



WHAT IS GEM?



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GEM Student Representative
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Purpose of GEM: Geospace Environment Modeling



- NSF Division of Atmospheric and Geospace Sciences
- Broad and community-initiated
- Concentration: Earth's magnetosphere
 - Coupling to atmosphere and solar wind
 - Dynamical and structural properties of geospace
- Global Geospace General Circulation (GGCM) model with predictive capability
- Strategy: series of campaigns and focus groups
 - Both theoretical and observational
 - Focusing on particular aspects of the geospace environment

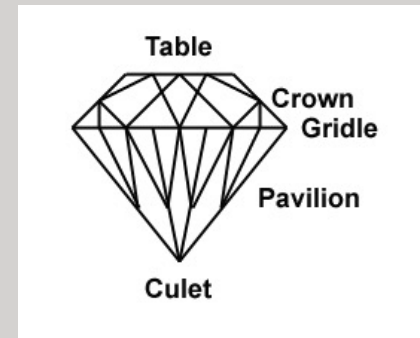
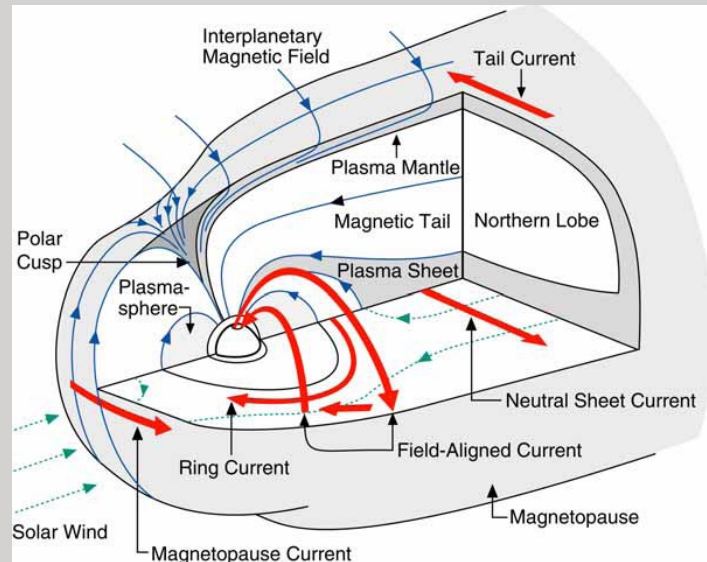


History of GEM



- Global Change Program was a mid-80's multidisciplinary, multiagency response to concerns about threats to the environment from anthropogenic activities
- GEM was the magnetospheric community's proposed contribution
- Proposed in a meeting with the NSF Director in September 1986
- Two options
 - narrowly focused on the contribution of solar irradiance to global change
 - thorough study of general circulation of the magnetosphere
- Physically distinct regions and processes would be studied in campaigns
- Campaigns were to be time limited (3 years) and three could run simultaneously
- A campaign could consist of multiple working groups
- The product of a campaign would be a module for a magnetospheric general circulation model

1. Dayside, including boundary layers and plasma/energy entry.
2. Inner magnetosphere and storms.
3. Tail, including plasma sheet and substorms.
4. Magnetosphere - ionosphere coupling, aurora.
5. Global General Circulation Modeling (GGCM).

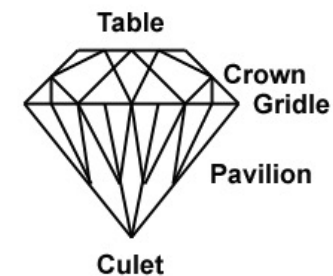
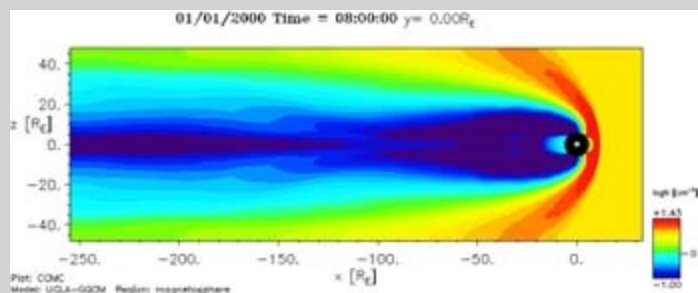
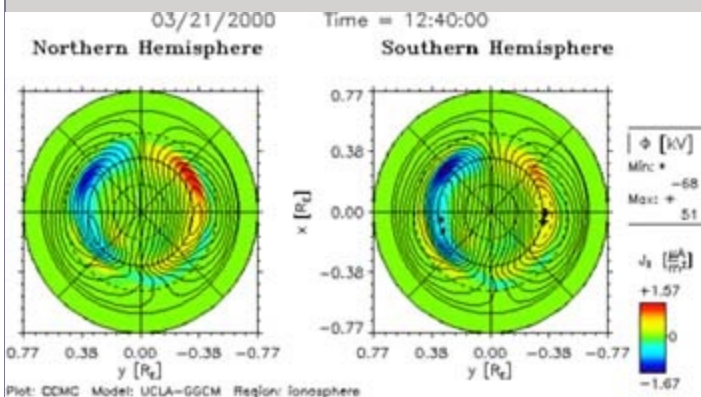




