startDate, stopDate)
ivmResults = pysat.ssn1.median2D(ivm, [0, 360, 24], 'apex_long',
                                [0, 24, 24], 'mlt', ['iv_mer'])
                                1D
                                Analysis Routine

# create CODMIC instrument object
cosmic = pysat.Instrument('cosmic2013', 'gps', 'ionprf',
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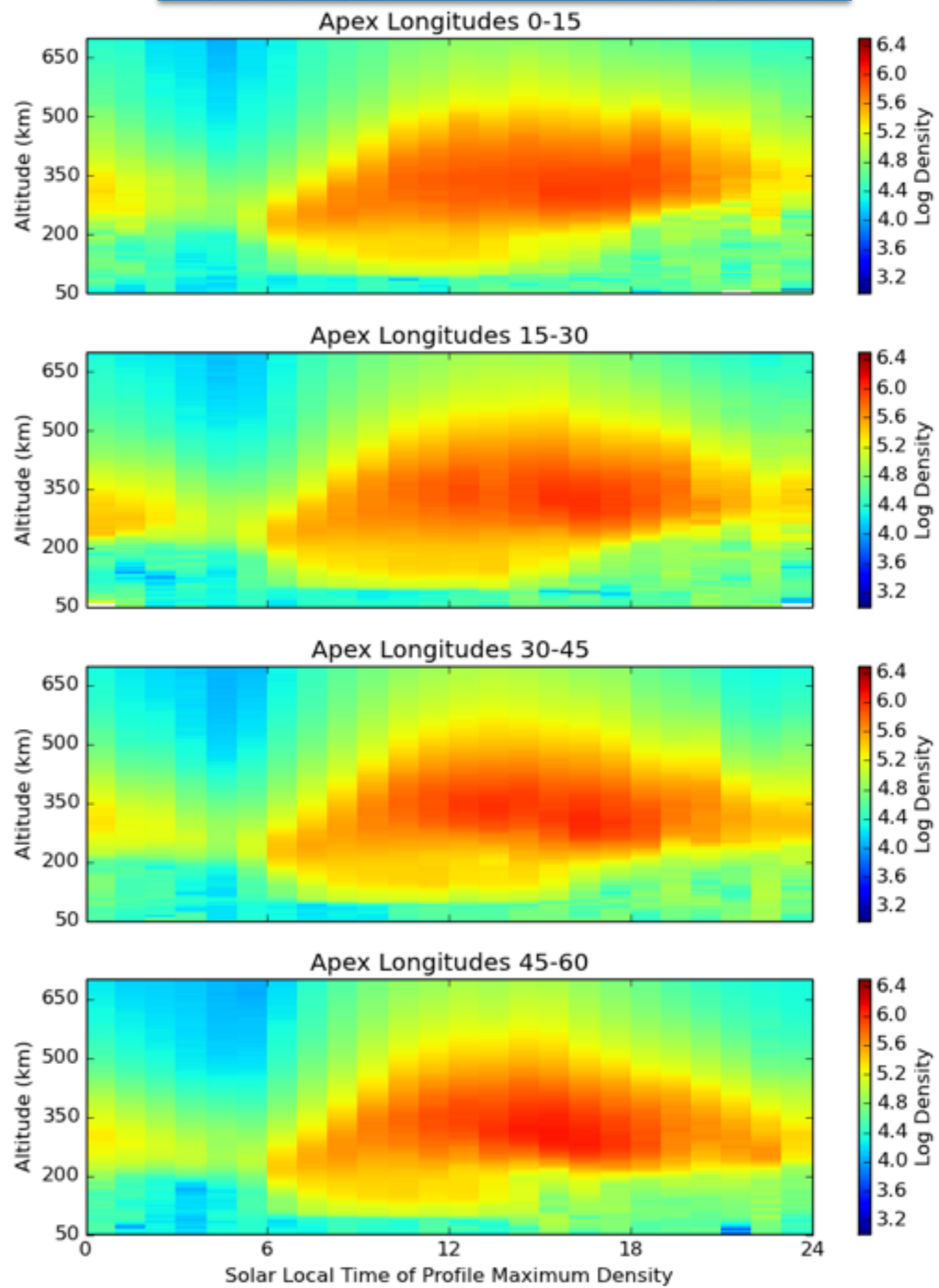
