

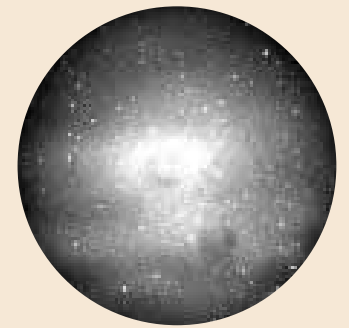
**Comparing momentum flux of mesospheric
gravity waves using different background
measurements and their impact on the
background wind field**

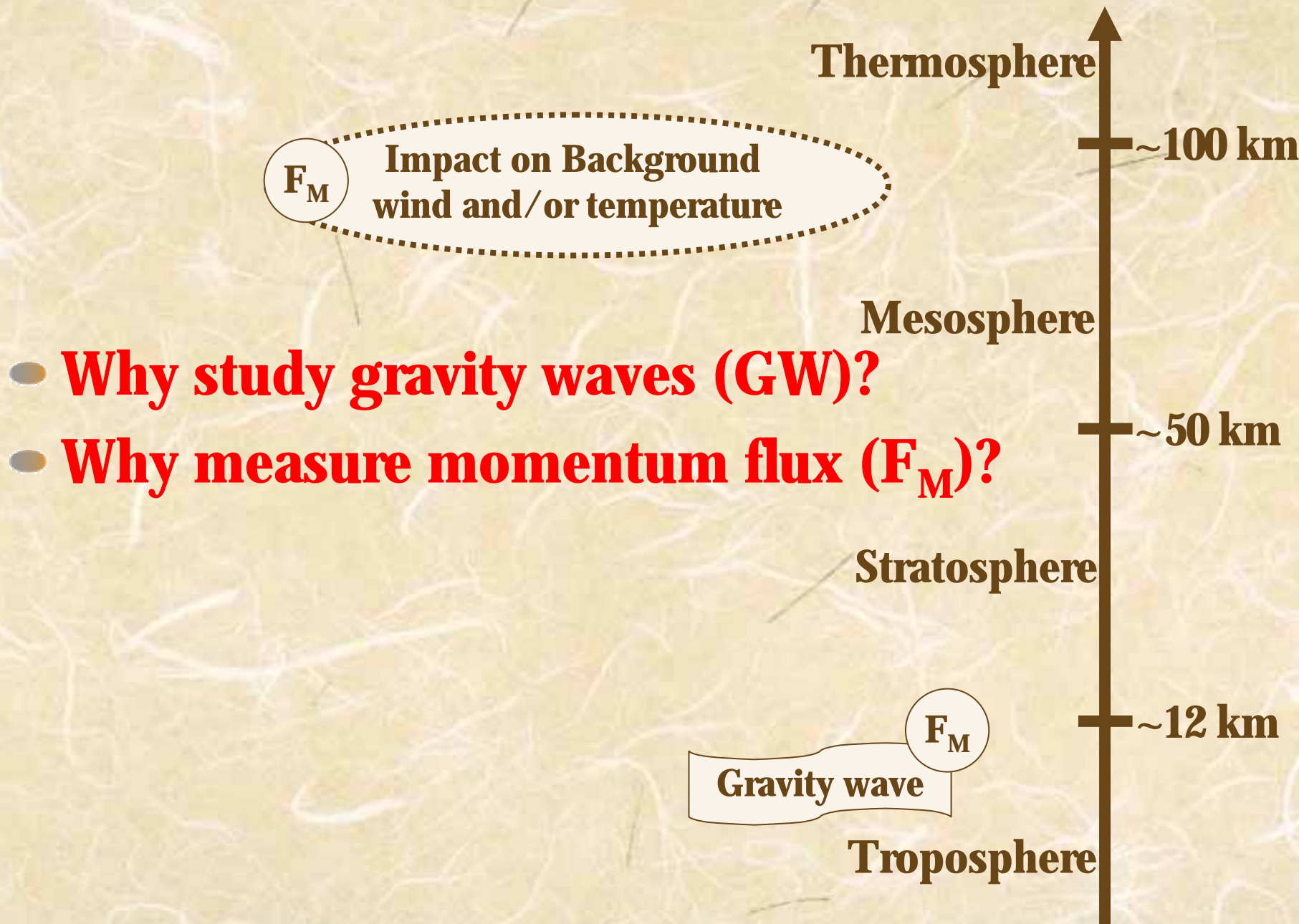
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Outline

- **Introduction**
- **Instrumentation at Maui-MALT, Hawaii**
 - **Mesospheric Temperature Mapper (MTM)**
 - **Na Wind/Temperature Lidar**
 - **Meteor Wind Radar**
- **Case study 1:**
 - Radar vs. Lidar for Momentum Flux Studies**
- **Case study 2:**
 - Gravity Wave-Critical Level Interaction**
- **Conclusions**





Momentum Flux Calculations for Quasi-Monochromatic Events

● Momentum Flux (F_M) equations:

▶ $F_M = \frac{k}{m} \cdot \frac{g^2}{N^2 \cdot CF^2} \cdot \left(\frac{I'}{\bar{I}}\right)^2$ ----- (1) assumption of " $\lambda_h \gg \lambda_z$ "
[Swenson and Liu, 1998]

▶ $F_M = \frac{k \cdot m}{(k^2 + m^2)} \cdot \frac{g^2}{N^2 \cdot CF^2} \cdot \left(\frac{I'}{\bar{I}}\right)^2$ ----- (2) assumption of " $\lambda_h \sim \lambda_z$ "

$$\left\{ \begin{array}{l} m^2 = \frac{N^2}{(c-u)^2} - k^2 - \frac{1}{4 \cdot H^2} \\ N^2 = \frac{g}{T} \cdot \left(\frac{dT}{dz} - \frac{g}{C_p} \right) \sim 4.0 \times 10^{-4} \text{ [rad/sec]} \\ H = RT/g \sim 6.0 \text{ [km]} \end{array} \right.$$

k: horizontal wave number
c: horizontal phase speed
I: Intensity

m: vertical wave number
u: background wind
T: background temperature
N: Brunt-Väisälä frequency
H: scale height
g: gravity acceleration
CF: Cancellation factor
C_p: adiabatic laps rate

Qu: What quality background data are needed for F_M ?

Plan

- **We compare results of F_M (for GWs observed by the MTM) calculated using:**
 - **Na wind/temperature lidar data (exhibiting high time and vertical resolution)**
 - **meteor wind radar (with lower resolution as compared to Na lidar but constant operation)**
 - **We investigate the advantages of each method for F_M estimations.**
- **Using a case study of GW dissipation associated with wind filtering at a critical level, we quantify the impact from the GW on the background wind field.**

Case Study 1: Radar vs. Lidar for Event Momentum Flux Estimates

Station

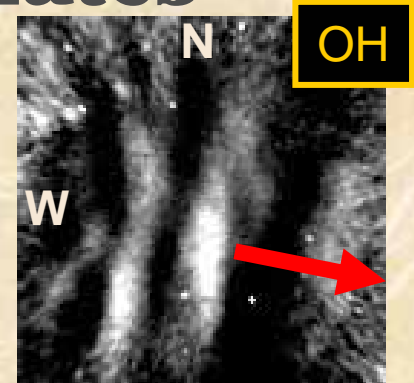
- Maui-AEOS Facility - Hawaii (20.8°N, 156°W)

Data

- Mesospheric Temperature Mapper (MTM)
 - *GW propagation parameters, intensity perturbations*
- Na Wind/Temperature Lidar → *Background wind and temperature*
- Meteor Wind Radar → *Background wind*

