

The MLT-X Grand Challenge Workshops, 2014-2017

Coupling and Transport Processes from the Upper Mesosphere through the Middle Thermosphere (80-200 km)

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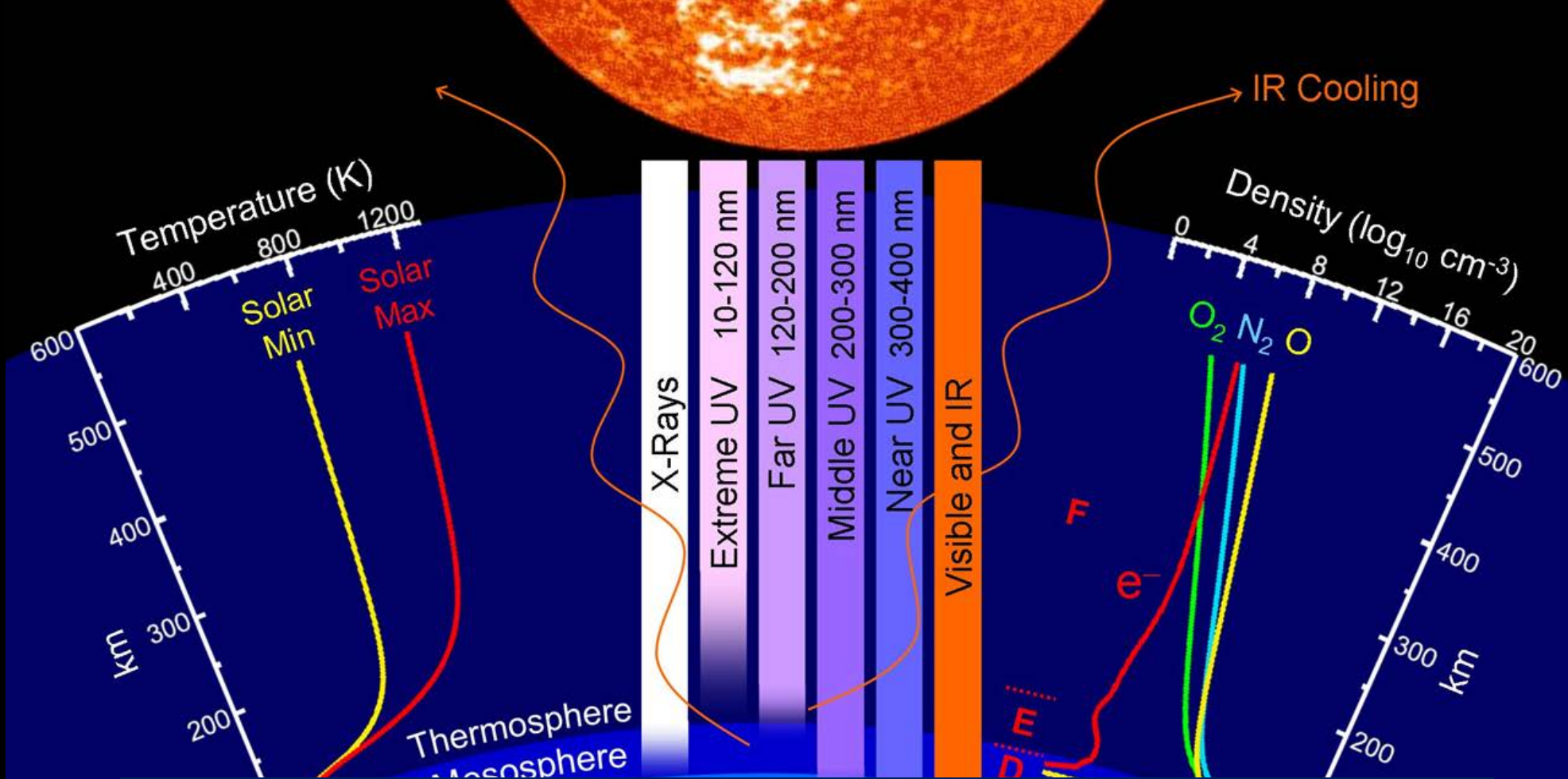
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CEDAR 2014-2017

- ❖ What is the role of the neutral gas in coupling with the plasma to establish the predominant state of the Earth's upper atmosphere and ionosphere between 80 and 200 km?
- ❖ How do wave-induced transport, dissipation and turbulence influence the structure, composition and circulation of Earth's upper atmosphere between 80 and 200 km?



