

# OpenGGCM-CTIM Study of Ionosphere and Thermosphere Energy Deposition under Northward IMF Condition

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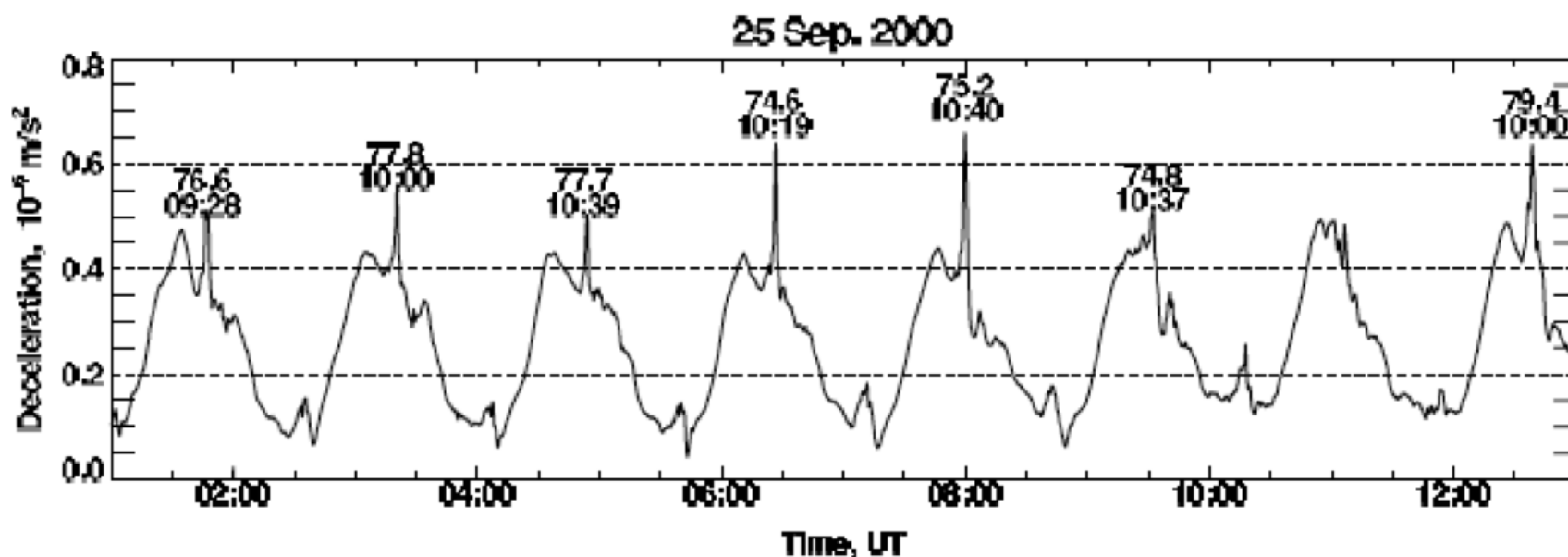
GEM-CEDAR  
*Boulder , 2013*

# Outline

1. Challenging Minisatellite Payload (CHAMP) satellite observation of thermospheric density enhancement anomaly.
2. DMSP observation of strong Poynting flux.
3. OpenGGCM results.

# High-latitude local thermosphere density enhancements

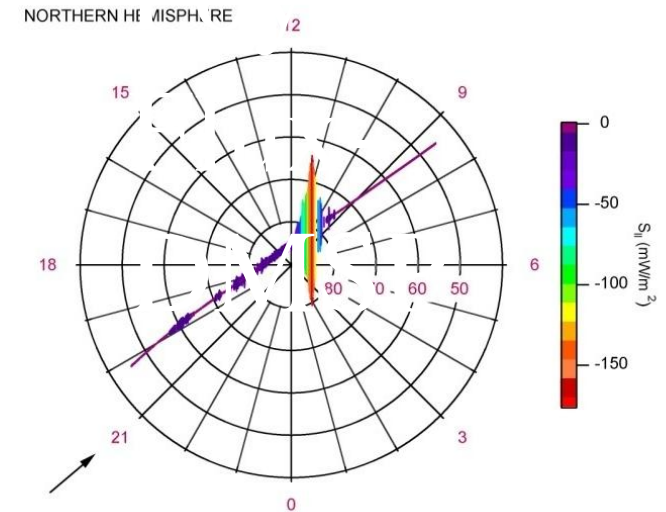
*Lühr et al.* [2004] showed that, under relatively quiet geomagnetic conditions, the CHAMP satellite often observes regions of enhanced density at ~ 400 km altitude in the noon sector at high latitudes correlated with small scale field-aligned currents (FACs) associated with the dayside cusp.



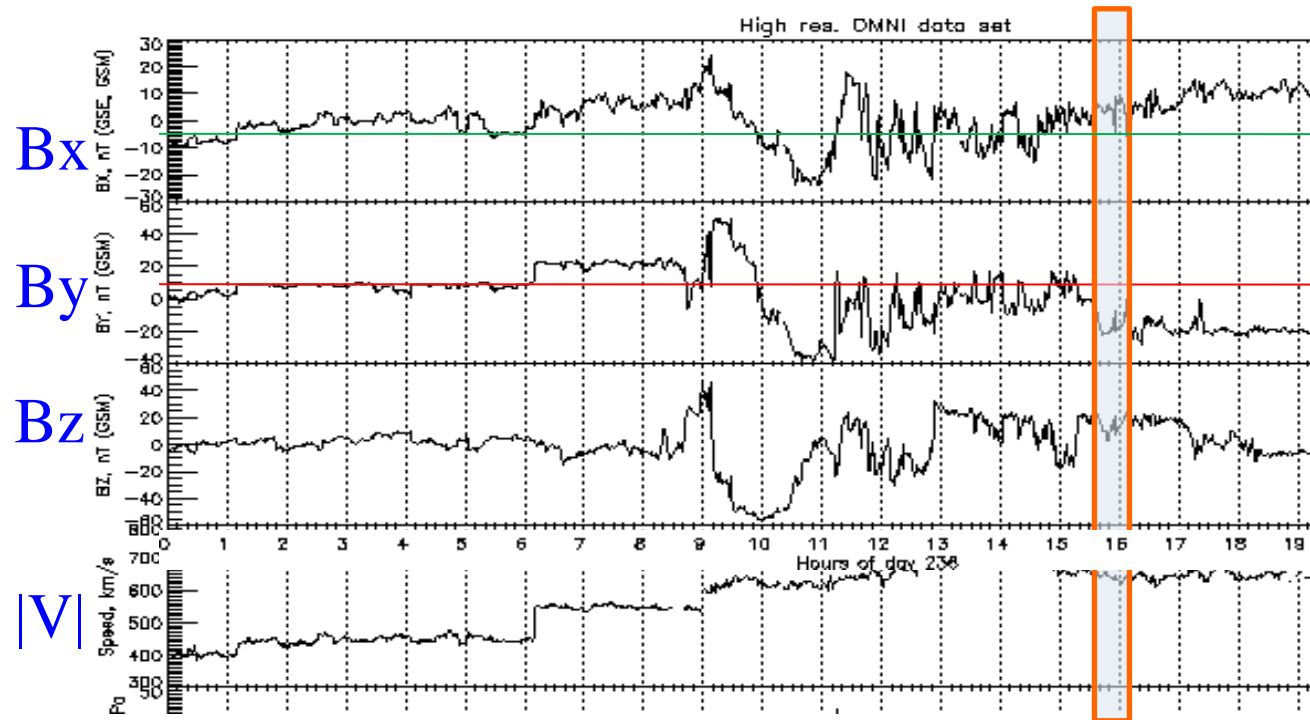
Air drag measured by the accelerometer on board CHAMP. The harmonic variations indicate the range of change over an orbit. Small-scale features are superimposed. The peaks in air drag are labeled by their corrected magnetic latitude and magnetic local time. (Adapted from *Lühr et al.*, 2004.)

# DMSP Enhanced Poynting Flux (Knipp et al., 2011)

- Strong, localized Poynting flux is observed,
- near cusp region,
- during northward IMF with strong  $B_y$  component (quiet magnetosphere,  $K_p < 2$ ).