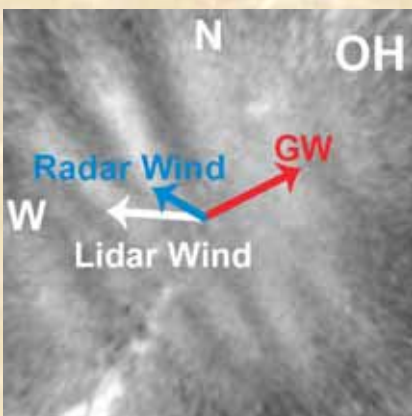
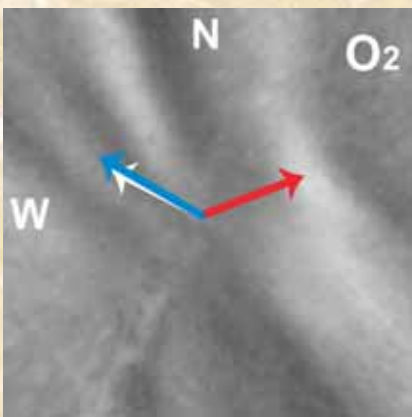
Simultaneous Observations

- 4 events (2 nights)

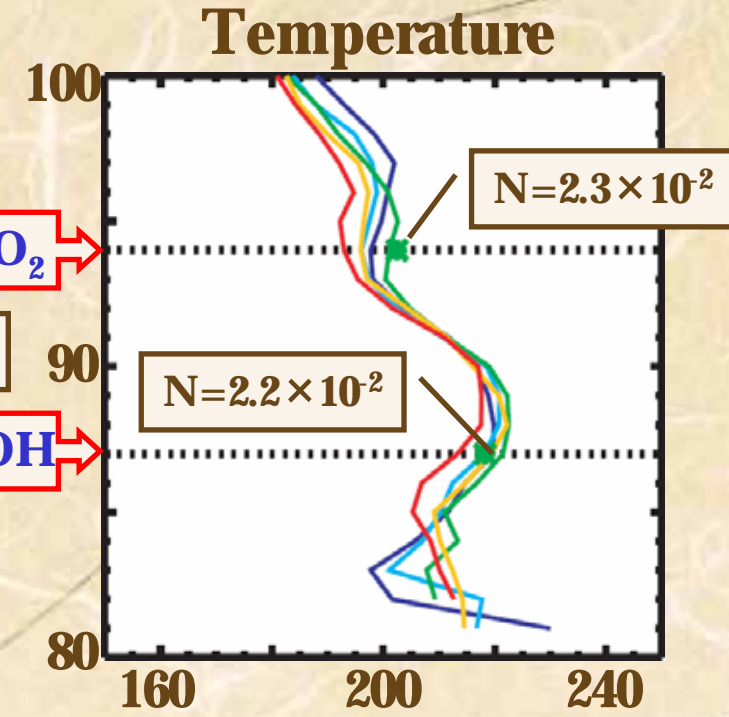
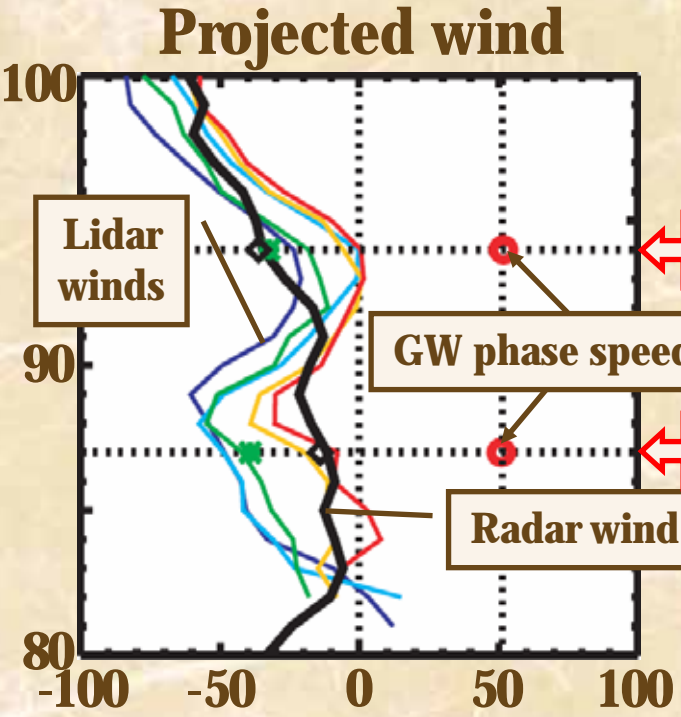


Date	UT	Emission	λ_h [km]	c [m/s]	Period [min]	Direction [°]	DI/I
Jul. 9, 2002	11:30	OH	35.5	51	11.6	63.7	9.5
		O ₂	52.8	51.8	17.0	69.4	12
	14:30	OH	41.3	44	15.6	99.8	12.5
		O ₂	42	46.8	15.0	100	8.9
Aug. 12, 2004	7:30	OH	26.9	38.3	11.7	57.9	5.3
		O ₂	-----	-----	-----	-----	-----
	13:00	OH	22.1	26.9	13.7	103.7	3.6
		O ₂	23.9	25.3	15.7	105.5	3.7

Event #1: July 9, 2002, 11:30 UT



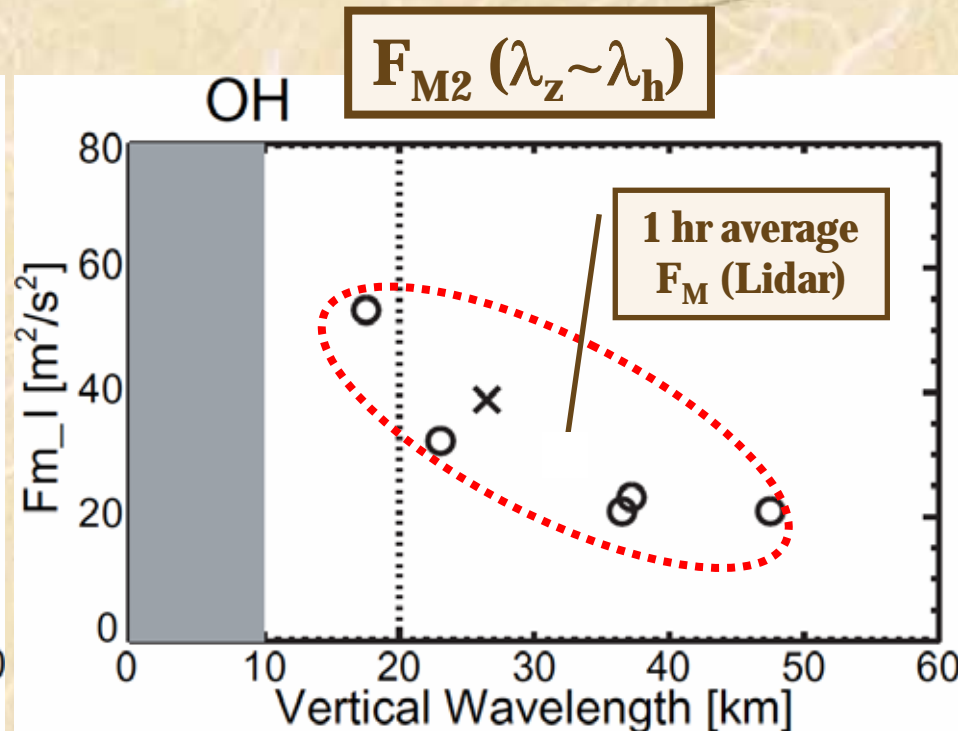
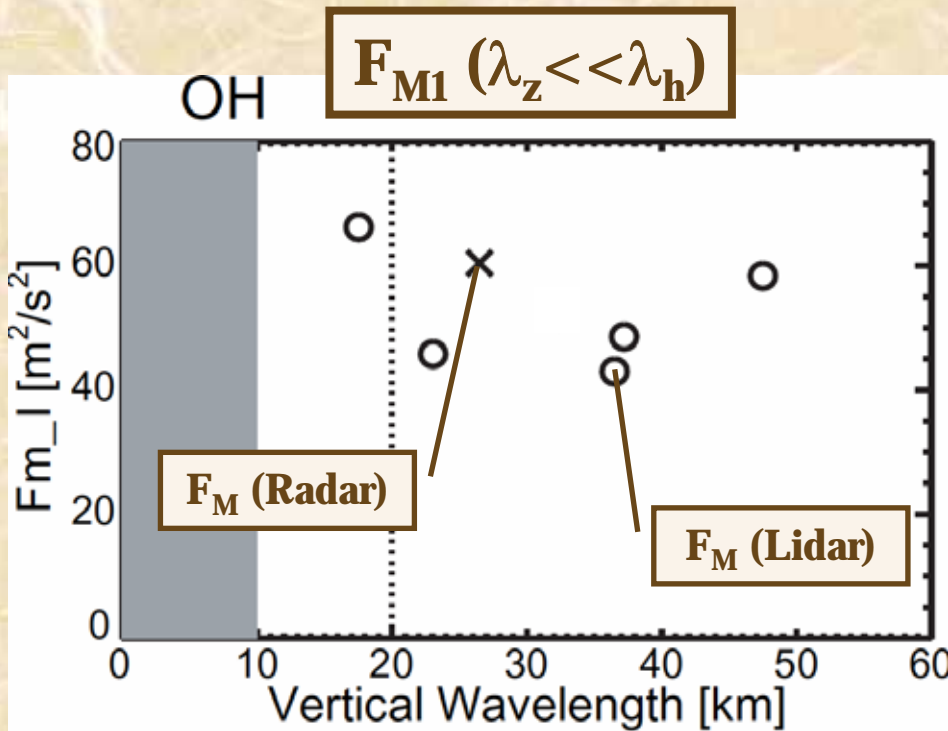
- Wave parameters**
- ▶ $\lambda_h \sim 40$ km
 - ▶ $c \sim 50$ m/s
 - ▶ $I'/I \sim 10\%$



