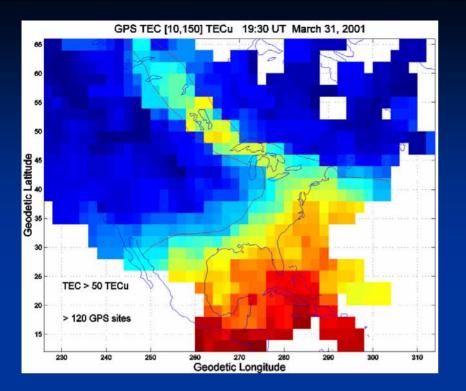
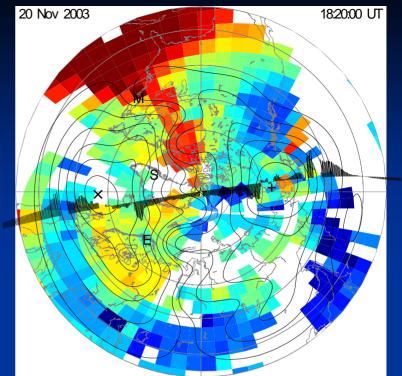
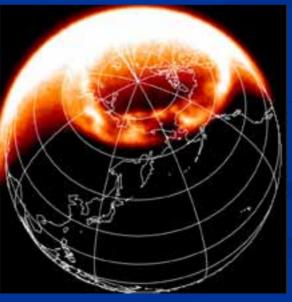


DAST Distributed Arrays of Small Instruments GPS Receivers **Optical Imagers** Interferometers Ionosondes Scinifikition and VLF Rx Tomography Receivers Solar Observations Magnetometers Passive & Active Radar Radio Receiver • - Riometers Netiton Monitors Technology: ITR, Miniaturization O IPS Arrays **PO Opportuntitie** Each Current Monitor