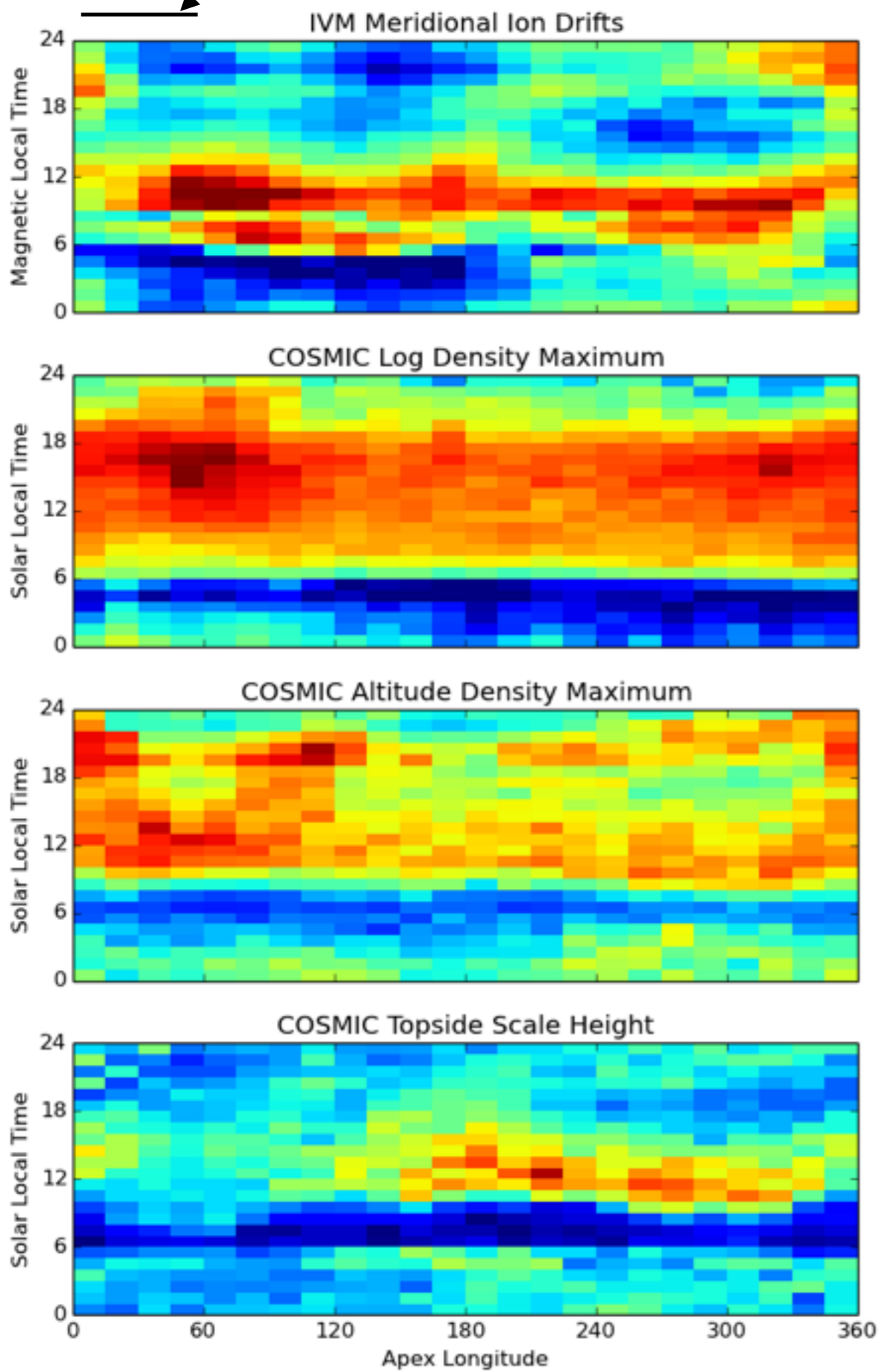
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                                   [0, 24, 24], 'edmaxlct', ['profiles', 'edmaxalt', 'lognm', 'thf2'])
                                   3D
                                   1D
                                   1D
                                   1D
                                   Analysis Routine
# plot commands
```