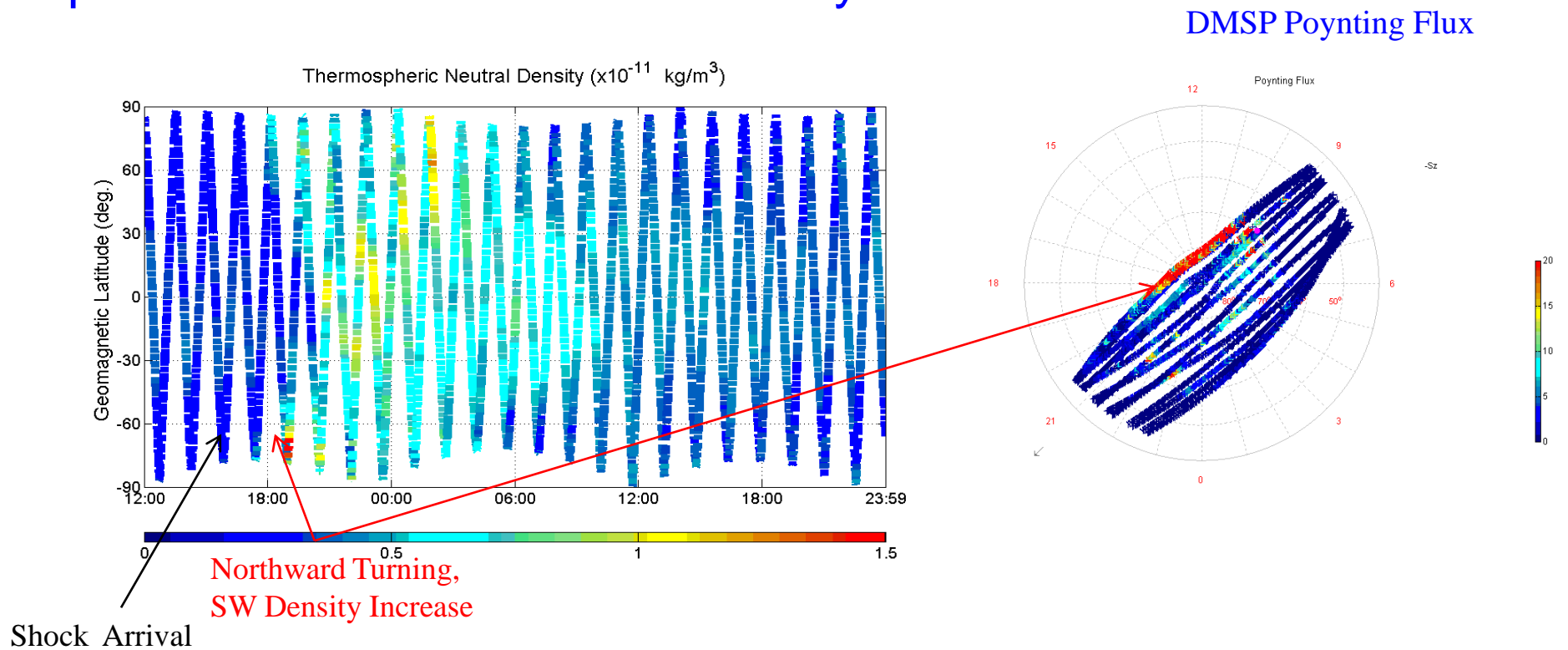


## IMF and Solar Wind



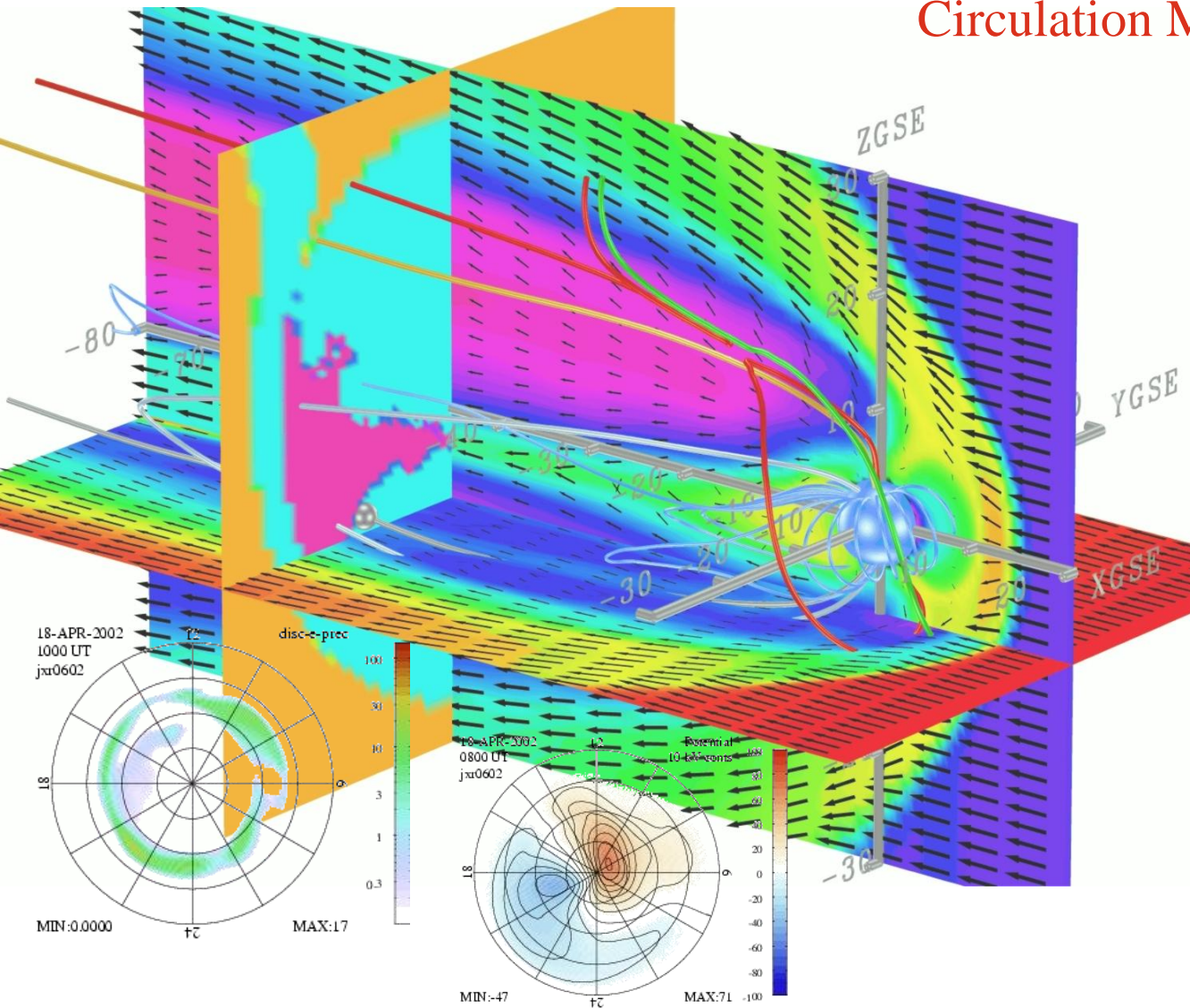
# CHAMP Data: Neutral Density Enhancement

- The localized energy input has a profound effect on neutral density.



# OpenGGCM: Global Magnetosphere Modeling

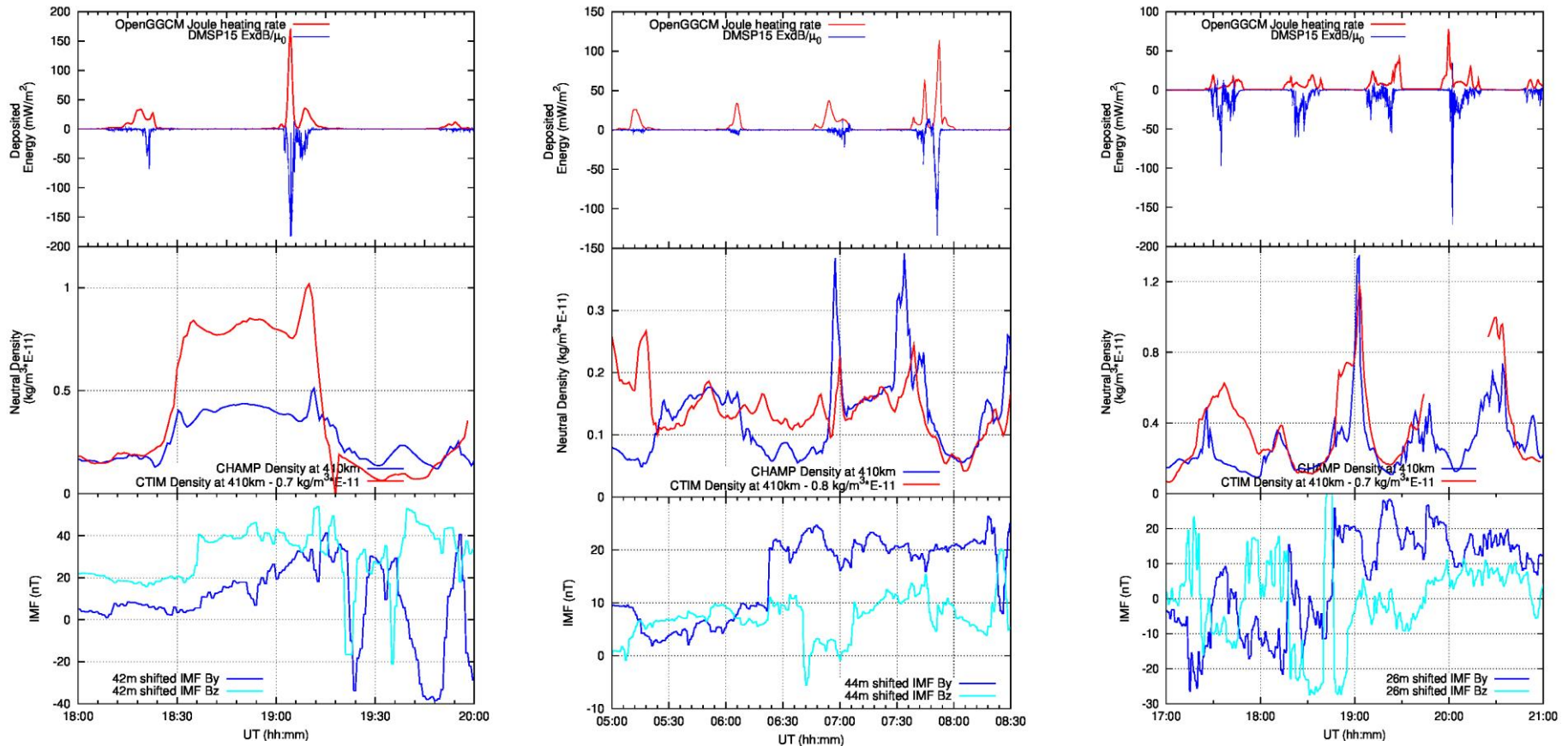
## Open Geospace General Circulation Model



- 3d Magnetohydrodynamic magnetosphere model.
- Coupled with NOAA/SEC (*Fuller-Rowell*) 3d dynamic/chemistry ionosphere - thermosphere model (CTIM).
- Coupled with inner magnetosphere / ring current models: Rice U. RCM, NASA/GSFC CRCM.
- Model runs on demand (>300 so far) provided at the Community Coordinated Modeling Center (CCMC at NASA/GSFC).  
<http://ccmc.gsfc.nasa.gov/>
- Fully parallelized code, real-time capable. Runs on IBM/datastar, IA32/I64 based clusters, PS3 clusters, and other hardware.
- Used for basic research, numerical experiments, hypothesis testing, data analysis support, NASA/THEMIS mission support, mission planning, space weather studies, and Numerical Space Weather Forecasting in the future.
- Funding from NASA/LWS, NASA/TR&T, NSF/GEM, NSF/ITR, NSF/PetaApps, AF/MURI programs.

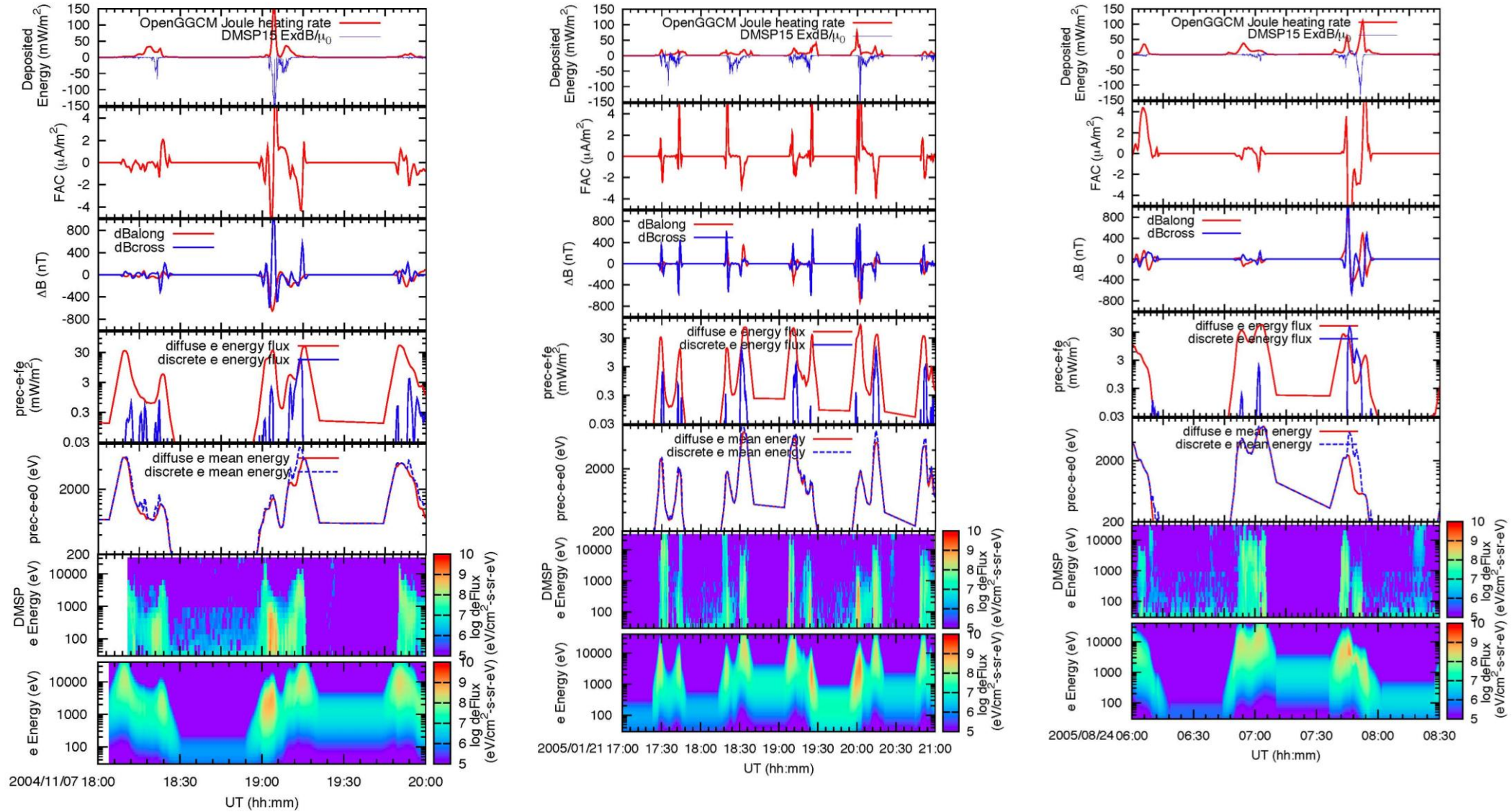
**Personnel:** J. Raeder, D. Larson, W. Li, A. Vapirev, K. Germaschewski, L. Kepko, H.-J. Kim, M. Gilson, B. Larsen, Y. Ge (UNH), T. Fuller-Rowell, N. Muriyama (NOAA/SEC), F. Toffoletto, A. Chan, B. Hu (Rice U.), M.-C. Fok (GSFC), A. Richmond, A. Maute (NCAR)

