

Local-time Variations of Low Latitude Lower Thermospheric SABER CO₂ during Equinoctial Solar Minimum



Cornelius Csar Jude H. Salinas^{1,2,3,*}, Loren C. Chang^{1,3}, Mao-Chang Liang^{1,4}, Jia Yue^{5,6}, Liying Qian⁷, James Russell III⁵ and Martin Mlynczak⁸



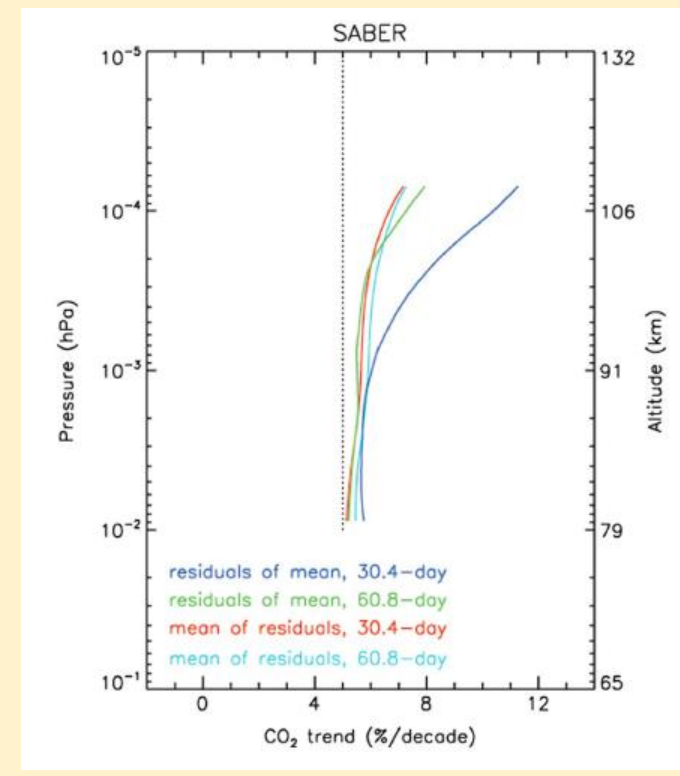
¹Taiwan International Graduate Program- Earth Systems Science, Academia Sinica, Taipei, Taiwan; ²Department of Atmospheric Sciences, National Central University, Zhongli, Taiwan; ³Center for Astronautical Physics and Engineering, National Central University, Zhongli, Taiwan; ⁴Institute of Earth Sciences, Academia Sinica, Taipei, Taiwan; ⁵Hampton University, Hampton, Virginia, USA; ⁶ESSIC, University of Maryland, USA; ⁷High Altitude Observatory, Boulder, Colorado, and ⁸NASA Langley Research Center, Hampton, Virginia, USA

