

Exploring the Sun and its effects on the
Earth's atmosphere and physical environment...

HIGH ALTITUDE OBSERVATORY

Dynamical Coupling in the Earth's Upper Atmosphere

Outstanding Challenges for DASI

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Outline

- Motivation
 - Context: thermosphere-ionosphere space weather
- Observational evidence of ionospheric variability during geomagnetically quiet conditions
 - 1- σ variations in NmF2 during quiet times
 - IMAGE & FORMOSAT-3/COSMIC observations of 4-peaked longitudinal ionospheric variations
- Concluding Remarks

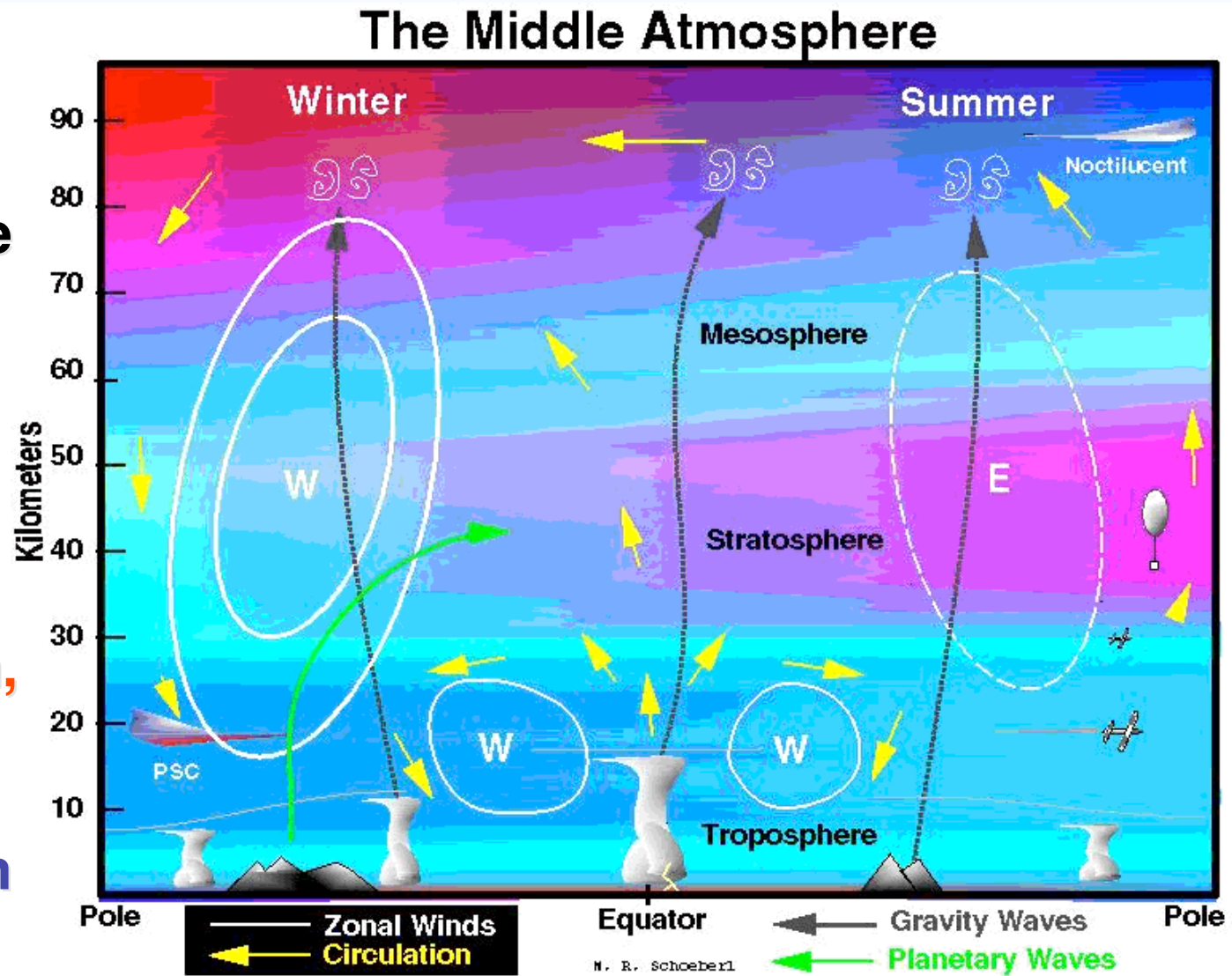
Motivation

- Quiescent variability of thermosphere-ionosphere
 - unresolved research topic
 - potential sources - lower atmosphere 🙌
residual disturbance effects
- Increasing insight into the impact of lower atmospheric sources on the mesopause region
 - 2007 CEDAR tutorials
- One system - the atmosphere doesn't "know" about the boundaries that we impose on it
 - models
 - funding
 - observations
- The mesopause is the "gateway" to the thermosphere-ionosphere

Macro-Scale Effects of Micro- & Meso-Scale Variations

The TI Gateway: The Mesopause Region

Gravity Waves: sources, evolution, dissipation, effects
 → ????
 penetration into the TI