MLT-X has the MOST SIGNIFICANT vertical gradients in nearly every property of the upper atmosphere

Universal Processes in the MLT-X

MLT-X Processing of External Inputs MAXIMIZE in the region

- ❖ Processing of EUV Solar Radiative Flux
- ❖ Processing of Solar Wind / Magnetosphere Energy Flux
- ❖ Processing of Cosmic Dust Flux
- ❖ Processing of Lower Atmosphere Wave Flux
- ❖ Processing of Constituent flux

MLT-X Internal Workings are HIGHLY DIVERSIFIED

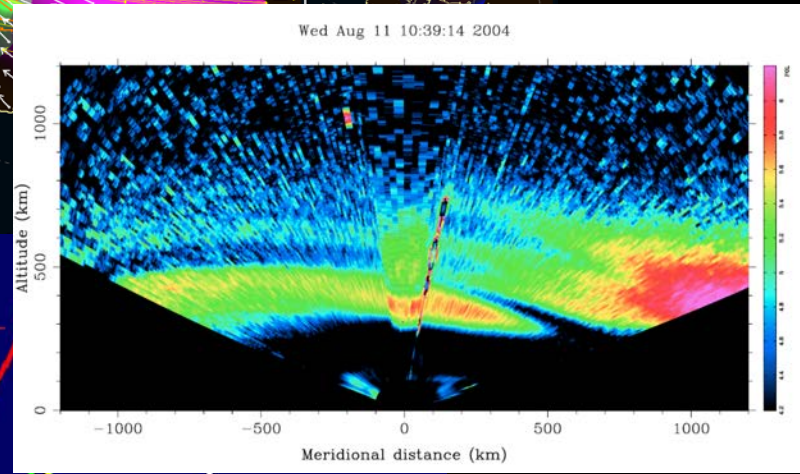
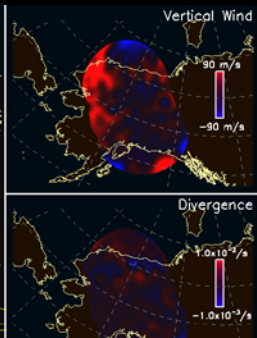
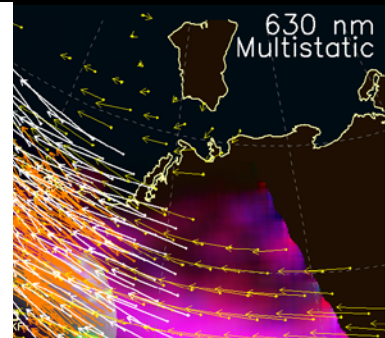
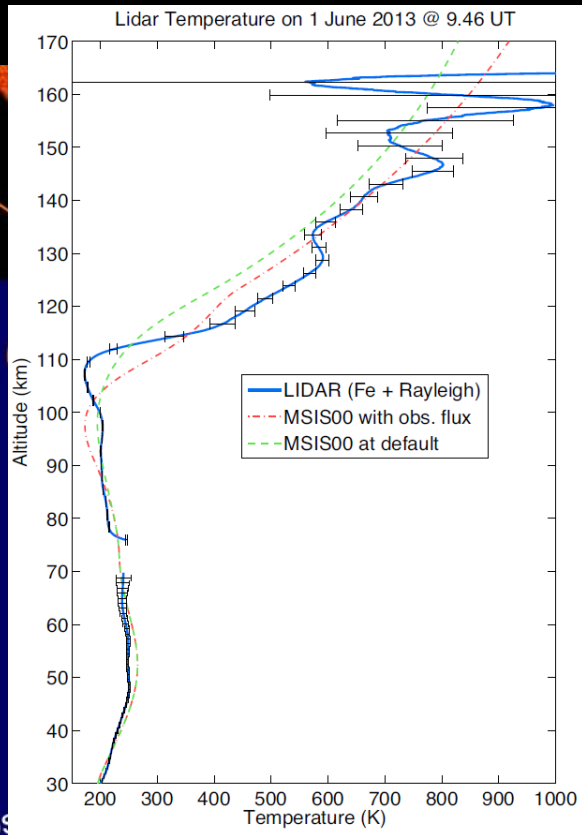
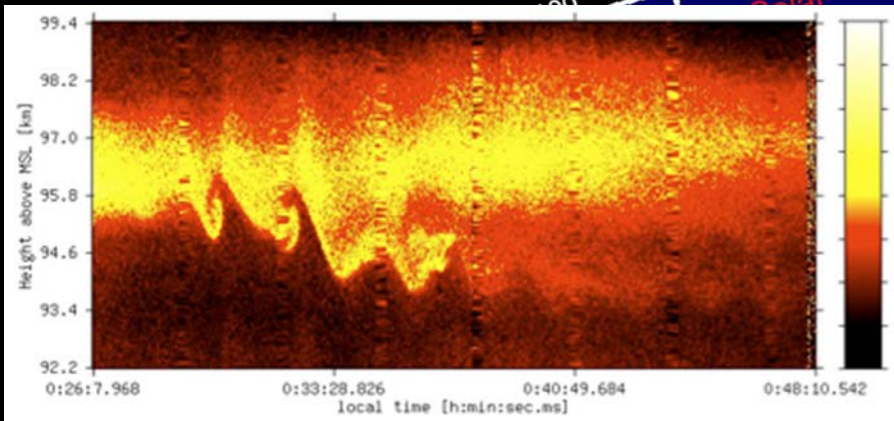
- ❖ Workings of Plasma-Neutral Interactions
- ❖ Workings of Wave Dissipation/Generation
- ❖ Workings of Instabilities and Turbulence
- ❖ Workings of Momentum, Energy and Constituent Transport
- ❖ Workings of Chemical Interactions

