

# Simulating Global-scale Observation of Limb and Disk's (GOLD) measurement of the July 2<sup>nd</sup>, 2019 total solar eclipse's effect on the Ionosphere-Thermosphere system

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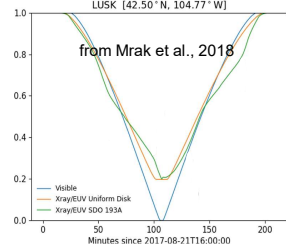
## Abstract

A total solar eclipse will occur in the southern hemisphere on July 2, 2019 from approximately 17 to 22 UT. It will cross South America where it's effects can be observed from a geostationary orbit by the Global-scale Observation of Limb and Disk (GOLD) instrument. GOLD will observe the changes in OI 135.6 nm and N<sub>2</sub> LBH band emissions, which will be used to derive changes in composition ratios, neutral temperatures on the dayside and peak electron densities on the nightside as a result of the eclipse. Simulations of the eclipse's effects on the Ionosphere-Thermosphere (I-T) system, as seen by the GOLD instrument, have been performed using the Thermospheric General Circulation Model (TIEGCM) and the GLOBal airGLOW (GLOW) model. Calculation show small but observable differences in the daytime airglow brightnesses (~10 R change) and temperatures (~30K change). The results obtained are being used to deduce the required measurement criteria for GOLD to successfully observe the expected changes. The magnitudes and temporal-spatial development of these changes can be compared with observations. This comparison will provide a test for our understanding of the eclipse's effects on the I-T and their origin.

## Background

Moon's shadow moving at supersonic speed blocks solar radiation → localized cooling and reduction in photoionization

Fig I. Measured attenuation of radiation during the August 21, 2017 eclipse at Lusk, WY. This attenuation leads to rapid cooling and a reduction in photoionization.



Large scale waves propagate away from the eclipse's path causing global scale changes

GOLD observation from a geostationary orbit will provide unique global scale measurement

## GOLD: Global-scale Observation of Limb and Disk

GOLD records spectra from ~132-163 nm as a function of slit height from a geostationary orbit above South America

LIMB scans will also be made but this work focuses on day and night disk scans

These measurement of OI 135.6 nm and N<sub>2</sub> LBH provide:

- O/N<sub>2</sub> and effective disk Temperature (day)
- Electron densities (night)

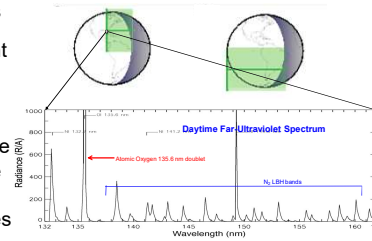


Fig II. Normal day scan mode of the spectrograph

## 2019 Eclipse and GOLD's view of the Eclipse

### Total Solar Eclipse of 2019 Jul 02

Geocentric Eclipation - 19:21:36.4 UT J.D. = 248667.30667  
 Greatest Eclipse - 19:22:53.0 UT J.D. = 248667.30751

from nasa.gov

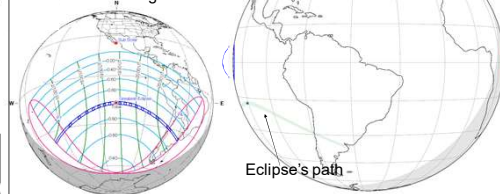
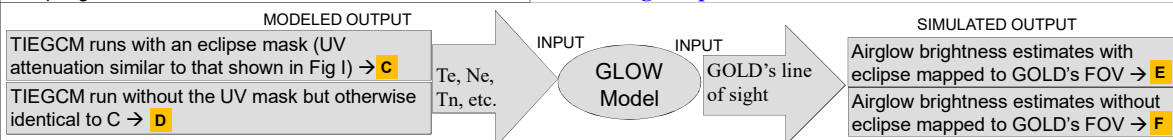


Fig III. Left: Totality (dark blue) and partial eclipse (sky blue) during July 2, 2019 and the times that they occur (green). Right: The eclipse's path of totality (shaded green) as seen by GOLD.