▶ $F_{M1} = \frac{k}{m} \cdot \frac{g^2}{N^2 \cdot CF^2} \cdot \left(\frac{I'}{\bar{I}}\right)^2$ ----- " $\lambda_h \gg \lambda_z$ " [Swenson and Liu, 1998]

▶ $F_{M2} = \frac{k \cdot m}{(k^2 + m^2)} \cdot \frac{g^2}{N^2 \cdot CF^2} \cdot \left(\frac{I'}{\bar{I}}\right)^2$ ----- " $\lambda_h \sim \lambda_z$ "

F_M Results (Event #1)



$F_{M1} \sim 40-70 \text{ m}^2/\text{s}^2$

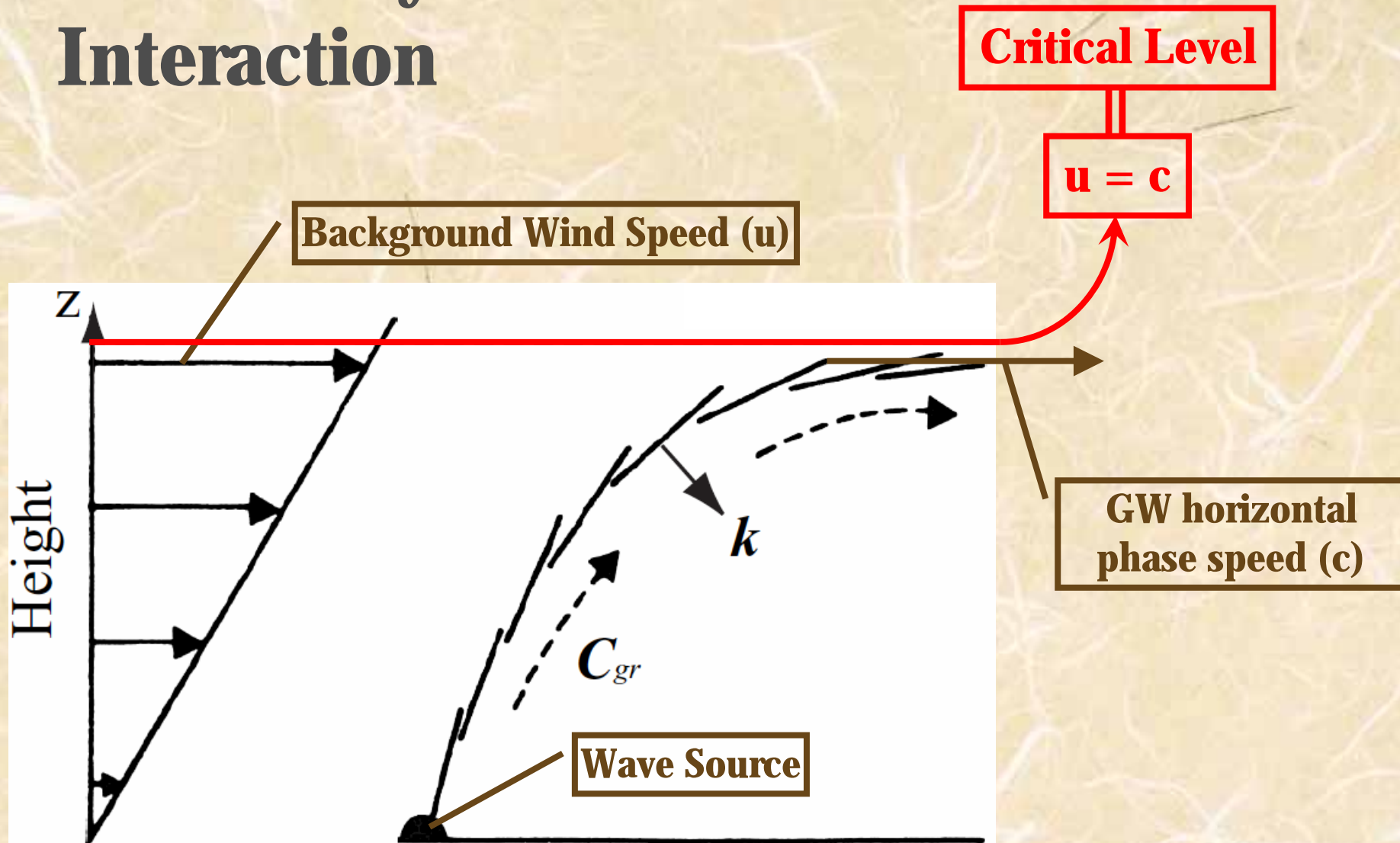
$F_{M2} \sim 20-50 \text{ m}^2/\text{s}^2$

► $F_M(\text{Radar}) \approx 1 \text{ hr average } F_M(\text{Lidar})$

Conclusion 1

- F_M calculated using revised assumption ($\lambda_z \sim \lambda_h$) gives significant lower values ($> \sim 30\%$) than method used in previous studies ($\lambda_z \ll \lambda_h$).
- Background wind profiles are more critical for estimating vertical wavelength and F_M than background temperature data.
- F_M (radar wind) \sim 1 hr average F_M (Na lidar wind & temp.)
- **Thus:**
Under typical mesospheric conditions, Meteor Radar wind data (combined with constant N value) produces reasonable estimates of the wave event F_M for long-term studies.