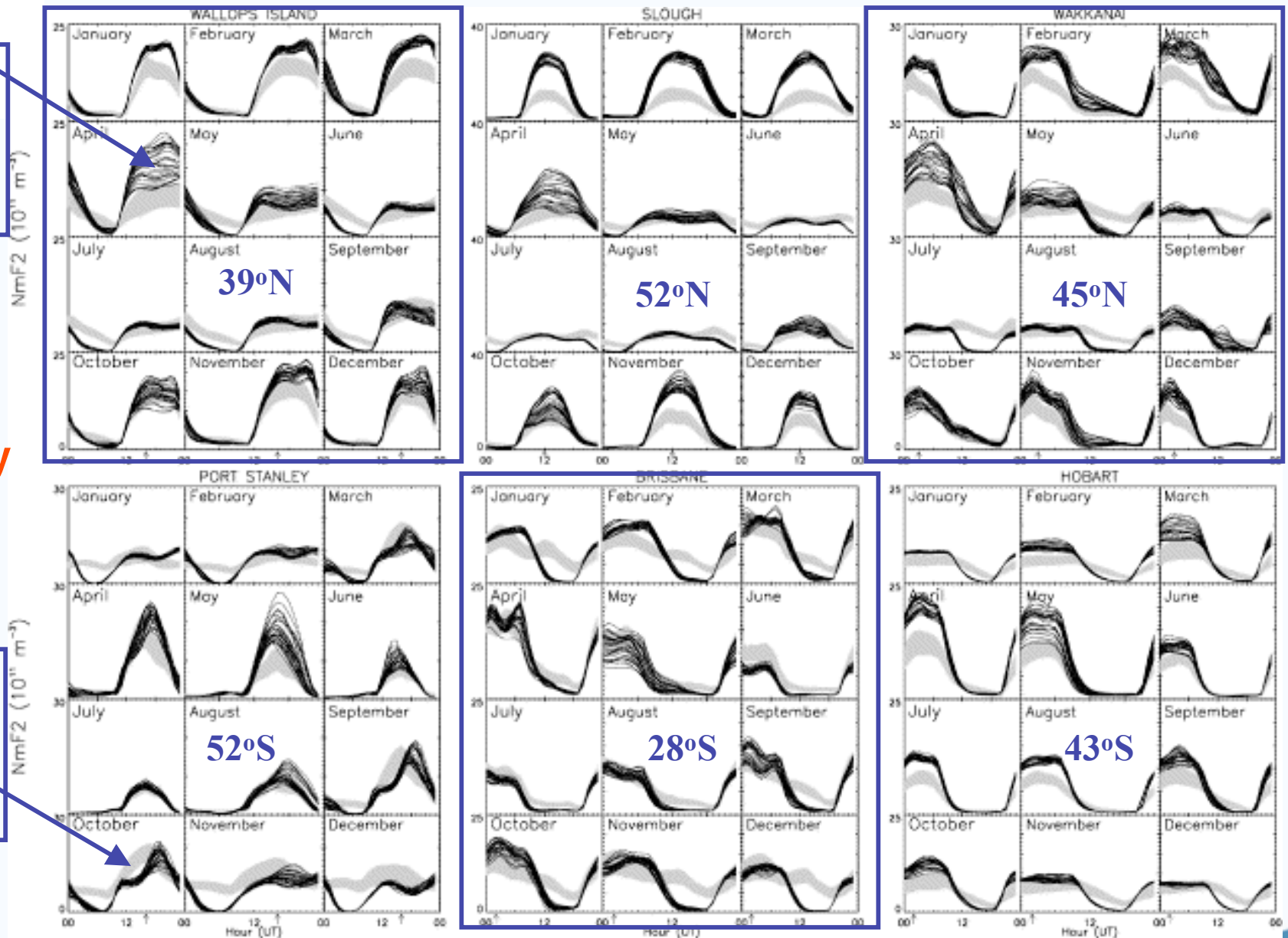


Natural Variability of the Quiescent Ionosphere (after Mendillo et al., 2002)

daily predictions coupled CCM3 TIME-GCM

peak density vs. UT

1 σ : monthly mean ionosonde data

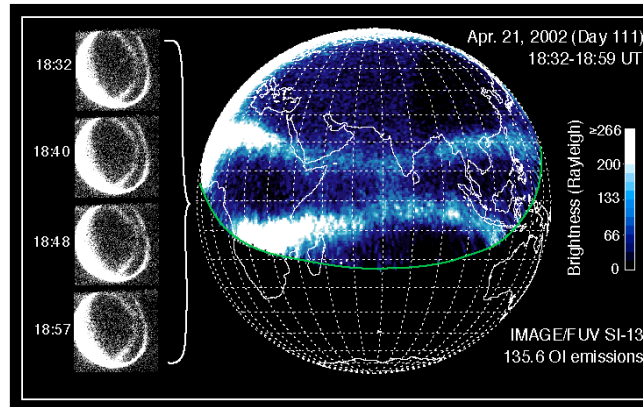


**IMAGE-FUV
observations
of the
equatorial
ionization
anomaly (EIA)
during
March 2002
equinox with
DE3 115-km
temperature
perturbations
from the NCAR
global-scale
wave model
(GSWM)**

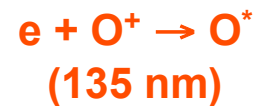
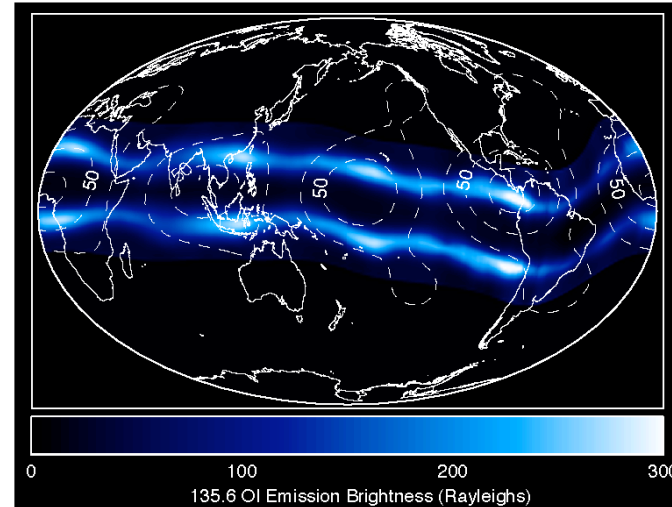
Geophysical
Research
Letters
10 August 2006
Volume 8 Number 88

Immel et al., 2006

Far-Ultraviolet Ionospheric Emissions



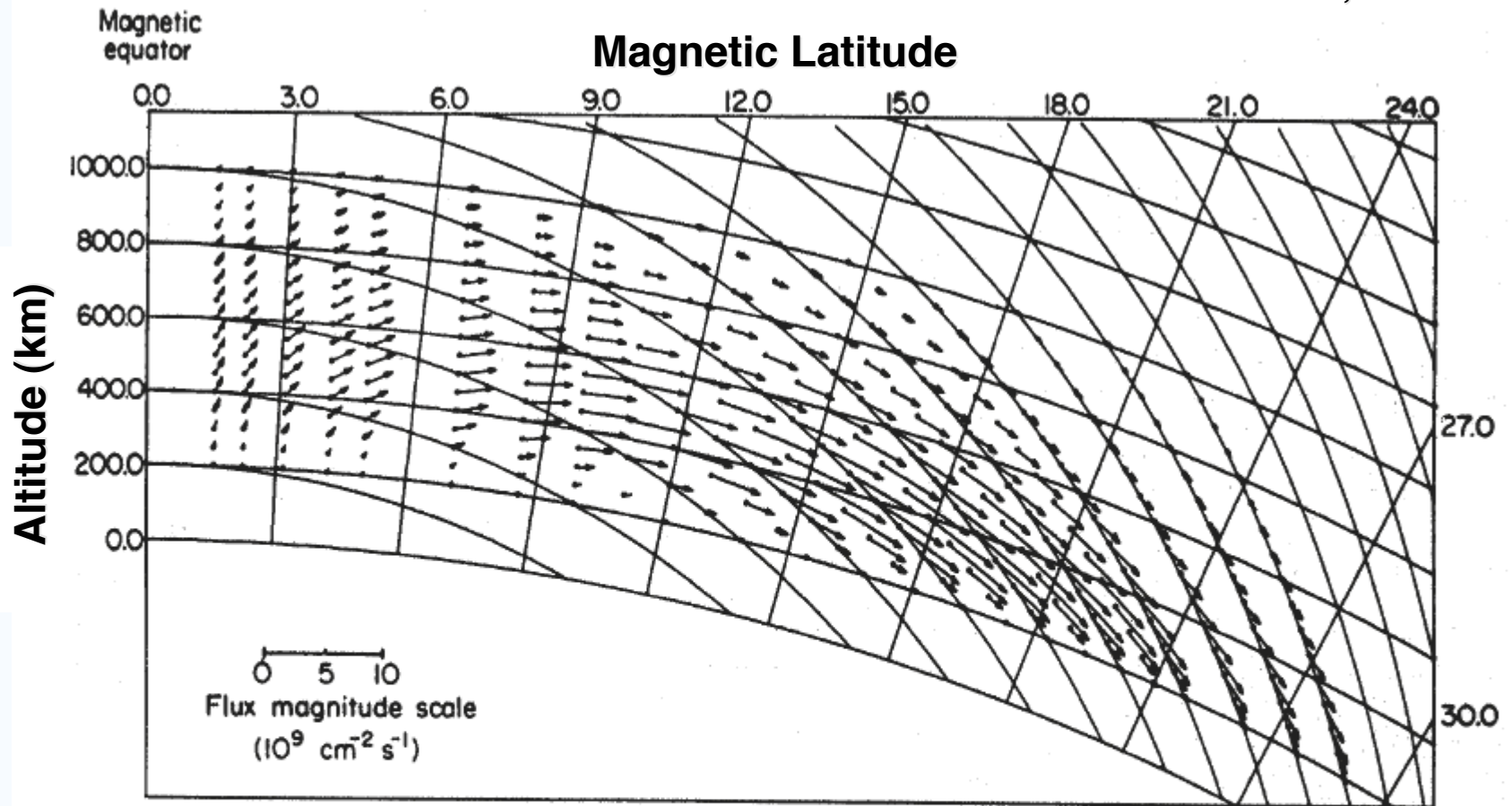
Diurnal Tidal Temperature Amplitude at 115 km Altitude (°K)