- DMSP (Knipp et al 2011) observations show very high Poynting flux near the cusps.
- CHAMP observes regions of strong, localized neutral density enhancements.
- IMF is northward, with large  $B_y$  component, geomagnetically quiet time.
- OpenGGCM-CTIM simulations of events reproduce both Poynting flux and neutral density density “hot spots” for all 3 cases.



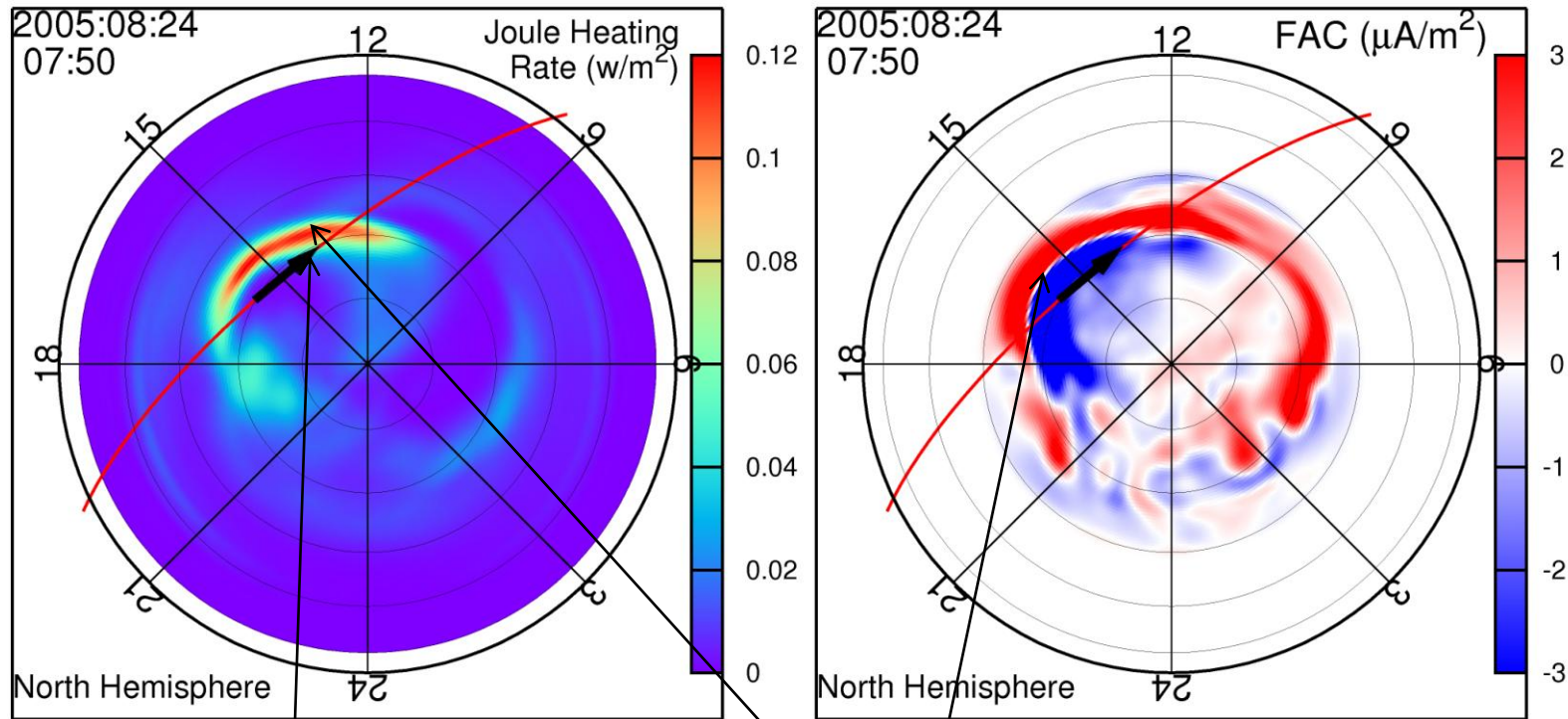
From top to bottom, the panels are: (negative downward) DMSP Poynting flux (blue) and OpenGGCM Joule heating rate (red, they should be equal in magnitude by Poynting's theorem), CHAMP density (blue) and OpenGGCM-CTIM density at CHAMP (red), and IMF  $B_y$  and  $B_z$ .

# OpenGGCM vs DMSP



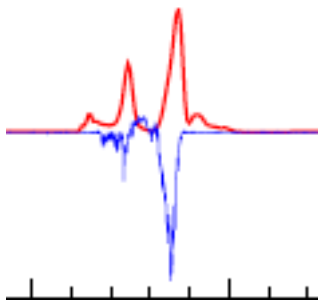


# Intense Joule heating and FAC regions

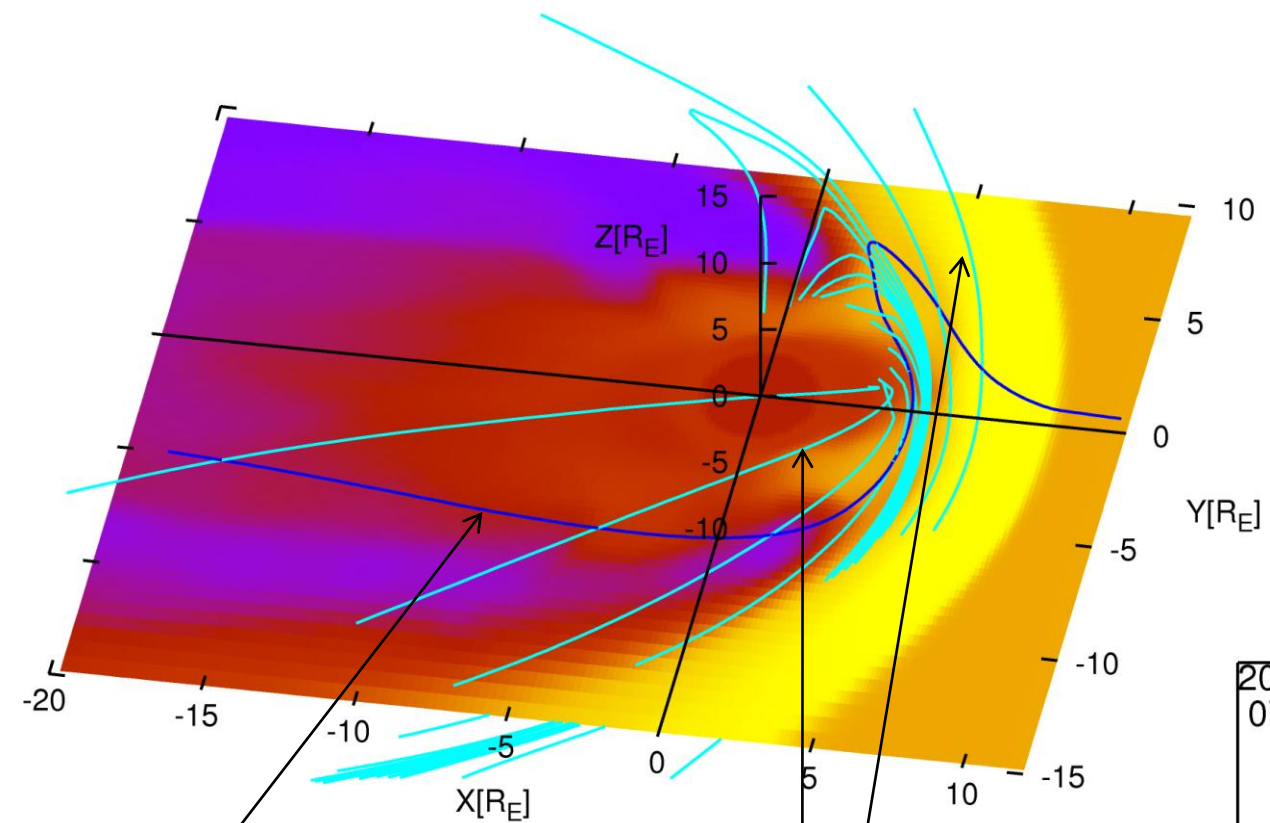


DMSPf15 at 07:50 UT

Intense Joule heating (Poynting flux) region flanked by a pair of opposite FACs (+:down)



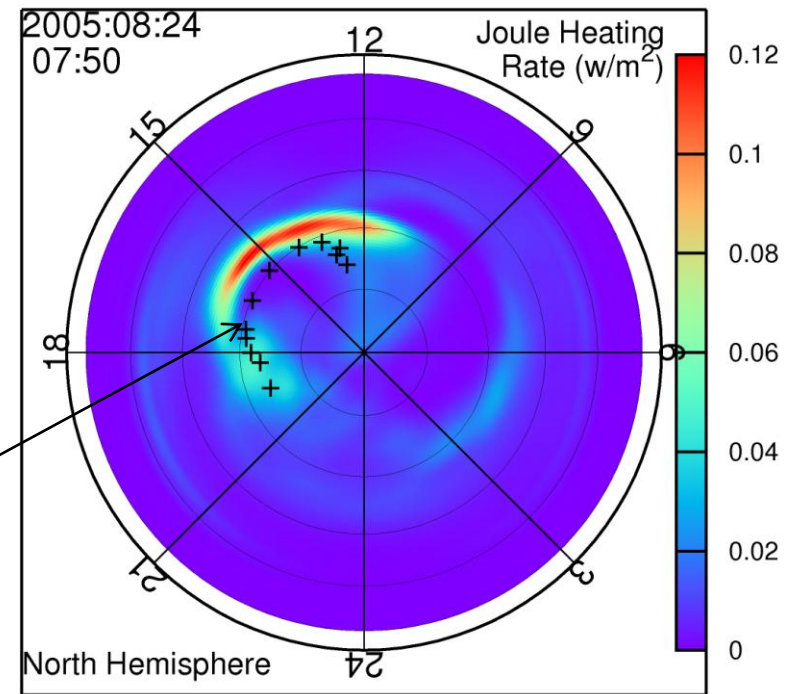
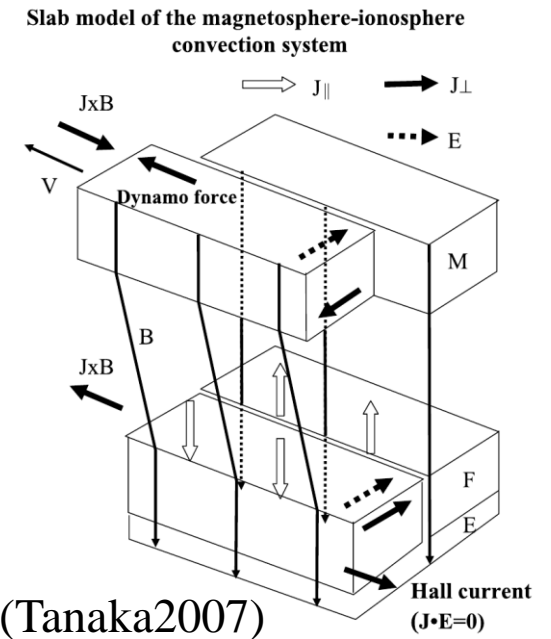
# Moving open field lines resulting from cusp reconnection