Case Study 2: GW-Critical Level Interaction



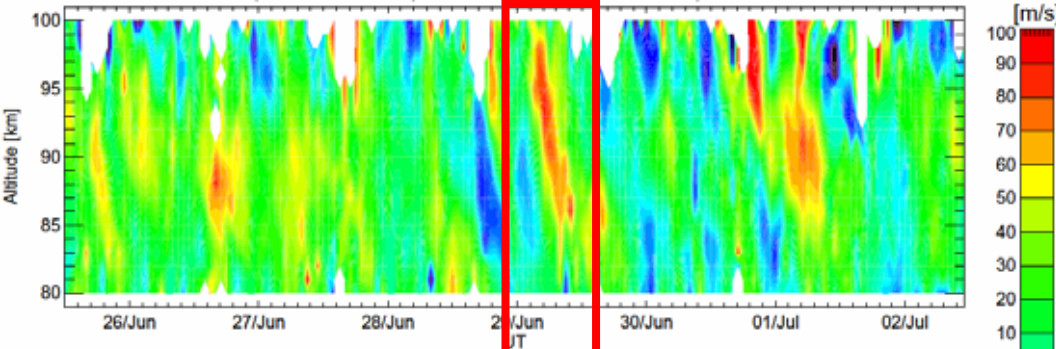
Qu: What happens to a GW at a Critical Level?

Ex. of GW-CL Interaction (June 29, 2003)

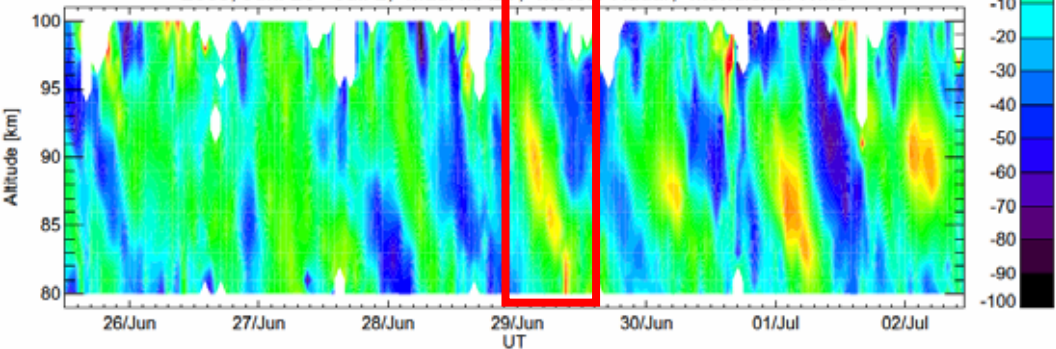
Meteor Radar Winds

MTM Images

2003, Jun. 26 - Jul. 2, Maui MALT, Meteor Rader, Zonal Wind

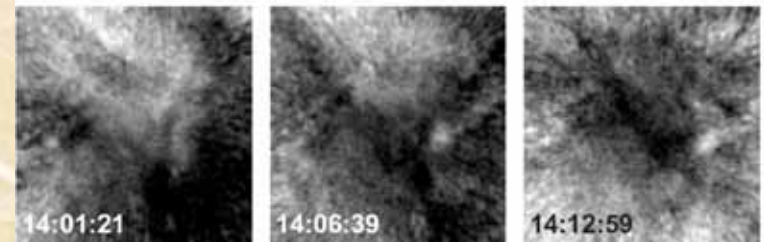
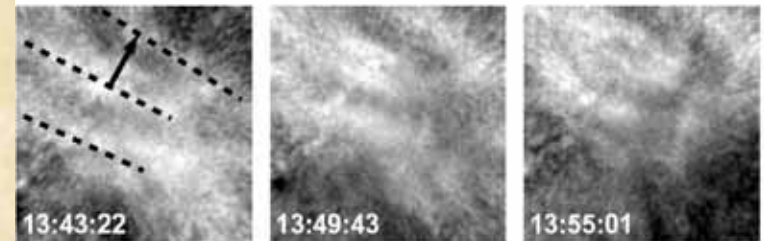


2003, Jun. 26 - Jul. 2, Maui MALT, Meteor Rader, Meridional Wind



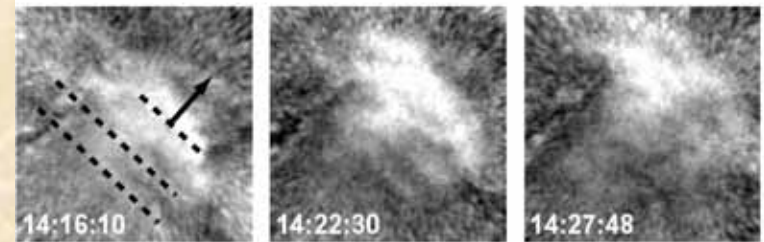
O₂ Band

-15 [%] 15



OH Band

-15 [%] 15

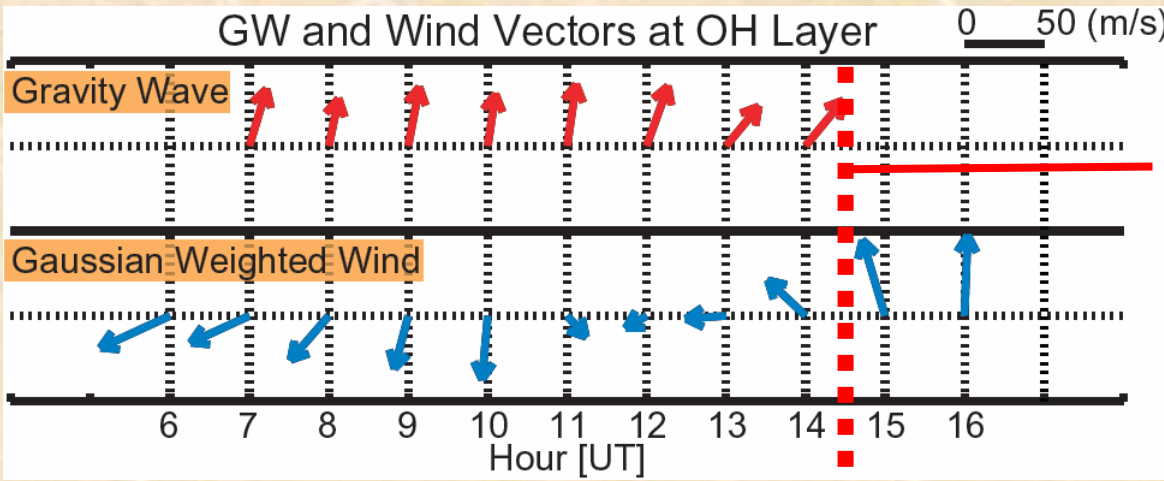


Observations:

- Strong Diurnal-Tide
- GW dissipation at O₂
- GW dissipation at OH

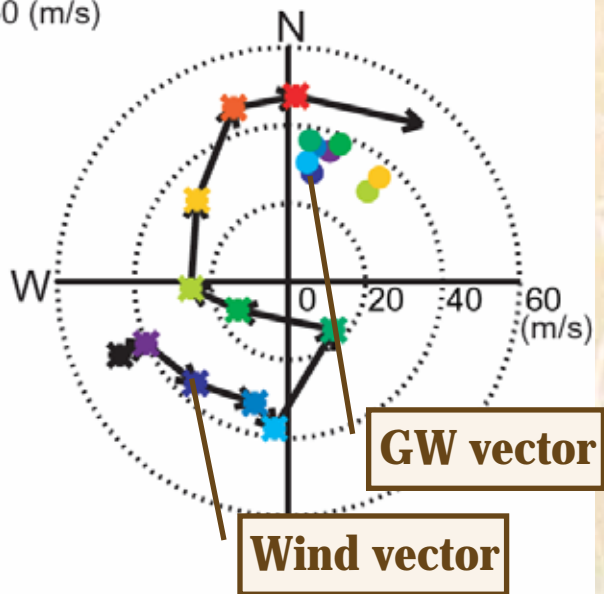
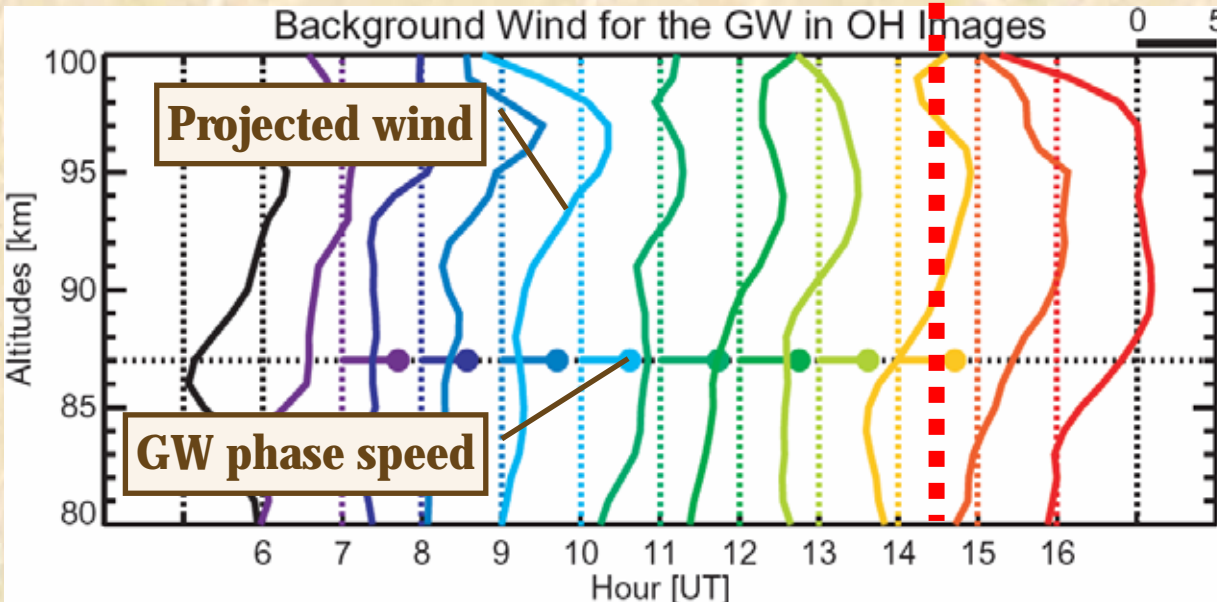
Cause of Wave Disappearance?