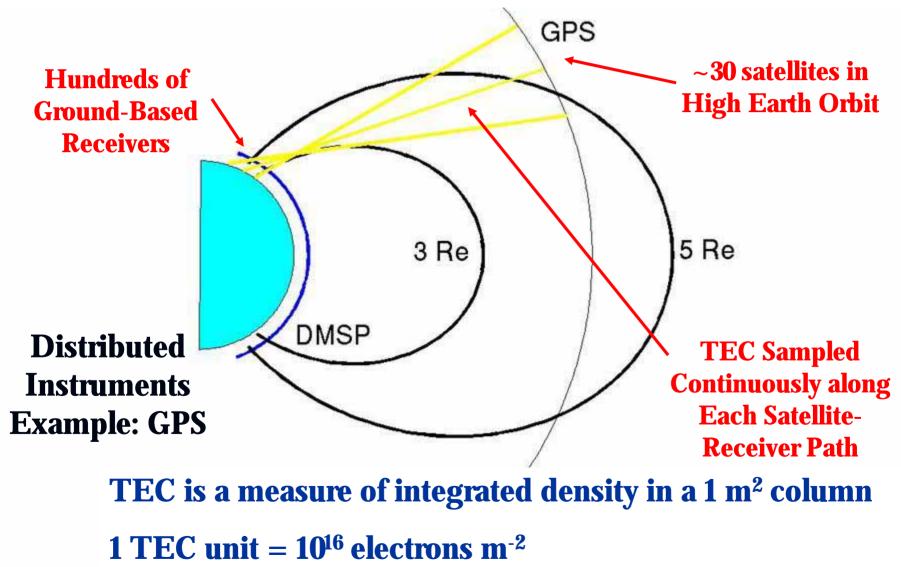


CEDAR/GEM Workshop Geospace System from Ground and Space

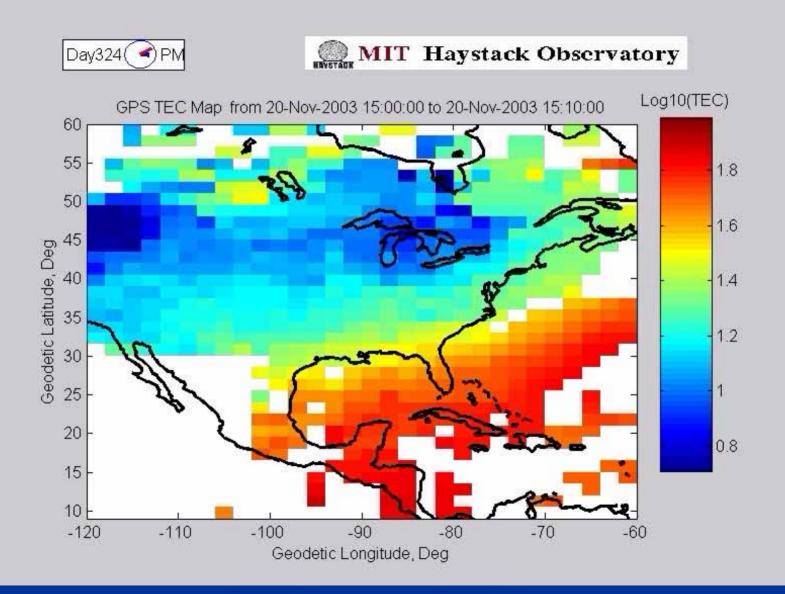
DASI Overview

- The NAS Solar and Space Physics Decadal Survey has recommended that the next major groundbased instrumentation initiative be the deployment of arrays of space science research instrumentation
- DASI arrays will provide continuous real-time observations of Earthspace with the resolution needed to resolve mesoscale phenomena and their dynamic evolution
- Ground-based arrays will address the need for observations to support the next generation of space weather data-assimilation models
- The time is right for DASI: developing technology and IT systems support a new science capability

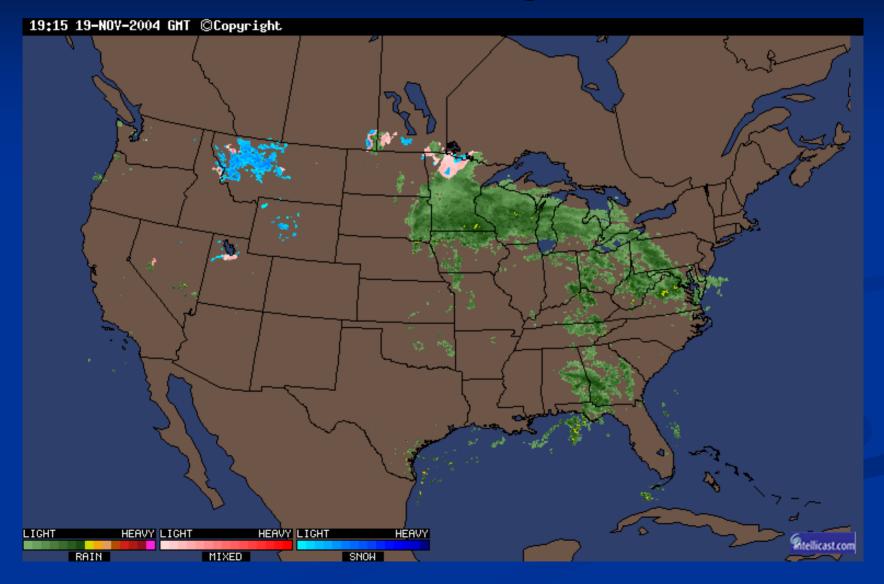
GPS samples the ionosphere and plasmasphere to ~20,000 km. Dual-frequency Faraday Rotation Observations give TEC (Total Electron Content)



