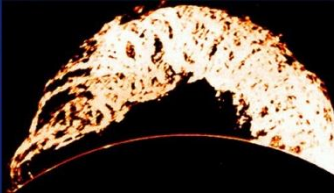
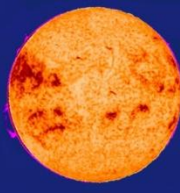
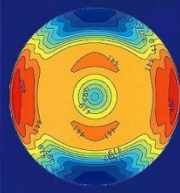


HAO



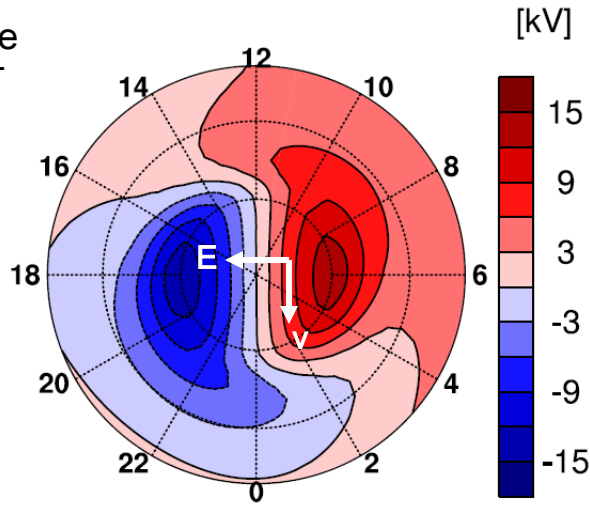
Properties of Electromagnetic Energy and Momentum Transfer Between the Magnetosphere and Upper Atmosphere

Art Richmond and Astrid Maute
NCAR High Altitude Observatory

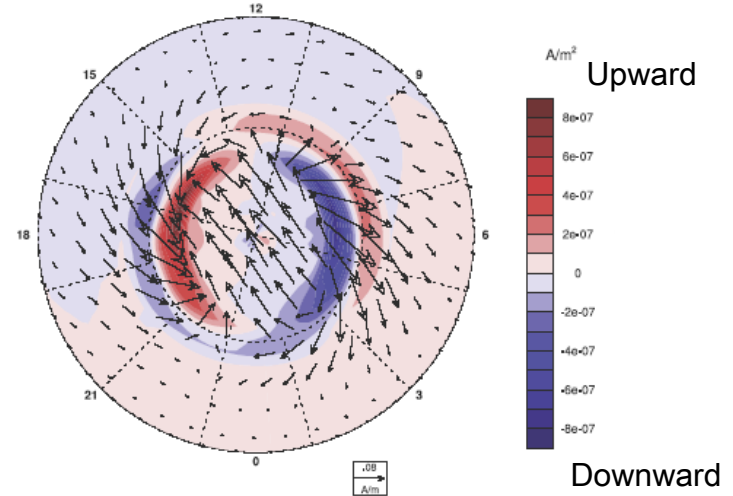
- Electromagnetic energy transfer on large and small scales
- Wind acceleration by Maxwell stress and feedback on plasma convection

Electric Potential

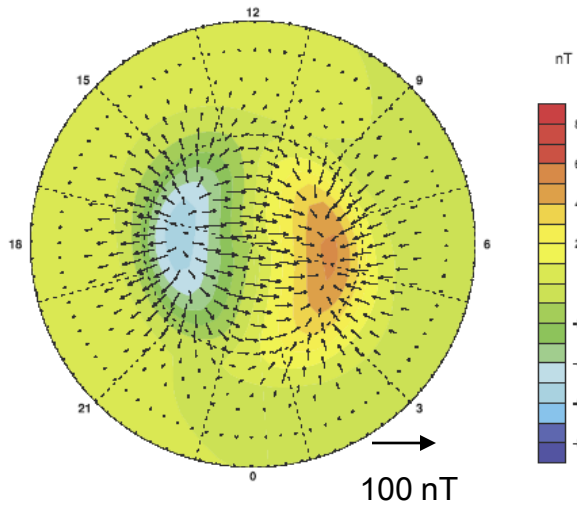
Northern hemisphere
June solstice, 17 UT