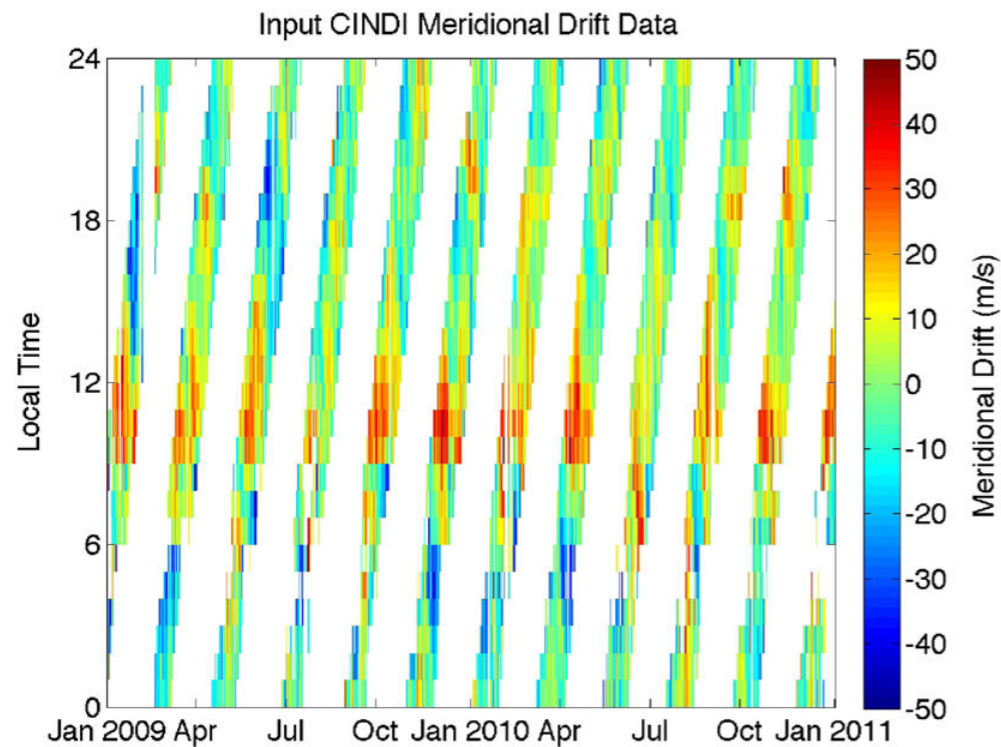
COSMIC profiles are aligned by altitude before averaging

Upward to downward drifts

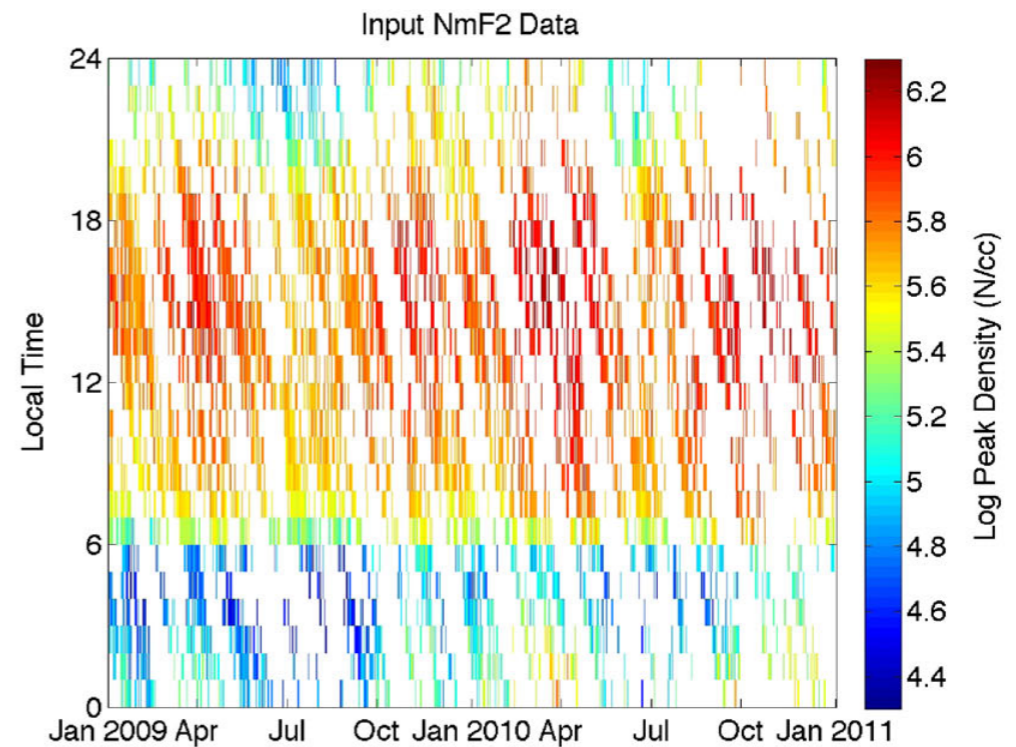
Influence of drifts observed in COSMIC density profiles



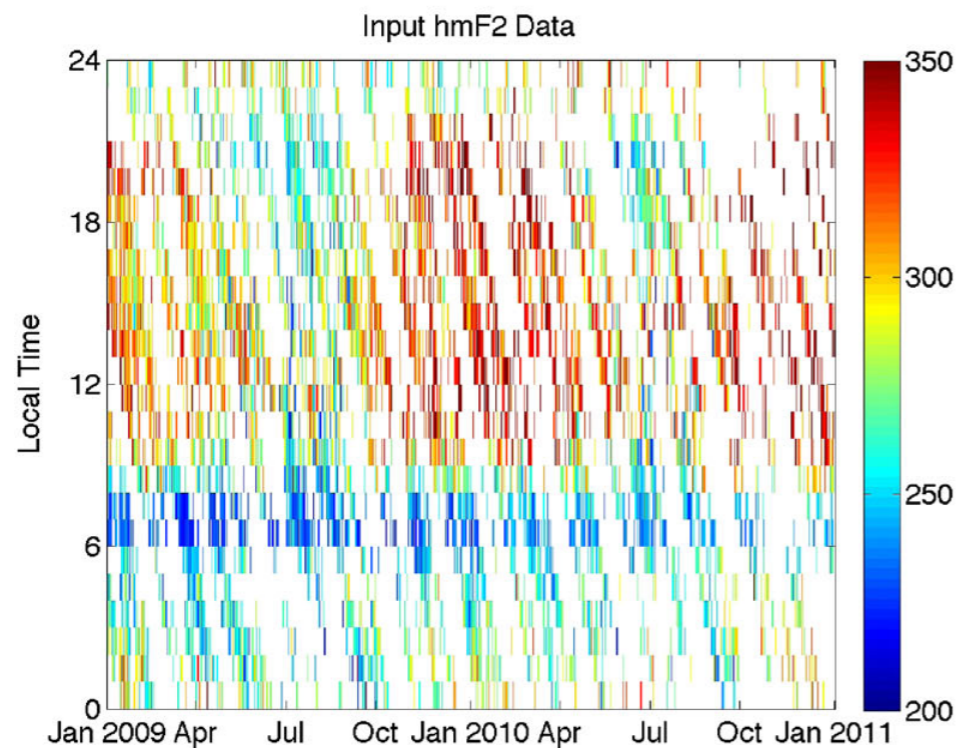
Equatorial Space Weather



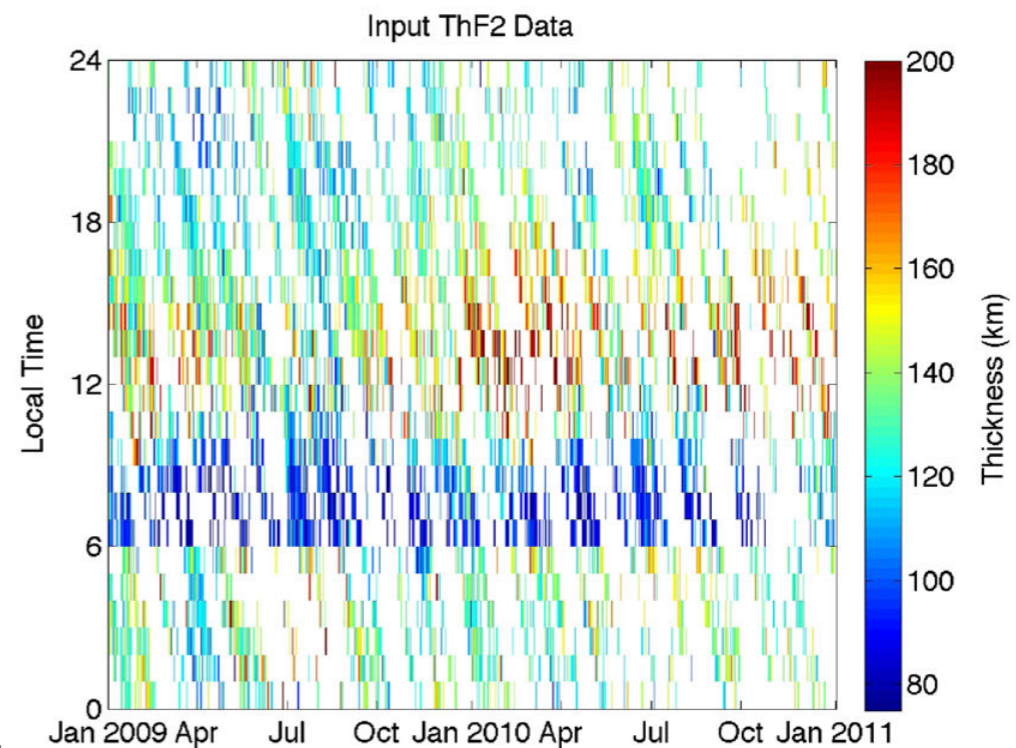
(a) CINDI Meridional Drifts



(b) COSMIC Peak Density