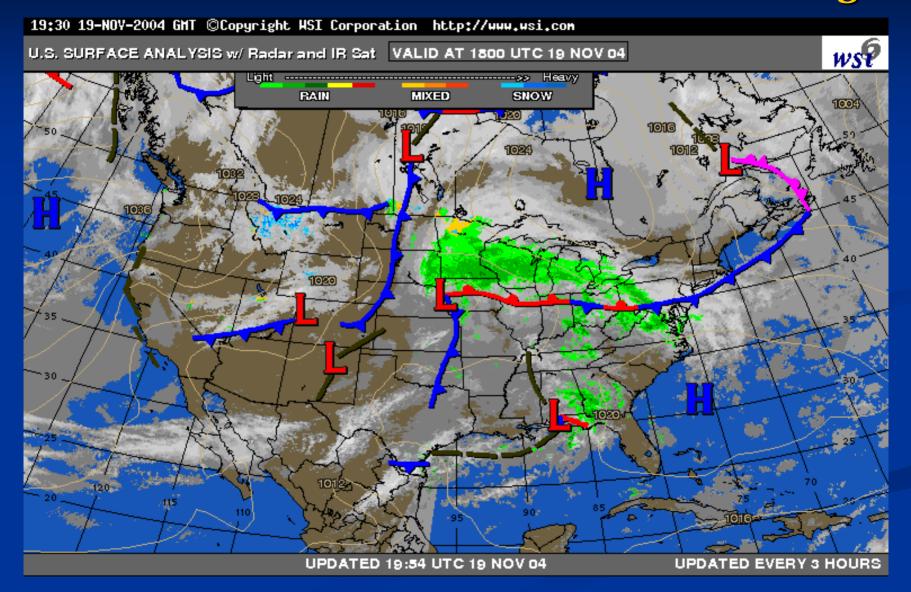
Observations: GPS TEC



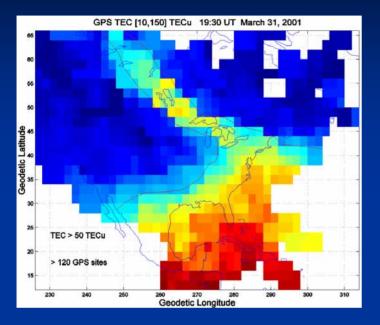
Today's Weather: NEXRAD Observations of Meteorological Storm Front



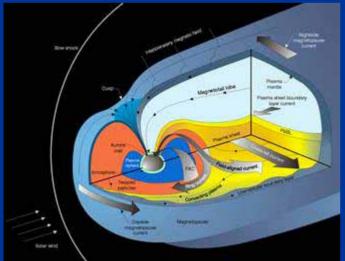
Analysis & Understanding are Well Developed Distributed Observations & Assimilative Modeling

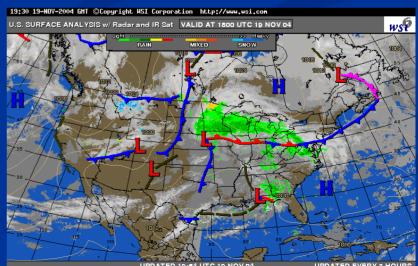


Imaging Meso-Scale Phenomena with Distributed Observations









UPDATED 19:54 UTC 19 NOV 04

UPDATED EVERY 3 HOURS

Report Outline (1)

Executive summary (2 pages)

1. Introduction

- Space Weather
- Distributed Arrays of Small Instruments The Next Logical Step

2.DASI Background (Decadal Survey)

- 1. Challenges
- 2. DASI

3. DASI Workshop

- Recurrent Themes
 - I. Insufficient Observations
 - **2. Geospace as a System**
 - **3. Real-Time Observations**

Persistent Themes

I Insufficient Observations

Observational space physics is data-starved, producing large gaps in our ability to both characterize and understand important phenomena. This is particularly true for Space Weather events, which often are fast-developing and dynamic, and extend well beyond the normal spatial coverage of our current sensor arrays.

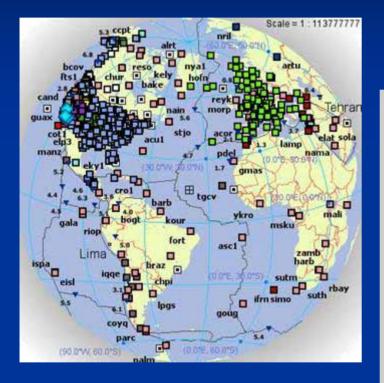
2 Geospace as a System

Geospace processes involve **significant coupling across atmospheric layers** and 'altitude boundaries', as well as coupling across **multiple scale sizes** from global (1000s km), to local (10s km), to micro-scale (meter-scale and smaller).

