

# GPS scintillation and irregularities at the front of a tongue of ionization in the nightside polar ionosphere over Svalbard

Christer van der Meeren<sup>1</sup>, Kjellmar Oksavik<sup>1,2</sup>, Dag Lorentzen<sup>2,3</sup>, Jøran Moen<sup>4,2</sup>, Vincenzo Romano<sup>5</sup>

<sup>1</sup> Birkeland Centre for Space Science, Department of Physics and Technology, University of Bergen, Norway.

<sup>2</sup> University Centre in Svalbard, Longyearbyen, Norway.

<sup>2</sup> Birkeland Centre for Space Science, University Centre in Svalbard, Longyearbyen, Norway.

<sup>4</sup> Department of Physics, University of Oslo, Norway

<sup>5</sup> Istituto Nazionale di Geofisica e Vulcanologia, Rome, Italy







#### Motivation

- Scintillations in polar cap well correlated with patch activity
- What about continuous TOIs?



[Gondarenko and Guzdar, 2004]



## Background

- Plasma irregularities of scale sizes 10m–1km cause scintillations on GPS signals
- Amplitude scintillations: 10–100m-scale irregularities, S<sub>4</sub> index

$$S_4^2 = \frac{\langle I^2 \rangle - \langle I \rangle^2}{\langle I \rangle^2}$$
,  $I = \text{power}$ 

• Phase scintillations: 100m–1km-scale irregularities, usually  $\sigma_{\phi}$  index

 $\sigma_{\phi}^2 = \langle \phi^2 \rangle - \langle \phi \rangle^2$ ,  $\phi$  = detrended phase



## Methodology

- Case study of scintillation on the leading edge of a TOI over Svalbard on 31 Oct 2011
- 3 GPS receivers (NovAtel GSV4004B GISTM) (2 in NYA, 1 in LYR)
- EISCAT Svalbard Radar
- 630.0nm ASI at NYA
- SuperDARN Hankasalmi radar





## TOI across polar cap





## TOI across polar cap





# TOI airglow over Svalbard





Geomagnetic overview







**GPS IPP locations** 







## Discussion – the problems of $\sigma_{\phi}$

[Forte and Radicella, 2002; Béniguel et al., 2004; Forte, 2005; Beach, 2006; Mushini et al., 2012]

- Phase scintillation does not necessarily imply irregularities
- May instead be «false» scintillations from the gradient itself
- 1. Poor data detrending
  - $\sigma_{\phi}$  index sensitive to low-frequency phase variations
  - Gradient = low-frequency phase variation
  - 0.1 Hz too low at high latitudes (gradients correspond to phase variations > 0.1 Hz)
- 2. No scintillation after the gradient
  - · Gradient drift instability (GDI) stable on leading edges
  - Irregularities would most likely develop from inside the structure toward the front
  - Uncertain vorticities, rotations making the front unstable to the GDI











#### Conclusion

- Case study of GPS scintillation at TOI front on 31 October 2011
- Notable phase scintillation at the leading gradient, no significant amplitude scintillation
- Increases in the  $\sigma_{\phi}$  index are most likely due to TEC gradient and poor data detrending
- No evidence in favour of structuring (significant enough to cause GPS scintillation) on the TOI front, or inside the high-density region immediately behind the gradient