## Modeling Eclipse's effect in the I-T

## Special campaign mode on July 2



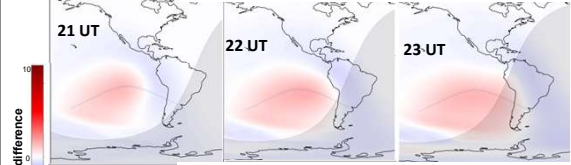
Two channels available (A & B):

A will perform normal full disk scan at 30 minutes cadence

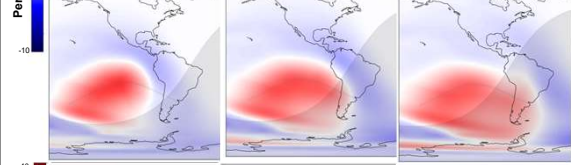
B scans south-western quarter of the disk at 30 minutes cadence until 18:10 LT (22 UT) before switching to normal night scans

At 19:10 LT (23 UT), A starts normal night scans

Difference: C and D O/N<sub>2</sub> change at pressure level -2 (~160 km): TIEGCM



O/N<sub>2</sub> change at pressure level 2 (~250 km): TIEGCM



Neutral temperature at pressure level -2 (~160 km): TIEGCM

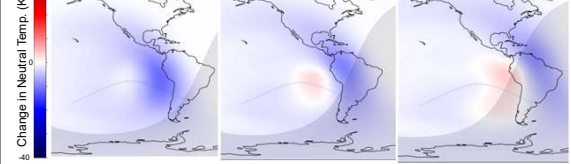
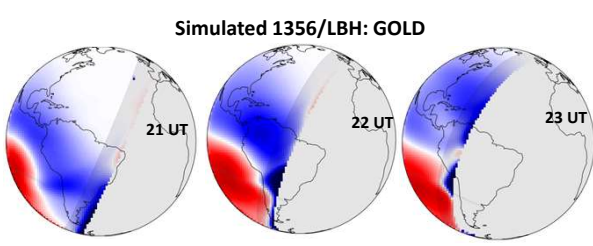
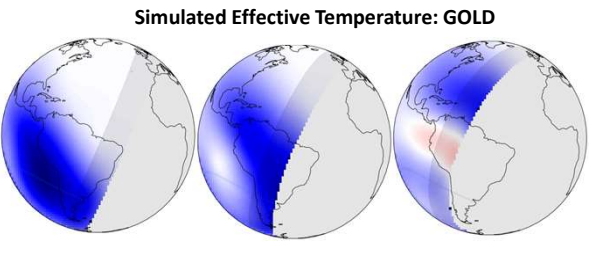


Fig IV. Left: Modeled change in O/N<sub>2</sub> at pressure level of -2 (top), 2 (middle) and neutral temperature at pressure level of -2 (bottom) due to the effect of the July 2, 2019 eclipse as simulated by TIEGCM at three different times, few hours after the eclipse. Right: Simulated changes in OI 1356/ N<sub>2</sub> LBH ratio (top) and effective temperature as seen by GOLD. Eclipse's path is shown in light gray.

Difference: E and F



Simulated Effective Temperature: GOLD



**Summary**

Eclipse induced changes in the I-T simulated in terms of changes in OI 135.6 nm and N<sub>2</sub> LBH bands as seen by the GOLD instrument

TIE-GCM simulation shows observable change in O/N<sub>2</sub> at pressure level of -2 (daytime GOLD observation)

Changes in O/N<sub>2</sub> at F<sub>2</sub> peak (pressure level 2) may affect the nighttime EIA strength (night-time GOLD observation)

Changes due to the lower atmosphere have an effect but are not included in this study. How much of a difference would it make?

GOLD observation of the July 2, 2019 eclipse will help address these questions