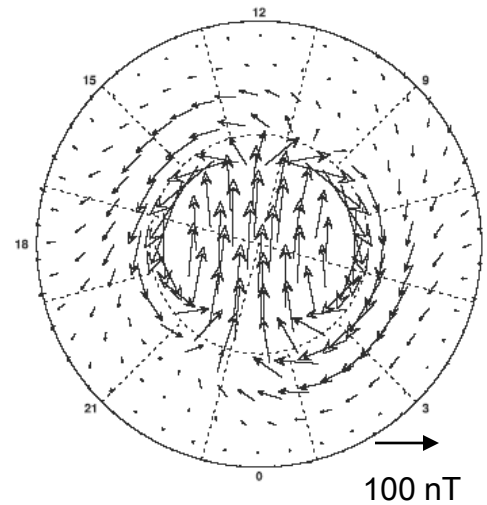
Ionospheric and Field-Aligned Current



Poloidal Magnetic Perturbation Field, 400 km



Toroidal Magnetic Perturbation Field, 400 km



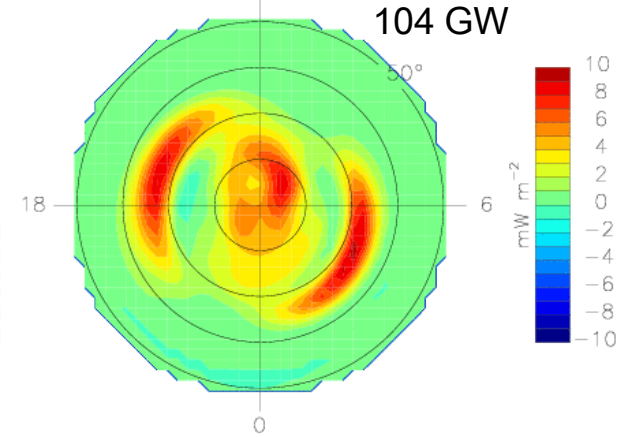
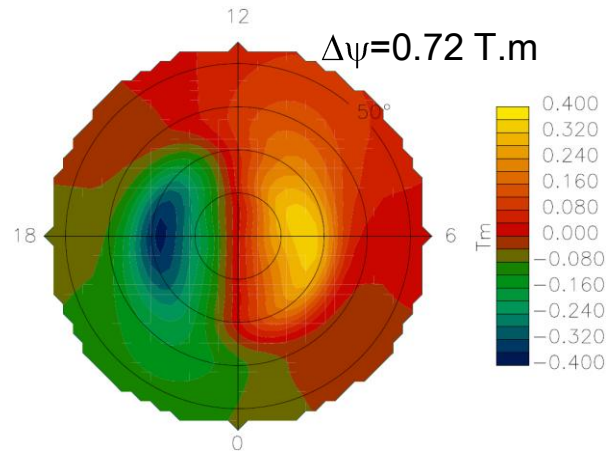
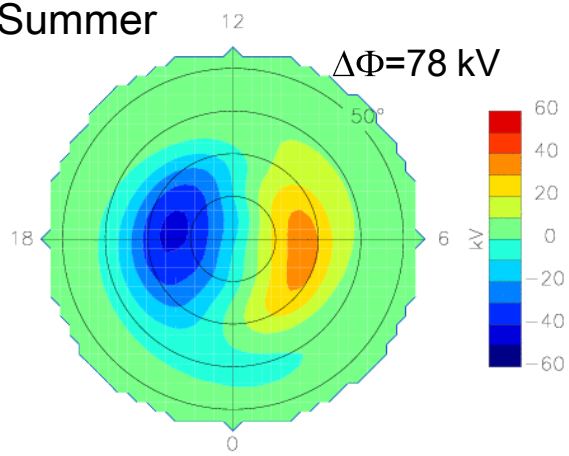
IMF $B_z = -5$ nT, $B_y = 0$