$$I = \alpha \int n_e^2 ds$$

Equatorial Ionospheric Anomaly - The Fountain Flux in the Magnetic Meridian Plane

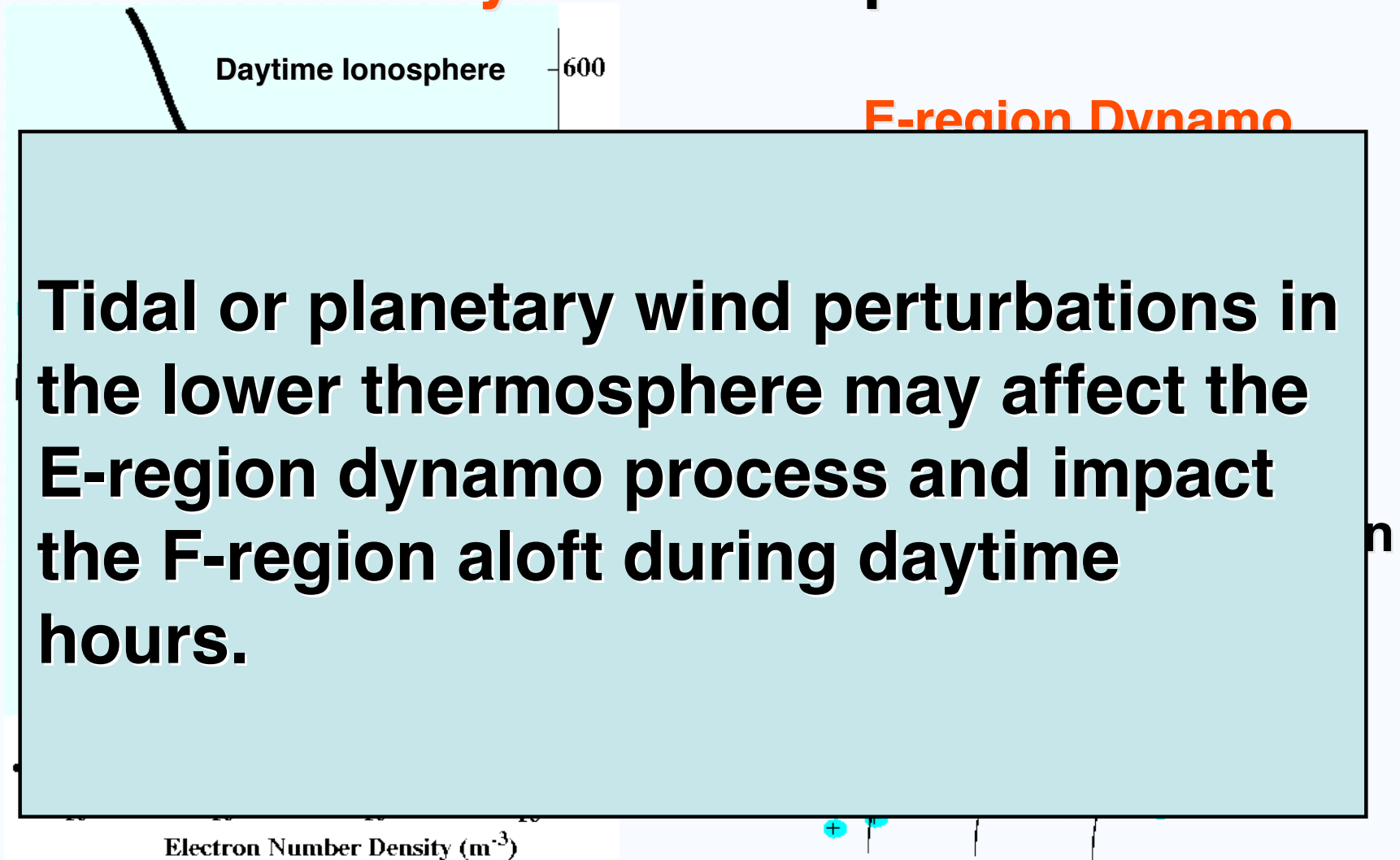
after *Hanson & Moffett, 1966*



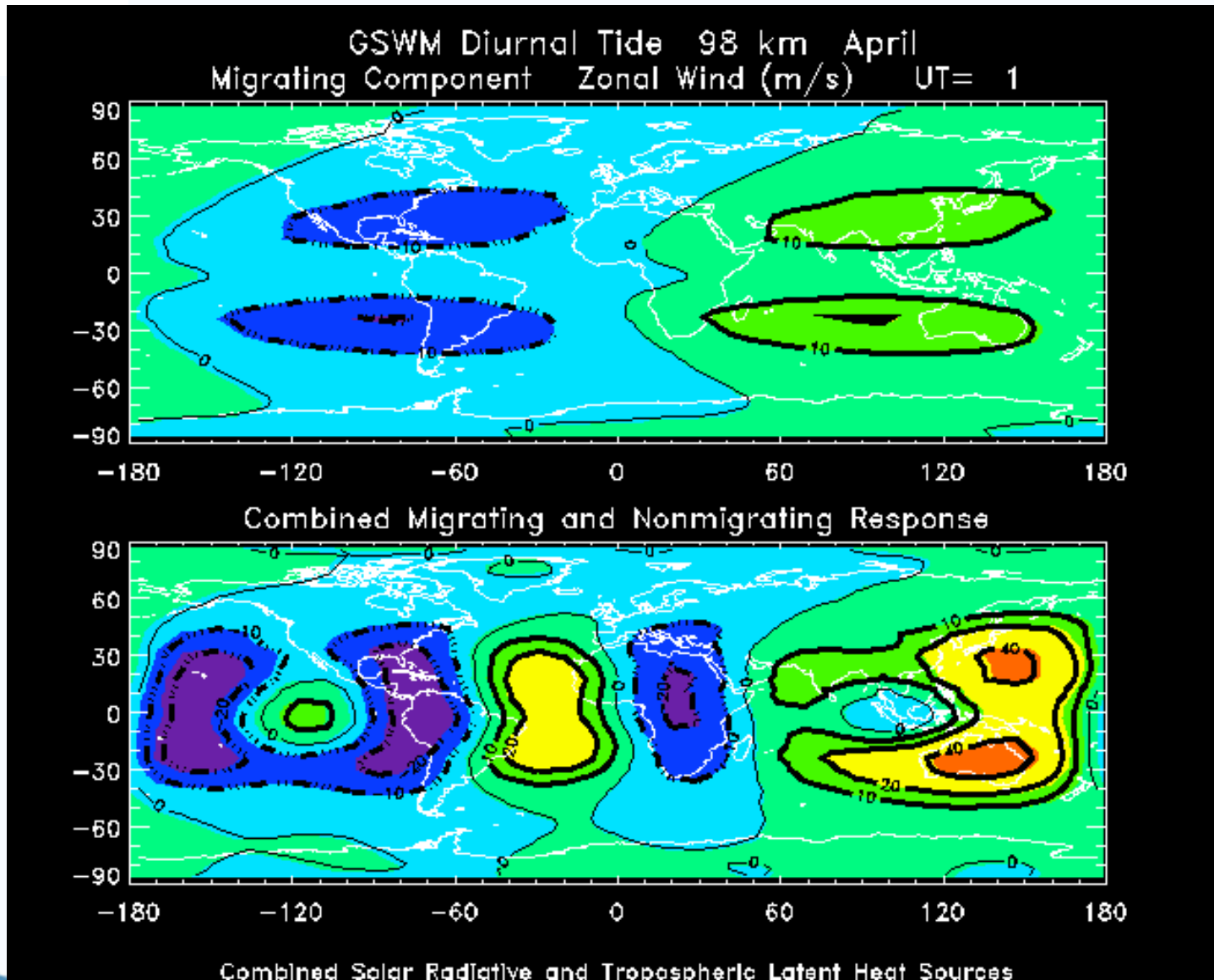
Combined effects of:

Upward drift (eastward E)
Plasma diffusion along B

Coupling into the Low & Middle Latitude Daytime Ionosphere



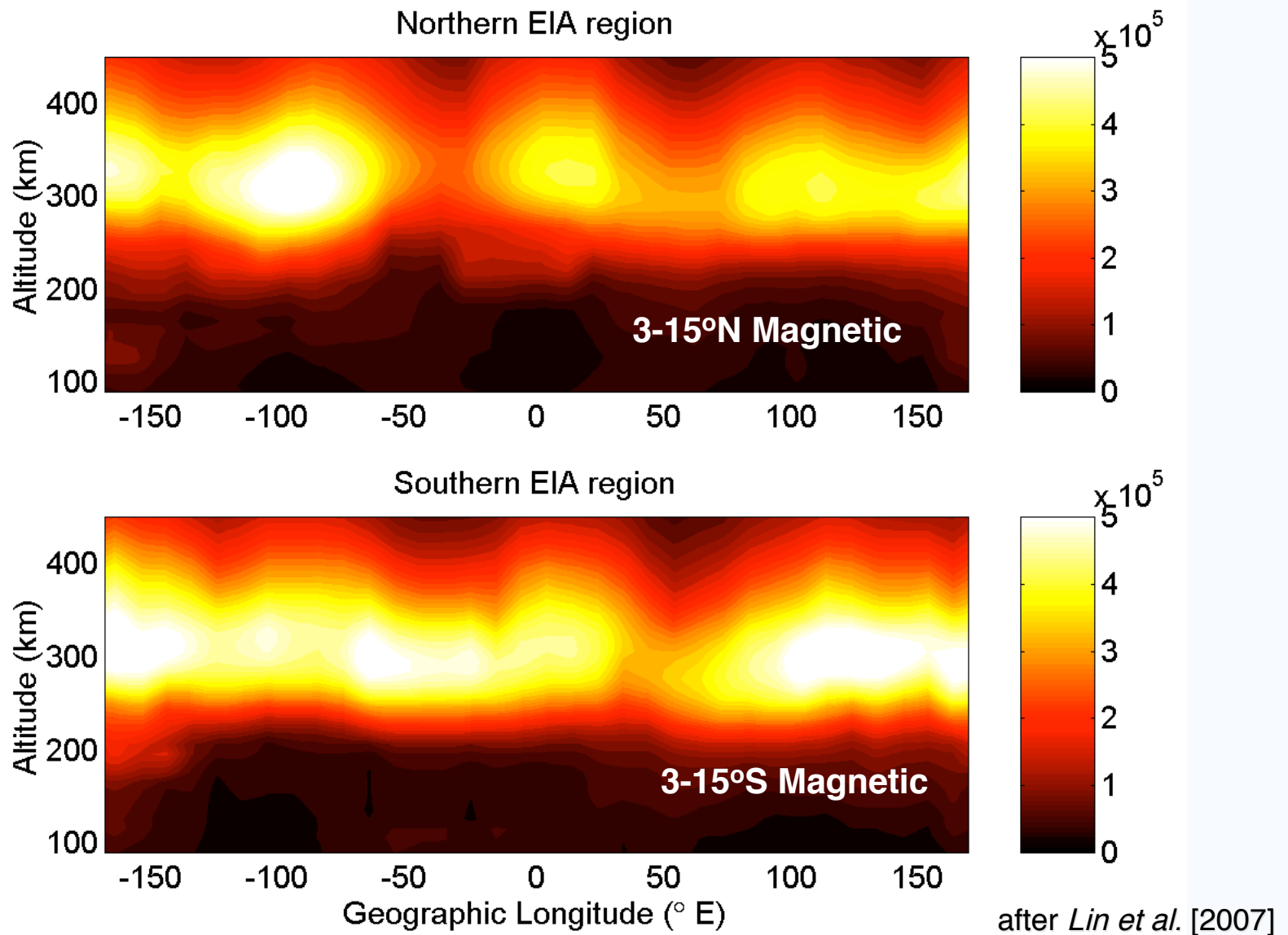
GSWM Diurnal Zonal Wind - 98 km



GSWM-00
migrating
tide

GSWM-02
=
GSWM-00
+
latent heat
response

FORMOSAT-3/COSMIC Electron Density 20-22 LST - Sep 2006



Snapshot

synoptic map - universal time (t_{UT}) perspective
tidal perturbation, f' , at a given latitude -

$$f' = \sum A_{s,n} \cos[\omega_n(t_{UT} - \phi_{s,n}) - s\lambda]$$

s = zonal wavenumber; λ = longitude

n = harmonic - 1=diurnal; 2=semidiurnal, etc.