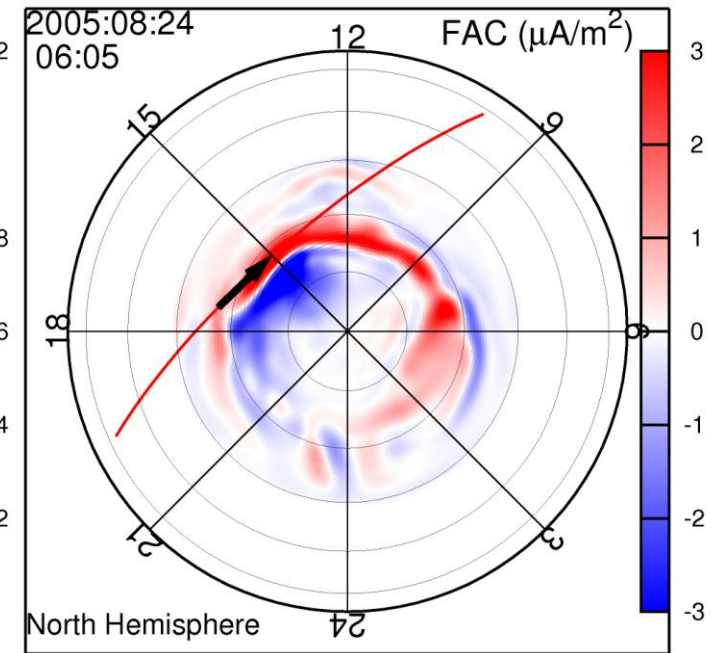
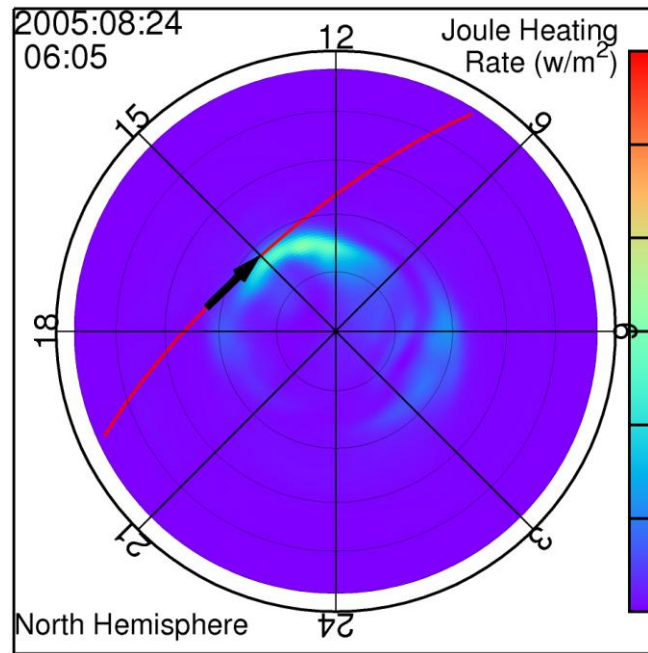
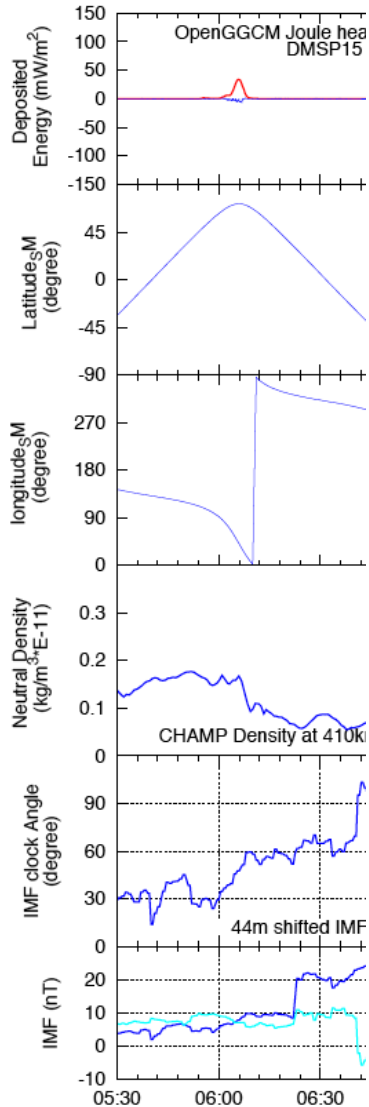
Path of a fluid element entering magnetosphere

Field line attached to this fluid element

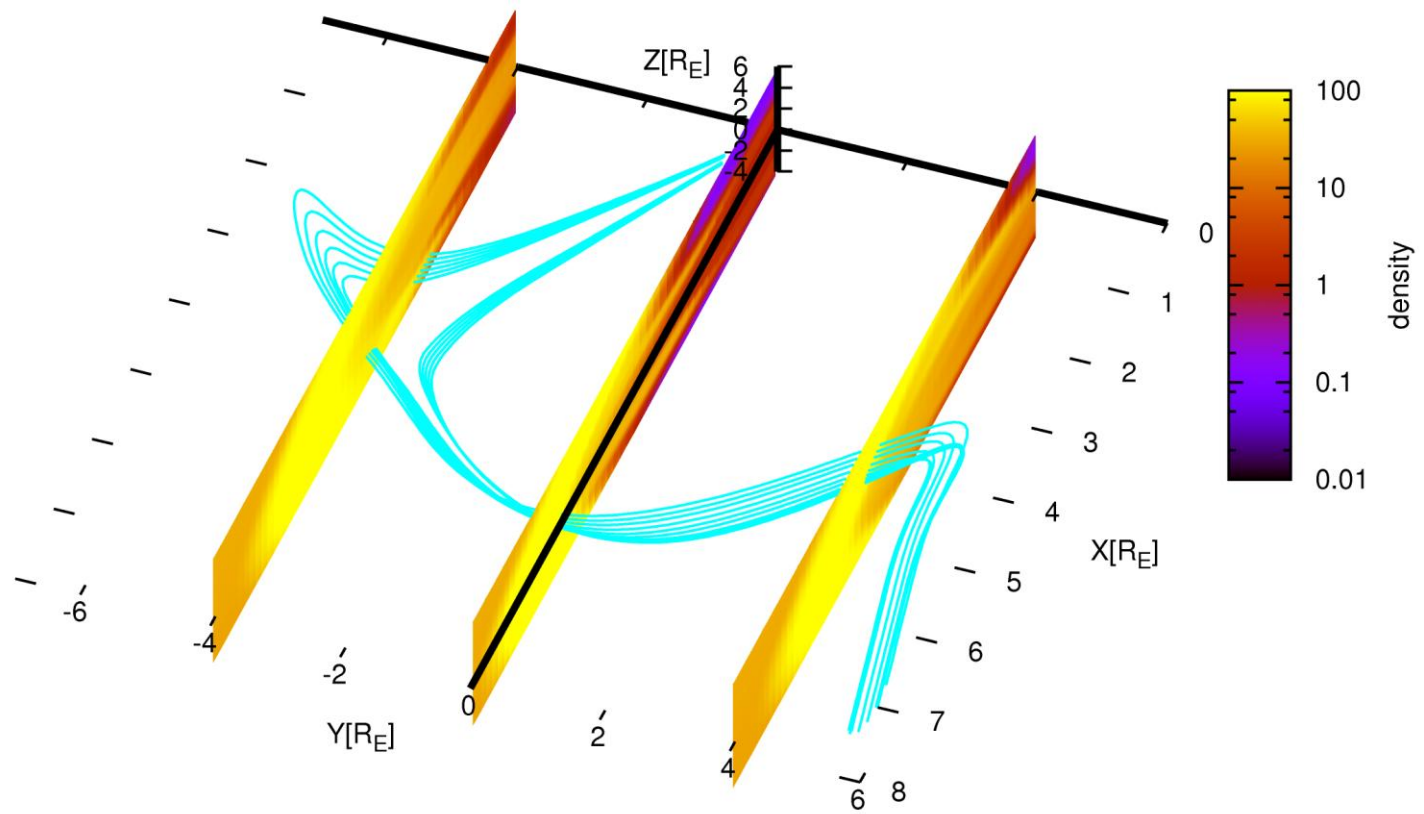
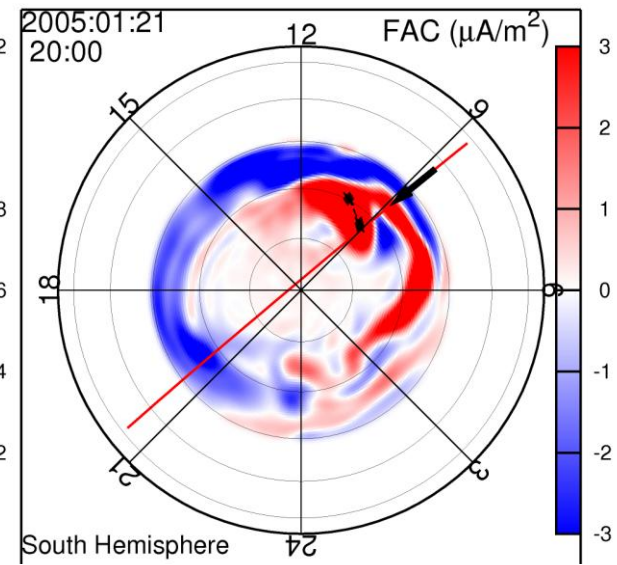
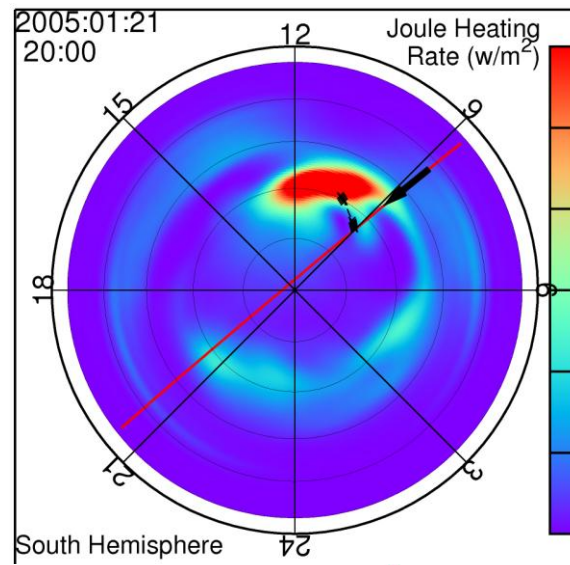
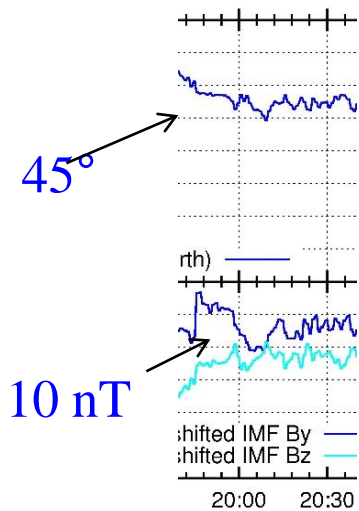
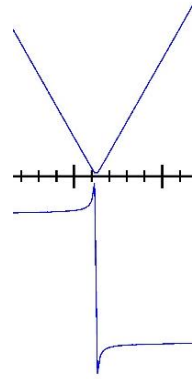
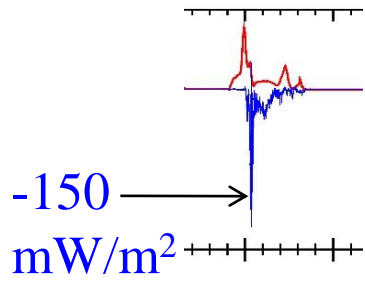
Foot points of these field lines