MLT-X GC Workshop Four-Year Plan

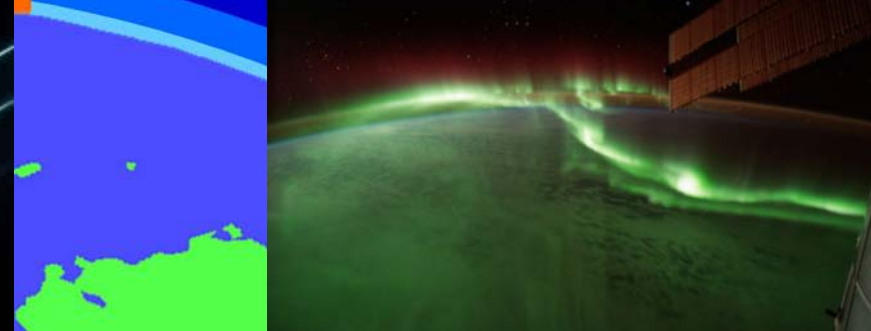
- ❖ *Year 1:* Hold a CEDAR workshop to **establish middle thermosphere science goals** related to the two questions posed by inviting theorists and experimentalists involved in middle thermosphere / ionosphere studies. Consider the diverse processes that occur at low, middle and high latitudes to identify the impact of the science.
- ❖ *Year 2:* Hold a CEDAR workshop to consider the necessary **measurement scenarios** to achieve the goals established in the first year and align existing and planned instruments to meet the science requirements.
- ❖ *Year 3:* Hold a CEDAR workshop that establishes **current numerical modeling capabilities** of understanding the MLT-X and its broad influence on upper atmosphere processes in plasma-neutral coupling and wave / turbulence.
- ❖ *Year 4:* Hold a CEDAR workshop to plan campaigns and **to motivate new measurements** in this critical, poorly observed region of the thermosphere.

Temperature (K)