*corresponding author: ccjsalinas@gmail.com

(I) IMPORTANCE OF LOCAL TIME VARIATIONS OF CO₂

Accurate Calculation of Long-term Trend from Satellite-based CO₂ Observations

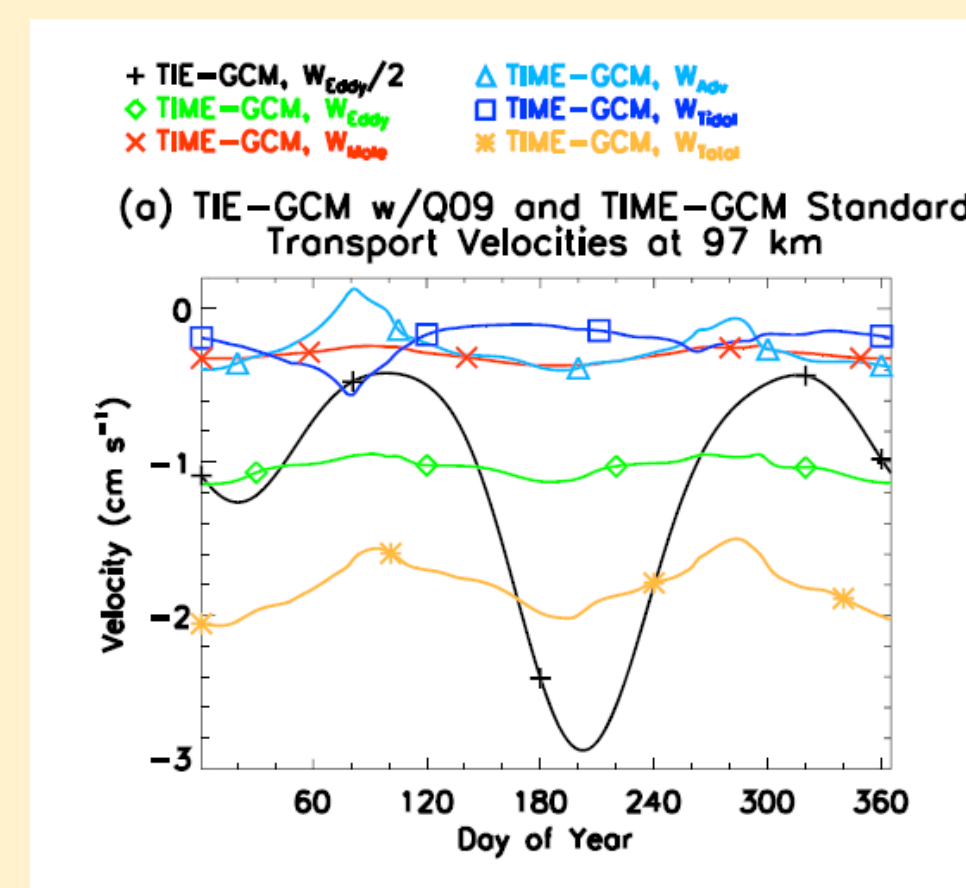
- Long-term trend of SABER CO₂ is dependent on sampling window.
- Dependence in sampling window implies a never before reported local-time variation in SABER CO₂.



[Qian et al, 2017]

Untangling Advection and Diffusion in the MLT Region

- Local-time variations in the MLT region are controlled by tides.
- CO₂ is an ideal tracer in the MLT region whose local-time variations can help estimate the complicatedly tangled tide-induced advection and diffusion.



[Jones et al, 2017]

Objectives of the Study:

- To present, explain and model, for the first time, the local time-variations of SABER CO₂.
- To determine the contributions of advection and diffusion to the local-time variations of SABER CO₂.
- To relate the local-time variations of SABER CO₂ and SABER temperature via vertical motion.

(II) DATASETS

SABER v2.0 CO₂ and Temperature Profiles

All profiles for March and April 2008 are gridded into local-time and latitude bins. Local-time coverage is between 0600 am to 1800 pm [Rezac et al, 2015].