Electric Potential

Magnetic Stream Function

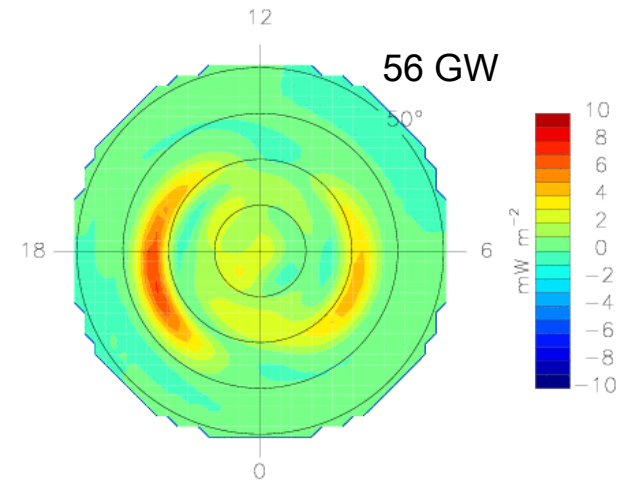
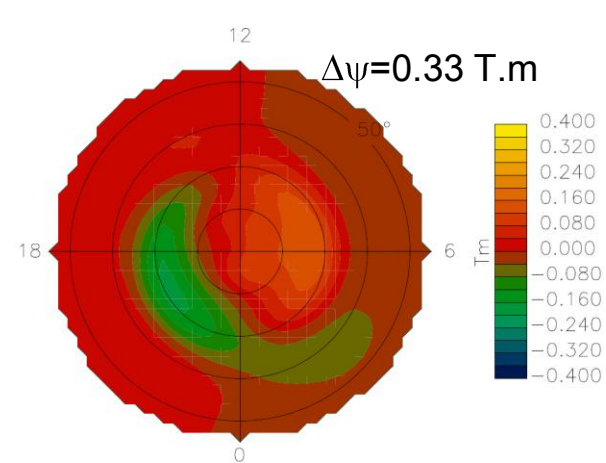
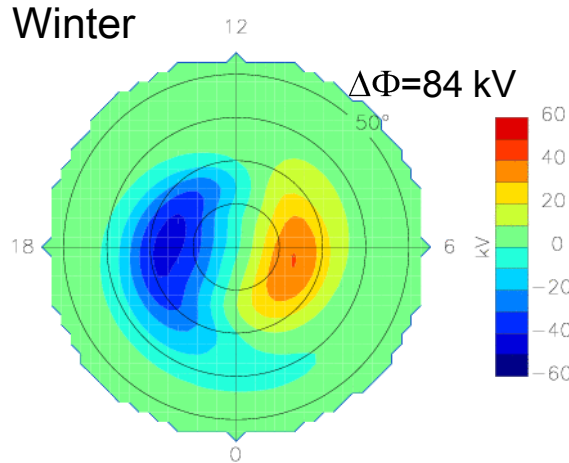
$\frac{\overline{\mathbf{E}} \times \overline{\delta \mathbf{B}}}{\mu_0} \cdot \mathbf{b}$

