

Dayside transients: Their importance and observational campaigns and missions to address them

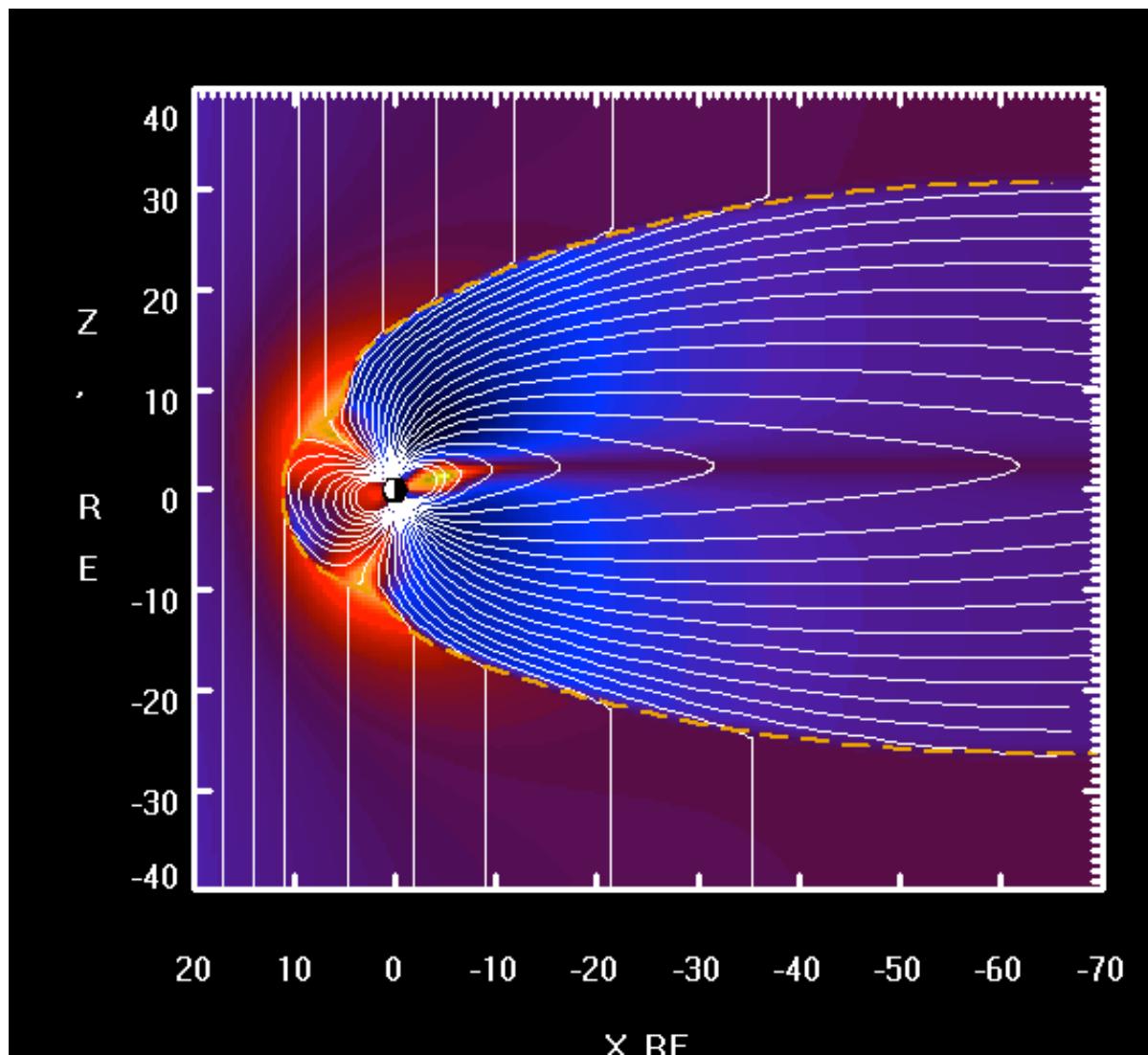
D. G. Sibeck

NASA/GSFC

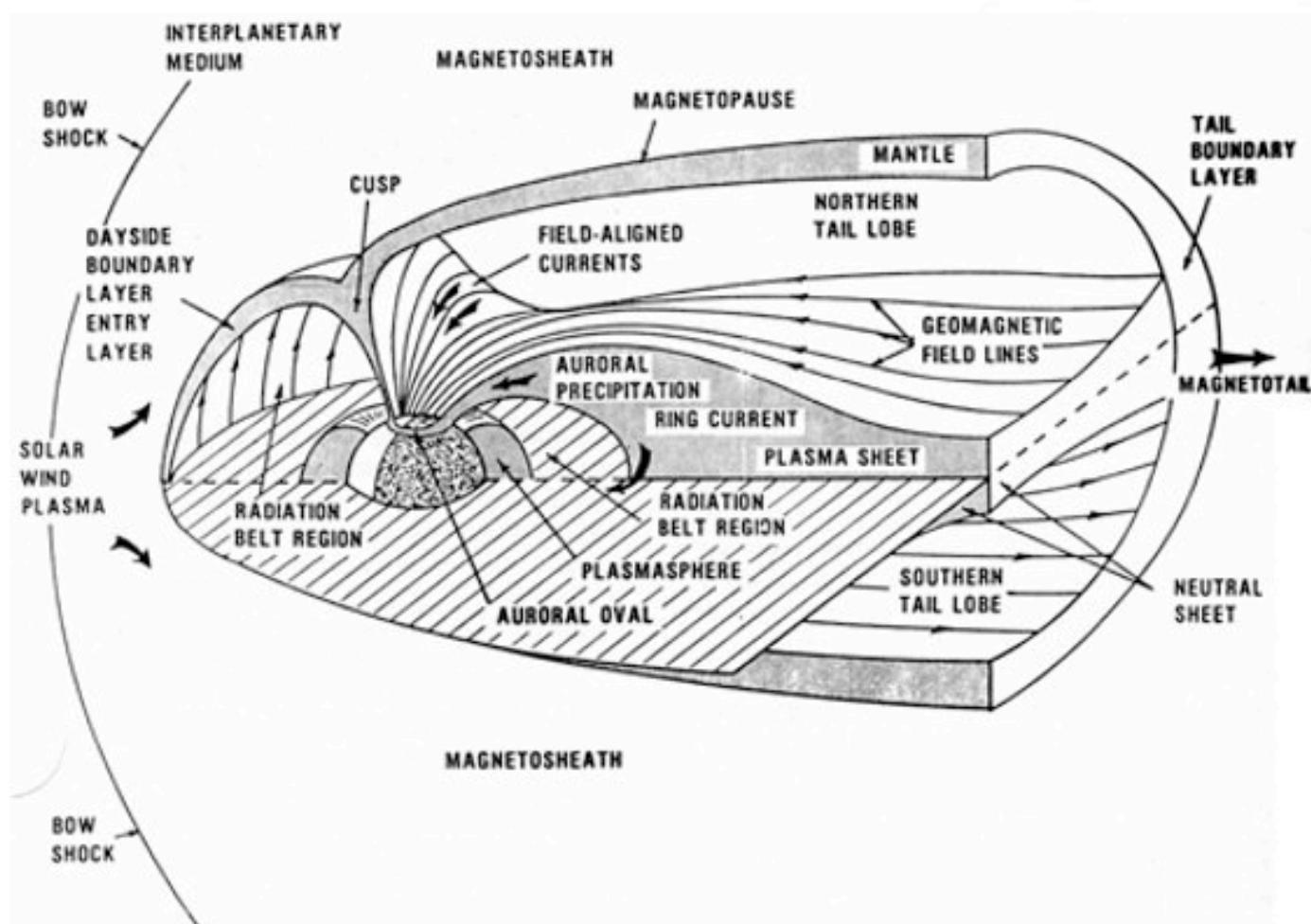
Outline

- 1. Steady and unsteady solar wind-magnetosphere interactions
- 2. Possible interpretations and their significance
- 3. Global Measurements for a Global Problem
- 4. Conclusion