$\omega_n = 2\pi n/24$ hours

$A_{s,n} / \phi_{s,n}$ = tidal amplitude / phase

Satellite Maps

asynoptic - local time (t_{LST}) perspective

$$t_{LST} = t_{UT} + 360^\circ \lambda / 2\pi 15^\circ$$

$$f' = \sum A_{s,n} \cos[\omega_n(t_{LST} - \phi_{s,n}) - (s-n)\lambda]$$

☞ observed zonal wavenumber = $|s-n|$

IMAGE & F-3/C Signatures

☞ wavenumber 4

$$s = -3, n = 1$$

☞ could be DE3

$$|s-n| = 4$$



TIME-GCM Simulations [*Hagan et al.*, 2007]

- Resolution: horizontal - $2.5^\circ \times 2.5^\circ$
vertical - 4 grid points per scale height
- Geomagnetically quiescent vernal equinox moderate solar conditions
perpetual day of year = 80 (March 21)
10.7-cm solar radio flux ($F_{10.7}$) = 150
hemispheric power = 8 GW
cross-polar-cap potential drop = 30 kV
- Lower boundary (~ 30 km) condition (LBC):
aggregate March global-scale wave model (GSWM) results
diurnal and semidiurnal perturbations - 13 wavenumbers
horizontal winds, temperature, geopotential height

Standard Run:

International Geomagnetic Reference Field (IGRF)

Diagnostic Runs:

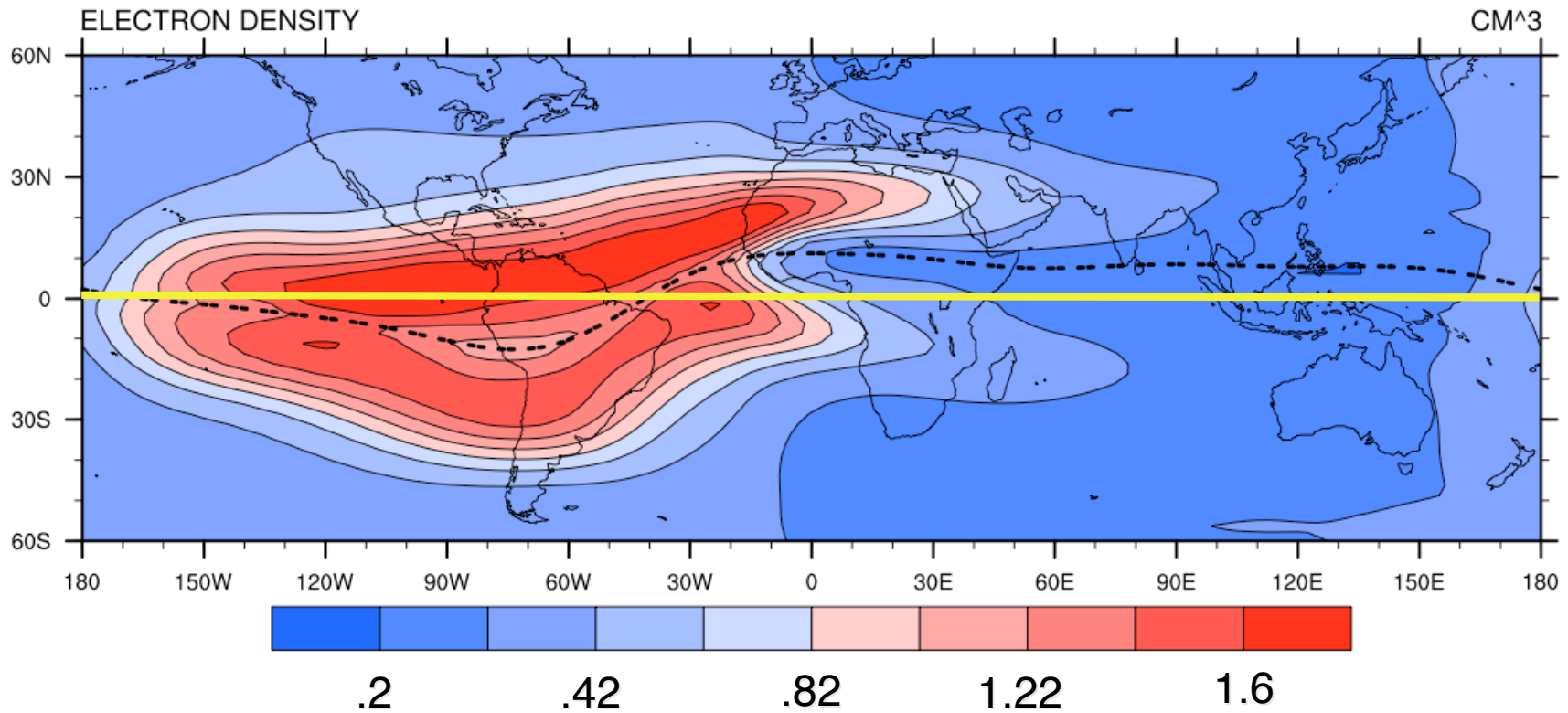
IGRF without tidal LBC (“without tropospheric tides”)

Aligned Dipole geomagnetic field

Aligned Dipole geomagnetic field without tropospheric tides



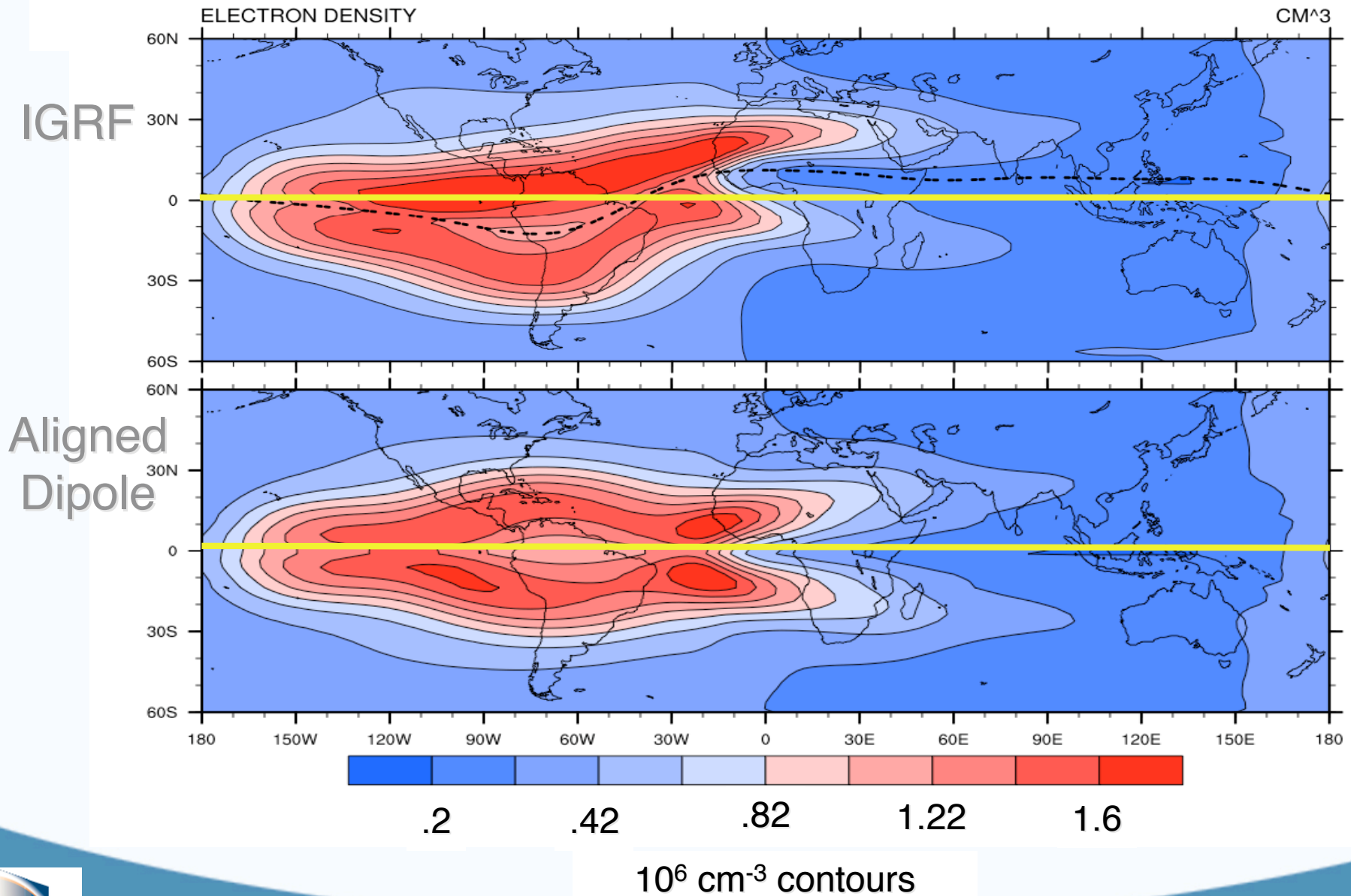
TIME-GCM Electron Density at 20 UT and 450 km