TIME-GCM Model Outputs

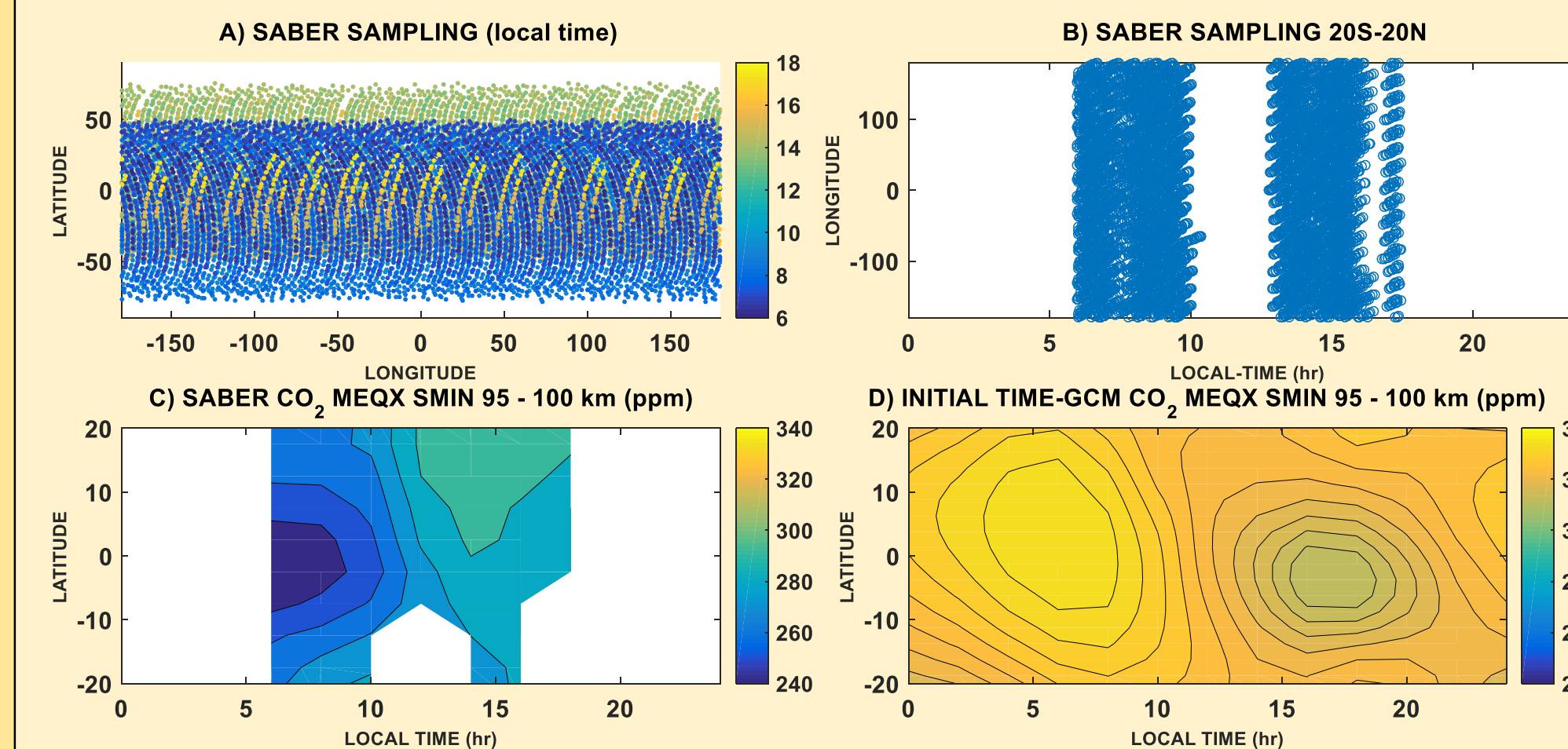
TIME-GCM was ran at double resolution for day 80 under solar minimum and geomagnetically quiet conditions with GSWM migrating diurnal and semidiurnal tides specified at the lower boundary of the model.

SD-WACCM Model Outputs

WACCM nudged with MERRA up to ~50 km. Outputs for March 2008 are used.

(III) LOCAL-TIME VARIATIONS OF SABER CO₂ AND ITS PHYSICAL MECHANISMS: TENDENCY ANALYSIS WITH TRANSFORMED EULERIAN MEAN (TEM) AND 3D CONTINUITY EQUATION FOR CO₂