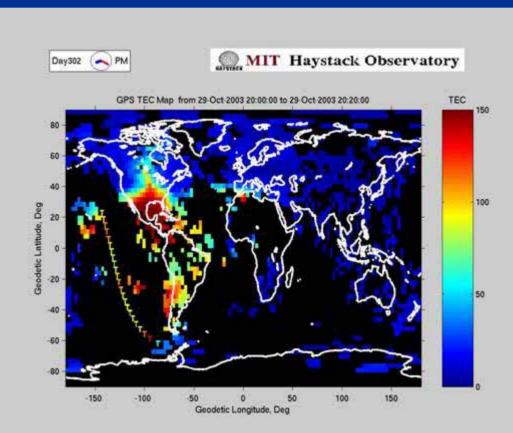
3 *Real-Time Observations*

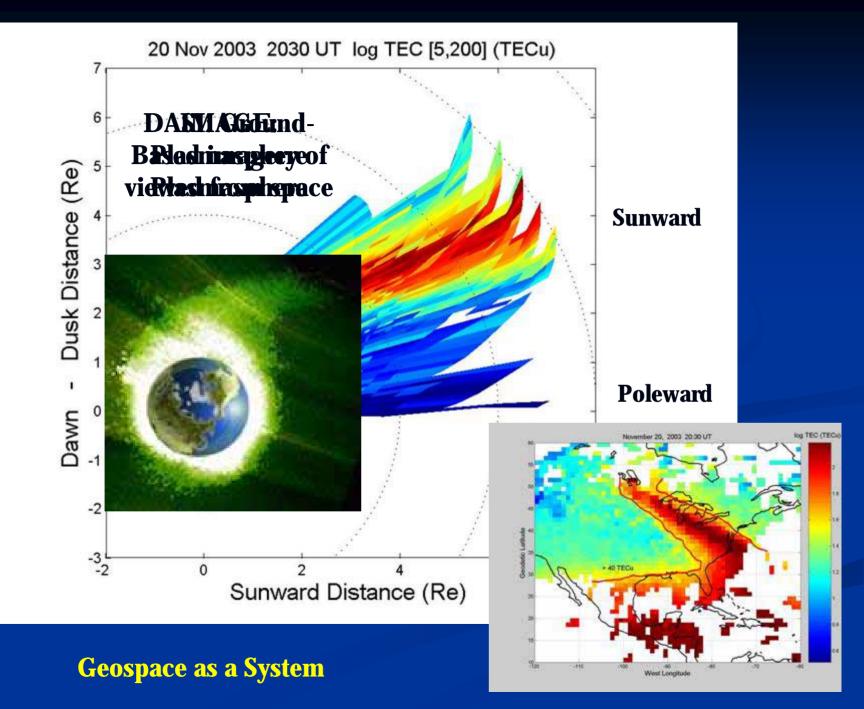
*E*lucidation of the fundamental coupling processes requires **continuous real-time measurements** from a distributed array of diverse instruments as well as physics-based data assimilation models.

Current Arrays: Limits on Global Coverage & Real-Time Access (e.g. GPS Receivers)



Issues: Logistics & International Participation

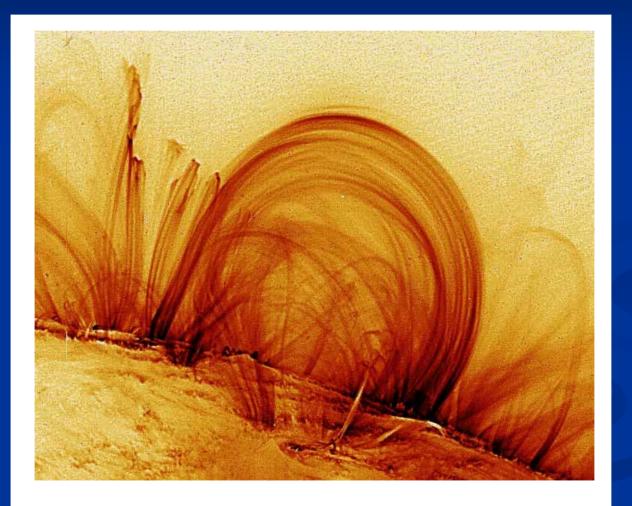




Report Outline (2)

4. Compelling Science I. Solar - Interplanetary ■ 1. Magnetic Fields ■ 2. Solar Activity ■ 3. Solar Variability • 4. Interplanetary Scintillations 2. Magnetosphere-Ionosphere I. Plasma Redistribution ■ 2. Particle Energization ■ 3. Effects of Preconditioning 3. Thermosphere-Ionosphere ■ 1. Ion-Neutral Coupling ■ 2. Waves and TI Variability

Solar Magnetic Fields & Variability



The Sun as the Driver of the Geospace Environment

Structure and Dynamics of the Sun's Interior

- local helioseismology

Solar Activity

- 1) understand CME formation
- 2) provide information about the magnetic structure of earth-directed CMEs and solar wind
- 3) predict the onset of large solar flares
- 4) learn how and where solar energetic particles are accelerated