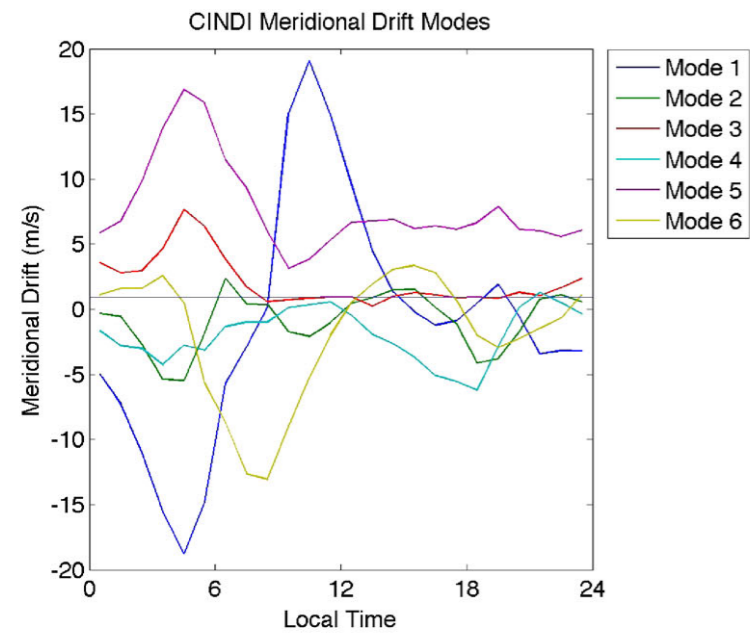


(c) COSMIC peak density height

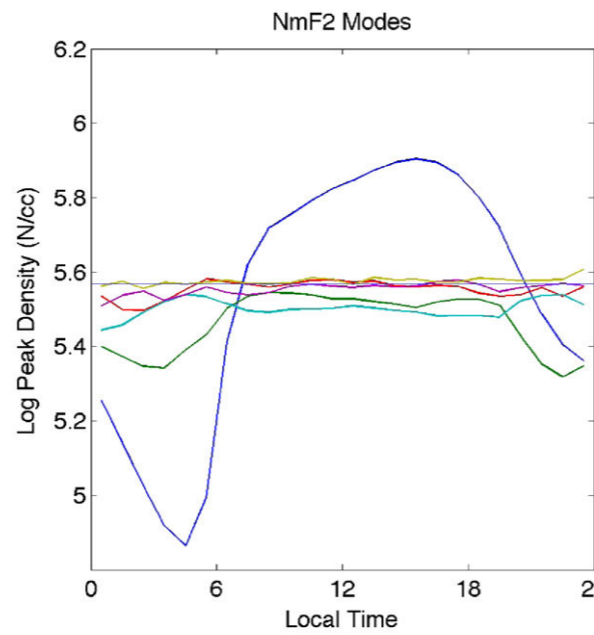


(d) COSMIC thickness

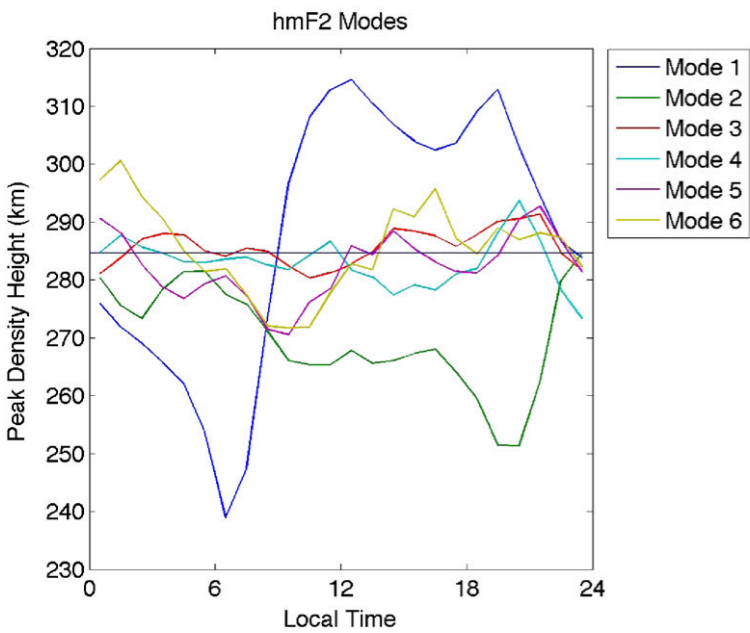
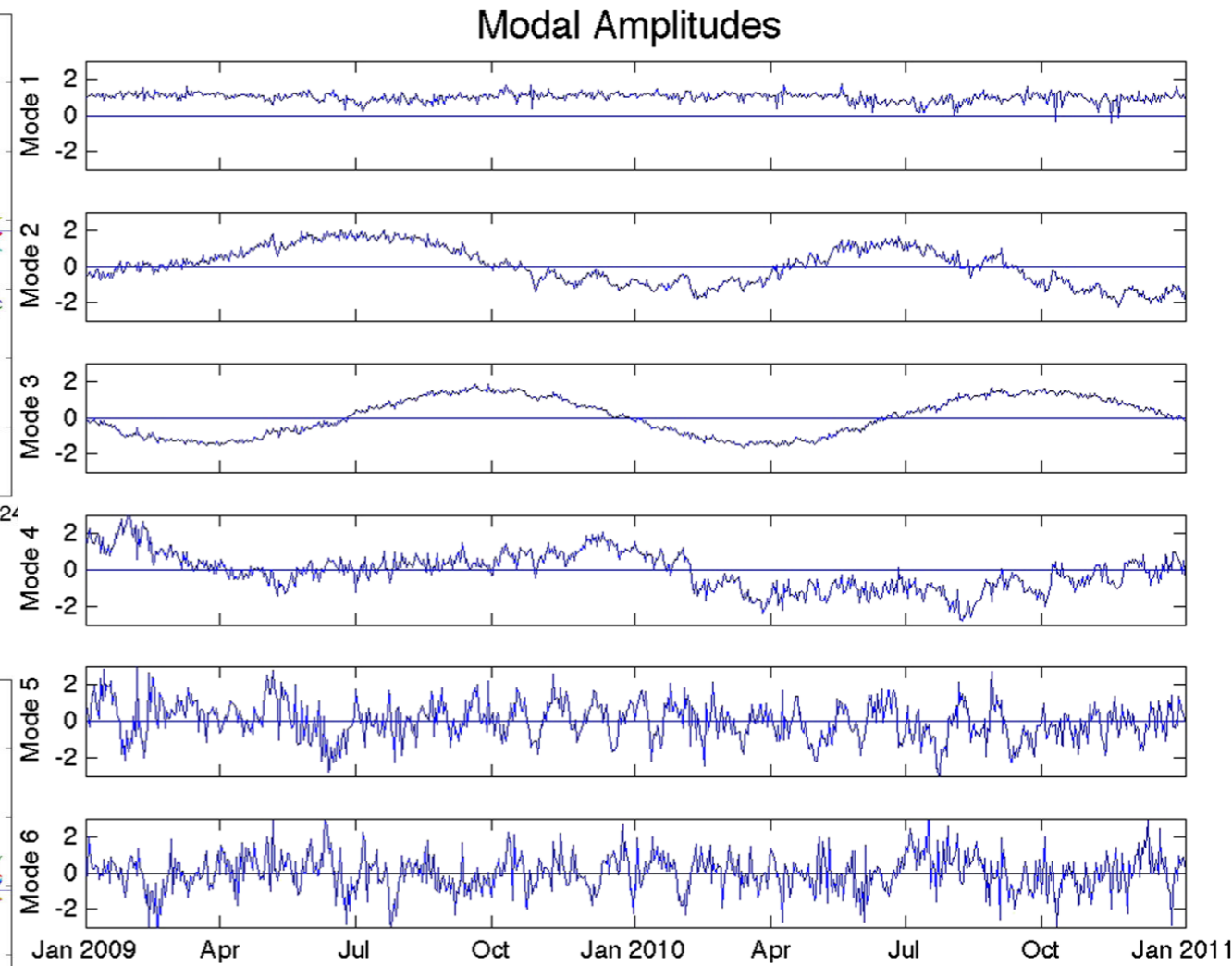
DINEOF Modes



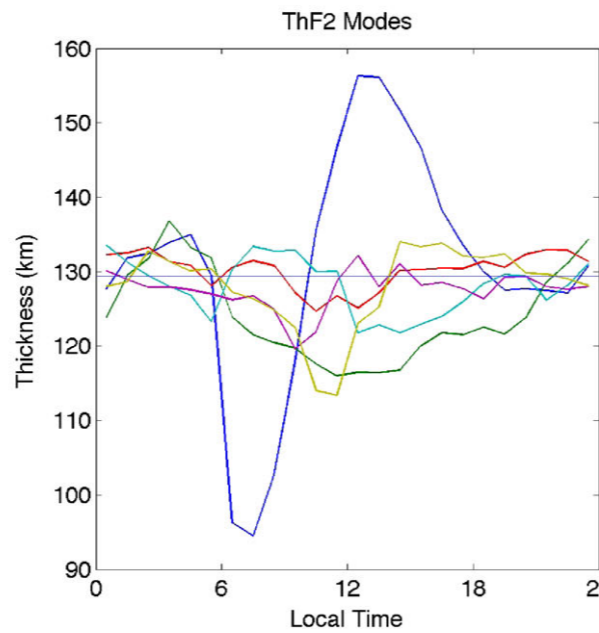
(a) CINDI Meridional Drifts



(b) COSMIC Peak Density

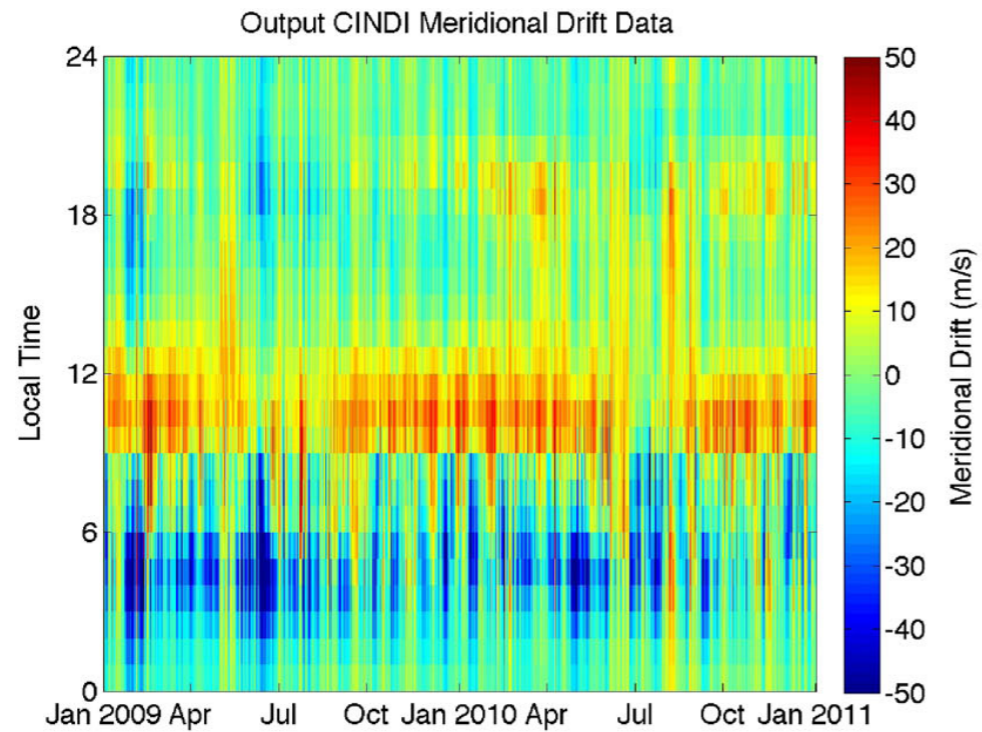


(c) COSMIC peak density height

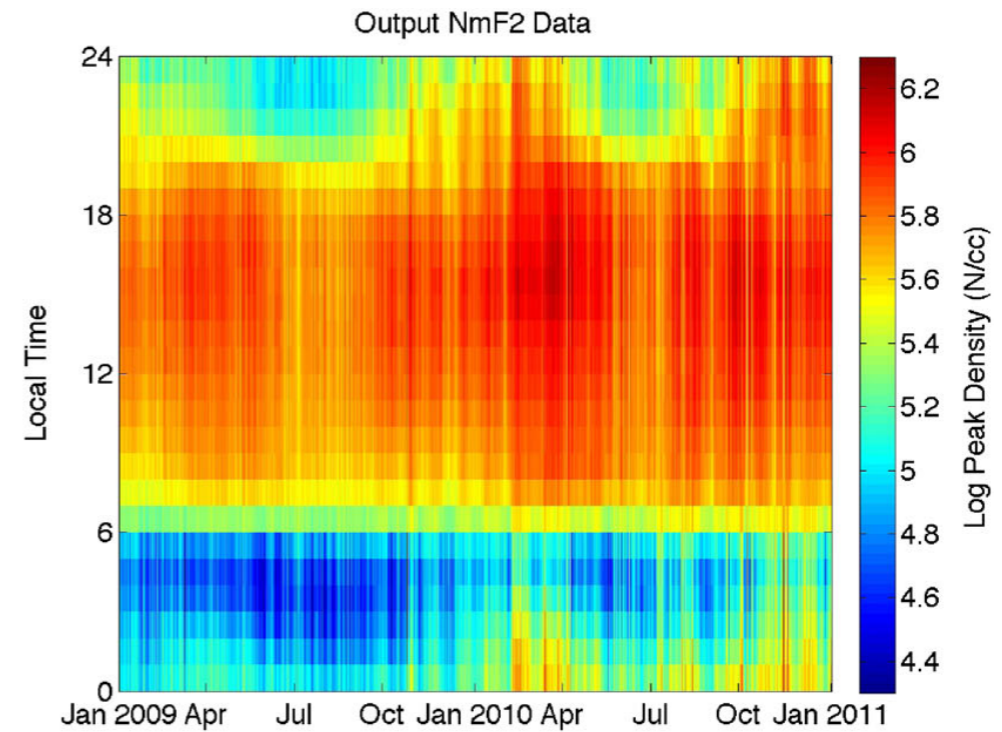