# Small IMF clock angle and moderate IMF magnitude



# Southern Hemisphere on 21 January 2005



# Joule Heating Hot Spot and Thermosphere Density Enhancement

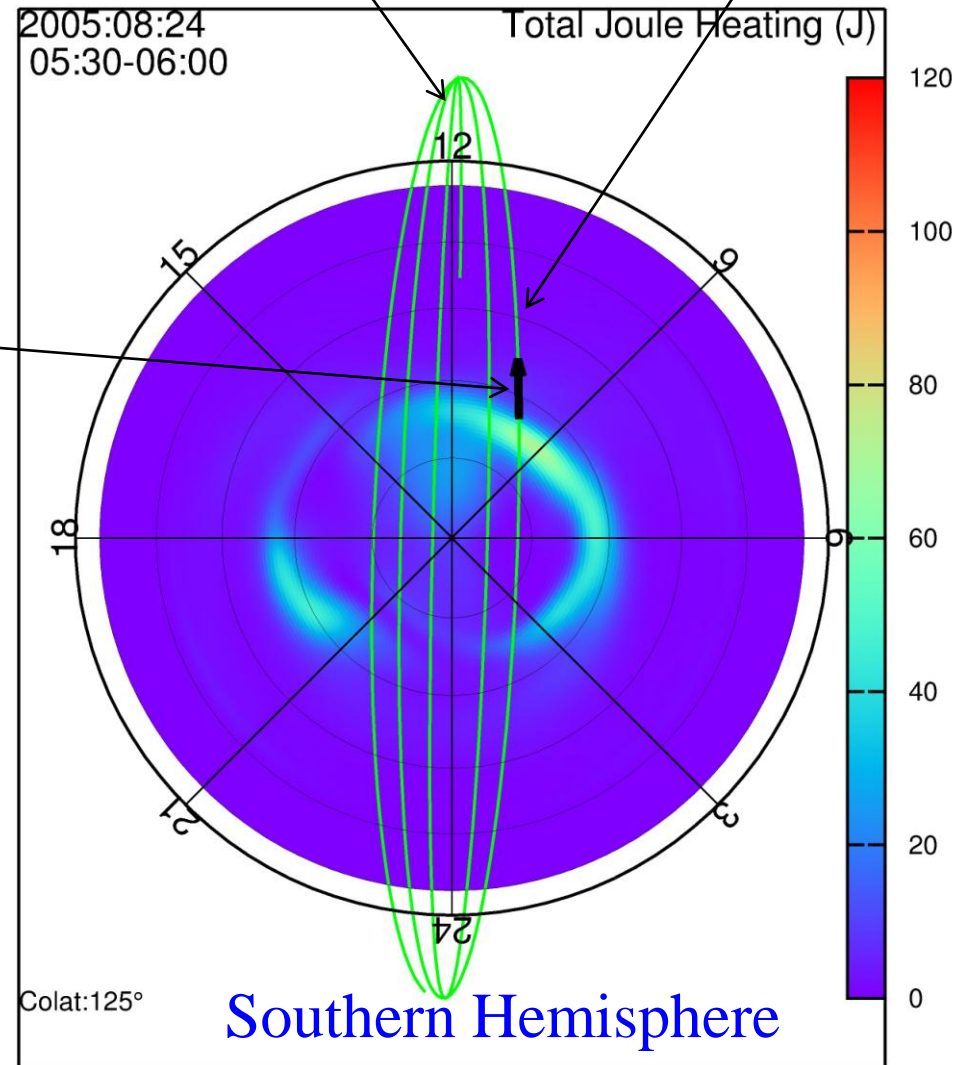
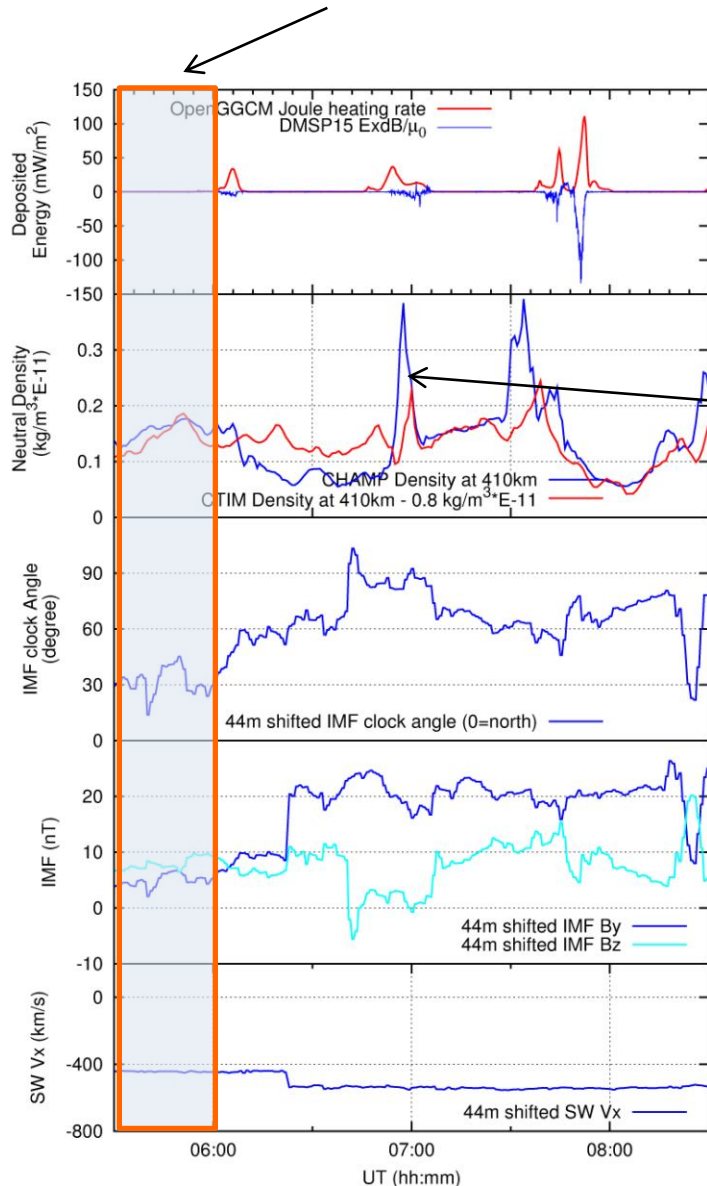
Event 2005-08-24

Moderate JH spot

Moderate N. IMF

CHAMP track UT 0400-0800

CHAMP southern pass



# Joule Heating Hot Spot and Thermosphere Density Enhancement

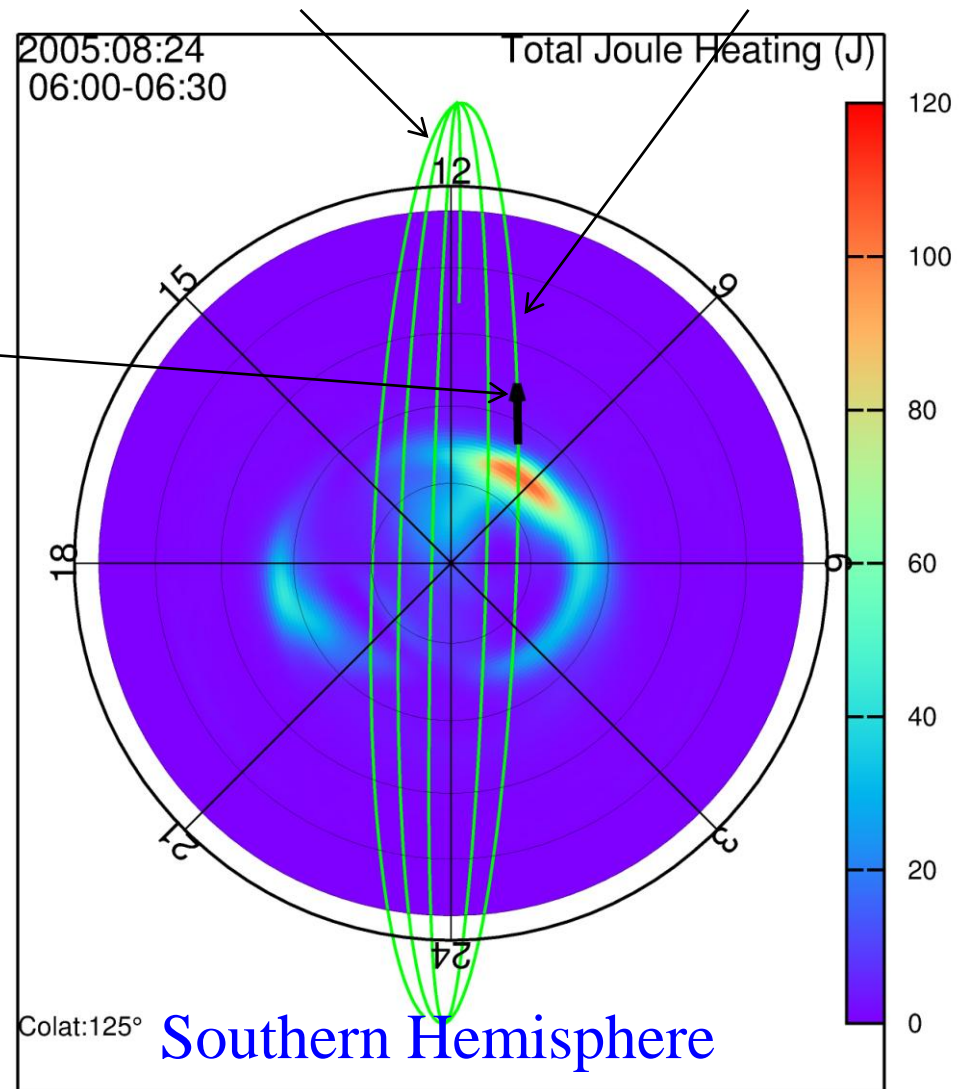
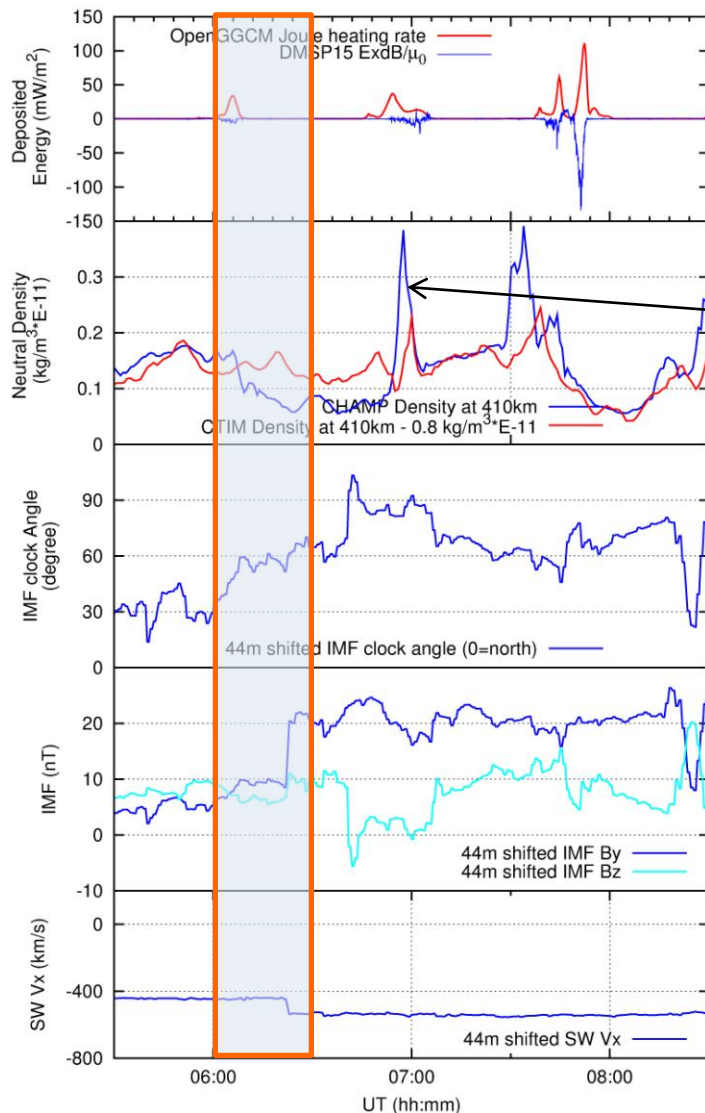
Event 2005-08-24