Heliospheric Structure

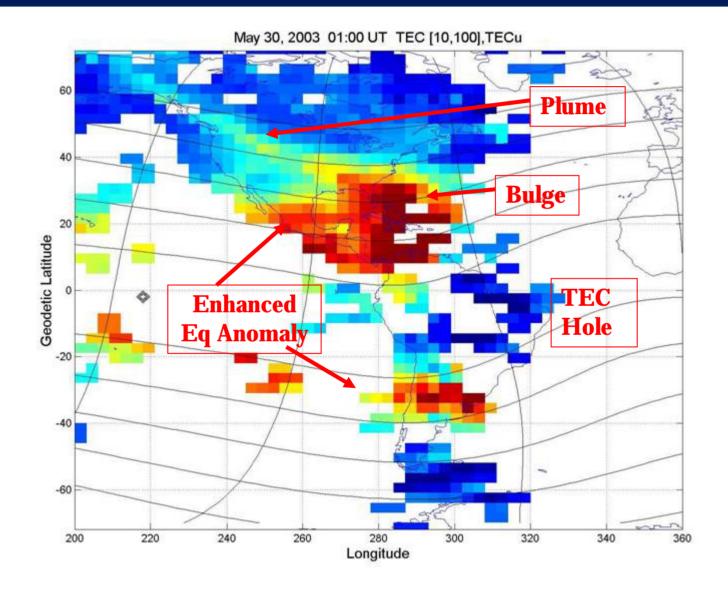
1) identify the magnetic processes that accelerate the solar wind

2) understand the interactions of CMEs and solar wind streams

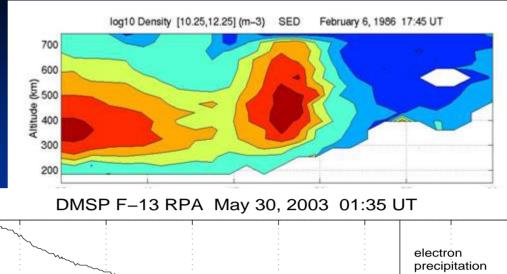
Science (1): Global Aspects of Localized SPW Effects (Plasmasphere Boundary Layer Science)

- Meso-scale SPW features which cross wide spans of latitude and couple low-mid-auroral-polar latitudes. Disturbance features with significant applications effects.
- Great **spatial/temporal variability in occurrence**. Rapid coupling across regimes of system. Spatial localization (e.g. SSA).
- Involve thermal plasma and its effects, electric fields, relationship to particle precipitation and wave growth and interactions.
- Questions: Reason for longitude effects; sources and characteristics of electric fields; cause of SPW features (e.g. gradients produced by 2ndary wave growth along boundaries); relationship to scintillation and intense wave growth (ERB). Thermal plasma effects on the development of M/S storm?

M-I Coupling Leads to Processes and Structures which Span and Interconnect the M-I-T System