(d) COSMIC thickness

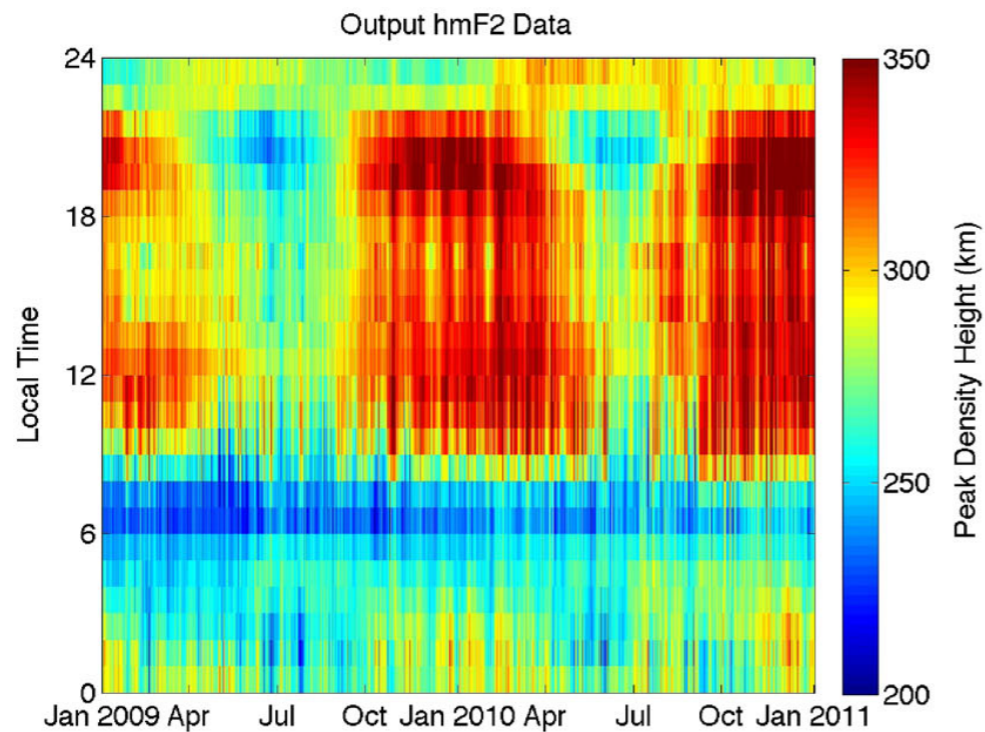
DINEOF Results



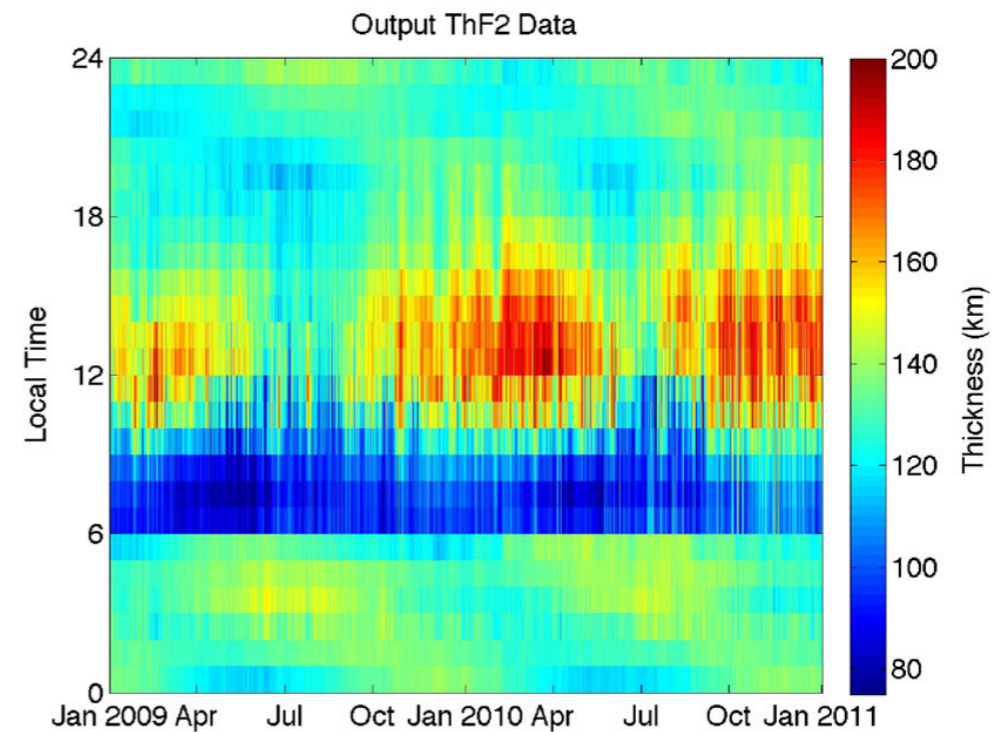
(a) CINDI Meridional Drifts



(b) COSMIC Peak Density



(c) COSMIC peak density height



(d) COSMIC thickness

System for System Science

Satellite based instruments

Models

Ground based instruments

pysat

Space Weather Drivers

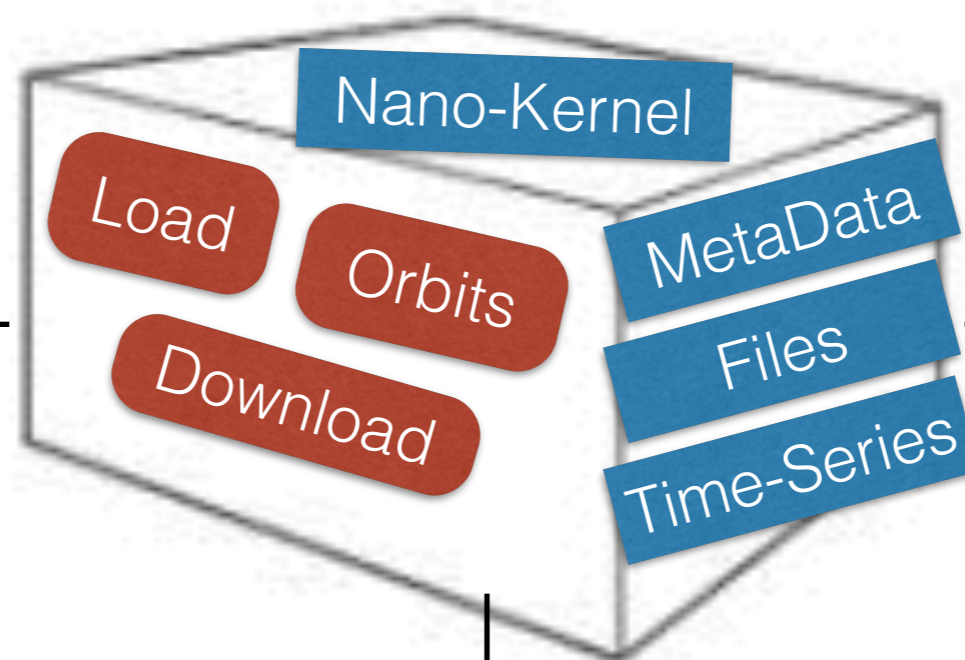