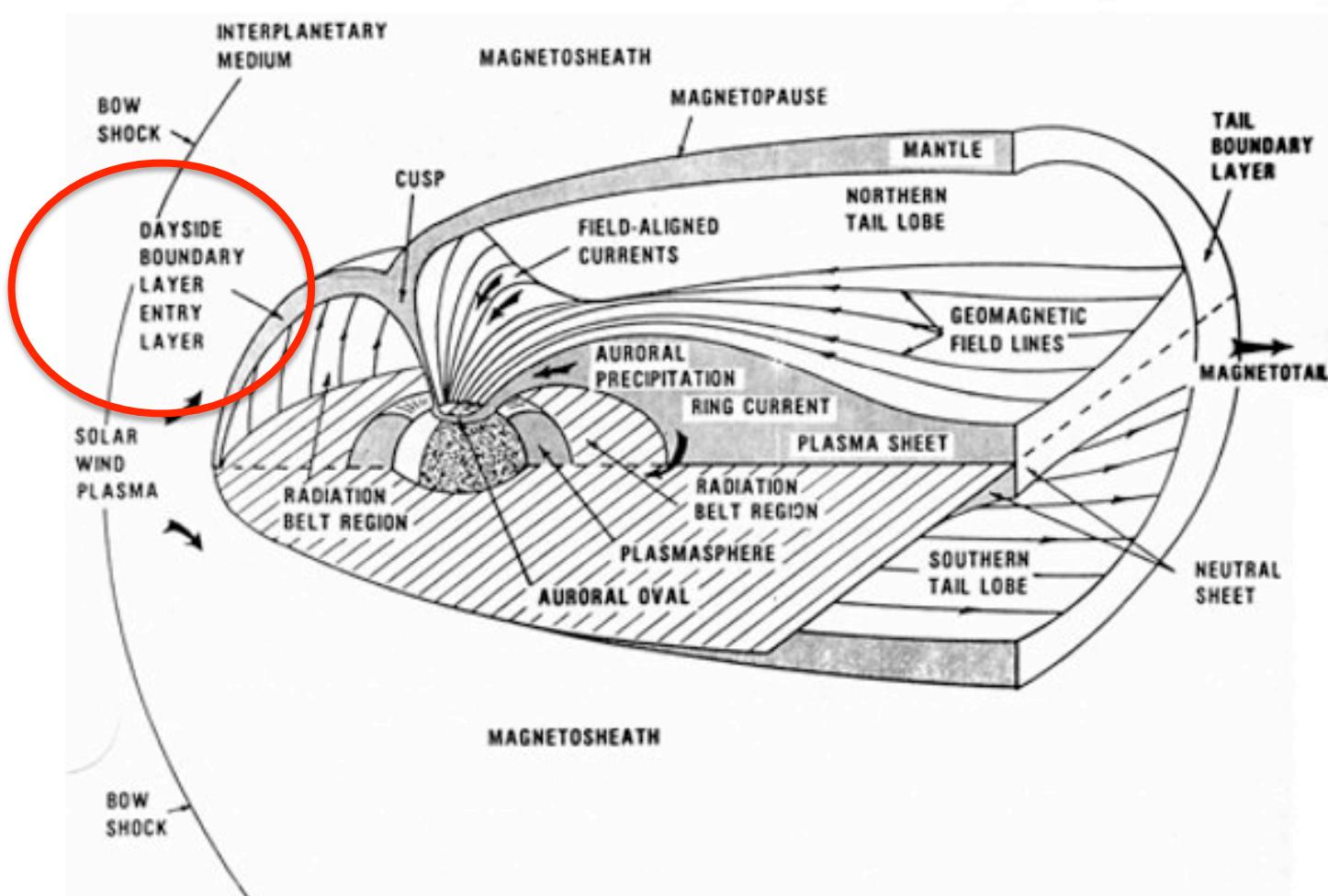
Steady Reconnection



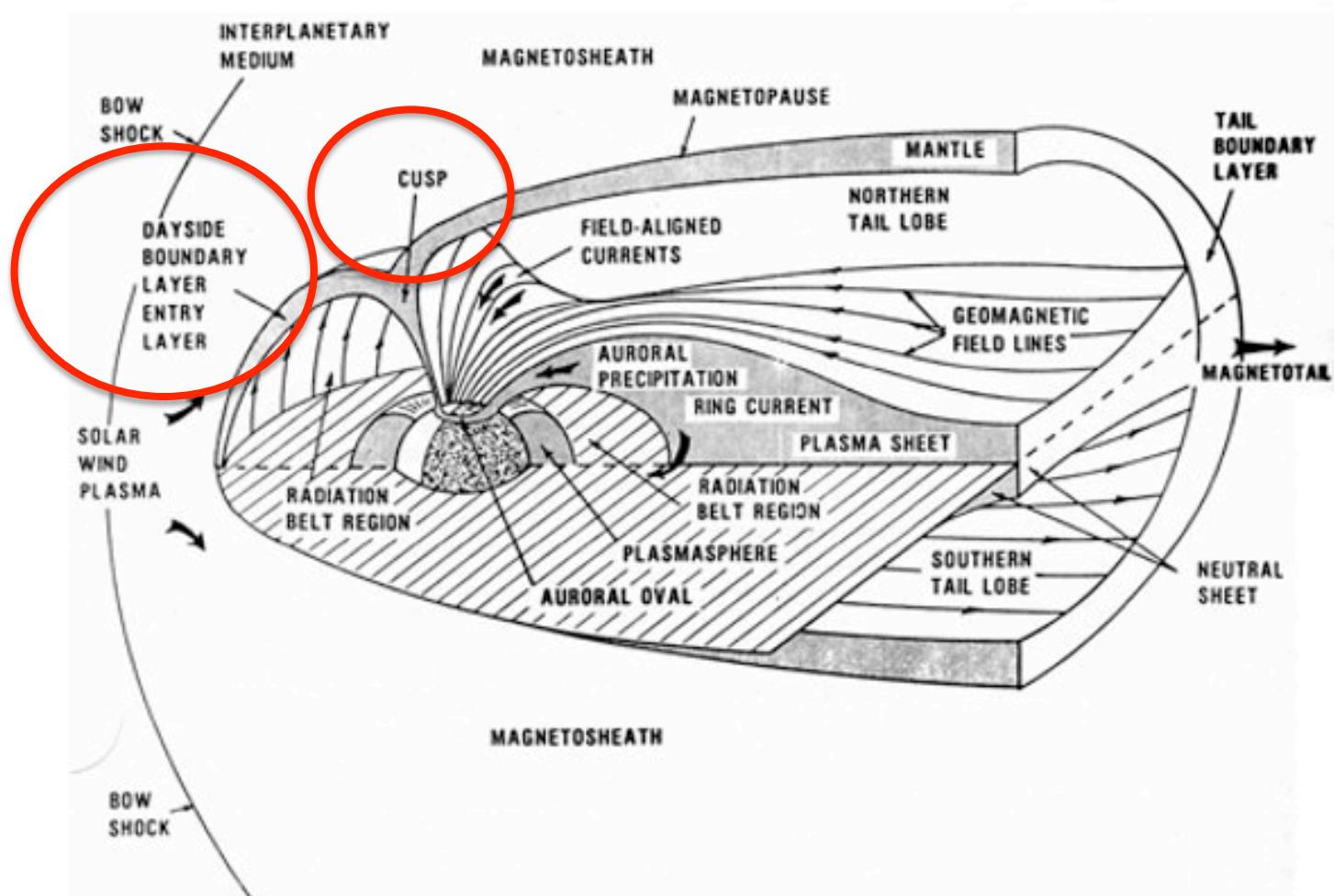
Steady-State Magnetosphere



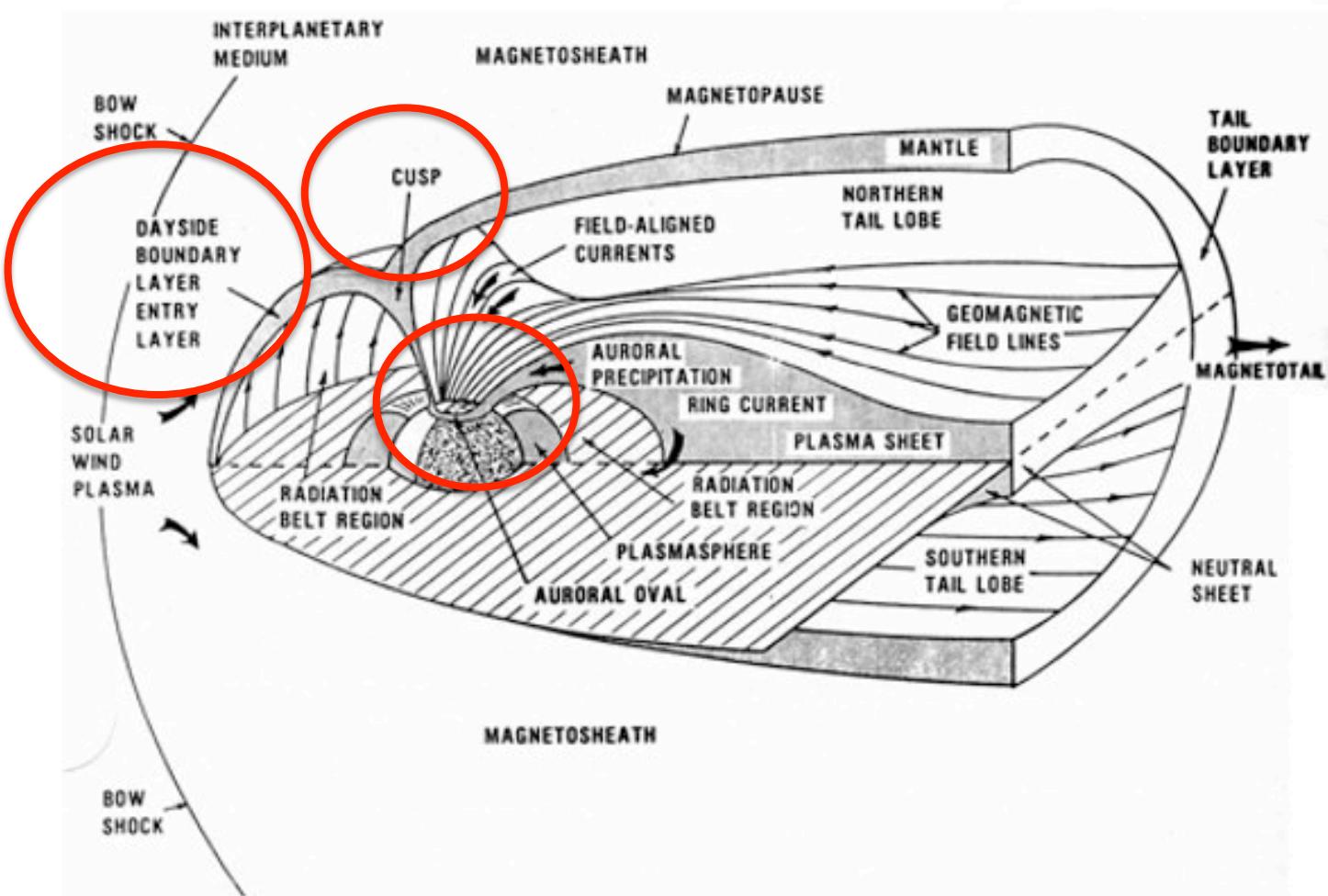
Steady-State Boundary Layers



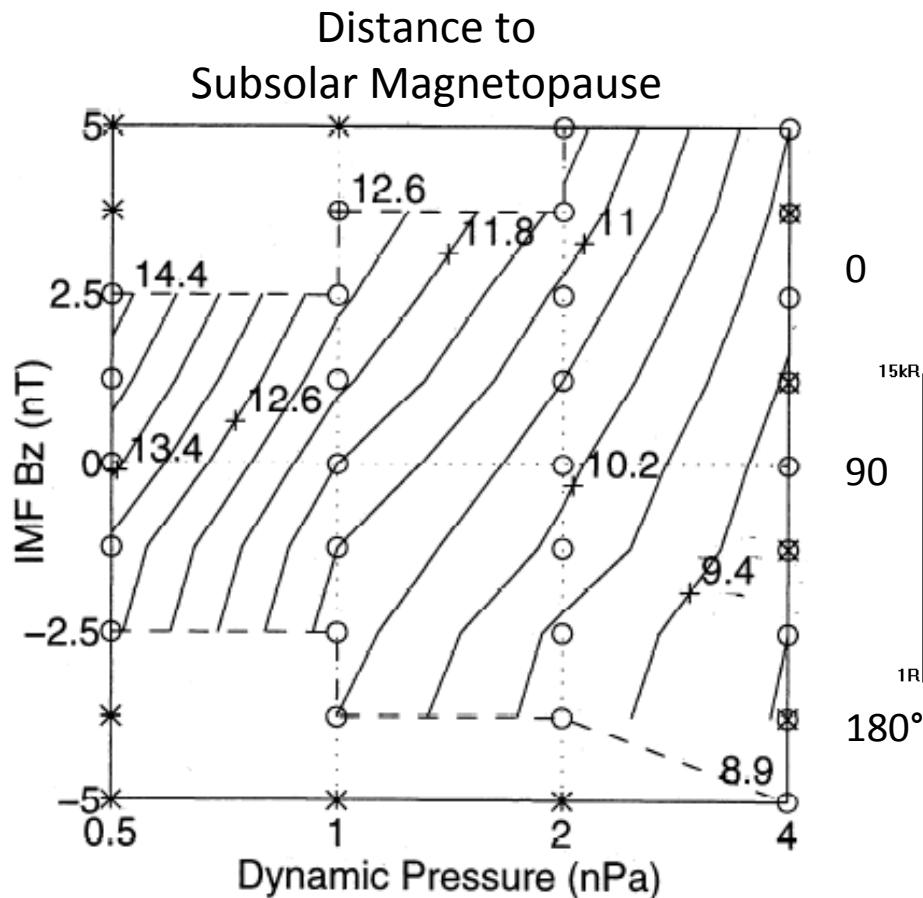
Entire Boundary Layer Maps to Cusp