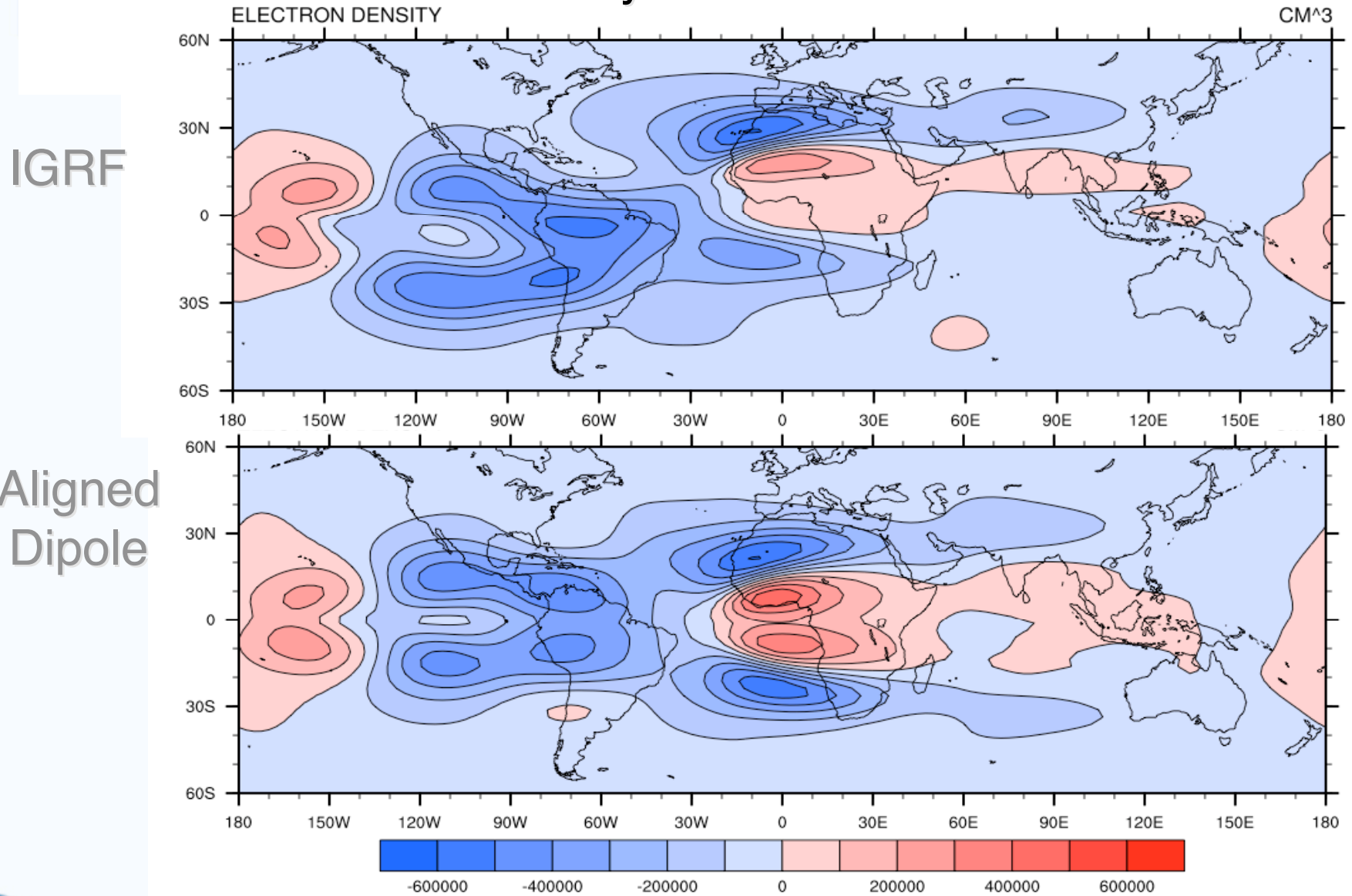
10^6 cm^{-3} contours



TIME-GCM Electron Density at 20 UT and 450 km



TIME-GCM Electron Density Differences* at 20 UT and 450 km

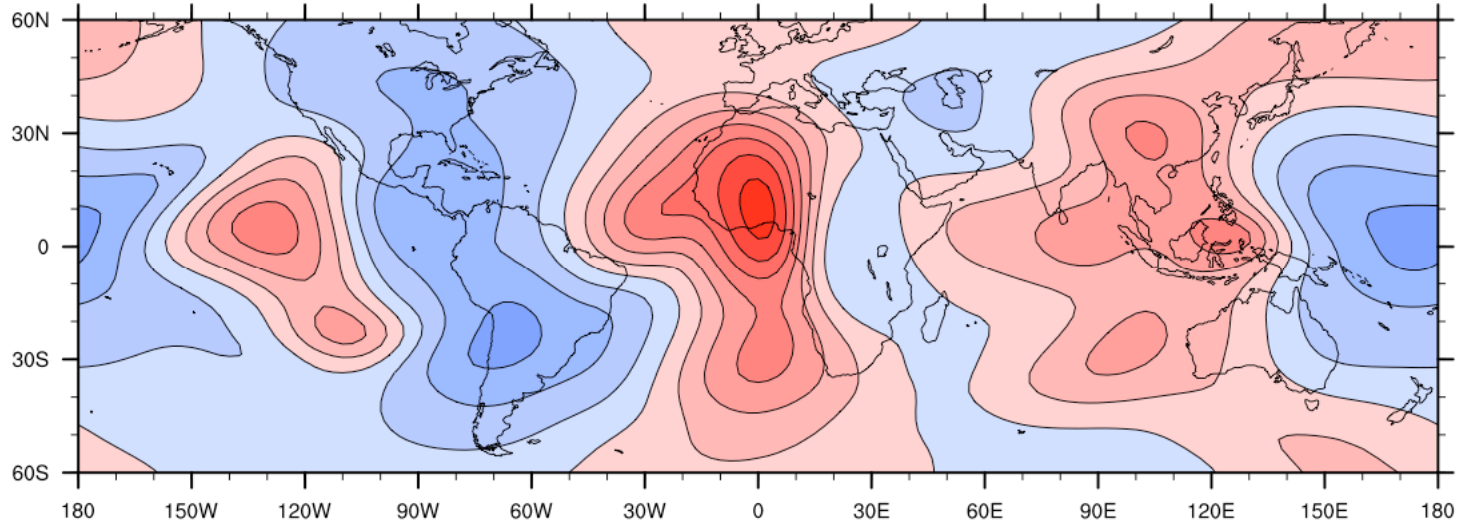


*with - without tropospheric tides

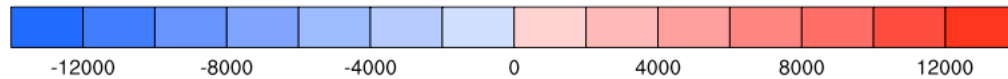
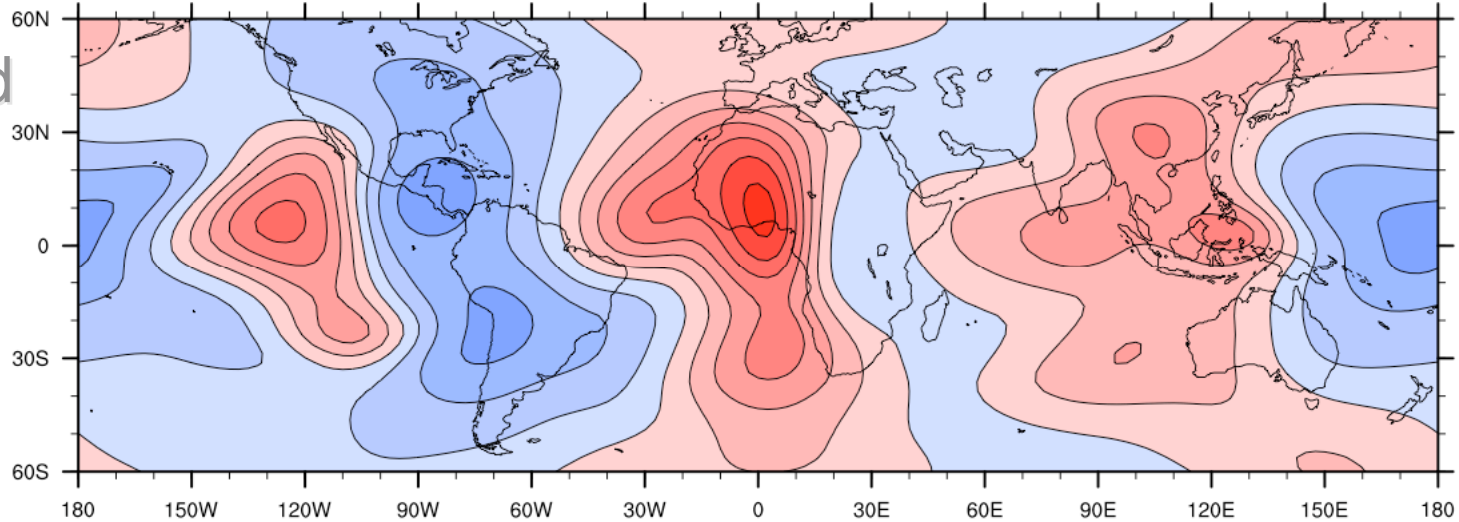