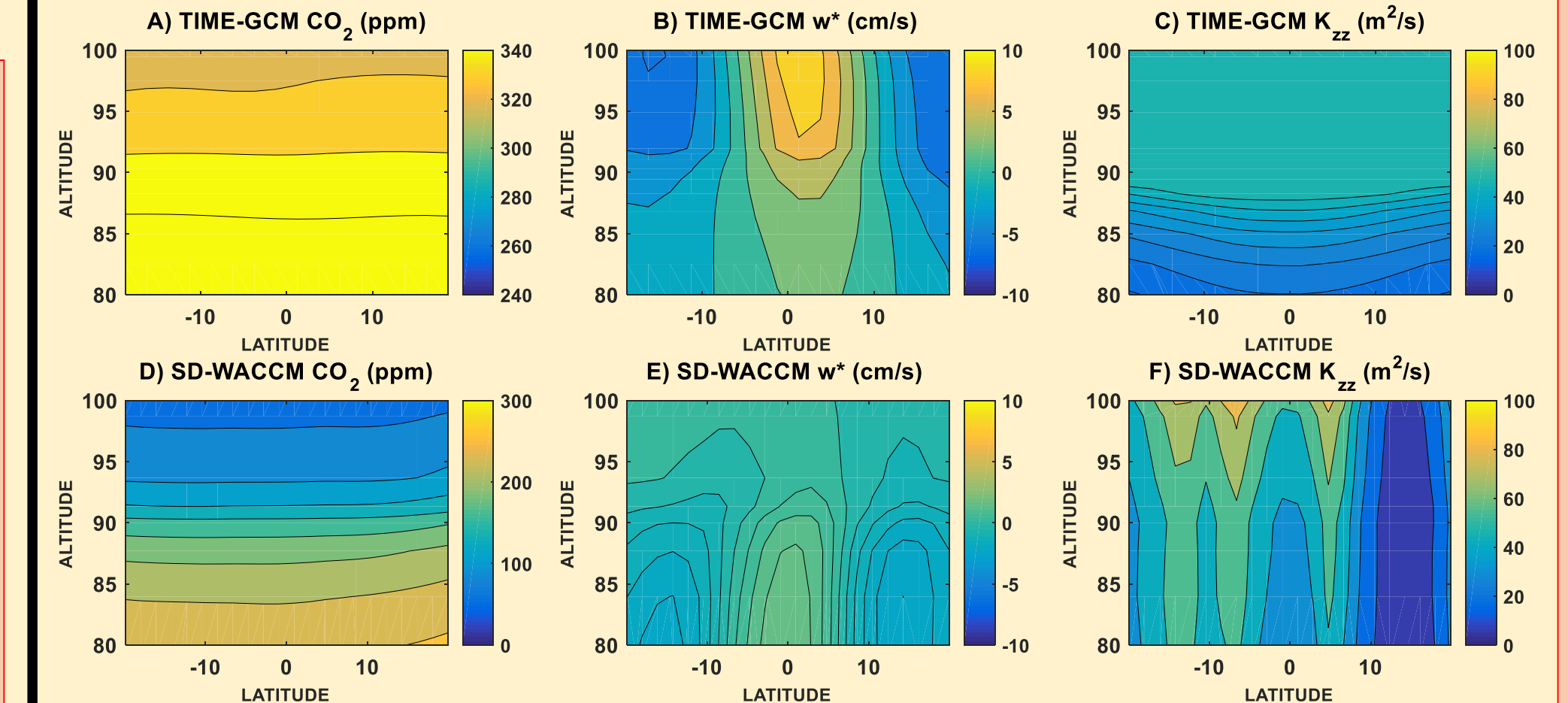
(A) SABER CO₂ vs TIME-GCM CO₂



HIGHLIGHTS

- SABER CO₂ has local min in morning and local max in afternoon.
- TIME-GCM CO₂ local-time variation is opposite of SABER CO₂.
- Zonal-mean TIME-GCM CO₂ is higher than zonal-mean SABER CO₂.