High Latitude Stations can Observe Cusp Remotely

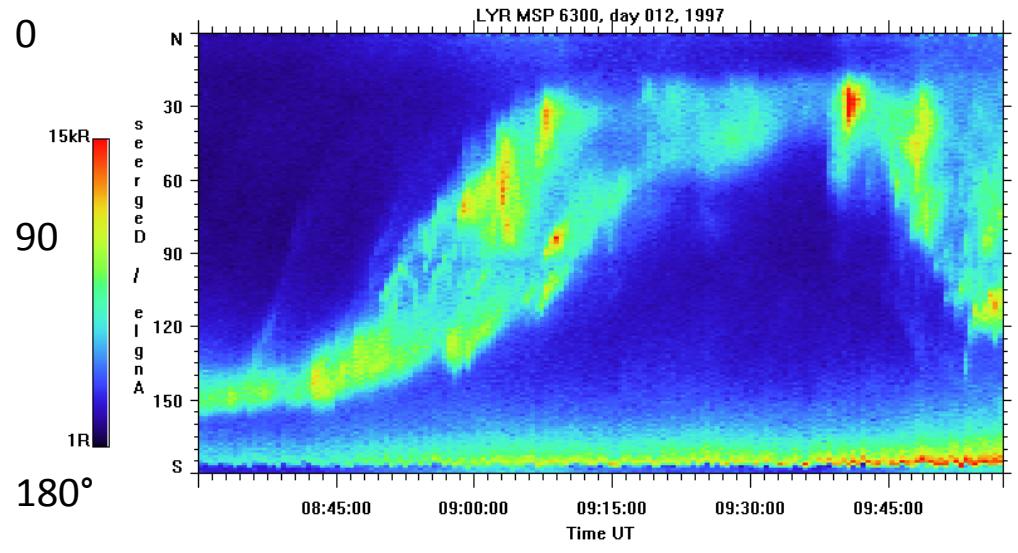


Some features are quasi-steady: Responses to Variations in SW P and Bz



Elsen and Winglee [1997]

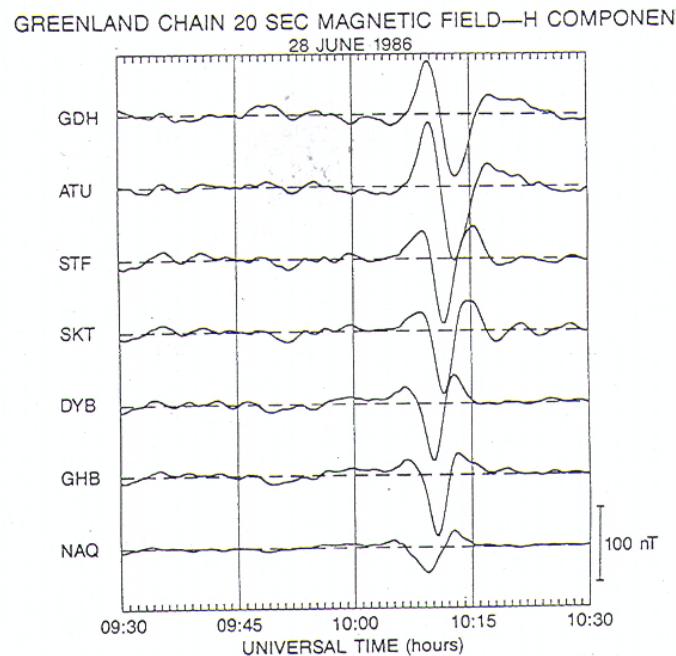
Location of Cusp Aurora:
Angle from Zenith