Summer



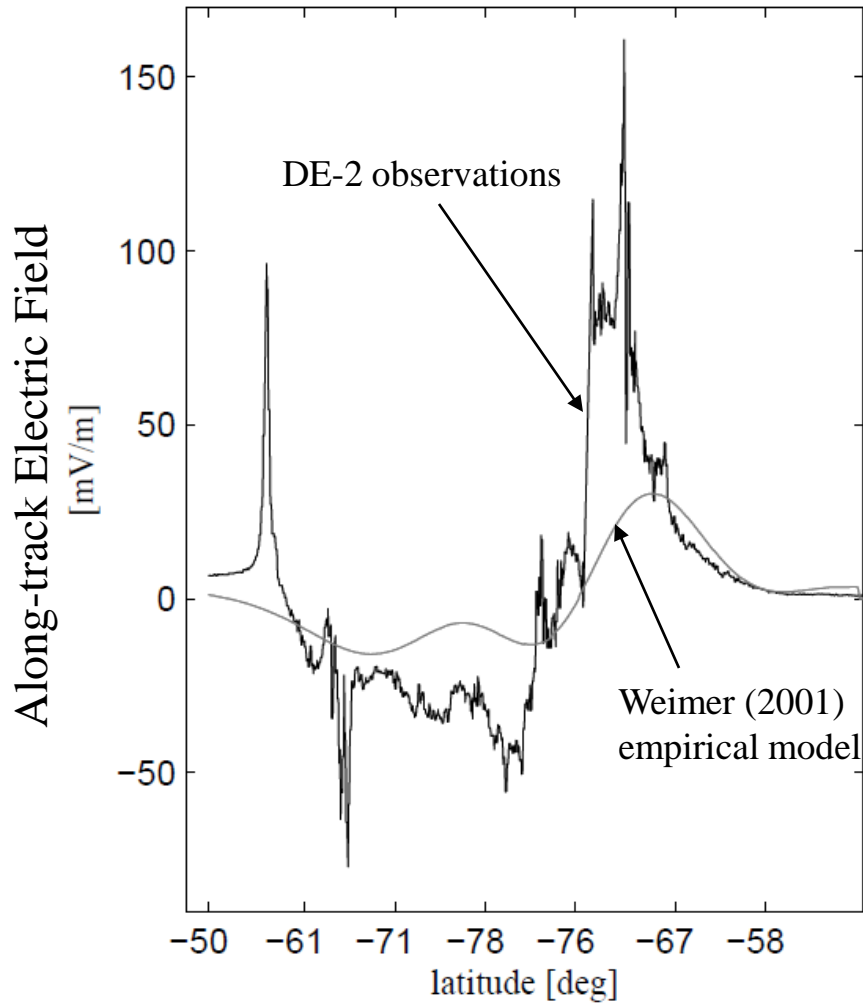
$\Delta\psi / (\mu_0 \Delta\Phi) = 7.4$ S

Winter



$\Delta\psi / (\mu_0 \Delta\Phi) = 3.1$ S

(a) DE2/IDM and Weimer



Matsuo and Richmond (2008)

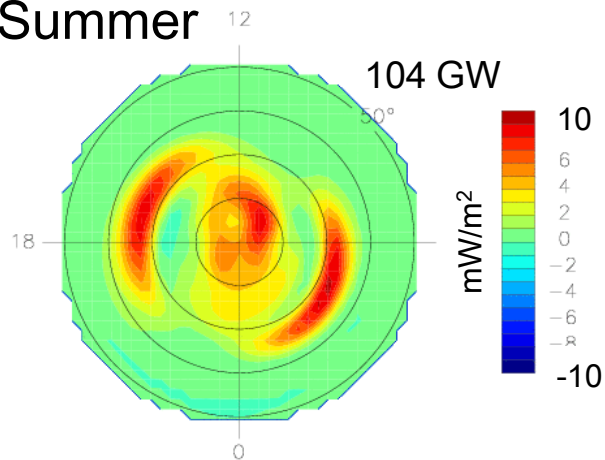
1981 September 26

1030-1055 UT

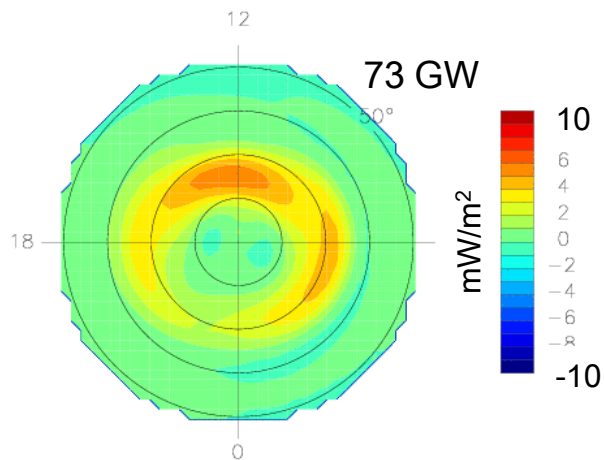
Field-Aligned Poynting Flux, IMF $B_z = -5$ nT, $B_y = 0$

$$\frac{\overline{\mathbf{E} \times \delta \mathbf{B}}}{\mu_0} \cdot \mathbf{b}$$