OH

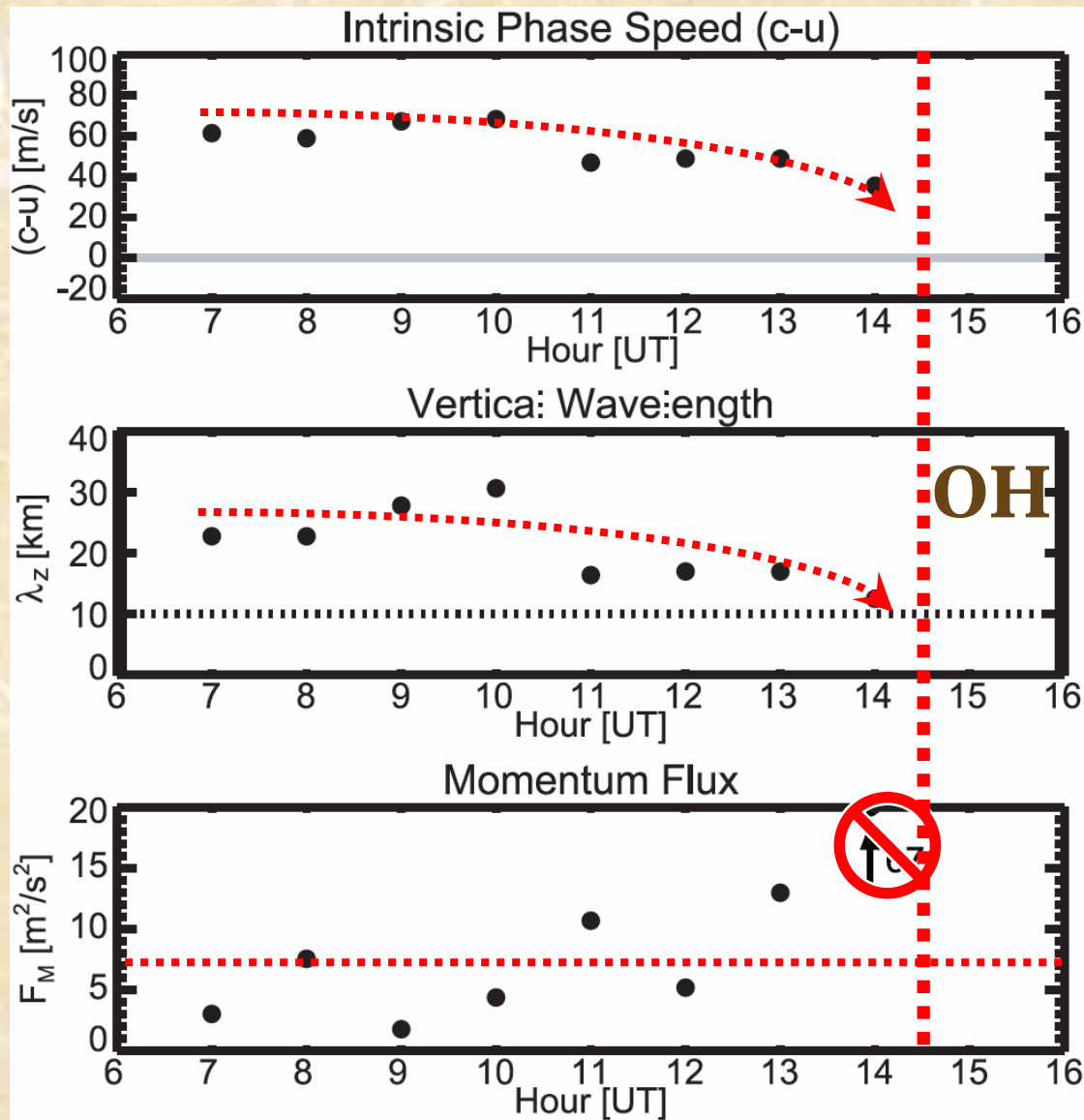


GW-CL interaction

Wave dissipation



Effect of Critical Level on GW Parameters and F_M



Intrinsic phase speed



decreases

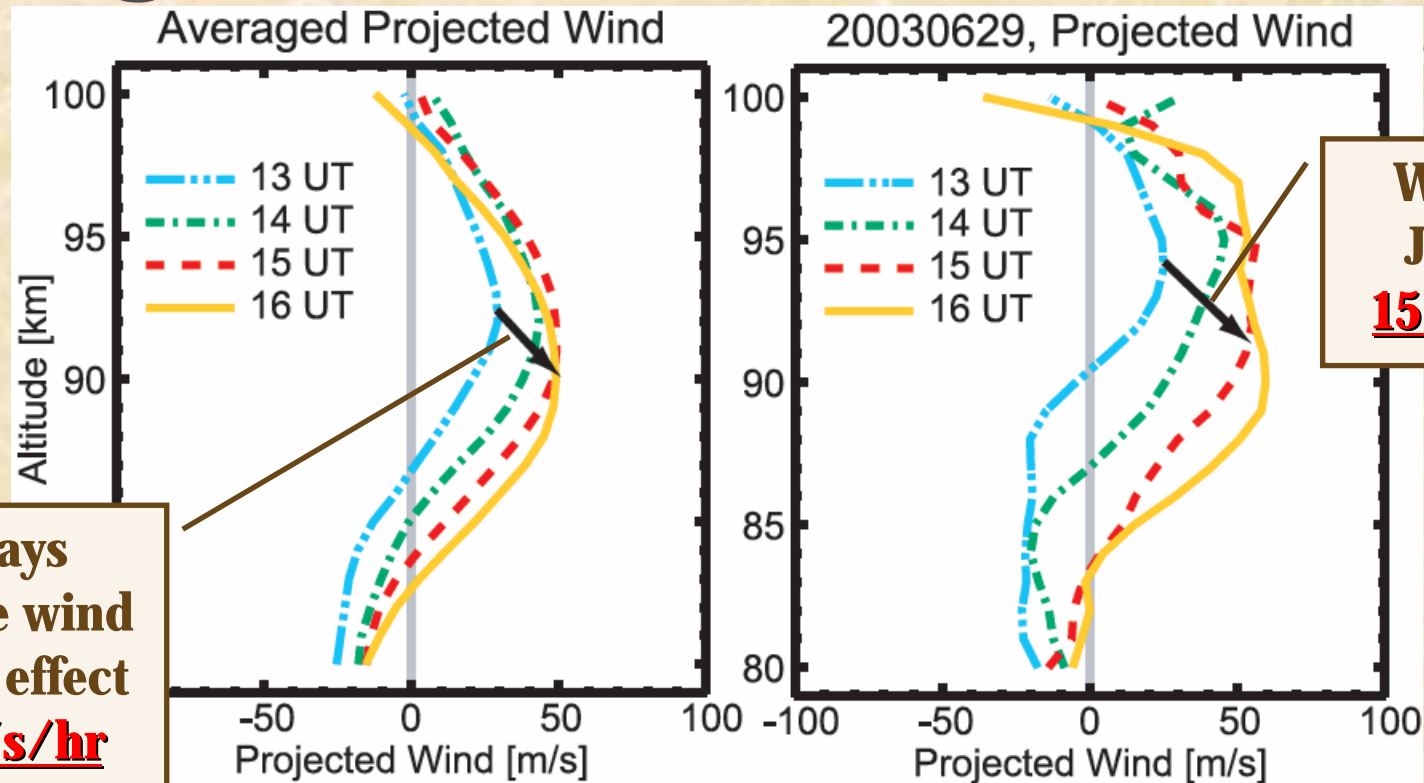
Vertical wavelength



decreases

Average $F_M \sim 7 \text{ m}^2/\text{s}^2$

Effects of Wave Dissipation on Background Winds?



30 days
average wind
~ tidal effect
10 m/s/hr

Wind on
June 29
15 m/s/hr

● Wind change caused by the wave dissipation:

● $\Delta u = F_M \times \Delta t / H \sim 5.8 \text{ m/s/hr}$ [Fritts et al., 2002]

Wave dissipation Δu at the CL $\sim 50\%$ of Tidal Δu

Conclusion 2

- **The GW dissipation was caused by wind filtering at a strong critical level (CL) that was generated by downward phase progression associated with the diurnal tide.**
- **The observed GW-CL interaction impacted the background wind (resulting in an acceleration), but not the background temperature (not shown).**
- **Comparison of acceleration due to the diurnal tide and due to GW dissipation at the CL suggests that F_M from short-period GWs (as observed by the MTM) has the capability to significantly accelerate the background wind field (in this case $\Delta u \sim 50\%$ of tidal effect).**
- **In general, the GW-CL interaction is quite an efficient mechanism, not only for wind filtering of GW propagation but also for wind acceleration by dissipating GW.**

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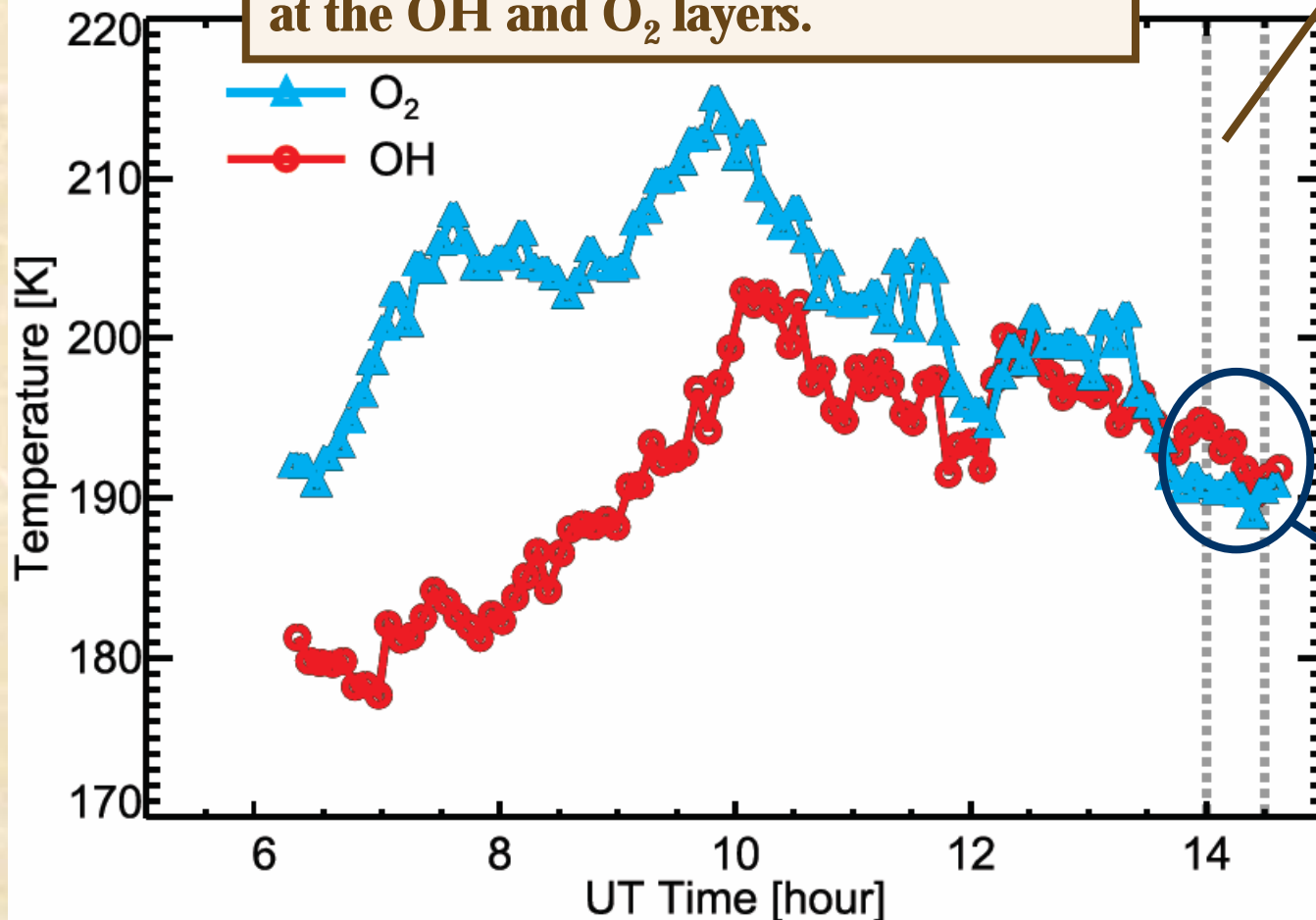


**See you
again!!**

Utah's sky

Effects of Wave Dissipation on Background Temperature?