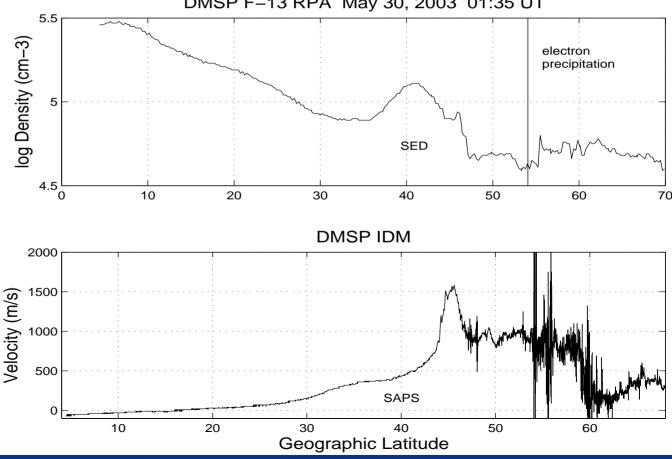


Regionally-Localized but Recurrent Space Weather Events

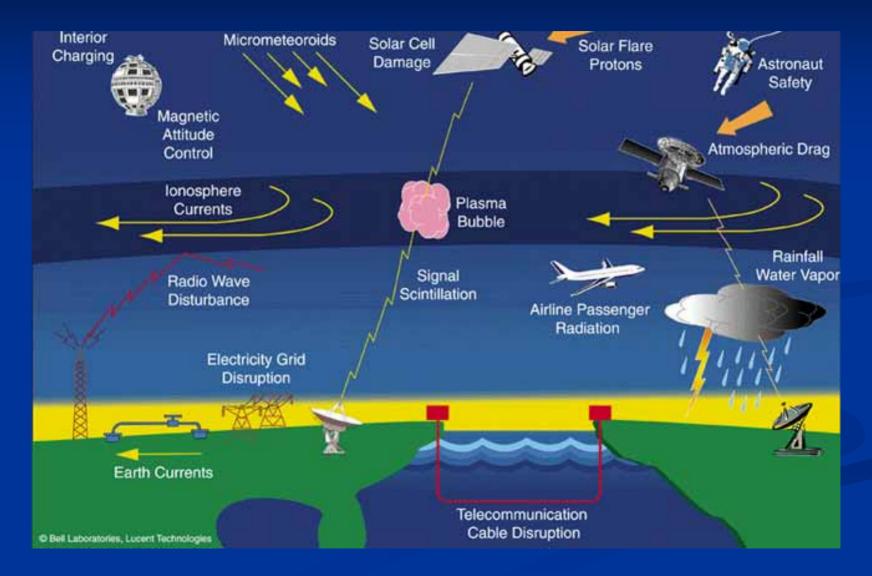




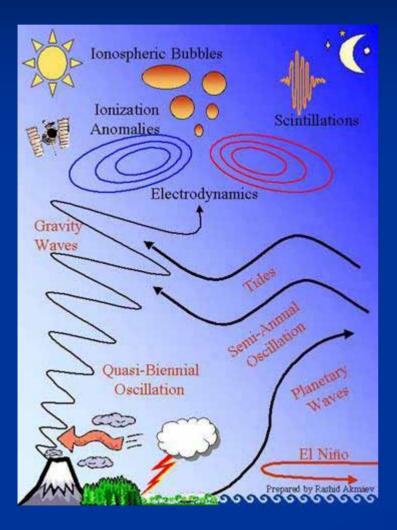
SAPS

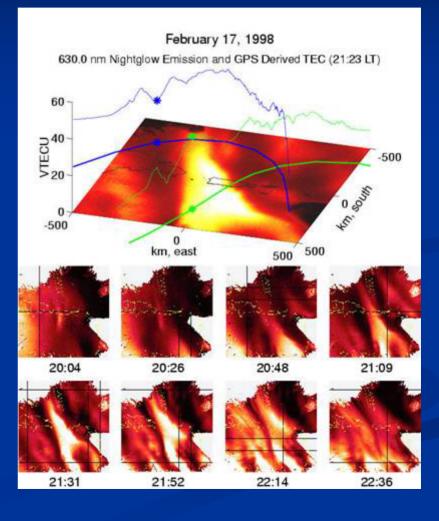


Space Weather



Waves, Processes, and Disturbance Response in the Thermosphere-Ionosphere System





Science (2): Quiescent State of the Ionosphere-Thermosphere (I-T) System

Geospace must be addressed as a system. The coupling among the principal region of geospace is not understood. Disturbances (e.g. SPW event) involve both the characteristics of the external drivers (e.g. IMF/SW) as well as the pre-existing state and structure of the T-I-M regions. The system response throughout SPW events depends in poorly-understood ways on the initial conditions of the geospace system. Multi-technique continuous observations are needed with a spatial resolution appropriate to characterize these systems. global coverage with near RT data. The coupling among the principal T-I-M regions (in both quiet and disturbed conditions) must be addressed and the feedback and/or control of one region on processes in the others is to be addressed.

 Questions: How does the global TI system respond to geomagnetic storms? How does the global response vary with altitude and time? What are the local/global responses to Solar Proton Events? How deep into the atmosphere do the local and global effects penetrate? How does the thermosphere affect the ionosphere?...the magnetosphere?

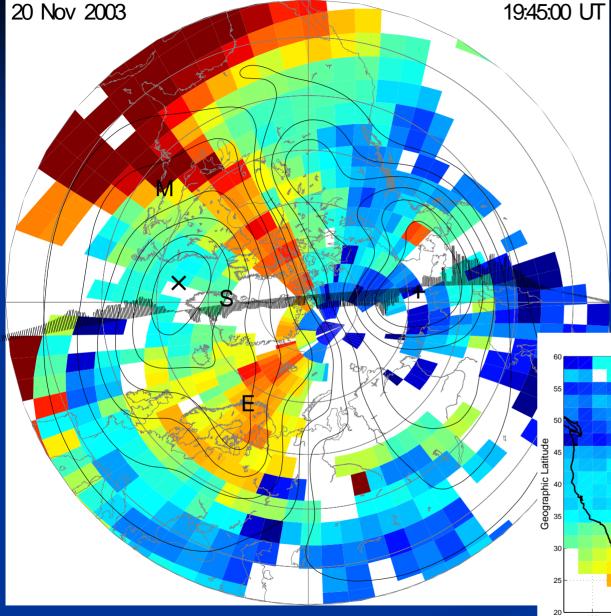
Report Outline (3)