Fasel, 2012

Others are More Transient: Cusp-Latitude Ground Magnetometers

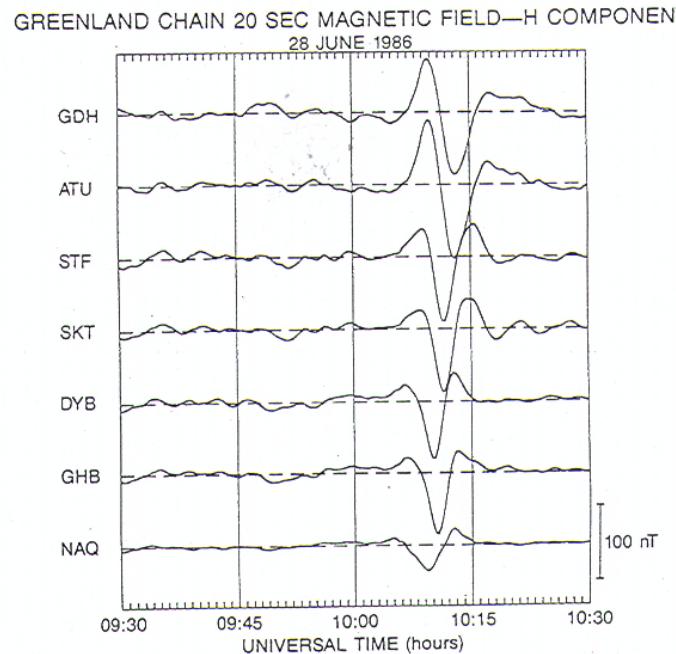
Magnetic Impulse Event or
Traveling Convection Vortex



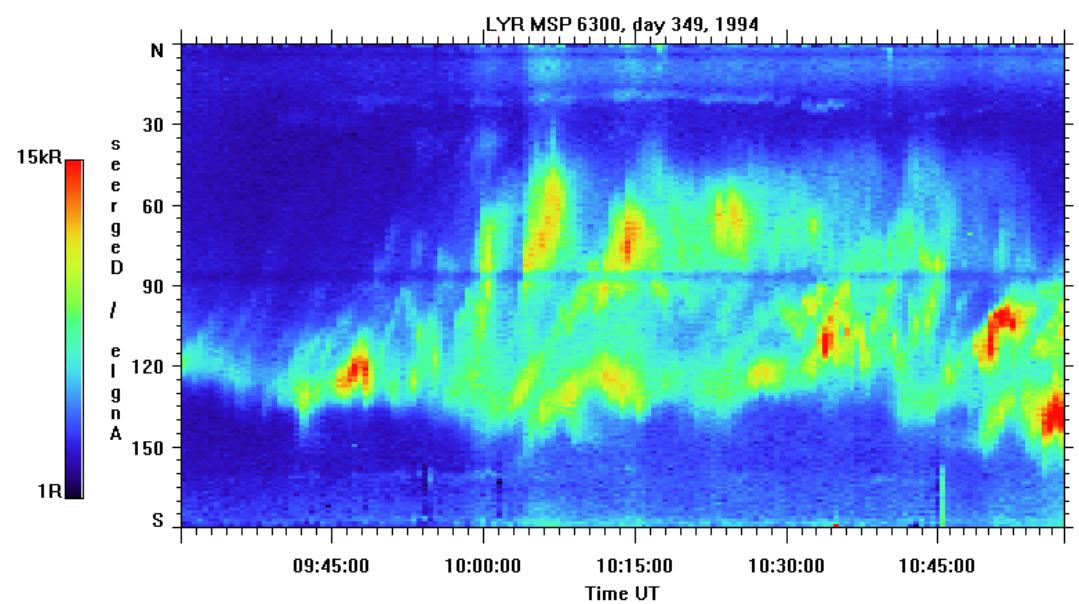
Friis-Christensen et al. [1988]

Transients are Common in Observations by Meridian Scanning Photometers

Magnetic Impulse Event or Traveling Convection Vortex



Poleward Moving Auroral Forms (including rebrightening)

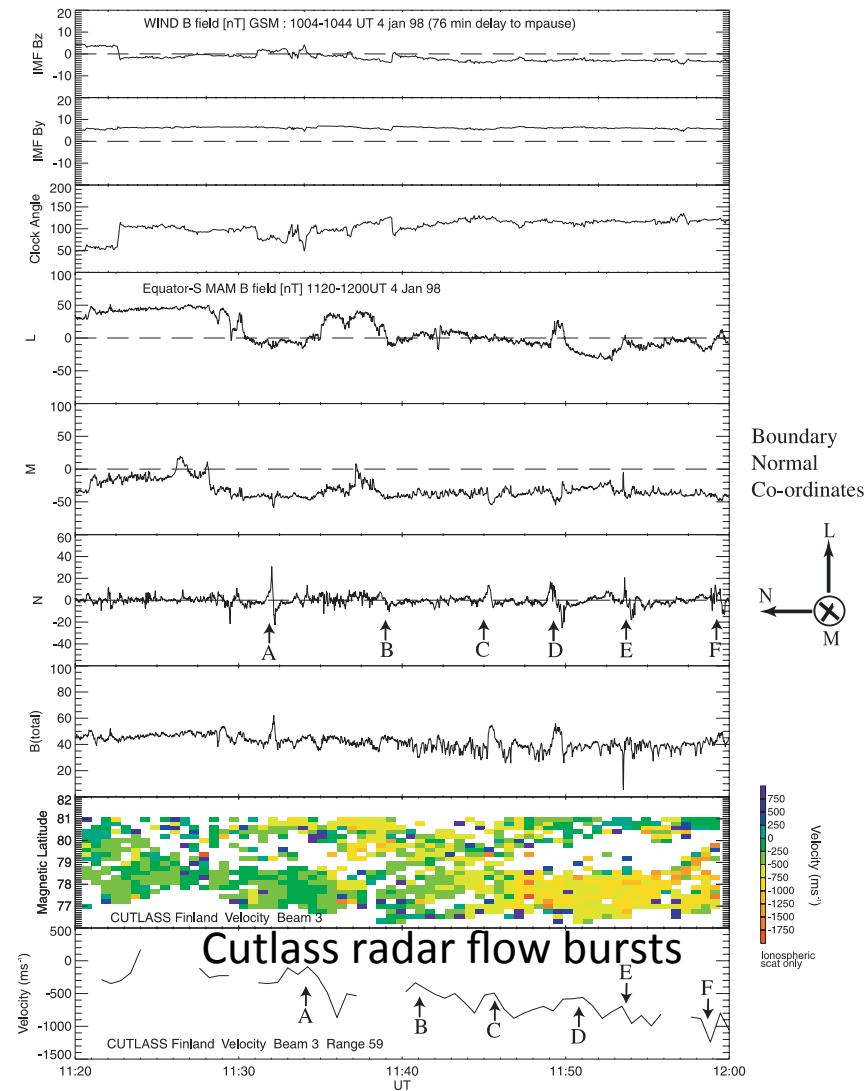


Friis-Christensen et al. [1988]

Fasel [2012]

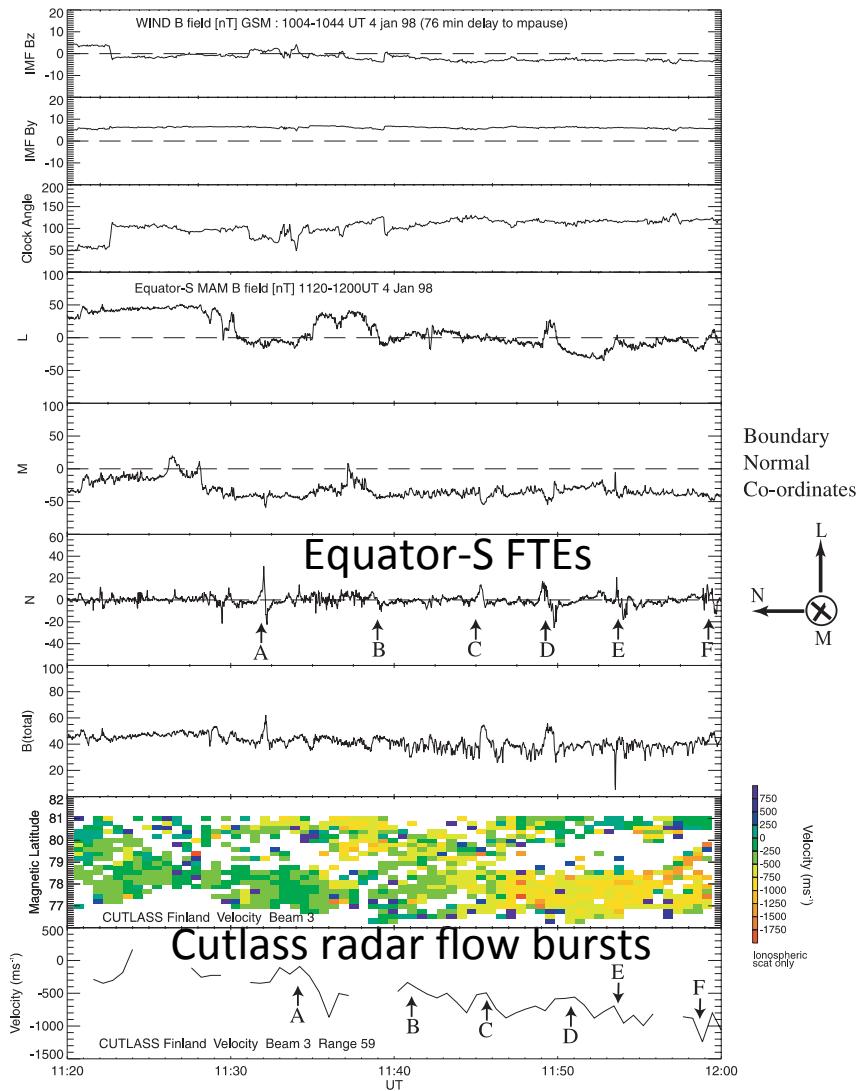
Transients are Common in Observations by Ground Radars