TIME-GCM Zonal Wind Differences* at 20 UT and 110 km

IGRF



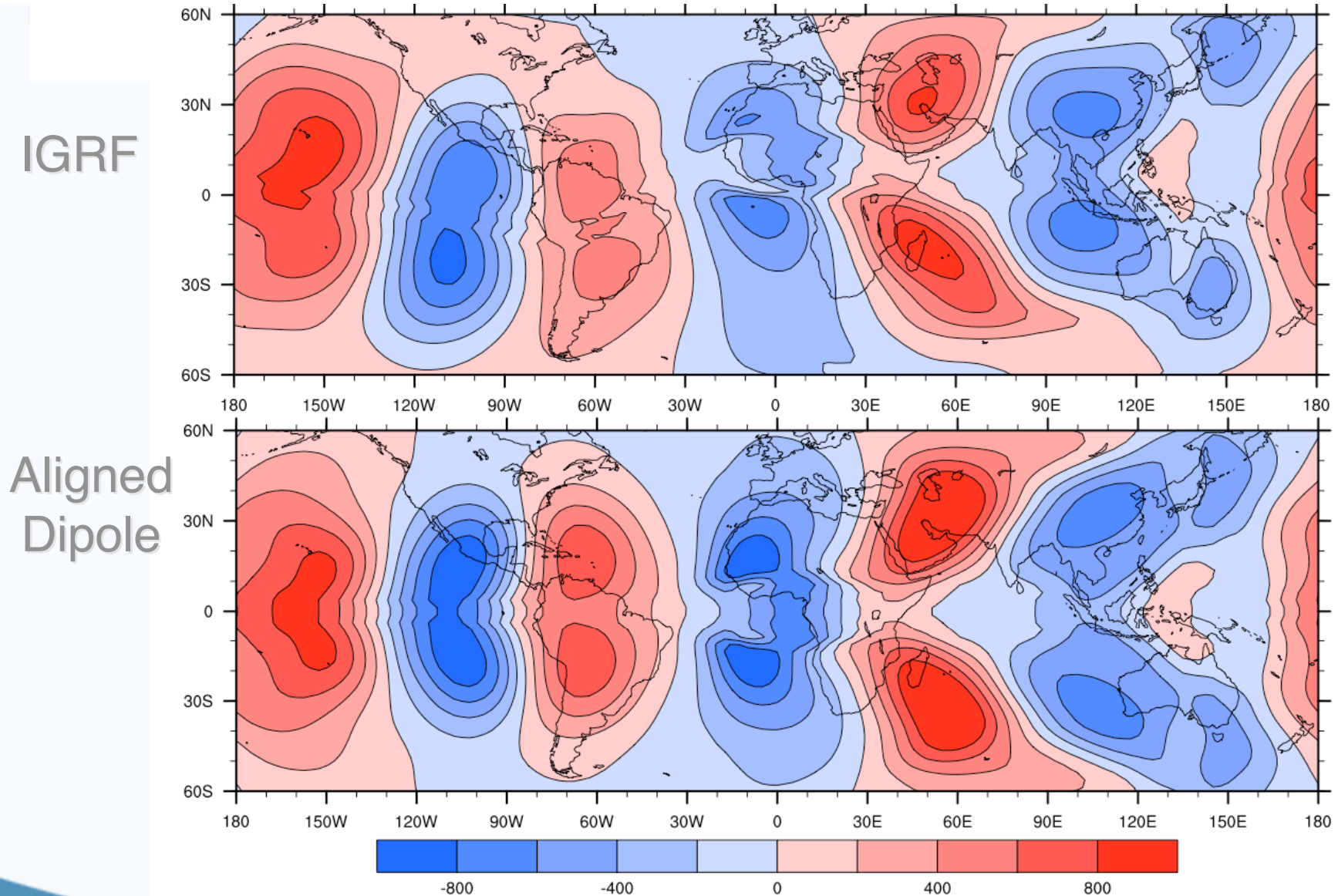
Aligned Dipole



*with - without tropospheric tides



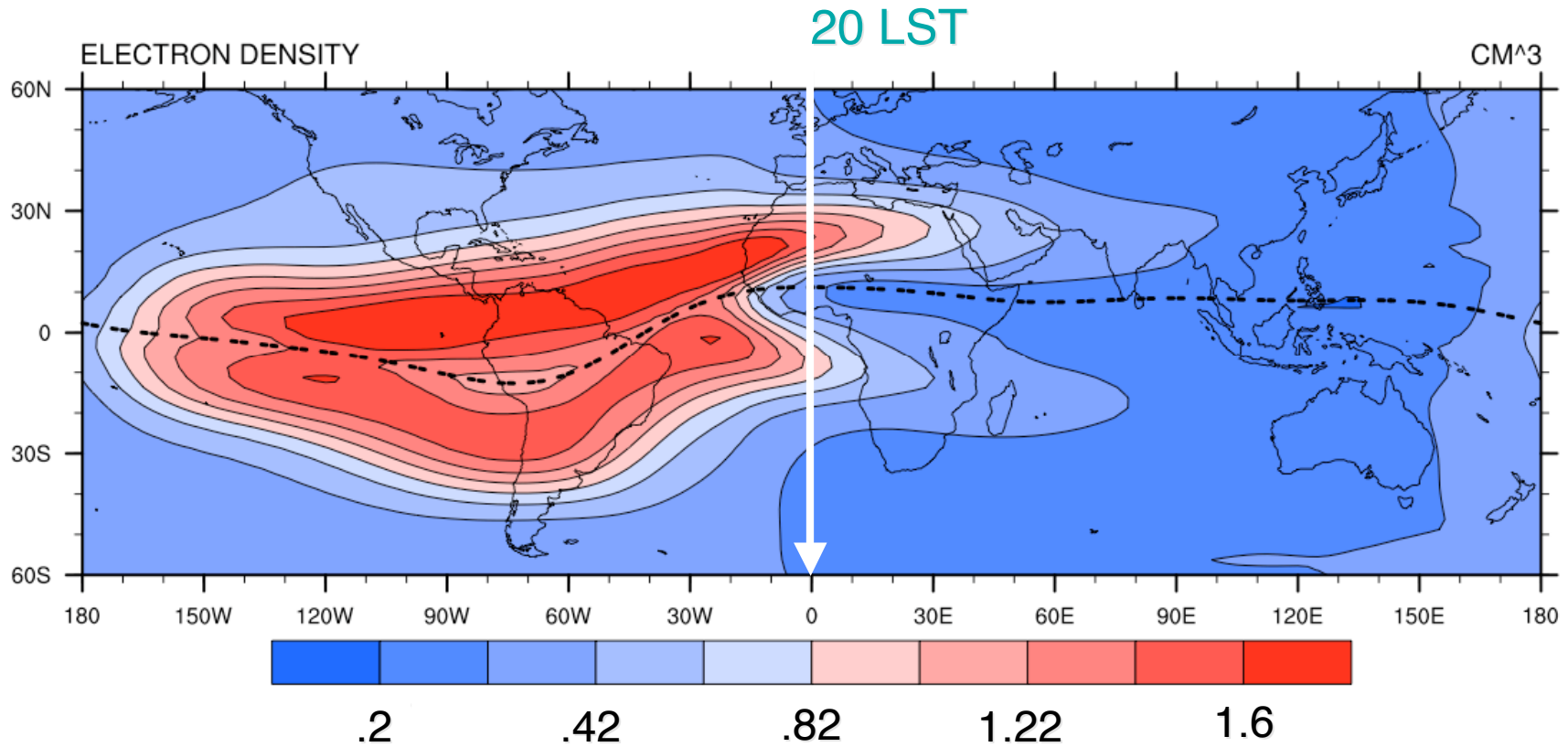