Zenith temperature variation
obtained by the MTM measurements
at the OH and O₂ layers.



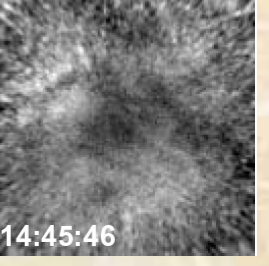
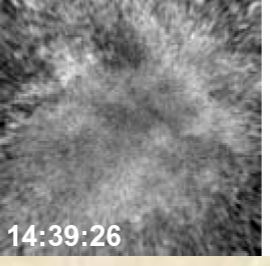
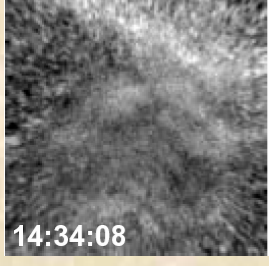
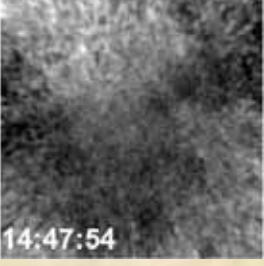
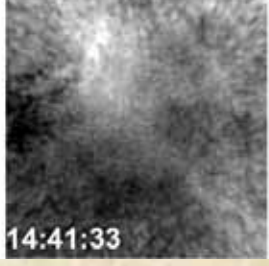
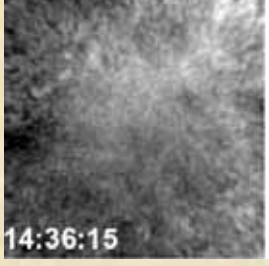
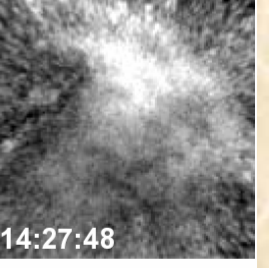
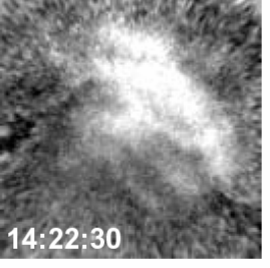
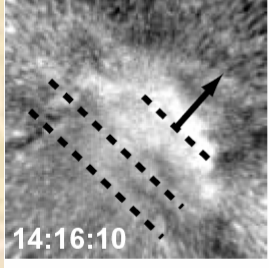
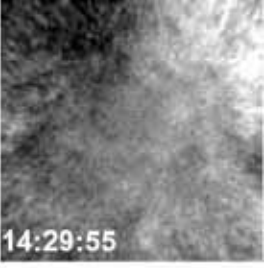
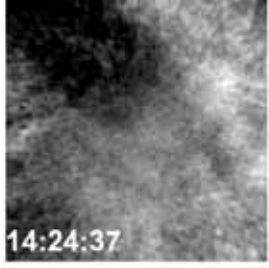
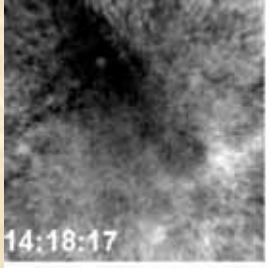
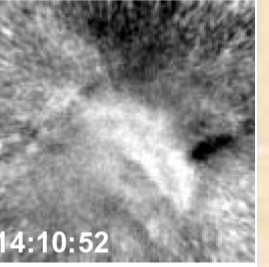
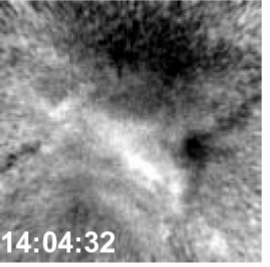
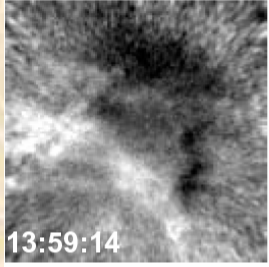
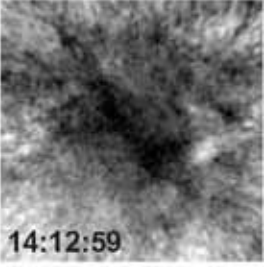
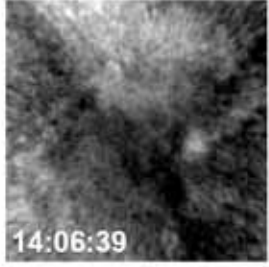
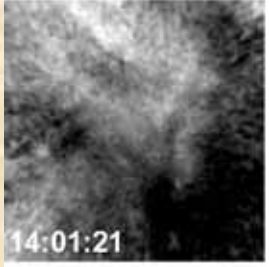
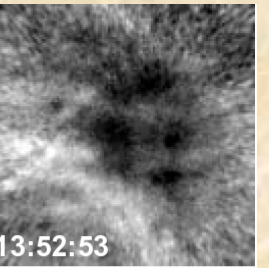
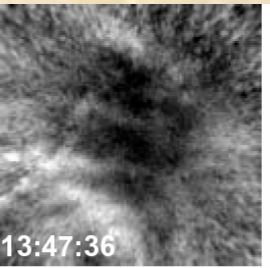
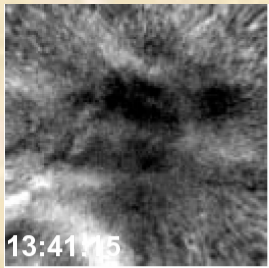
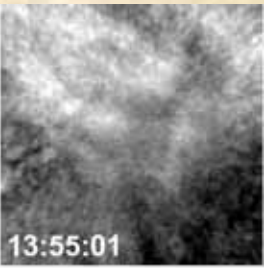
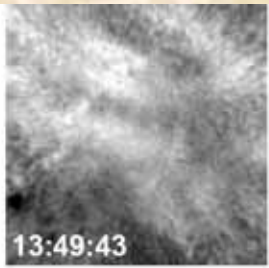
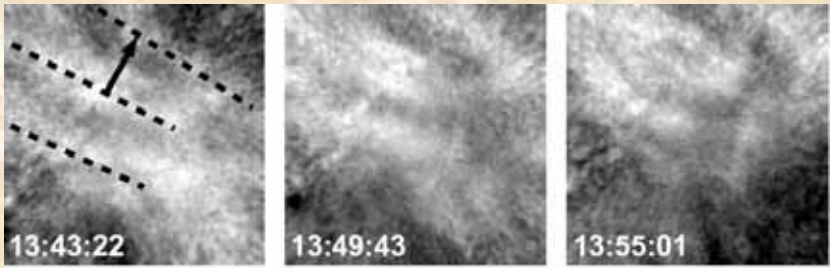
**Critical Level
Region**