Neudegg
et al. [2000]

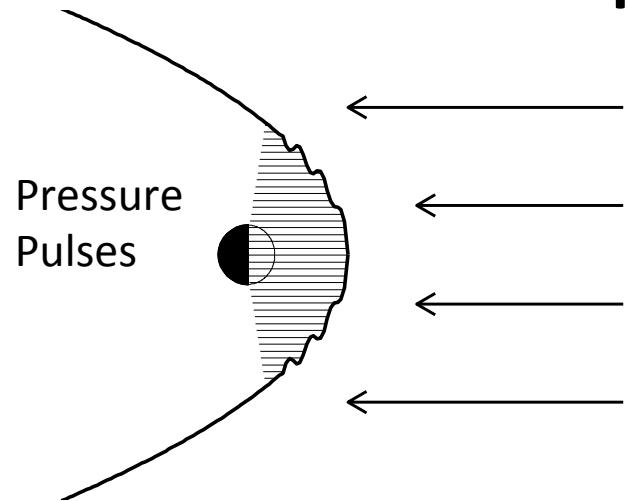


There Appears to Be a Magnetopause Connection

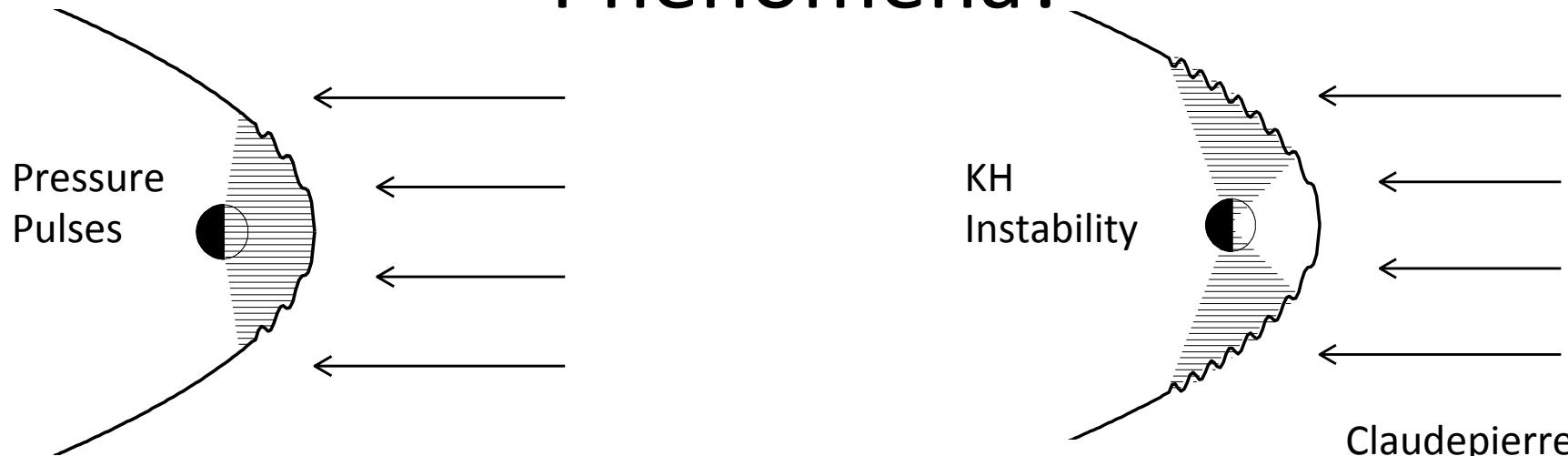
Neudegg
et al. [2000]



But to what Magnetopause Phenomena?