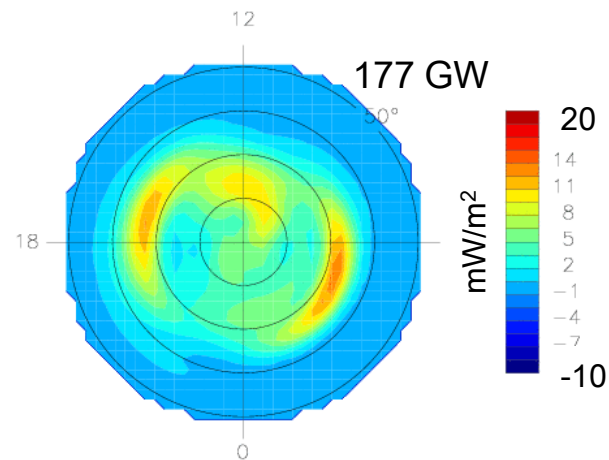
Summer



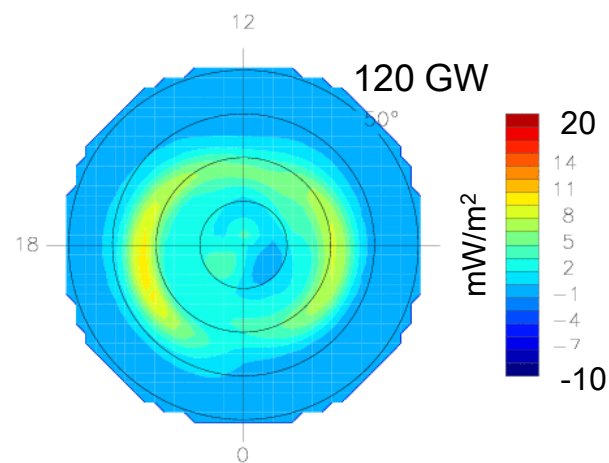
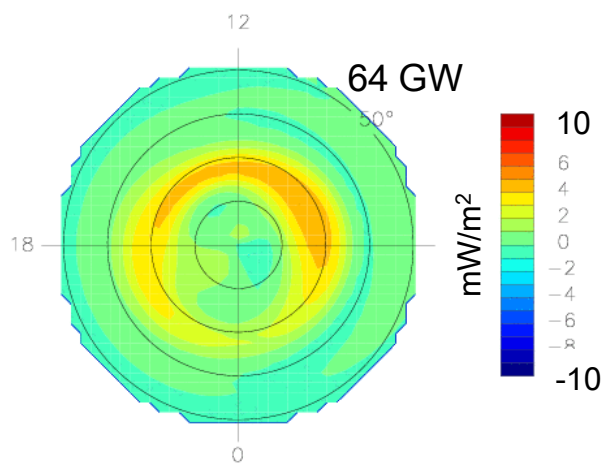
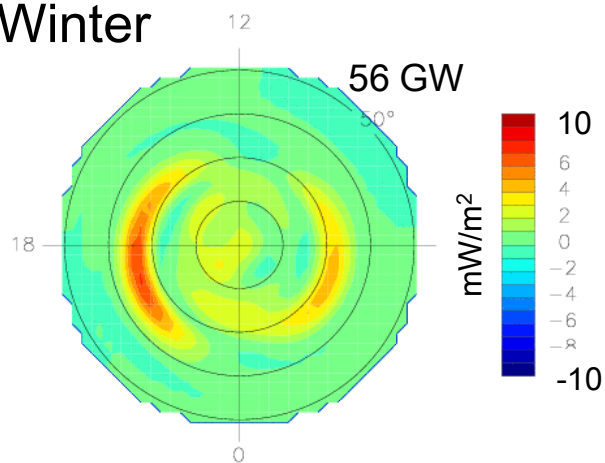
$$\frac{\overline{\mathbf{E}' \times \delta \mathbf{B}'}}{\mu_0} \cdot \mathbf{b}$$