(B) PHYSICAL MECHANISM OF ZONAL-MEAN CO₂

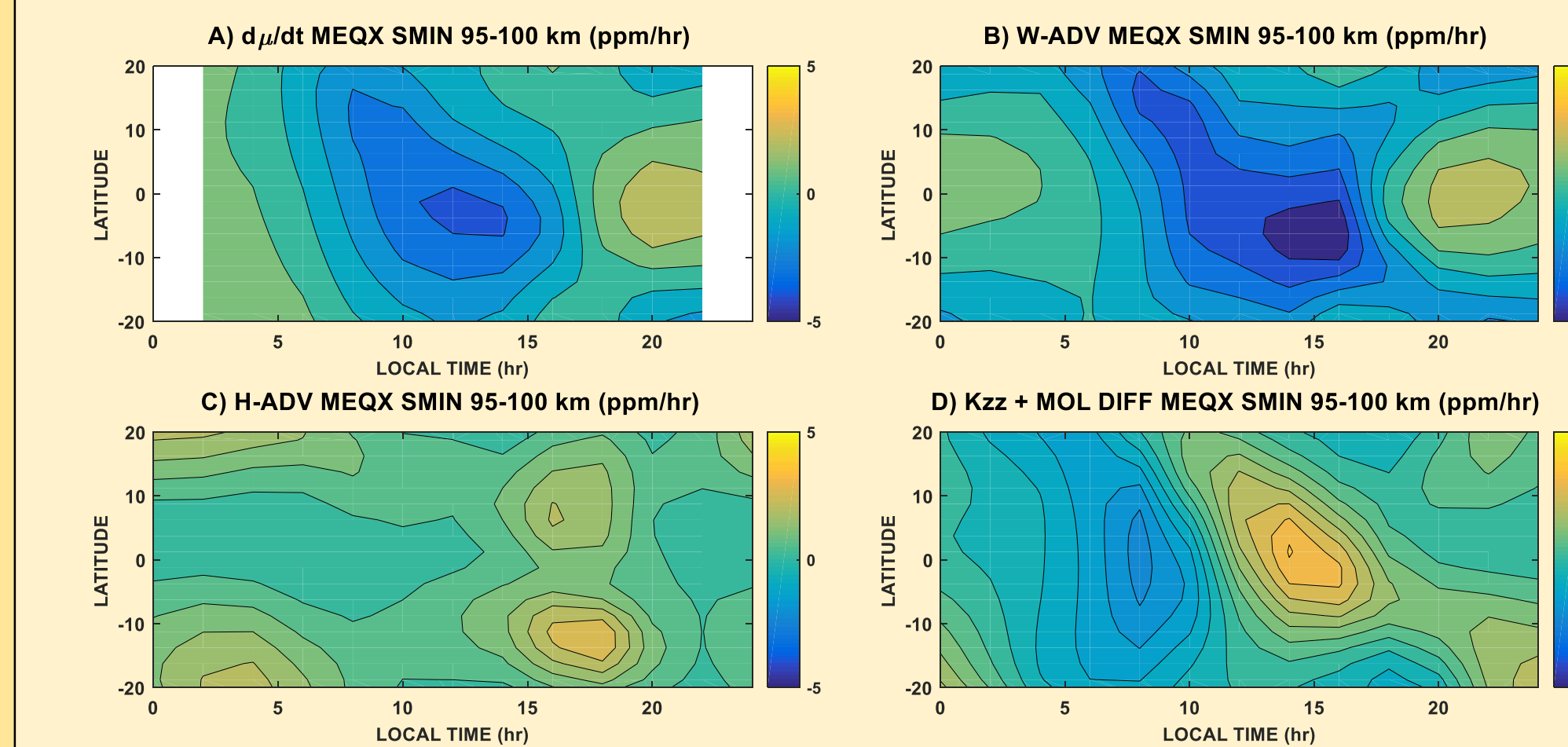


HIGHLIGHTS

- Zonal-mean TIME-GCM CO₂ is predominantly controlled by TEM vertical advection.
- TIME-GCM CO₂ has higher zonal-mean because of the presence of TEM upwelling.
- SD-WACCM zonal-mean CO₂ and TEM circulation are more realistic.