Median Averaging Routine onto Grid

Data Assimilation Method

System for System Science

Satellite based instruments

Models



Ground based instruments

Space Weather Drivers

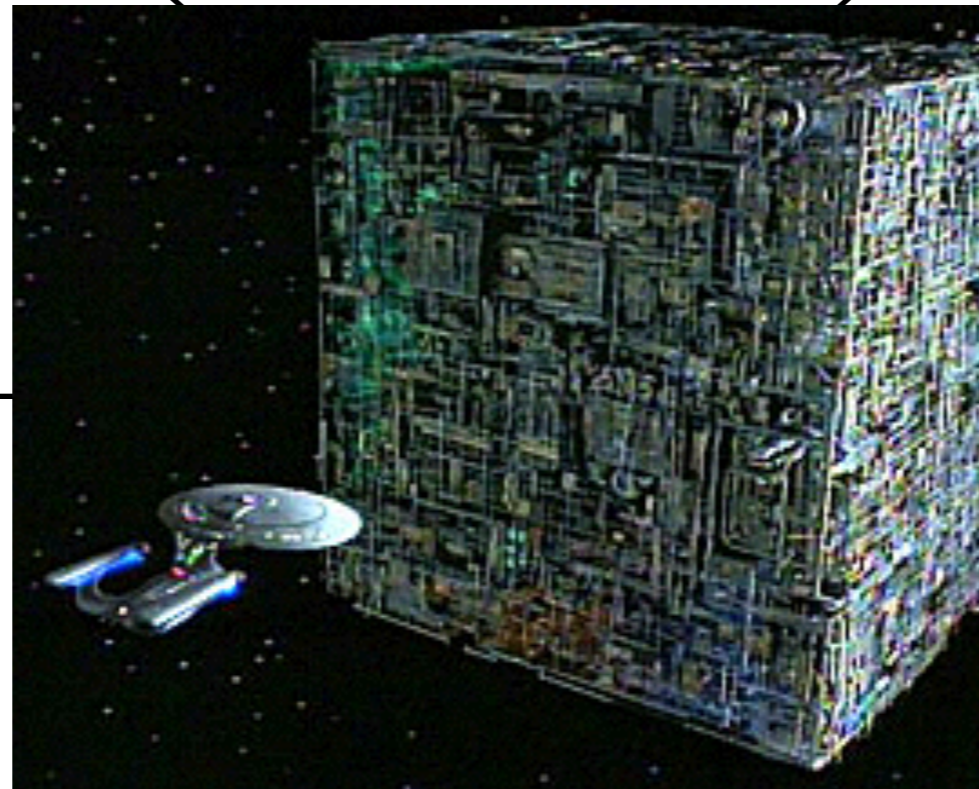
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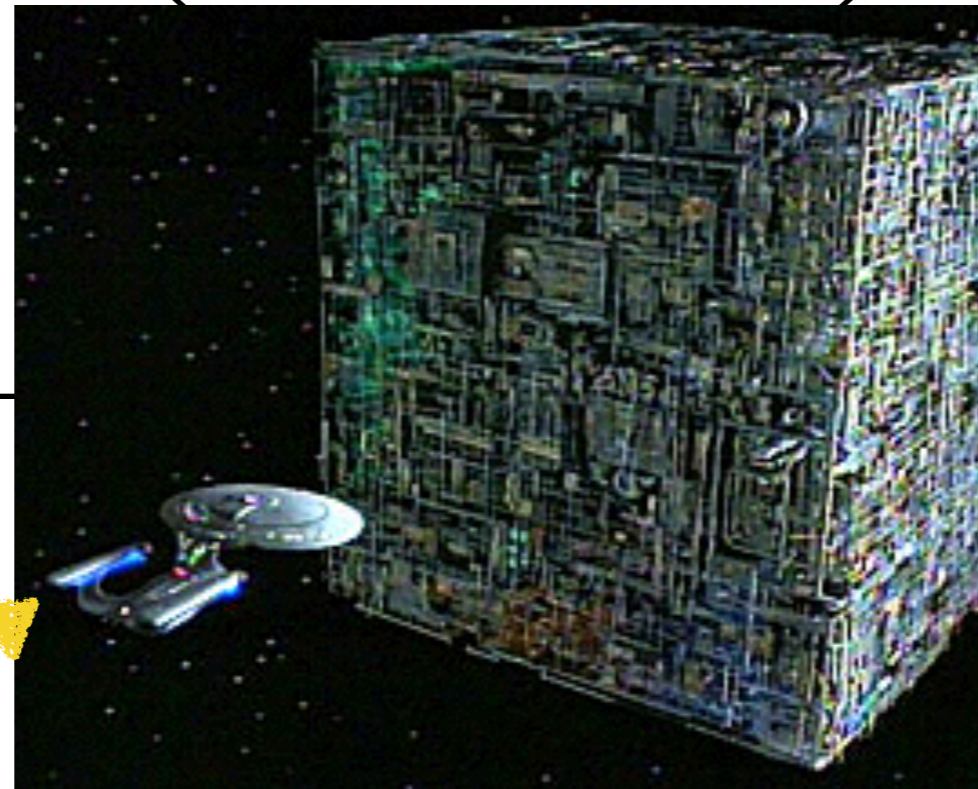
Satellite based instruments

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Ground based instruments

Space Weather Drivers

New Data Set



Median Averaging Routine onto Grid

Data Assimilation Method

Suggestion

Grad Students : You should use python. There are no good reasons to still use IDL (or even Matlab)

- Installation, at terminal command prompt:
pip install pysat Requires CDF, netCDF, HDF libraries as appropriate
- Full Documentation, tutorials, installation, API reference:
<http://rstoneback.github.io/pysat/>
- Stay up-to-date with latest code
git clone <https://github.com/rstoneback/pysat.git>
- Getting science python
Enthought (Canopy); Continuum Analytics (Anaconda);
PyCharm (full featured IDE) from JetBrains free for education

Question

What unique qualities does your data have that need to be accounted for by a general system for system science?

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