Hot JH spot

Large N. IMF By

CHAMP track UT 0400-0800

CHAMP southern pass



# Joule Heating Hot Spot and Thermosphere Density Enhancement

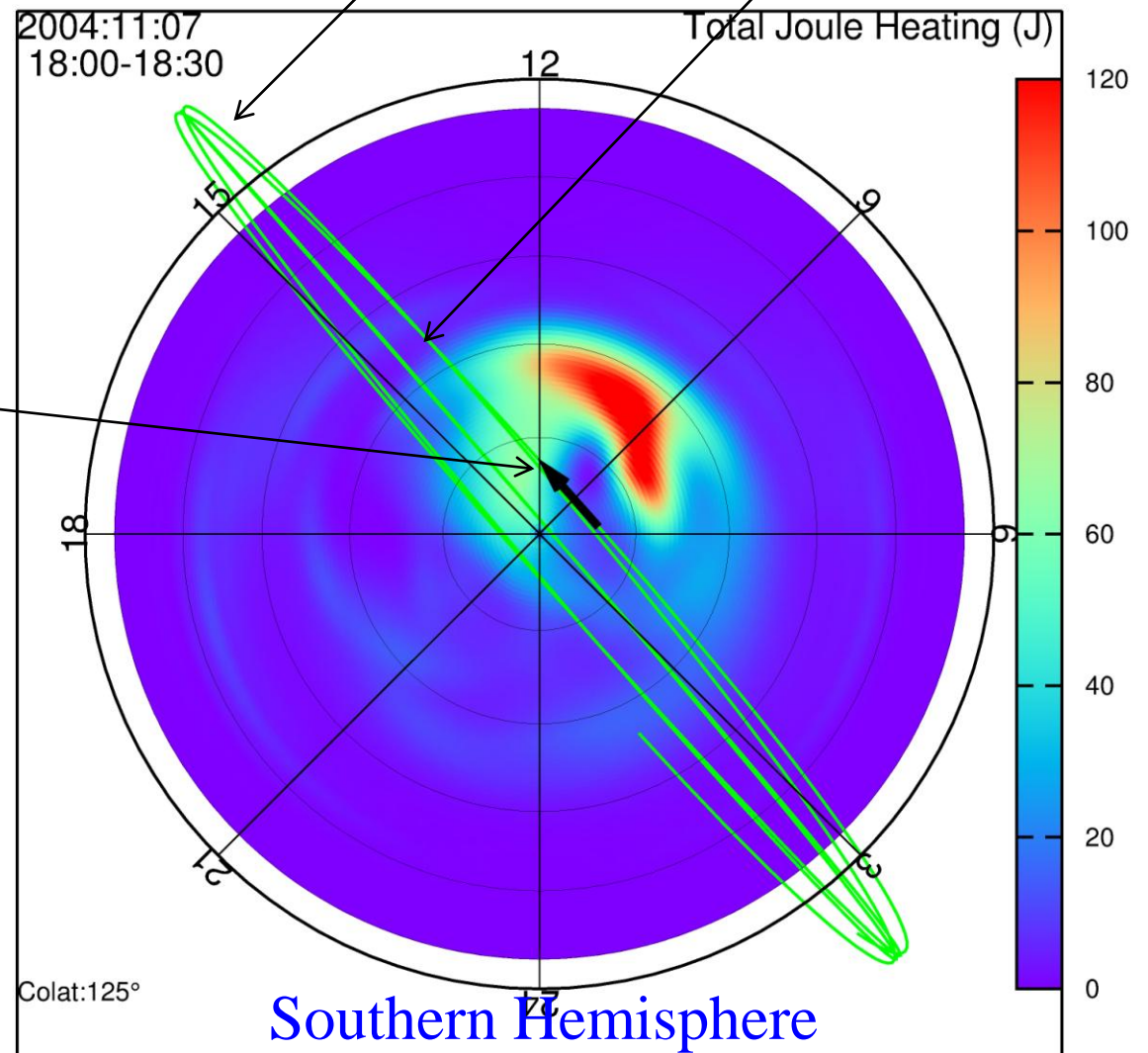
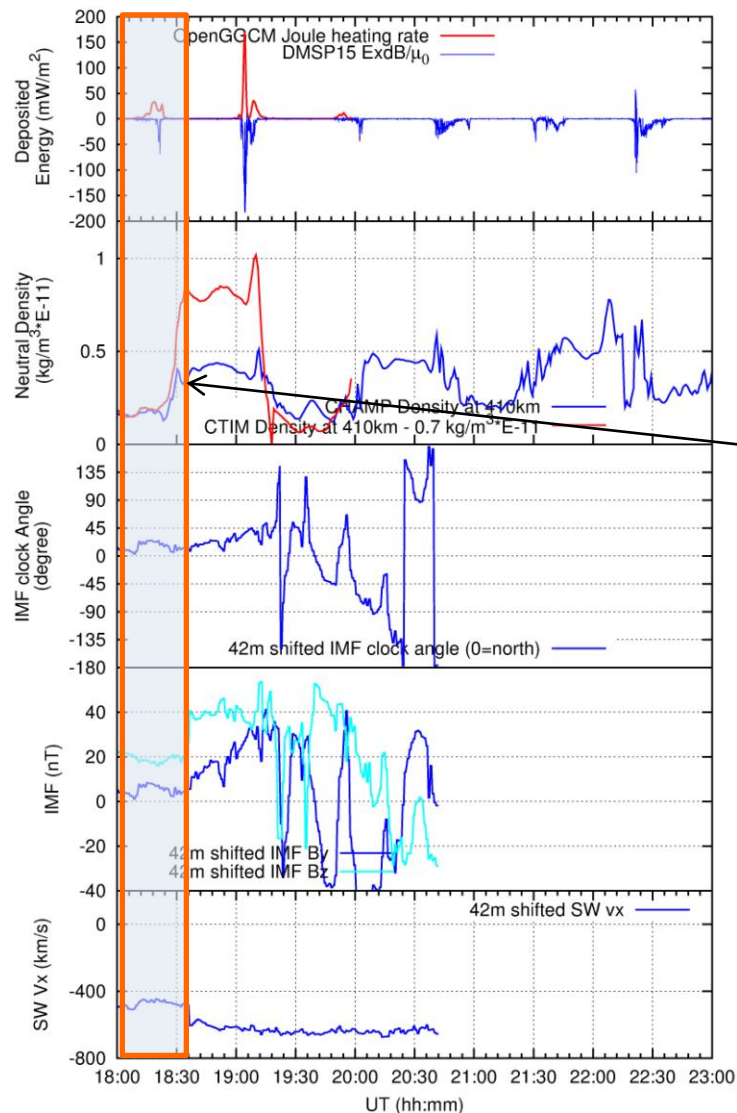
Event 2004-11-07