5. Instruments

- 1. Radiowave Instruments
- 2. Magnetometers
- 3. Optical Instruments
- 4. Computer Models
- **6. Existing Arrays**
- **7. Infrastructure Issues**
 - 1. Information Technology
 - 2. Instrument Deployment and Logistics
- 8. Education Opportunities
- 9. Summary and Conclusions
 - Next Steps
- **10.** References



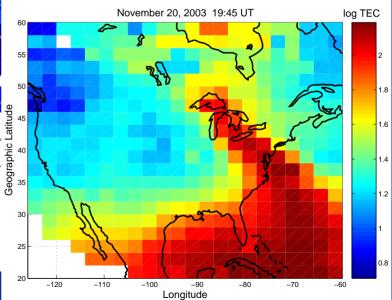
Distributed Instruments (HF Radar: SuperDARN)



Multi-Instrument Clusters to Address Physical Processes

Polar TOI

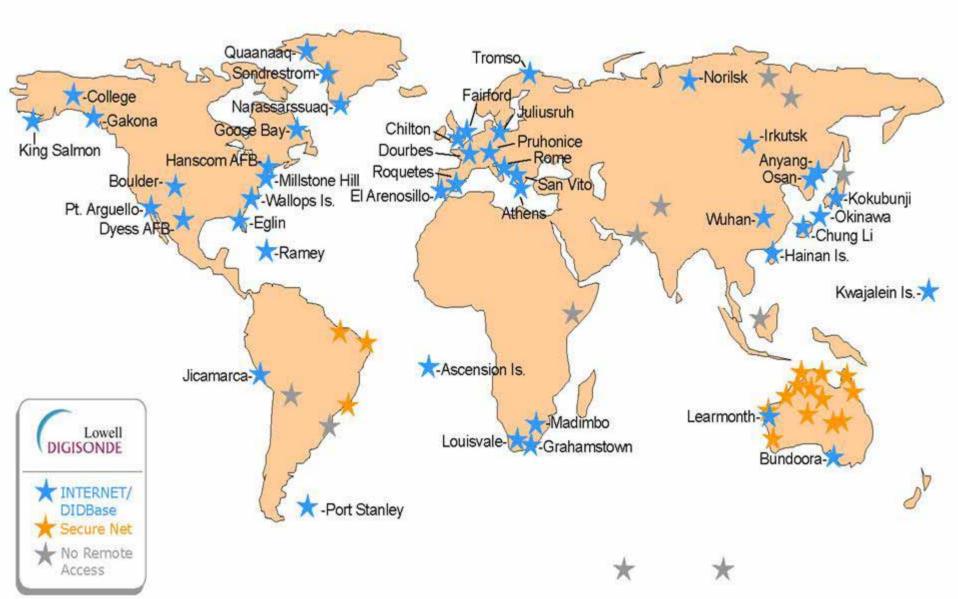
GPS TEC Map

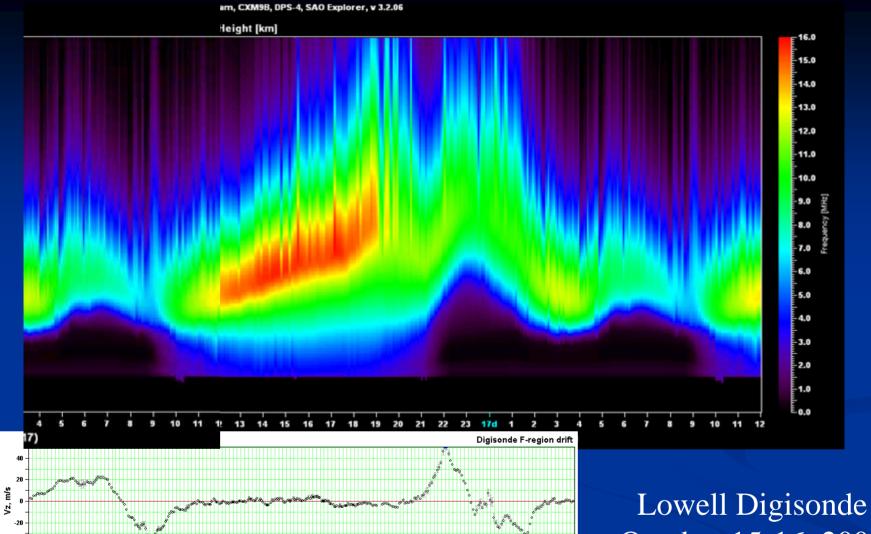


Merged SuperDARN/DMSP Convection

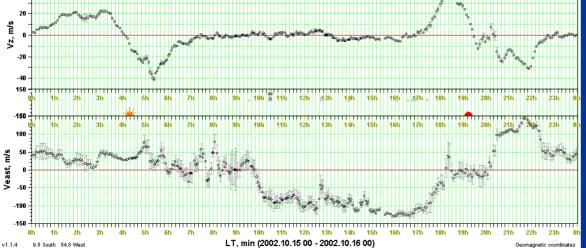
[Foster et al., JGR 2005]