$$\overline{\mathbf{S}} \cdot \mathbf{b}$$

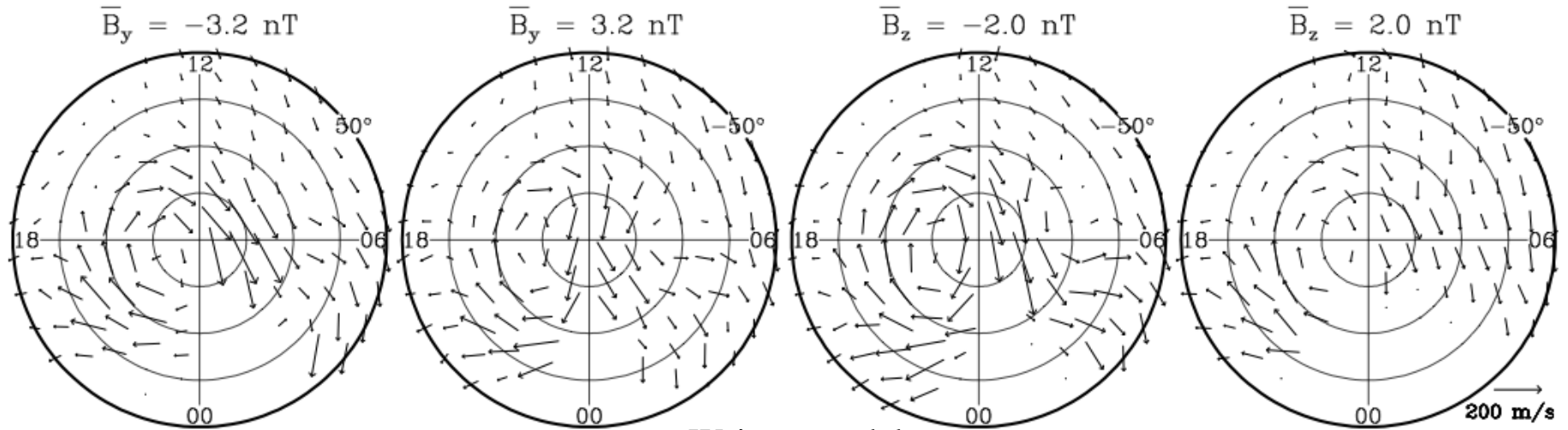


Winter

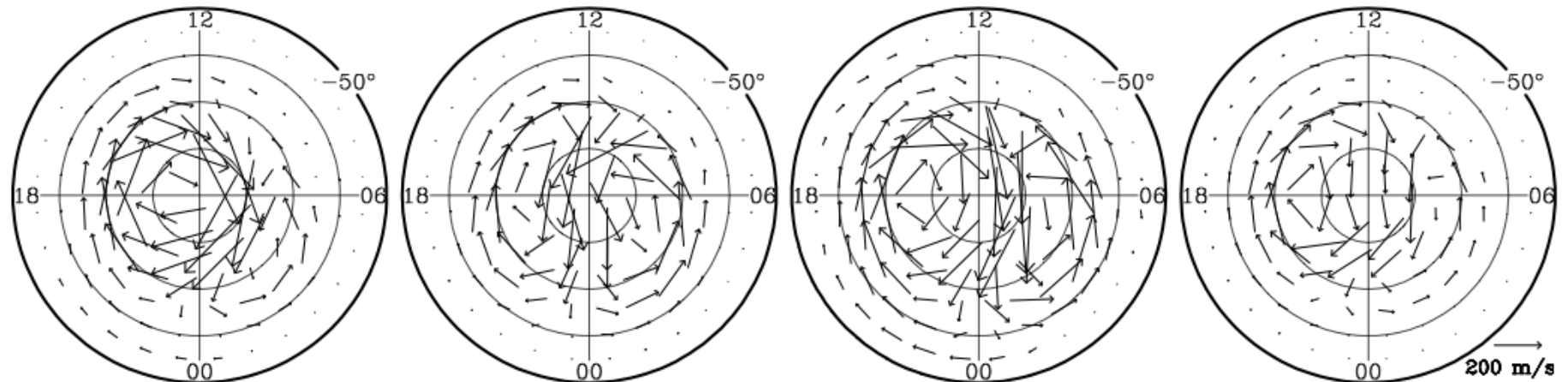


Southern Hemisphere, Summer

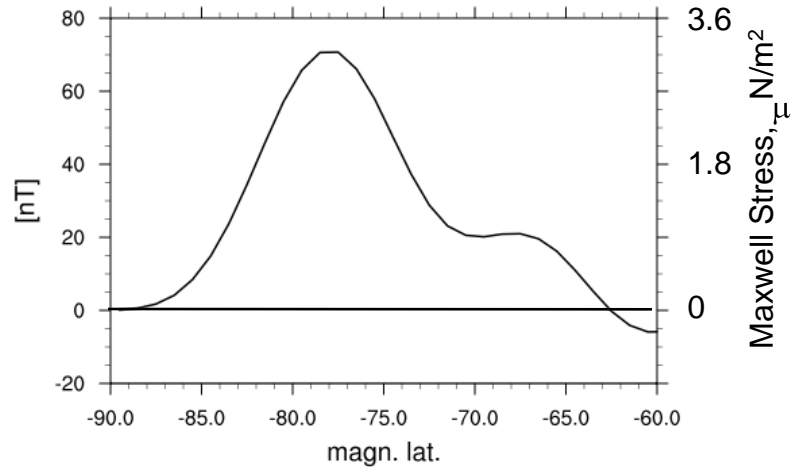
WINDII Observations
Pedersen-Weighted Neutral Wind



Weimer model
Convection Velocity

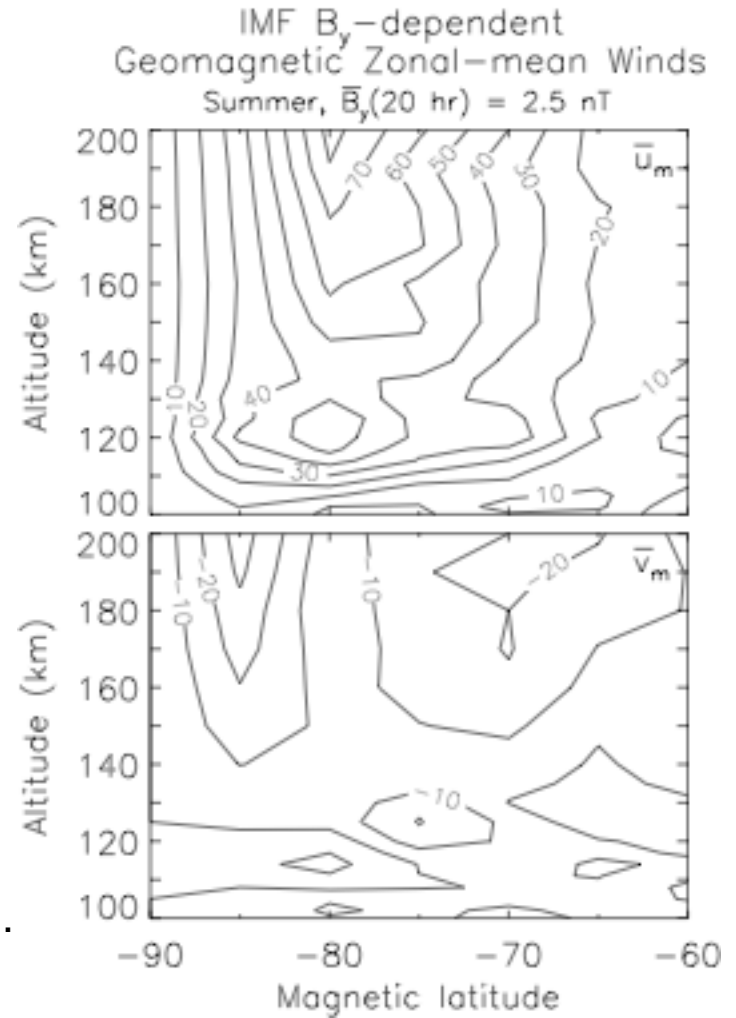


Zonal Mean Eastward Magnetic Perturbation Above Ionosphere (Summer, IMF $B_y = 2.5$ nT)



Atmospheric mass $\sim 10^{-3}$ kg/m² above 110 km

2×10^{-6} N/m² causes 2×10^{-3} m/s² acceleration,
which produces 40 m/s wind after 2×10^4 s (~ 6 hours).



Some Questions

- What are the distributions of energy and momentum flux during dynamic conditions?
- How are small-scale convection velocities related to ionospheric conductivity structures in the E and F regions?
- What is the dynamic response of winds to changing convection on different temporal and spatial scales?
- How does the magnetosphere respond to the winds?