Hot JH spot

Strong N. IMF

CHAMP track UT 1800-2300

CHAMP southern pass



# Joule Heating Hot Spot and Thermosphere Density Enhancement

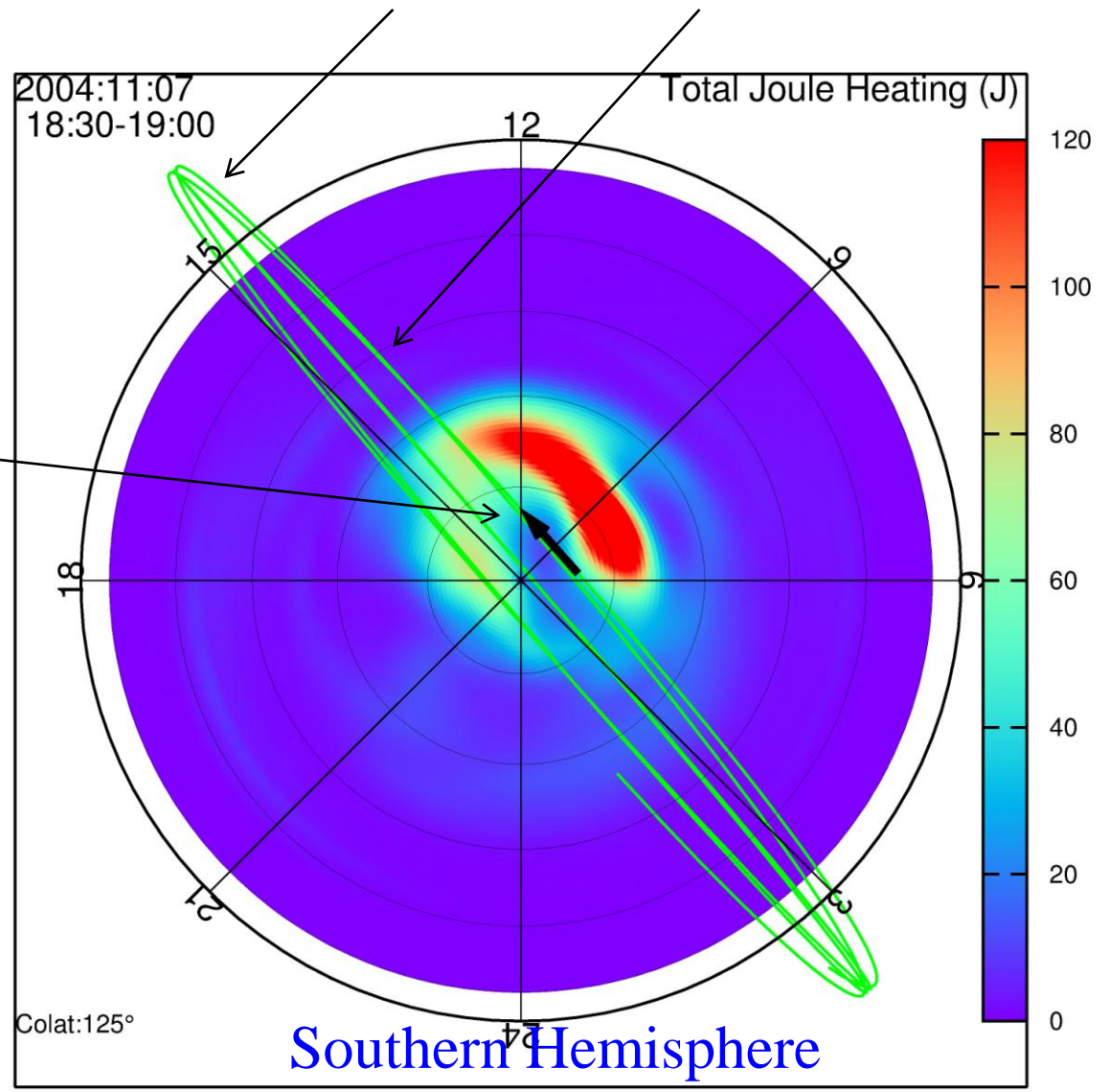
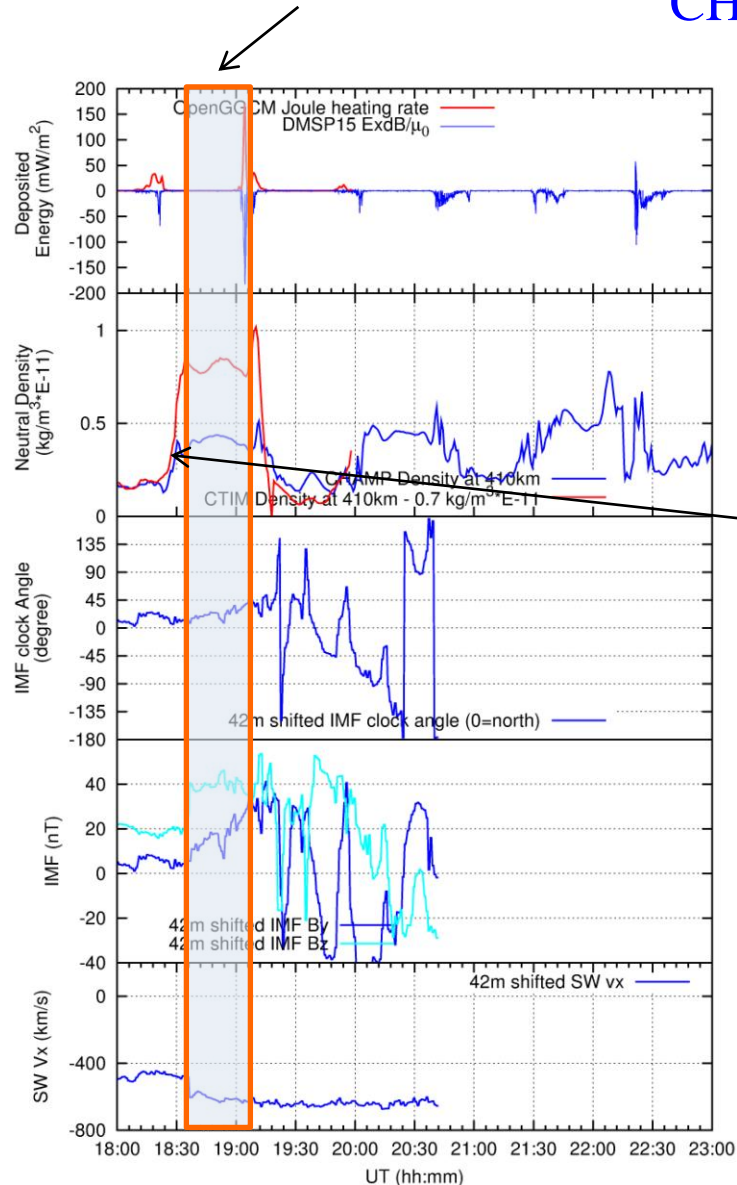
Event 2004-11-07

Hot JH spot

Extremely strong N. IMF

CHAMP track UT 1800-2300

CHAMP southern pass





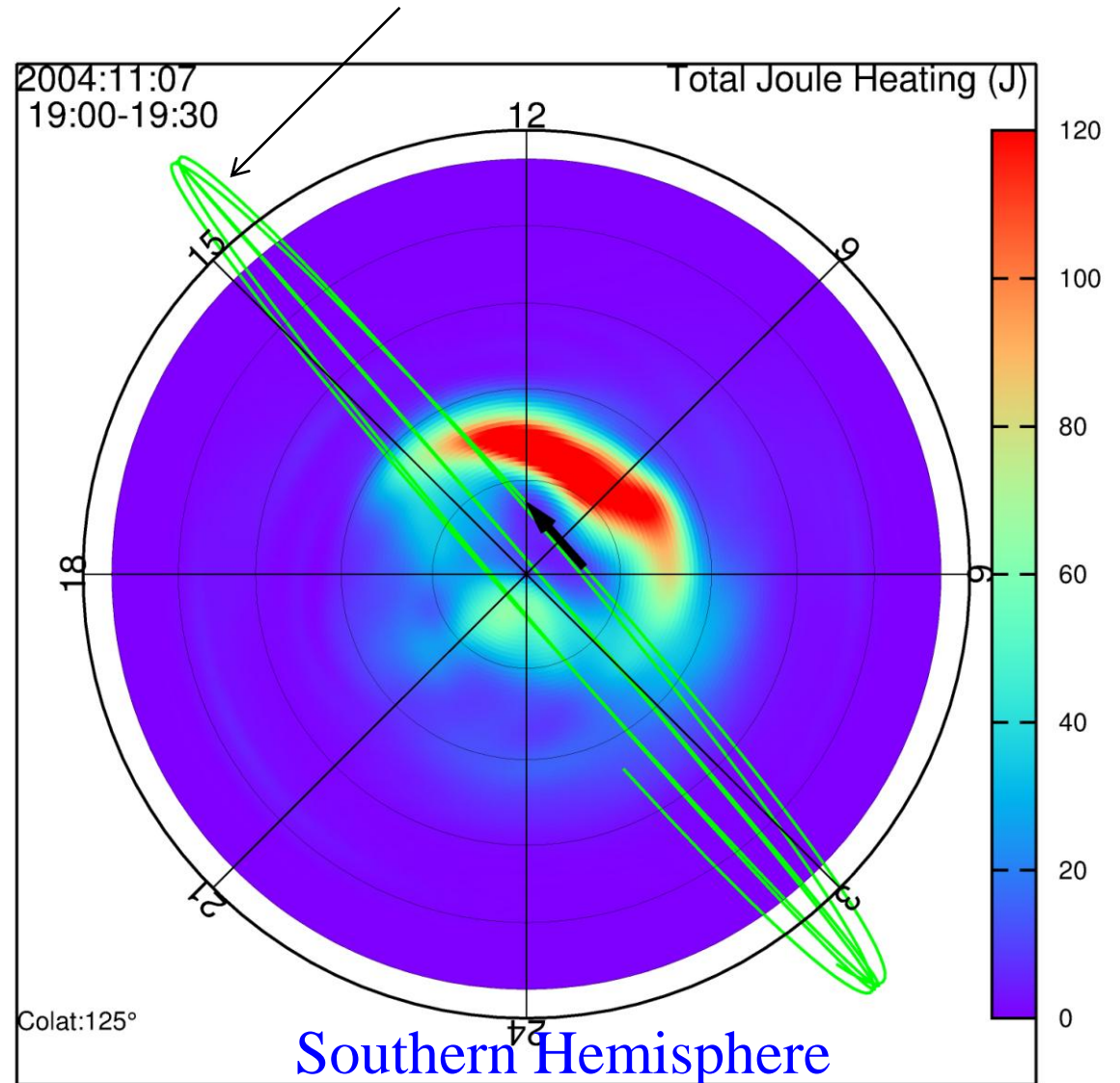
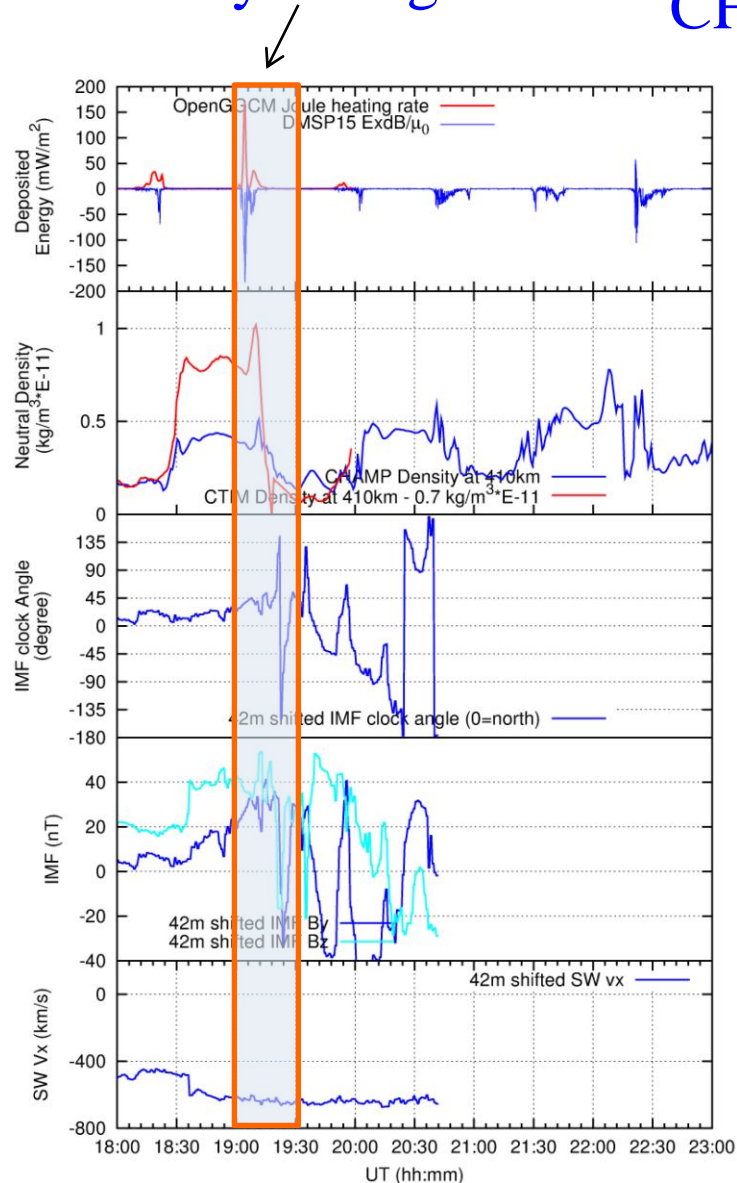
# Joule Heating Hot Spot and Thermosphere Density Enhancement