800 1200

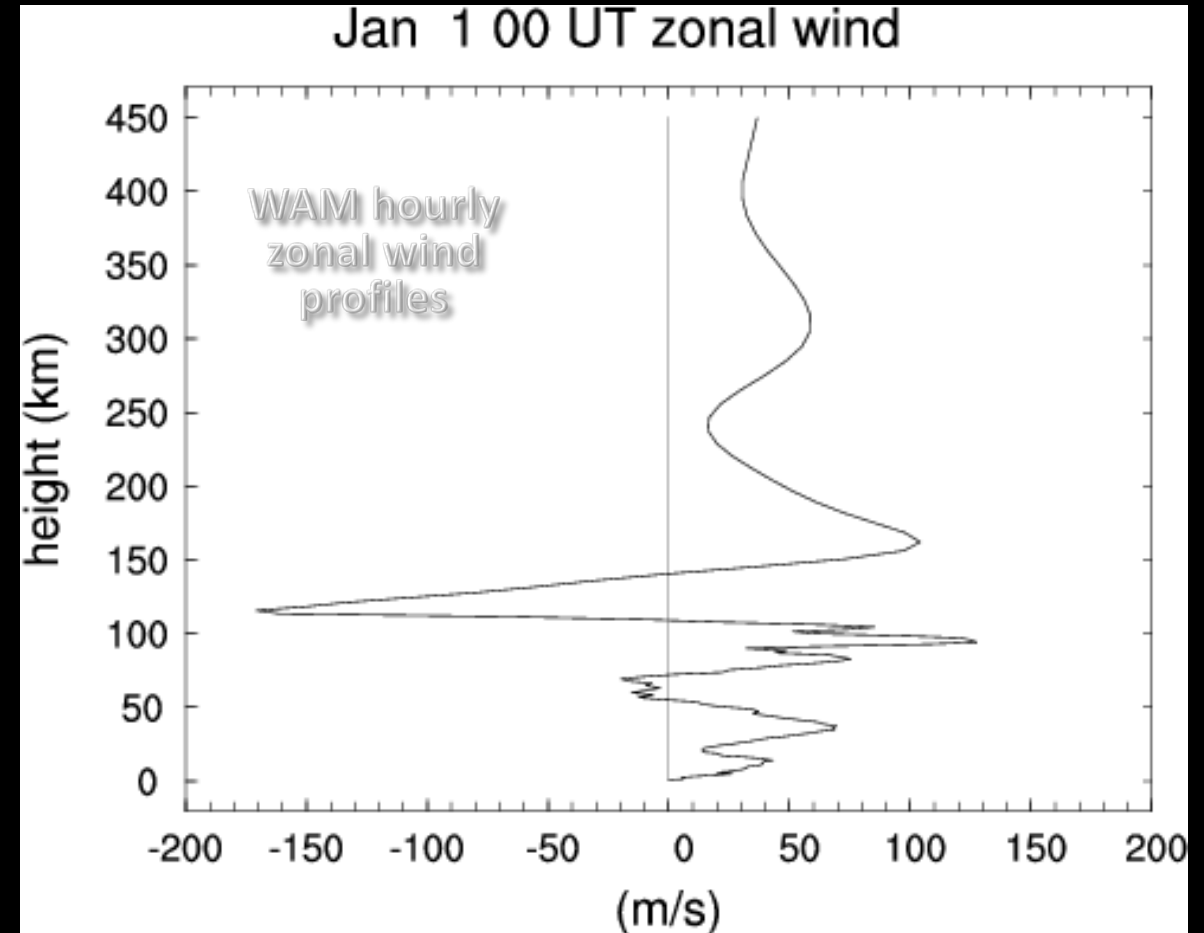
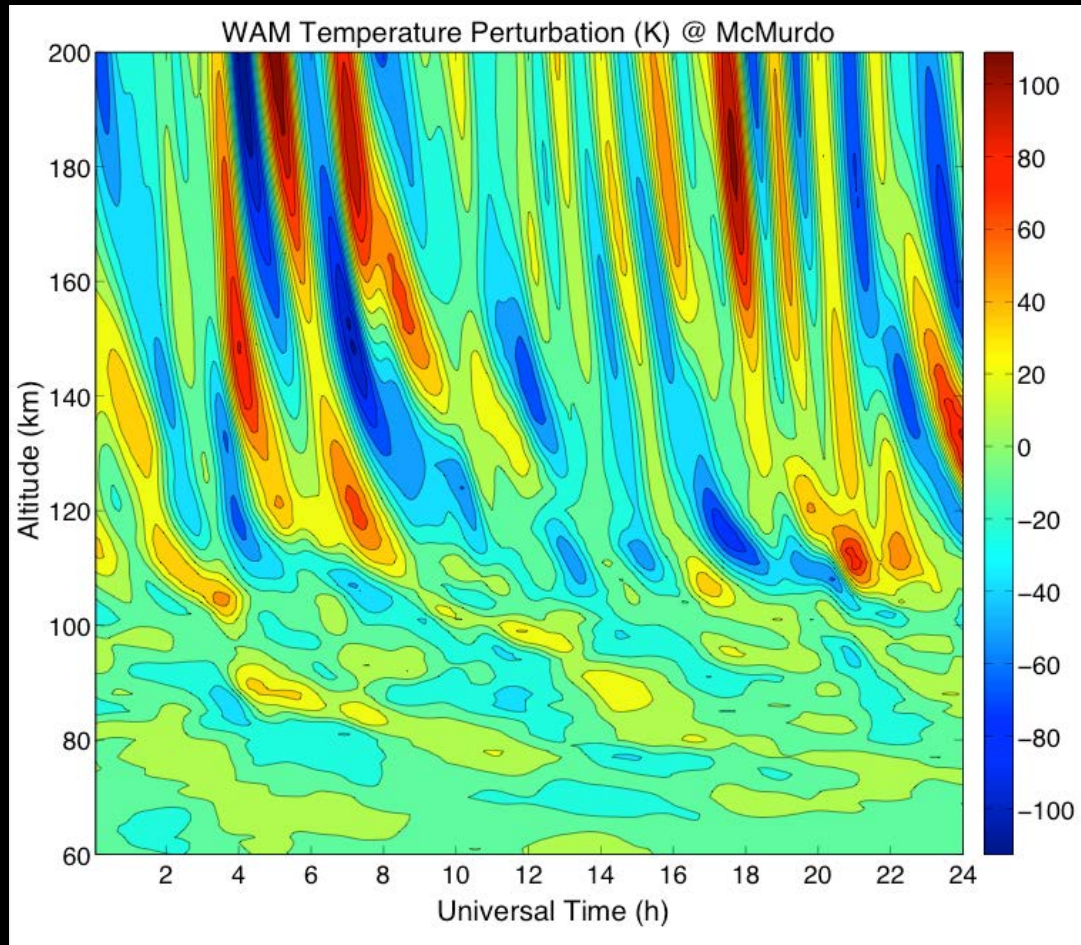