$$\frac{\partial \bar{\mu}}{\partial t} + \bar{v}^* \frac{\partial \bar{\mu}}{\partial \phi} + \bar{w}^* \frac{\partial \bar{\mu}}{\partial z} - \frac{1}{\rho_0} \frac{\partial}{\partial z} \left(\rho_0 K_{zz} \frac{\partial \bar{\mu}}{\partial z} \right) - \frac{1}{\rho_0} \frac{\partial}{\partial z} \left(\rho_0 D_{\mu} \frac{\partial \bar{\mu}}{\partial z} \right) + \frac{1}{\rho_0} \frac{\partial}{\partial z} \left(\rho_0 \mu W_D \right) = 0$$

(C) PHYSICAL MECHANISM OF LOCAL-TIME VARIATIONS IN CO₂

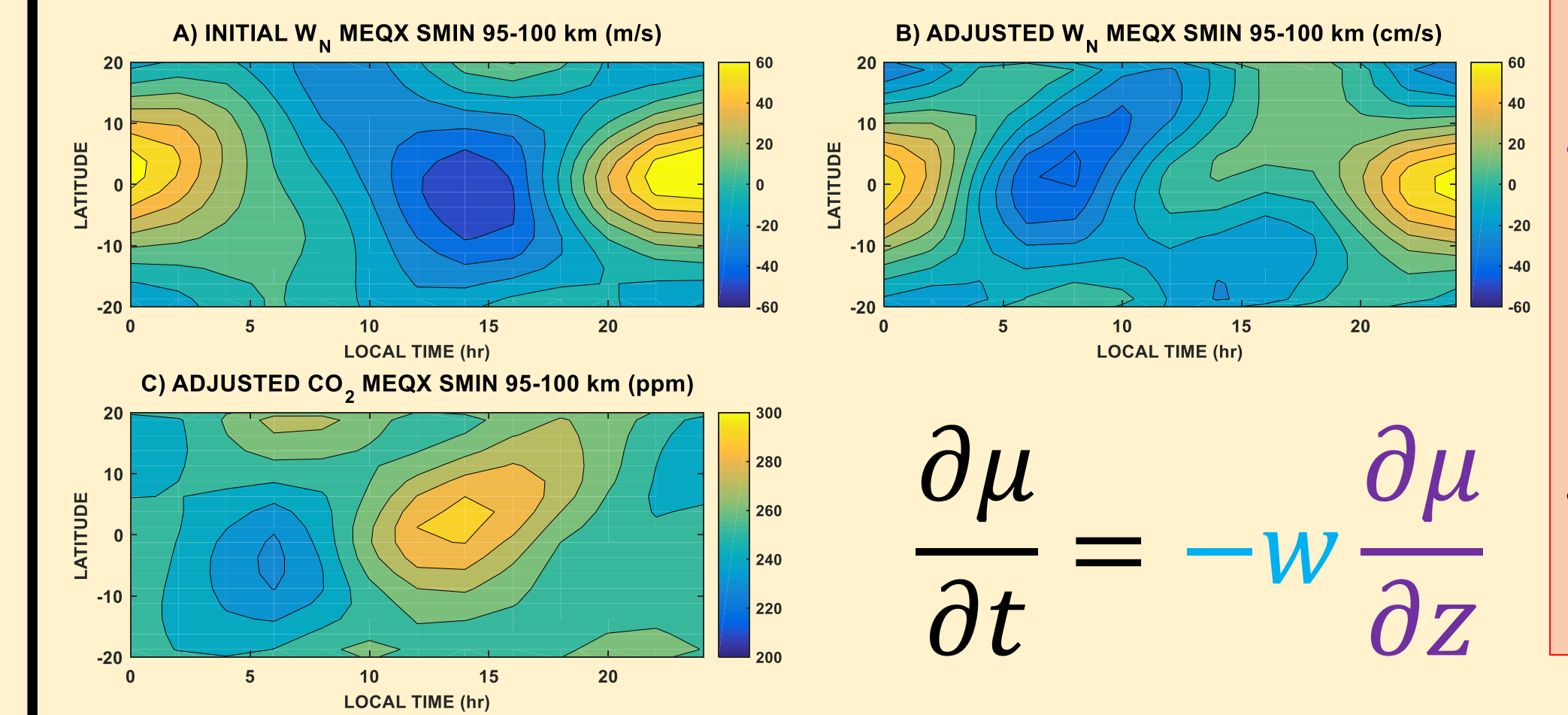


HIGHLIGHTS

- Local-time variations of low latitude lower thermospheric SABER CO₂ are predominantly controlled by vertical advection.
- Vertical motion values can be derived from low latitude lower thermospheric CO₂ during equinox.

$$\frac{\partial \mu}{\partial t} + \left(\frac{u}{a \cos \phi} \right) \frac{\partial \mu}{\partial \lambda} + \left(\frac{v}{a} \right) \frac{\partial \mu}{\partial \phi} + w \frac{\partial \mu}{\partial z} - \frac{1}{\rho_0} \frac{\partial}{\partial z} \left(\rho_0 K_{zz} \frac{\partial \mu}{\partial z} \right) - \frac{1}{\rho_0} \frac{\partial}{\partial z} \left(\rho_0 D_{\mu} \frac{\partial \mu}{\partial z} \right) + \frac{1}{\rho_0} \frac{\partial}{\partial z} \left(\rho_0 \mu W_D \right) = 0$$