Event 2004-11-07

Hot JH spot

Extremely strong N. IMF

CHAMP track UT 1800-2300



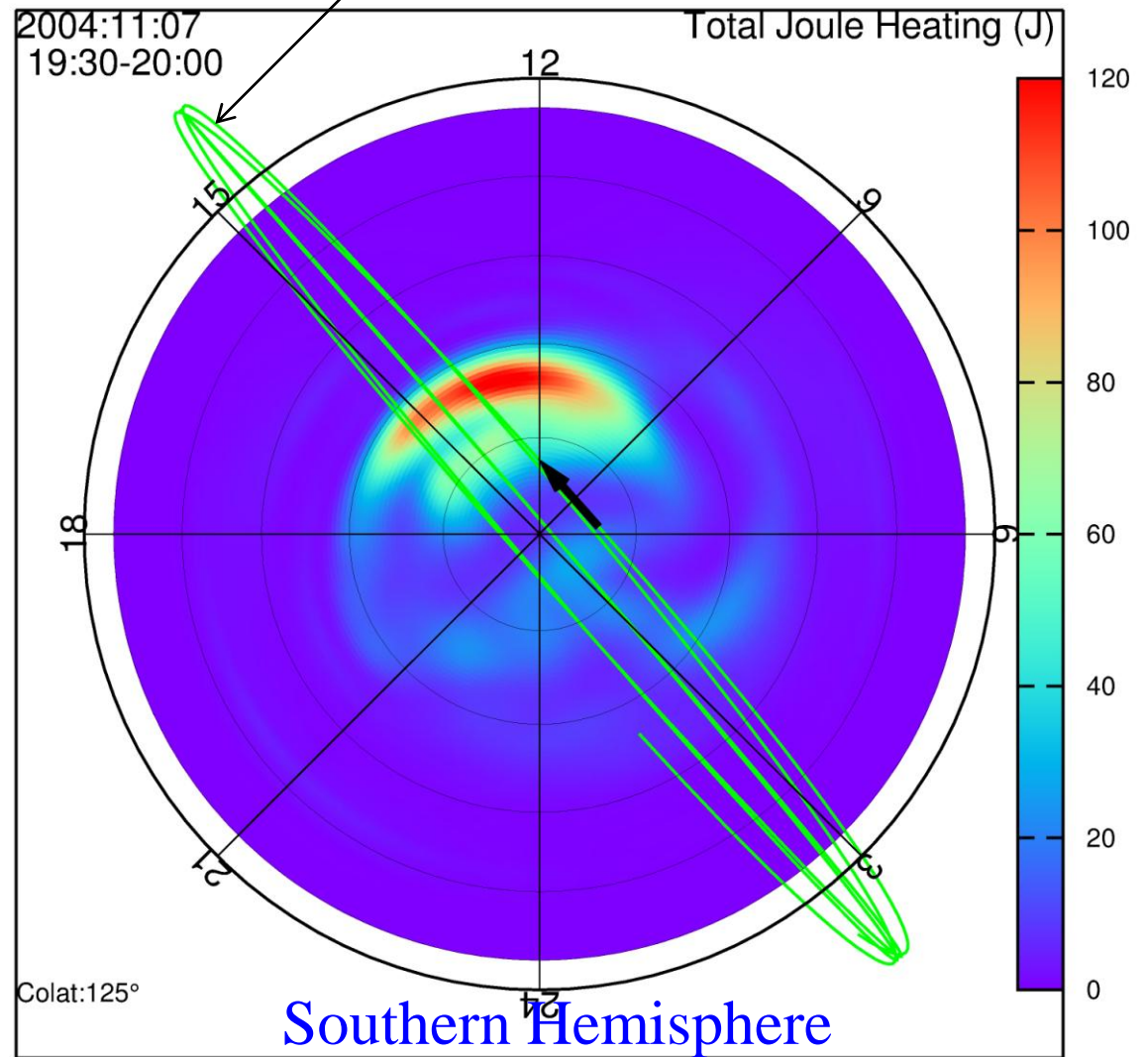
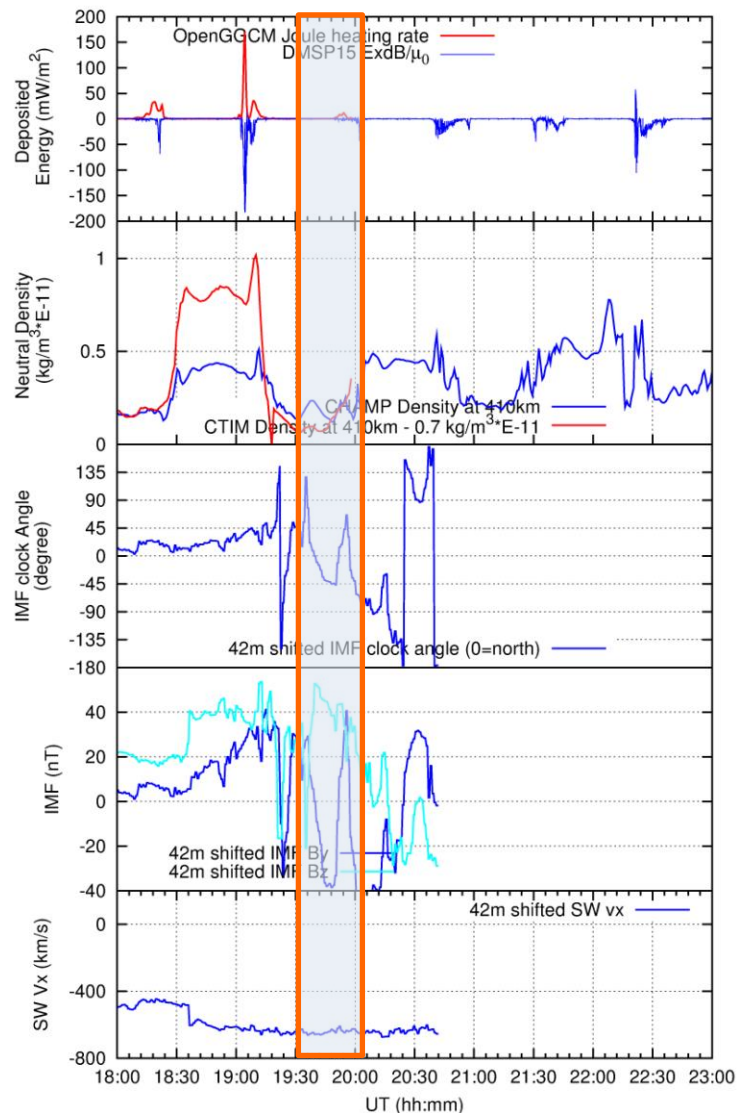
# Joule Heating Hot Spot and Thermosphere Density Enhancement

Event 2004-11-07

Hot JH spot

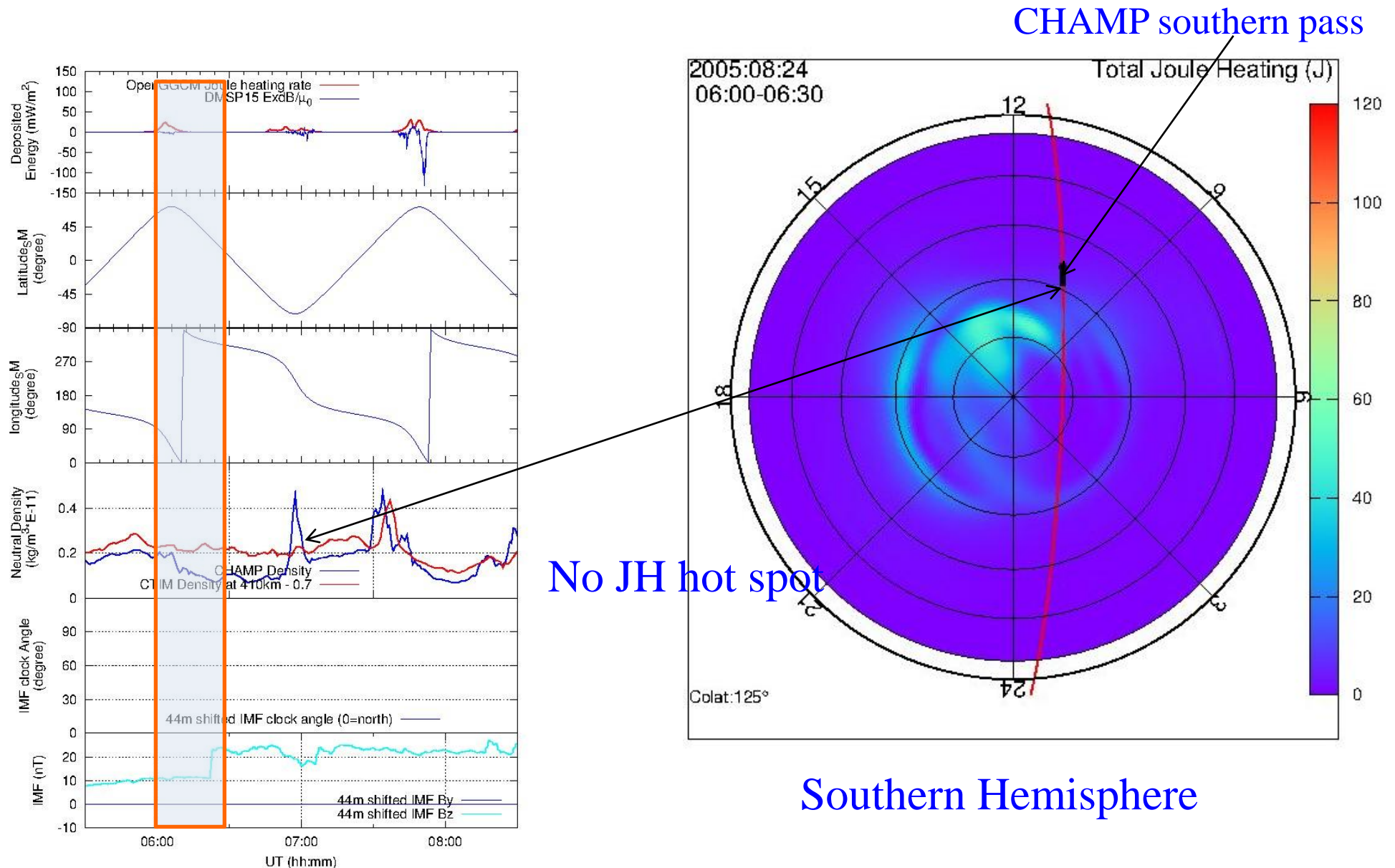
strong N. IMF

CHAMP track UT 1800-2300



# Joule Heating Hot Spot and Thermosphere Density Enhancement

Event 2005-08-24, simulated with 0 clock angle



# Summary

- Strong northward IMF with large clock angle ( $> \sim 40^\circ$ ) causes an extended latitudinal intense Joule heating (Poynting flux) and FAC region mostly in the dayside.
- This region is caused by the movement of the newly created open field lines resulting from cusp reconnection.
- The local high-latitude thermosphere density enhancements are highly correlated with the intense ionosphere Joule heating caused by cusp reconnection when the IMF is northward and has strong  $y$  component.

# References

Wenhui Li, Delores Knipp, Jiuhou Lei, and Joachim Raeder, "The relation between dayside local Poynting flux enhancement and cusp reconnection", JOURNAL OF GEOPHYSICAL RESEARCH, VOL. 116, A08301, doi:10.1029/2011JA016566, 2011

D. Knipp, S. Eriksson, L. Kilcommons, G. Crowley, J. Lei, M. Hairston and K. Drake, "Extreme Poynting flux in the dayside thermosphere: Examples and statistics", GEOPHYSICAL RESEARCH LETTERS, VOL. 38, L16102, doi:10.1029/2011GL048302, 2011