Thermosphere
Mesosphere
Stratosphere



Whole atmosphere simulations of a rich spectrum of waves propagating from the lower atmosphere impacting the thermosphere and ionosphere

A new paradigm in thermosphere-ionospheric modeling



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An Observational Transformation is Required

- ❖ A new measurement scheme of neutral properties through the MLT-X region is required
 - ❖ The MLT-X is the least understood and observed, yet the most dynamic and structured region in the upper atmosphere (strong vertical gradients and rapid variability)
 - ❖ Progress in ionosphere physics is shackled by the lack of neutral measurements
- ❖ A new level of correlated measurement is required
 - ❖ Wave fluxes are needed to quantify the lower atmosphere influence on the MLT-X
 - ❖ Neutral density, plasma density, neutral winds, and ion motion
- ❖ A new means to test and advance Whole Atmosphere Models is required
 - ❖ No current measurements adequately verify whole atmosphere model predictions in the MLT-X

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Observatory for Atmosphere Space Interaction Studies (OASIS)



From the desolate observational domain of the
MLT-X domain there lies on the horizon an **OASIS**

*Where thirsty aeronomers
can come and drink from its
data pool and be
enlightened.*



OASIS - An Evolution in Observations of the MLT-X

❖ Fundamental MLT-X Properties

- ❖ Neutral gas density, temperature, winds, and composition resolved in space and time (strong vertical gradients and temporal variations)

❖ Electrodynamic Measurements

- ❖ Plasma and neutral properties (conductivity, dynamo winds, frictionally-heated neutral temperatures)

❖ Flux Measurements

- ❖ Wave momentum, heat and constituent transport $\overline{w'u'}$, $\overline{w'v'}$, $\overline{w'T'}$, $\overline{w'\rho'}$
- ❖ Meteoric influx
- ❖ Eddy coefficients of thermal, constituent, and momentum diffusivity

