Equatorial Spread F

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- The Aeronomy Problem
- Techniques and Taxonomy
- Physical Explanations
- Practical Consequences
- Outstanding Questions

The Aeronomy Problem



The Aeronomy Problem



Ionosondes: Equatorial Spread F

Jicamarca, Perú 1930 LT 14 Sep (257) 2006



Miller CEDAR: ESF

Ionosondes: Equatorial Spread F

Jicamarca, Perú 2200 LT 14 Sep (257) 2006



Miller CEDAR: ESF

Transionospheric Radio: Scintillation

136-MHz scintillation at Natal, Brazil (\sim 18 dB full scale, from Yeh and Liu, 1982)





• Fresnel-scale $(\sqrt{\lambda z})$ irregularities.

•
$$S_4^2 = \frac{\langle I^2 \rangle - \langle I \rangle^2}{\langle I \rangle^2}$$

HF/VHF Radio: Transequatorial Propagation



from Röttger, 1973.

HF/VHF Radio: Transequatorial Propagation



SALISBURY- LIMASSOL

VHF Radar: Field-Aligned Irregularities

- Bragg-scale $(\lambda_{rad}/2)$ irregularities.
- Strong returns where $\mathbf{k} \perp \mathbf{B}$.



⁽Show movie from Koki if Dave didn't.)

In Situ: Equatorial Plasma Depletions



from Roddy, et al, 2010 (Reproduced with permission of AGU).

Optical: Equatorial Plasma Bubbles



Optical: Equatorial Plasma Bubbles



A Composite Picture

19 August (232) 2004 CXI/CNFI





Occurrence (Day-to-Day)



Occurrence (Day-to-Day)



Occurrence (Local)



Miller CEDAR: ESF

Occurrence (Global/Local)



From Gentile, et al, 2006.

African TE Propagation Rates 1978-1980



from *Cracknell, et al*, 1981 (Nov 1981 *QST*, reproduced with permission of ARRL).

Occurrence (Global)



- Differences? Similarities? DMSP surveys...
 - a narrow window of times (18–22 LT, most at 20 LT).
 - bubbles that reach the 840-km orbit altitude.
- Considerable coincidence to terminator "locus of meridionallity." (*Tsunoda*, 1985)

From Gentile, et al, 2006.

Physical Explanations: Dynamo Transition



Physical Explanations: Instabilities



- Interchange ("Rayleigh-Taylor") instability \rightarrow widely understood to be responsible for the largest-scale (larger than \sim 10s–100s m) irregularities.
- Irregularities responsible for VHF backscatter are likely due to a different instability process that is pumped/driven by the interchange instability, neutral wind, or both.

Physical Explanations: Precursors



After Makela and Miller, 2008.

- Large-Scale Wave Structure (visible in ISR and imaging); relation to PRE?
- Bottom-type layers

 → patchy
 irregularities that
 grow within
 LSWS/PRE
 structure
- Depletions always grow out of crests; crests do not always produce depletions

Practical Consquences

- Time transfer and timing on trans-ionospheric radio circuits (e.g., Global Navigational Satellite Systems (GNSS)).
- Loss of lock on coded/spread communication links.



- Role of neutral dynamics in seeding.
 - Neutral wind.
 - Gravity wave sources.
- Role of *E*-region density in seeding \rightarrow hard to observe.
- Distribution of irregularities by scale and location.
- Conjugacy (**E**-field mapping).
- Forecasting.



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