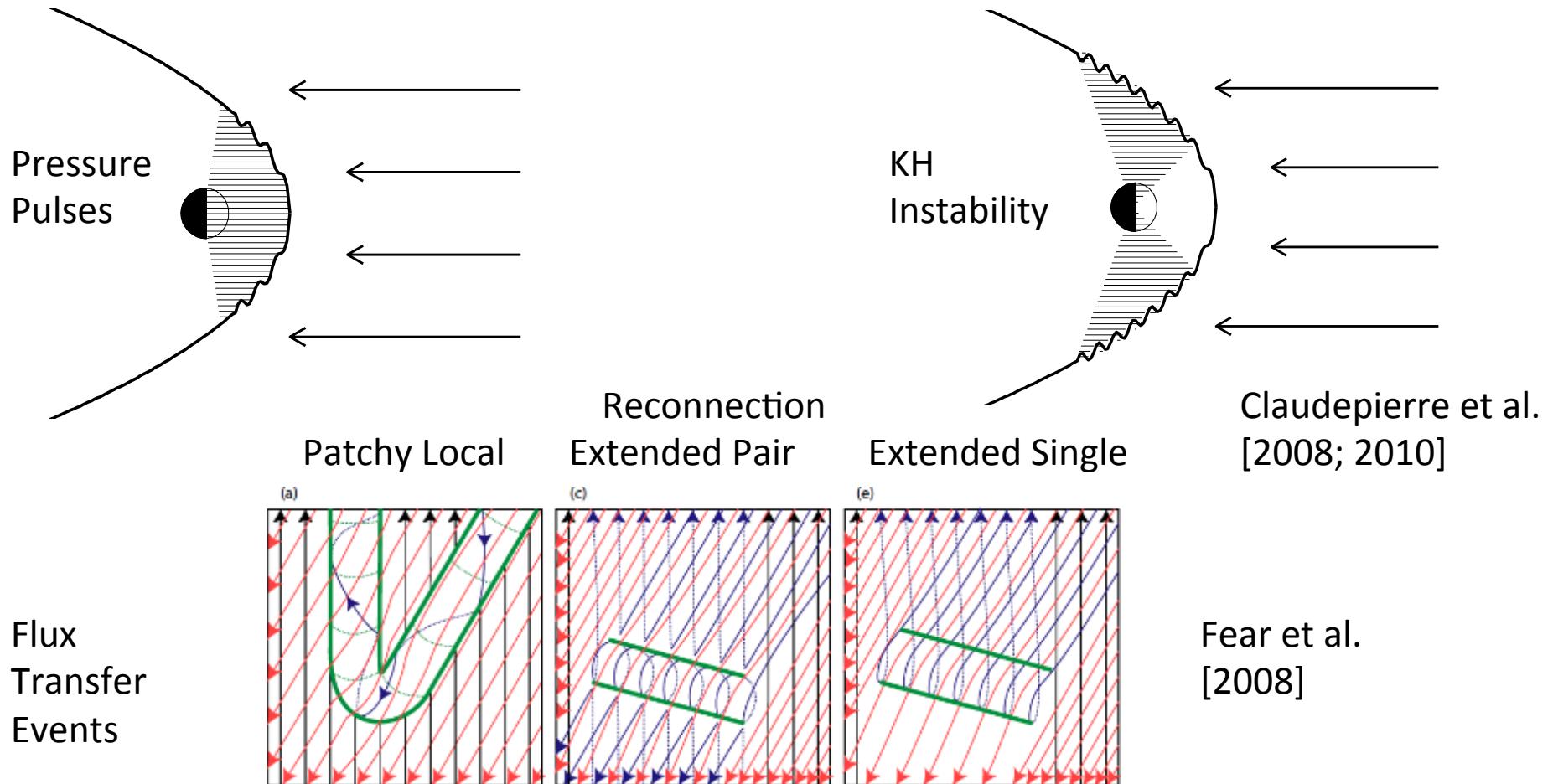


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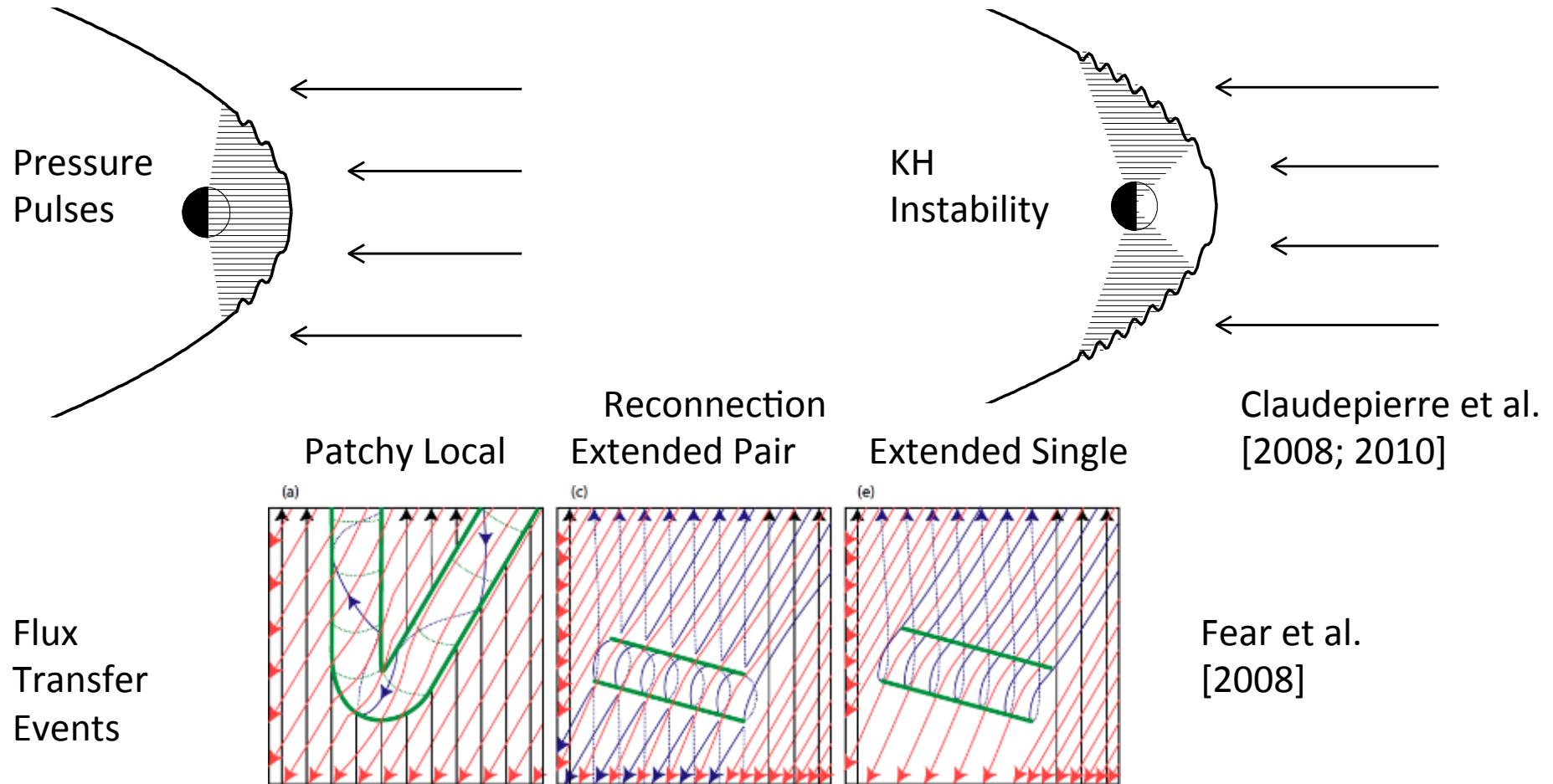


Claudepierre et al.
[2008; 2010]

Magnetopause Phenomena



Why it Matters

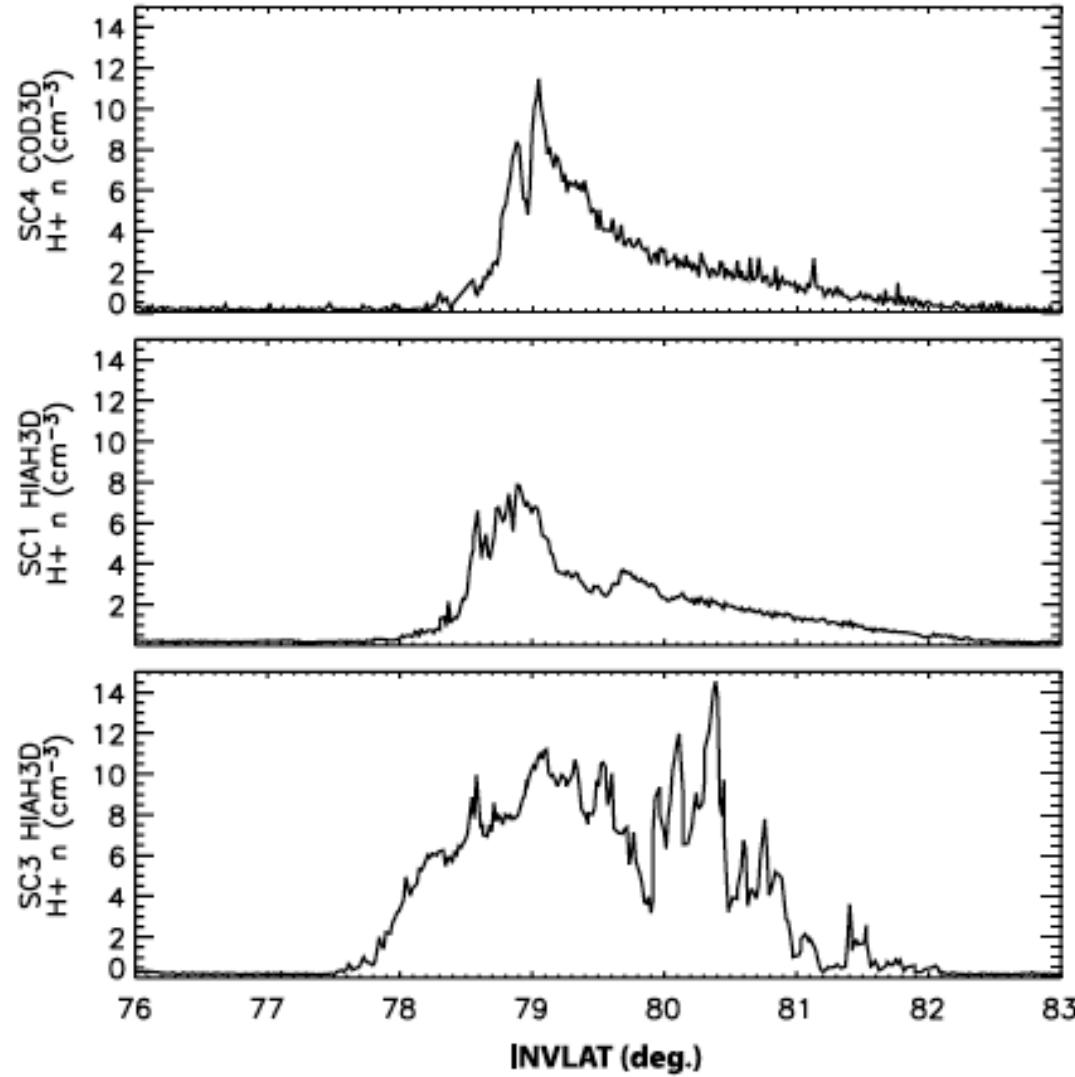


Significant: May dominate the solar wind-magnetosphere interaction
30% of the interaction [Lockwood et al., 1990]
100% of the interaction [Lockwood et al., 1995]

Whatever the Cause, There are IT Connections

- 1. Precipitating particles →
heating, conductivity, cosmic noise absorption, aurora
- 2. Field-aligned currents → magnetometer perturbations
- 3. Electric fields → convection
- 4. VLF/ELF/EMIC wave entry path

The Mid-Altitude Cusp Connection



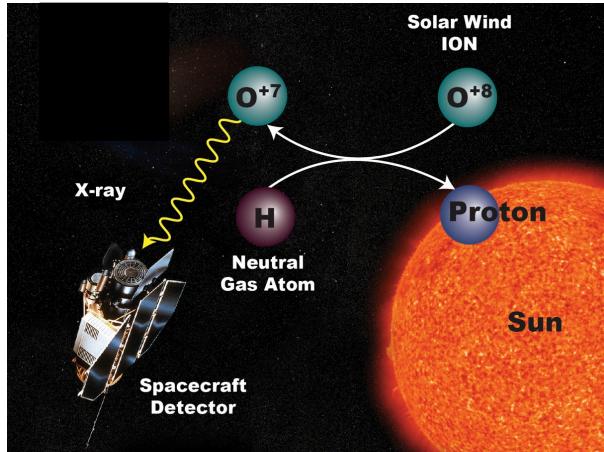
Plenty of evidence for
Structures and/or
Time-dependence:

Density Profiles for
Three Nearly
Simultaneous Cluster
Passes through the
Cusp

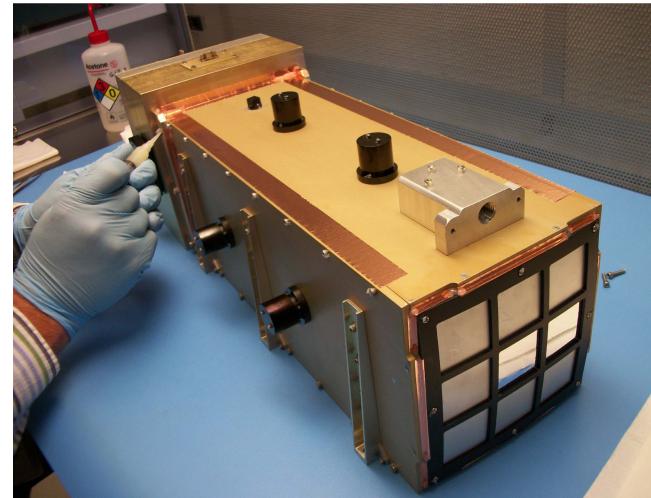
Wanted (to determine size, recurrence rates)

- 1. Global Images of the Magnetopause and Cusp
- 2. Global Images of the Dayside Auroral Oval
- 3. Extensive Arrays of Ground Observatories

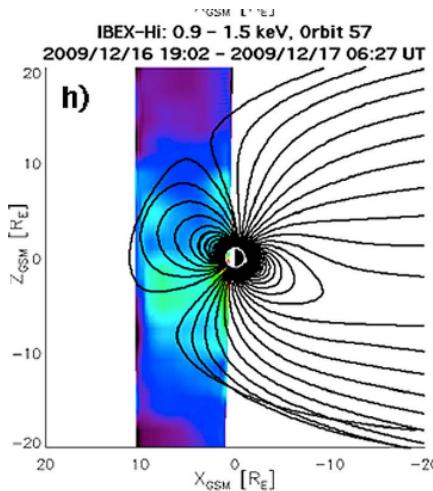
Global Imaging: Fact or Fiction?



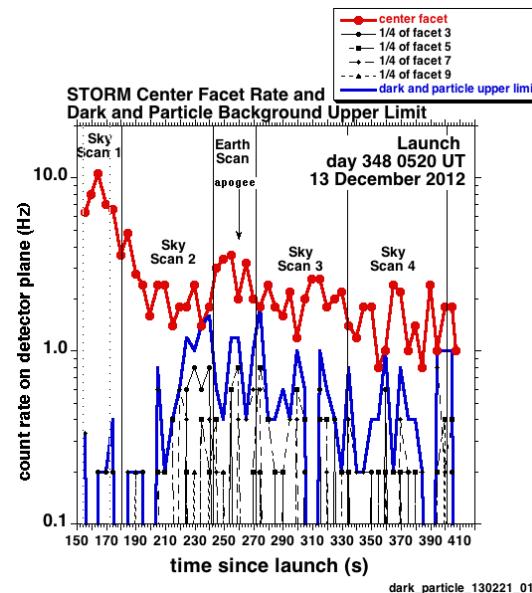
Charge
Exchange
Generates
Neutral AND
Soft x-rays



Prototype Soft X-Ray Imager flew on
DXL rocket December 12, 2012



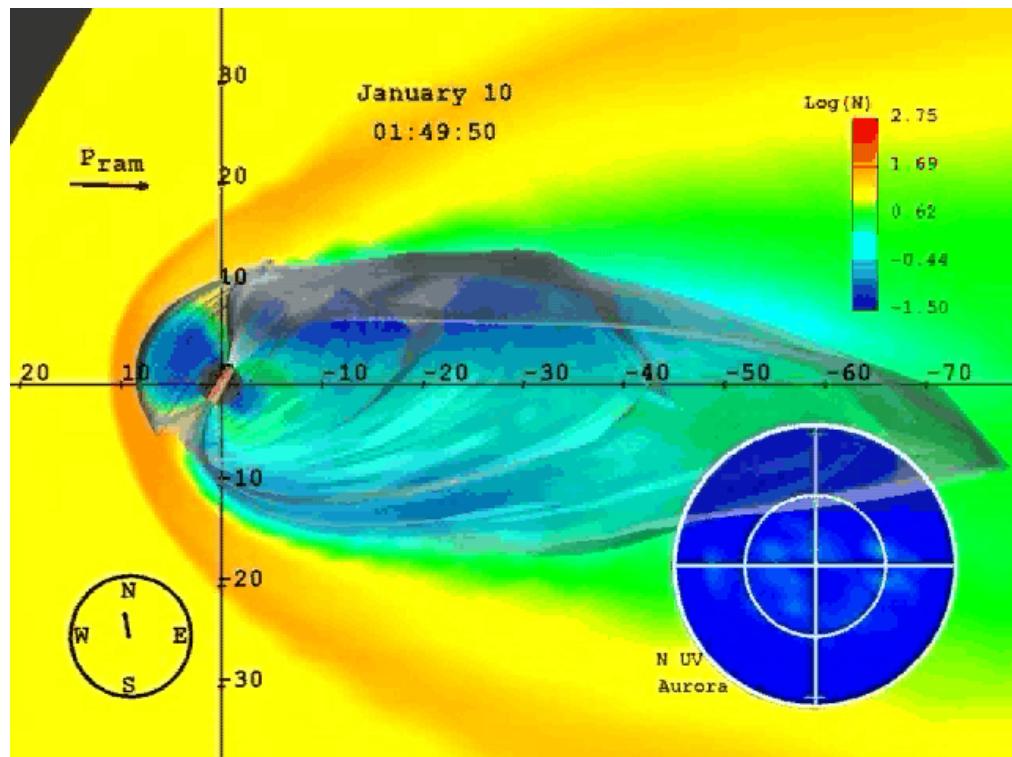
IBEX ENA Images of the Cusp
Petrinec et al. [2011]



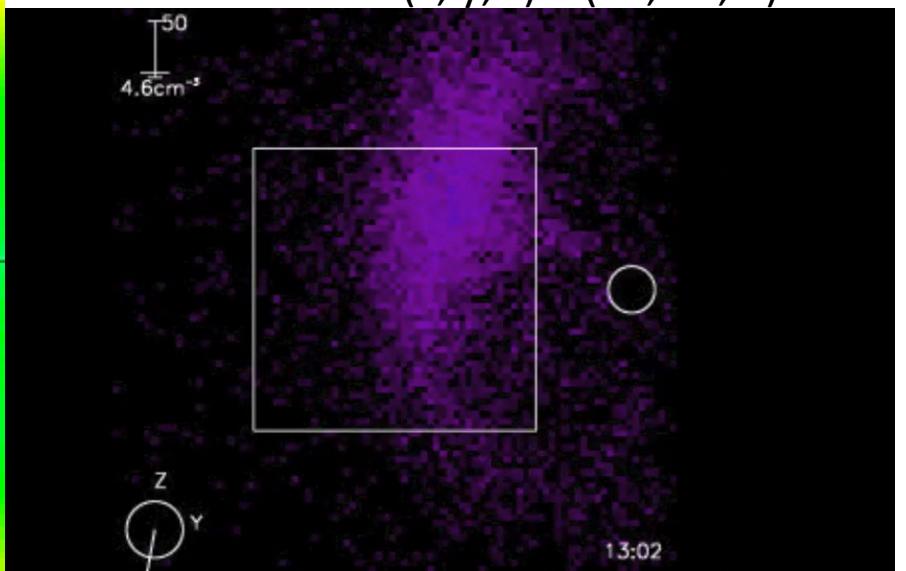
Light curve
looking nearly
along axis on
spinning rocket
(towards dawn
terminator)

Soft X-Ray Simulation

Global MHD Simulation (C. Goodrich)