“Gridle” of GEM: GGCM



- Developing, validating, and comparing analytic, first-principles numerical, and data-based models of geospace environment
- Span more than one region
- Global description of the geospace
- Multiple and/or coupled models
- Modeling "challenges" from focus groups

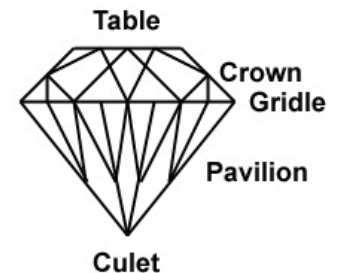




“Pavilion” of GEM: Focus Groups



- Focus Groups (FG) are at the heart of GEM
- Proposed by the community
- Must include deliverable
- Exist within a Research Area
- No more than 5 years
- Target number of active FGs: ~12
- ~2-3 new FGs selected every year





“Pavilion” of GEM: Focus Groups



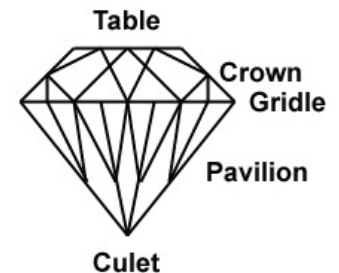
#	Name	Years	RA
1	Plasma Entry and Transport into and within the Magnetotail	2006-2011	Tail
2	Near Earth Magnetosphere: plasma, fields, and coupling	2007-2012	IMS, Tail
3	Space Radiation Climatology	2006-2011	IMS
4	Diffuse Auroral Precipitation	2006-2011	MIC, IMS
5	Plasmasphere-Magnetosphere Interactions	2008-2013	IMS
6	Substorm Expansion Onset: The First 10 Minutes	2008-2013	Tail
7	Modes of Solar Wind-Magnetosphere Energy Transfer	2008-2013	Tail
8	Dayside FACs and Energy Deposition	2010-2012	Dayside
9	Radiation Belts and Wave Modeling	2010-2014	IMS
10	The Magnetosheath	2010-2014	Dayside
11	Metrics and Validation	2011-	GGCM
12	The Ionospheric Source of Magnetospheric Plasma— Measuring, Modeling, and Merging into the GEM GGCM	2011-	
13	Scientific Magnetic Mapping and Techniques	2011-	



“Table” of GEM: Annual Meeting



- First GEM Meeting: September 23-25, 1991 at UCLA
 - Happy 20th Anniversary!
- Annual week-long summer meeting (usually Snowmass, CO)
- Student tutorial workshop on Sunday
- Banquet on Wednesday (Monday this year)
- Poster sessions Tuesday and Thursday
- Daily activities:
 - Morning tutorials
 - 3 focus group session slots
 - 3-4 focus groups at a time
- Steering committee meets Friday afternoon





“Cutlet” of GEM: Between Meetings



- GEM Mini-workshop in December
 - Sunday before AGU
 - 2 hr working group meetings for FGs
 - Steering Committee discusses and votes on proposed FGs
- Communication
 - Messenger (electronic). Brief Notices to the GEM community.
 - ✦ Email Majordomo@igpp.ucla.edu with the message “subscribe gem”
 - ✦ <http://heliophysics.blogspot.com/>
 - GEMstone. Annual summary of GEM activity centering on the June meeting. Attempts to summarize activities of all working groups. GEM’s most detailed historical archive
 - ✦ <http://aten.igpp.ucla.edu/gemwiki/index.php/GEMstone>