Digisonde Network



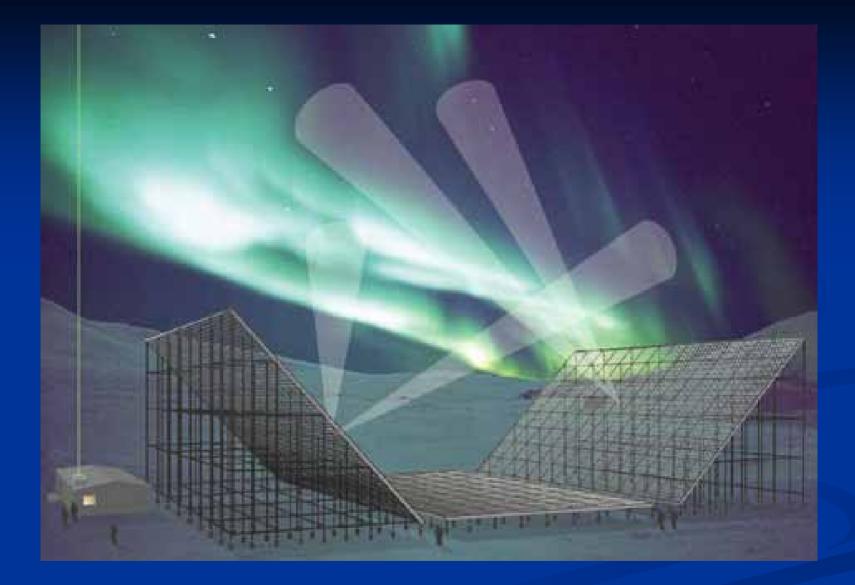








Auroral Processes: Distributed Imagers (Themis)



Large Instruments (Radar/Lidar: AMISR)

New Instruments: ISIS (Intercepted Signals for Ionospheric Science)



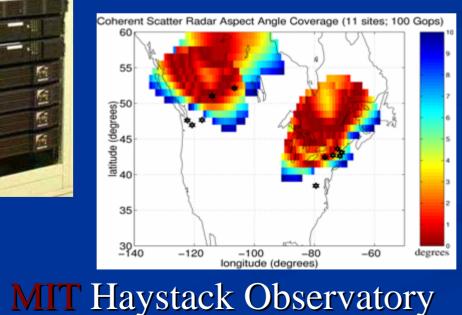
Multirole Coherent Software Radio Network

Multistatic Active and Passive Radar, Radio Scintillation Studies, RF monitoring

Cluster Computer Operational!

ISIS Array Node Assembly Has Started!

First Node Deployment to Greenbank Radio Observatory Summer 2005; Deployment Supported by MIT Lincoln Lab Multistatic Active Radar with MIT Millstone Hill Radar 140 Foot Telescope at Greenbank







Issues for Instrumentation and Arrays

Instrument types and their Deployment should be driven by the needs of the science

 Different processes have 'correlation lengths' which involve different scale sizes. Therefore, DASI sub-arrays will have differing spatial extent.

State of the art IT systems will be needed to realize the DASI architecture

DASI Report: Other Themes

- Education: Distributed instruments and R-T publiclyaccessible data provide extensive opportunities.
- State of the art IT systems will be needed to realize the DASI architecture
- SYNERGY: RT access to different types of data will enable new science arising from their combination.
- Instrument types and their Deployment should be driven by the needs of the science
- Different processes have 'correlation lengths' which involve different scale sizes. Therefore, DASI subarrays will have differing spatial extent.

Where is DASI Headed?

- The time is right for DASI: developing technology and IT systems support a new science capability
- Community involvement and planning workshops are needed to evolve a DASI implementation plan.
- State of the art IT systems will be needed to realize the DASI architecture – DASI must participate and keep abreast of IT development
- A phased, scaled approach to DASI should be adopted