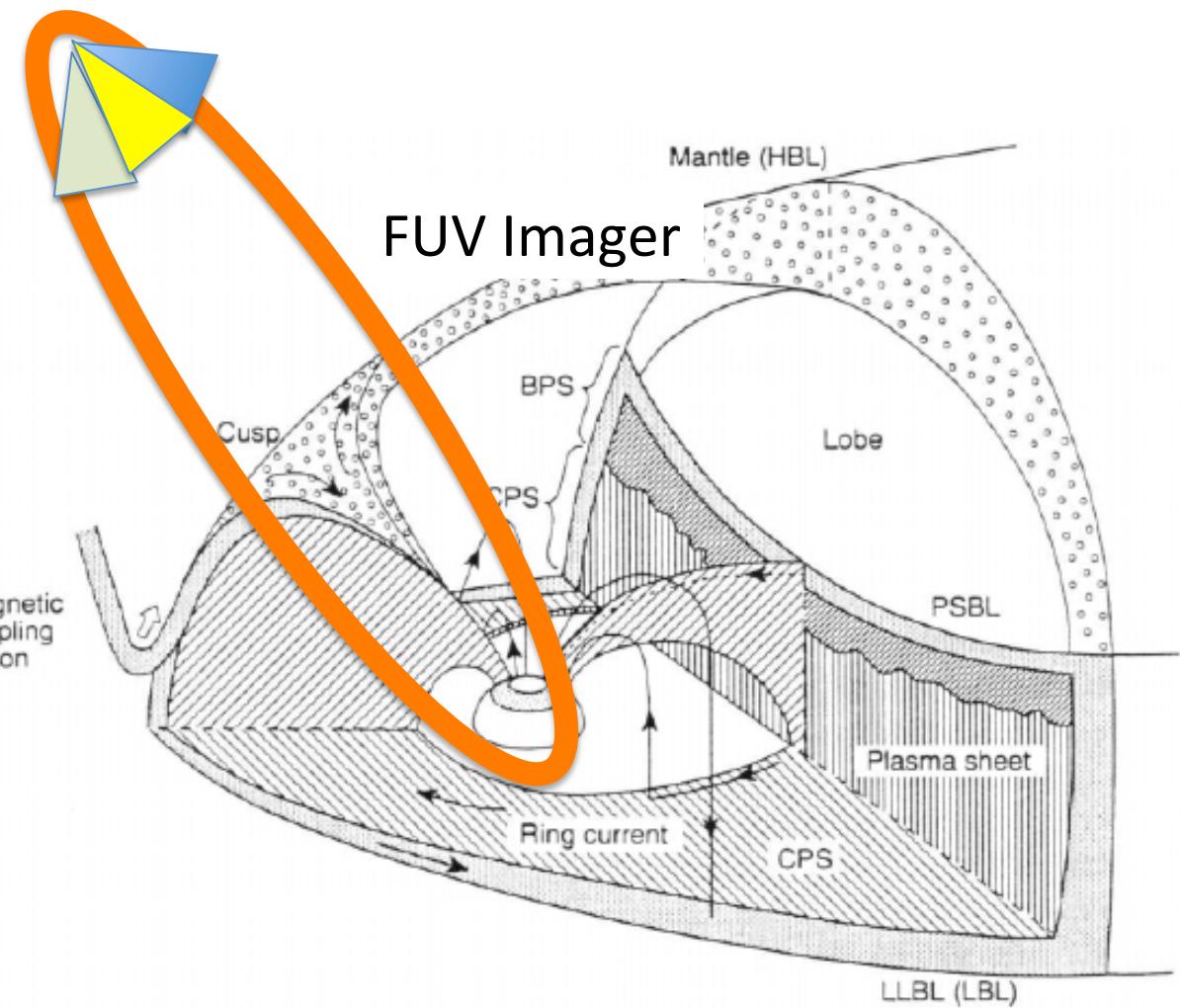
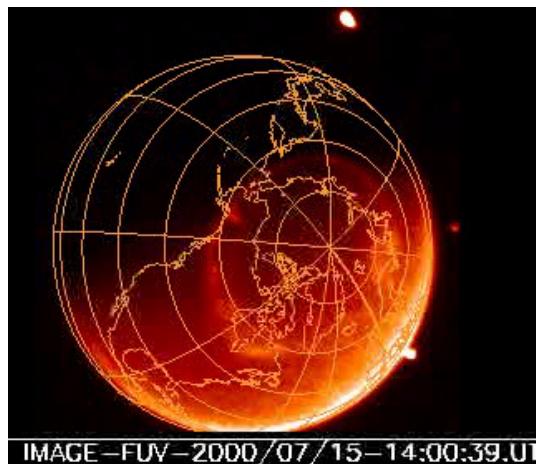


Soft X-Ray simulation
View from $(x, y, z) = (30, 10, 0)$ RE

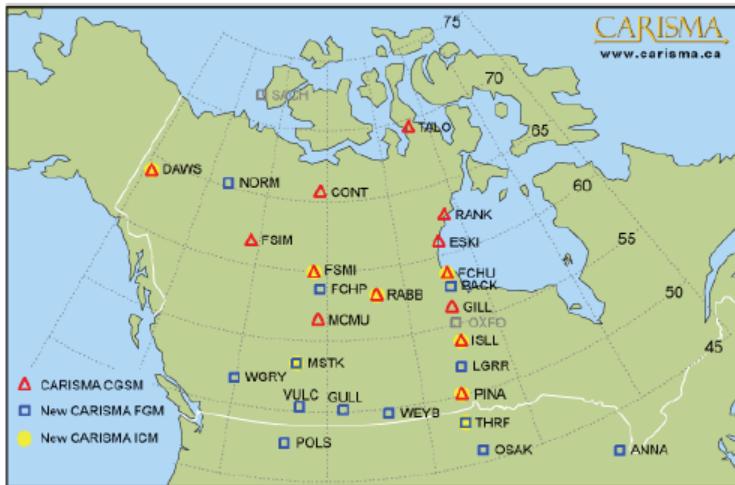


Magnetospheric Imaging from a High-Latitude High Inclination Orbit

Soft X-ray, ENA,
And FUV Imagers

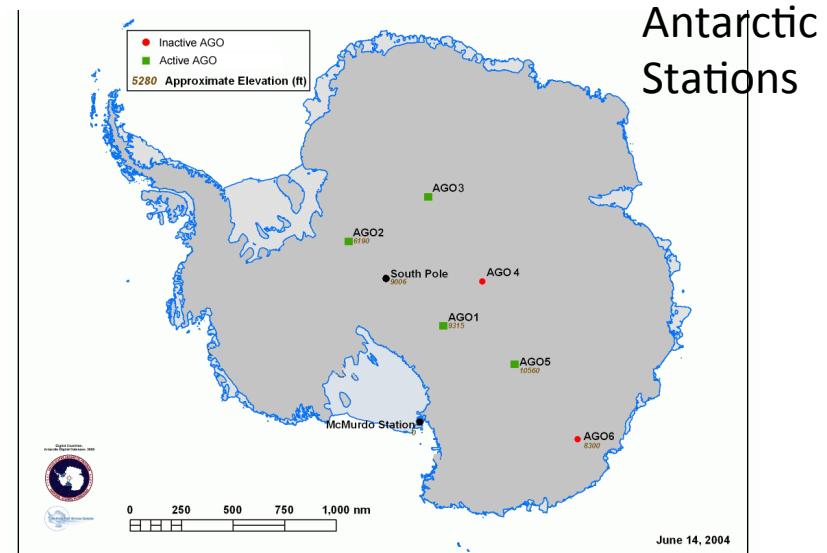


Relevant Ground-Based Observatories are in place

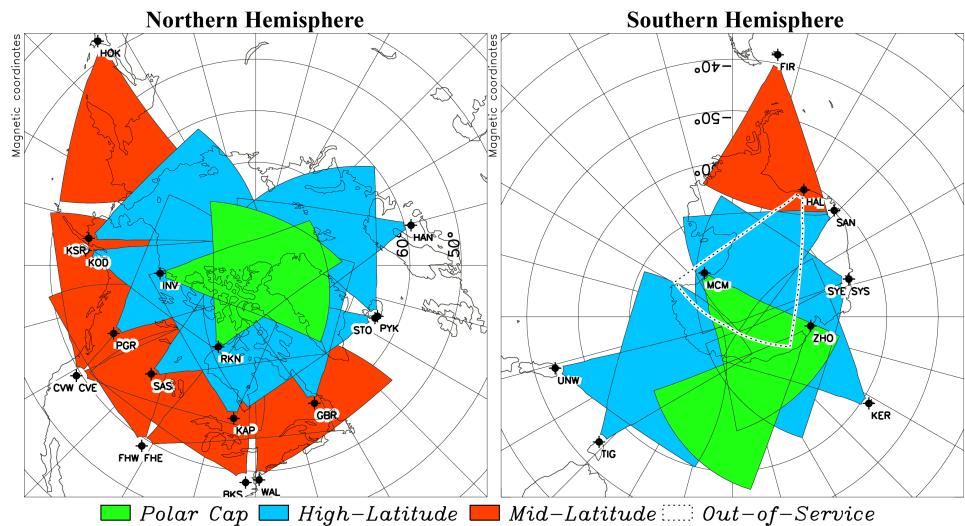


Canadian Ground-Based Arrays

SuperDARN
radar



Antarctic
Stations



Prognosis

- 1. Global images of the magnetosphere, cusp, and ionosphere
 - +
- 2. Extensive Arrays of ground-based observatories
 - =
- 3. Tremendous opportunities to determine the dimensions, recurrence rates, and therefore importance of the structures generated by the solar wind-magnetosphere interaction