(D) SIMPLIFIED TRANSPORT MODELING OF SABER CO₂



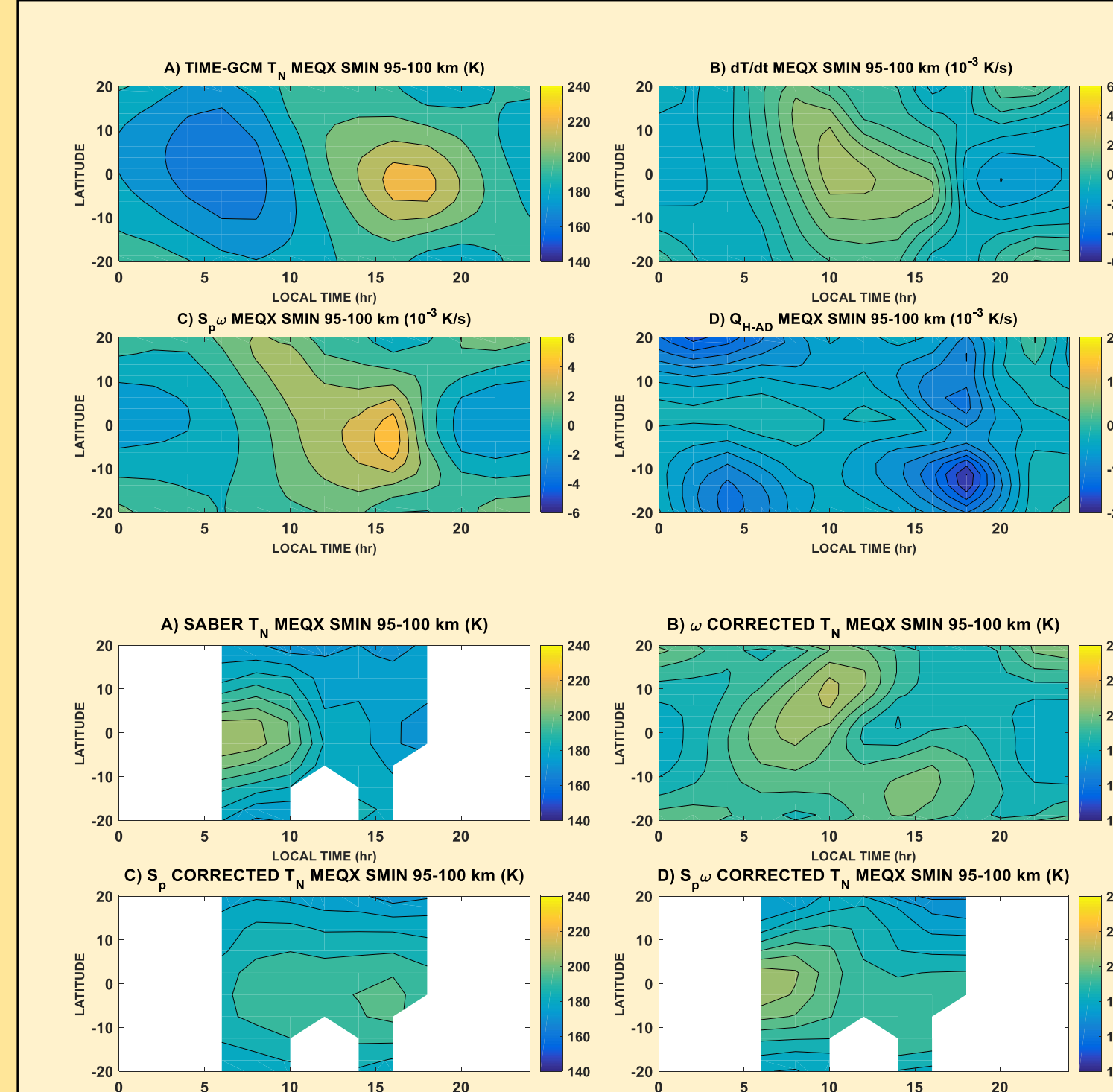
HIGHLIGHTS

- SD-WACCM monthly-mean CO₂ and adjusted TIME-GCM vertical velocity is used as input into our simplified transport model (STM).
- STM captured general features of the local-time variations of SABER CO₂

$$\frac{\partial \mu}{\partial t} = -w \frac{\partial \mu}{\partial z}$$

$$w = \bar{w} + w'_{DW1} \cos \left(\frac{2\pi t}{24} - \psi_{DW1} \right) + w'_{SW2} \cos \left(\frac{2\pi t}{12} - \psi_{SW2} \right) + w'$$

(IV) SABER CO₂ vs SABER T_n



HIGHLIGHTS

Vertical motion significantly affects low latitude lower thermospheric temperatures. This enables relating the SABER CO₂-derived vertical motions to SABER temperatures.

$$\frac{\partial T}{\partial t} = -\frac{w S_p \rho_0}{g}$$

Local-time variations of SABER CO₂-derived vertical velocity and SABER T_n-derived static stability and neutral density induces general features of the local-time variations of SABER T_n. Local-time variations of SABER CO₂ is consistent with that of SABER T_n.

(V) DISCUSSIONS AND CONCLUSION

Upper Atmosphere Climate Change:

- SABER CO₂ has a significant local-time variations that needs to be accounted for in calculating long-term trends. The significant local-time variations of CO₂ may help explain recently suggested local-time variations in upper atmosphere trends [Danilov, 2015].

Tide-induced Advection and Diffusion

- To simulate the local-time variations of SABER CO₂, TIME-GCM TEM vertical velocity above 95 km needs to be weakened while the local-time variations of vertical velocity need to be adjusted. Eddy diffusion plays a minor role.
- Local-time variations of SABER CO₂-derived vertical velocity also drives the local-time variations of SABER T_n. Local-time variations of SABER T_n are well accepted [Zhang et al, 2006; Mukhtarov et al, 2009; Pancheva et al, 2009; Xu et al, 2009; Sakazaki et al, 2012].

This work concludes that there is a significant local-time variation in low latitude lower thermospheric SABER CO₂ during equinoctial solar minimum that is driven predominantly by TEM and tide-induced vertical advection.