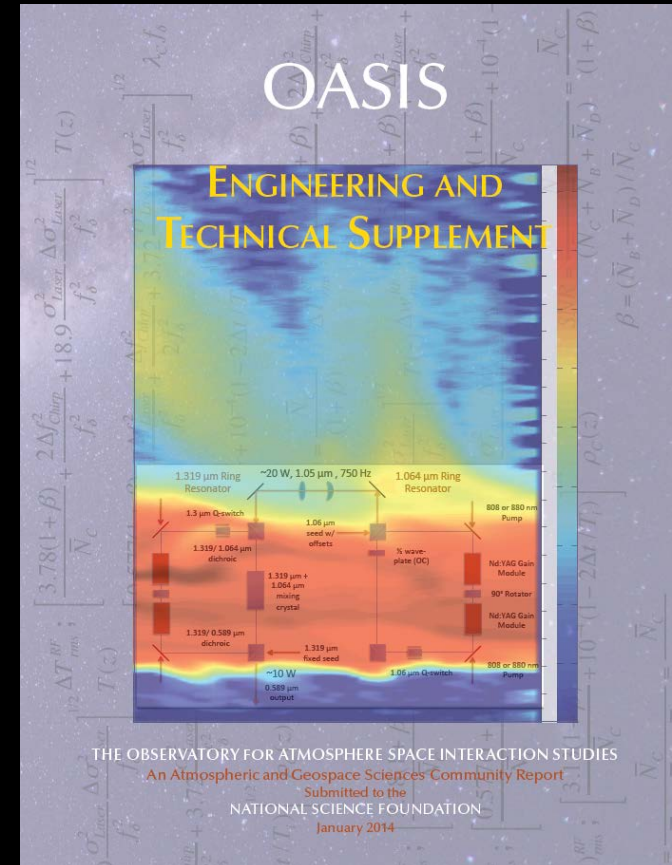
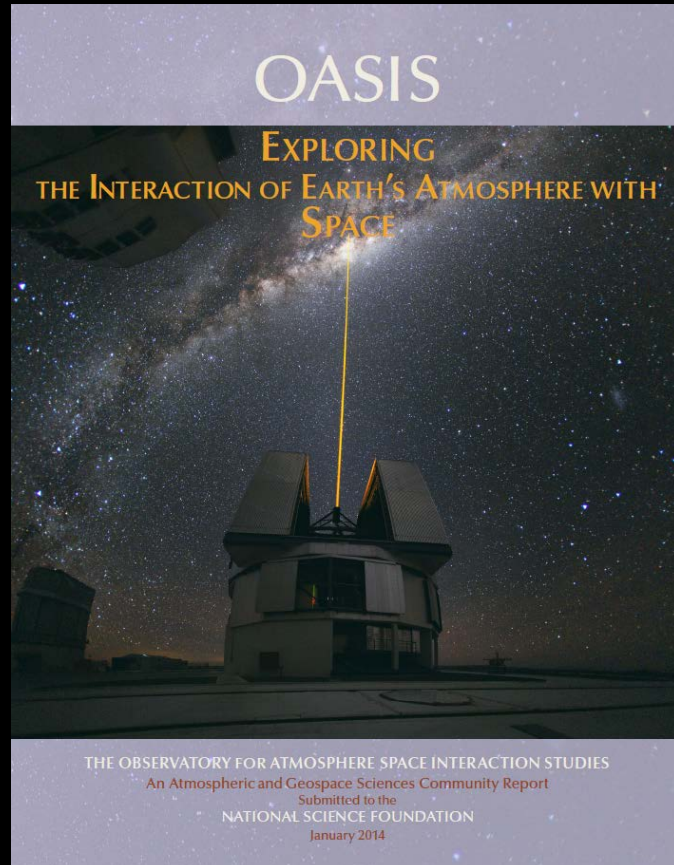
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OASIS

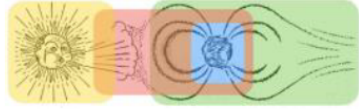
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EXPLORING THE GEOSPACE FRONTIER: QUO VADIS?

Science topic	Observational gap	Potential Facility Component(s)
Generation: Coronal and chromospheric Magnetism	Magnetic and energetic inputs at base of system.	CoSMO ¹ , FASR ² , SPRING ³
Propagation: Interplanetary Space	The “93-million mile gap”	VHF Space Weather Radar ⁴
Transference: Geospace plasma and energy sources	Undersampled in space and time	Ground and Space Distributed Sensor Networks ⁵
Dissipation: Upper Atmosphere plasma-neutral interactions	Neutral and plasma observations above 100 km	OASIS ⁶

OASIS

Observatory for Atmosphere Space Interaction Studies

- ❖ **A Transformational Observatory**
 - ❖ Able to observe neutral properties (T , u , v , w , ρ) from virtually the surface to 1000 km
 - ❖ Able to resolve wave fluxes into the thermosphere
 - ❖ Able to observe neutral and plasma interactions from 80 to 200 km
 - ❖ Able to detect meteor entry, ablation and fluxes
- ❖ **A new kind of GeoSpace Facility**
 - ❖ A centerpiece facility for lidar observing: the neutral gas equivalent to incoherent scatter radar plasma capability
 - ❖ Regional Coverage by: Incoherent scatter radar (AMISR), other radars, imagers, interferometers, spectrometers and, perhaps, in situ measurement capabilities using balloon and rocket payloads