**No
significant
change**

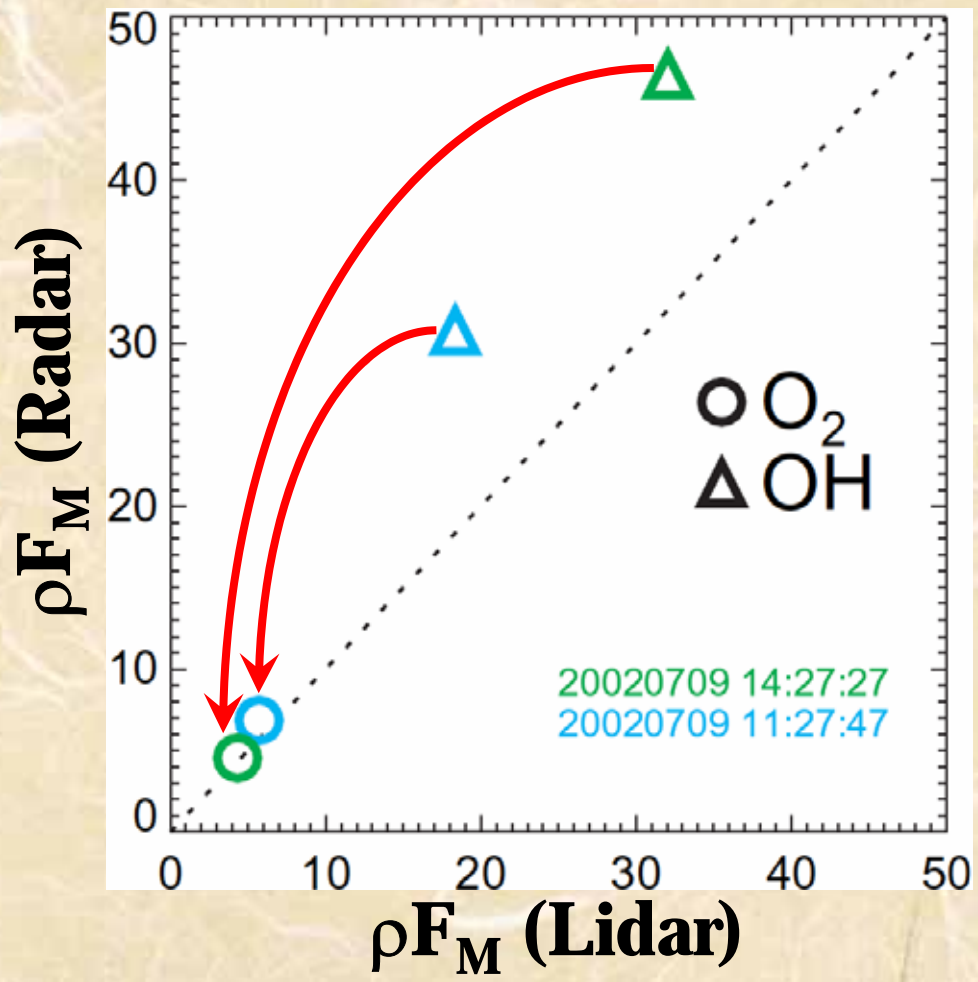
Critical Level Event

● O₂

● OH



Comparison of ρF_M between OH and O₂

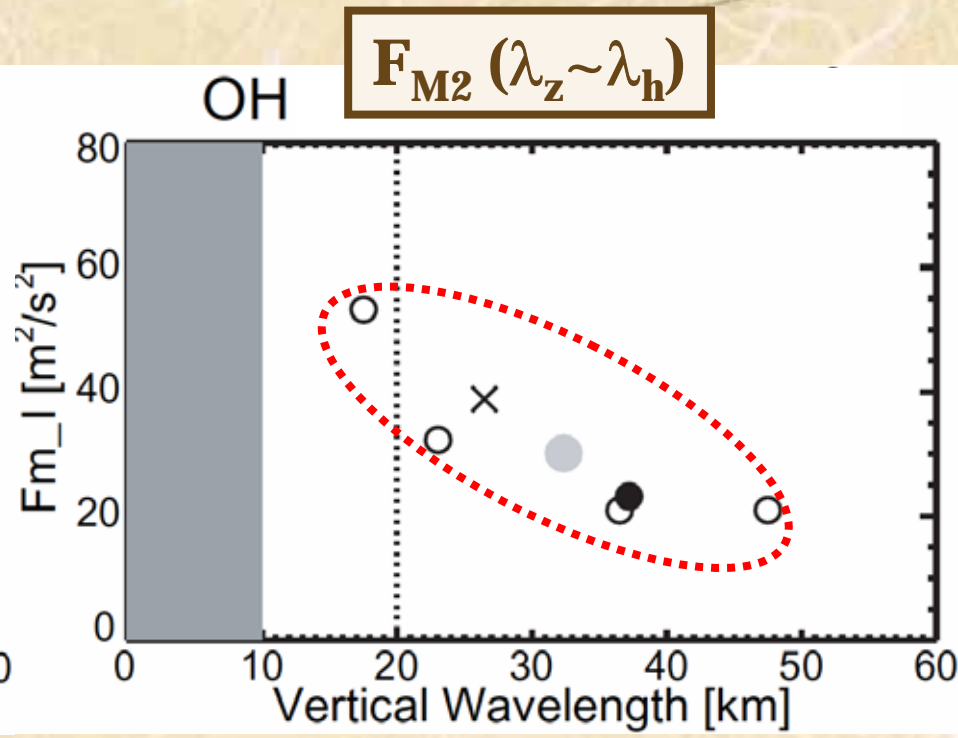
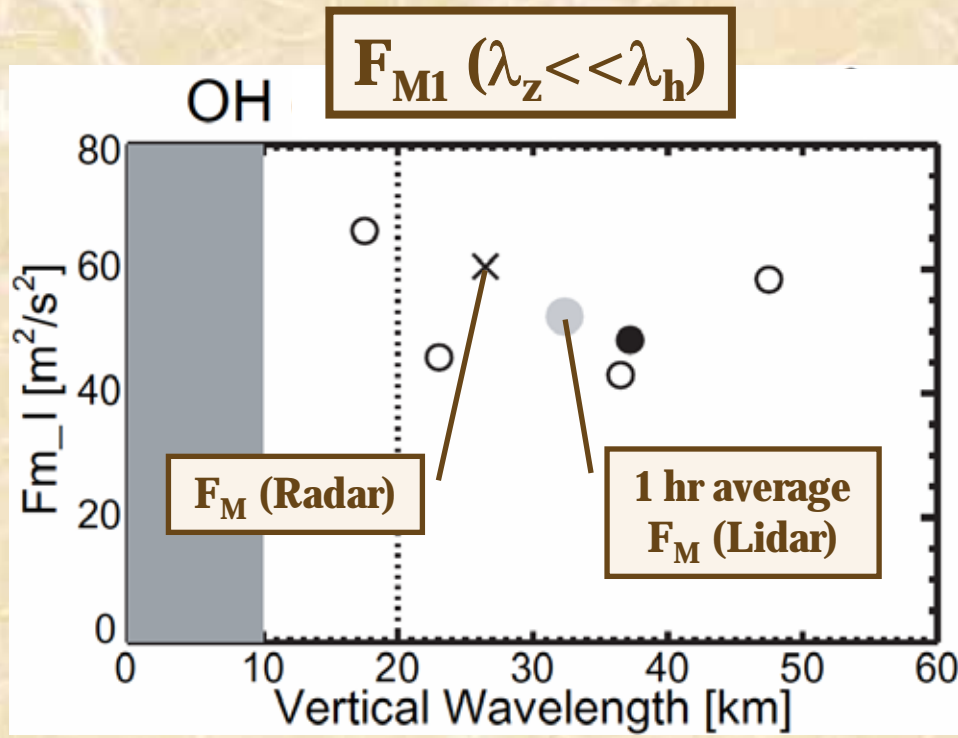


→ GW lose F_M during upward propagation.

Comparing momentum flux of mesospheric gravity waves using different background measurements and there impact on the background wind field

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F_M Results (Event #1)



$F_{M1} \sim 45-60 \text{ m}^2/\text{s}^2$

$F_{M2} \sim 20-40 \text{ m}^2/\text{s}^2$

F_M (Radar) \approx 1 hr average F_M (Lidar)