TIME-GCM Vertical ExB Drift Differences* at 20 UT and 110 km



*with - without tropospheric tides



TIME-GCM Electron Density at 20 UT and 450 km



10^6 cm^{-3} contours

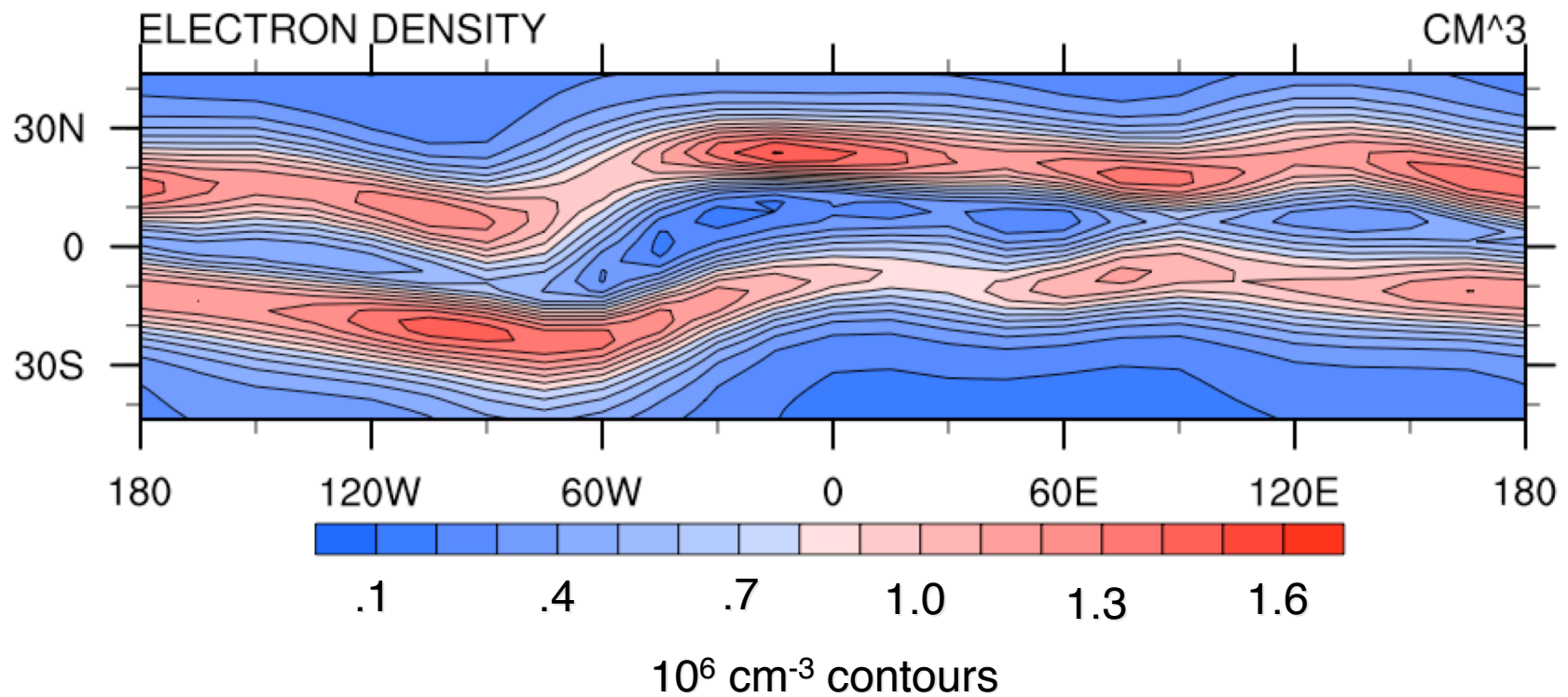


NCAR

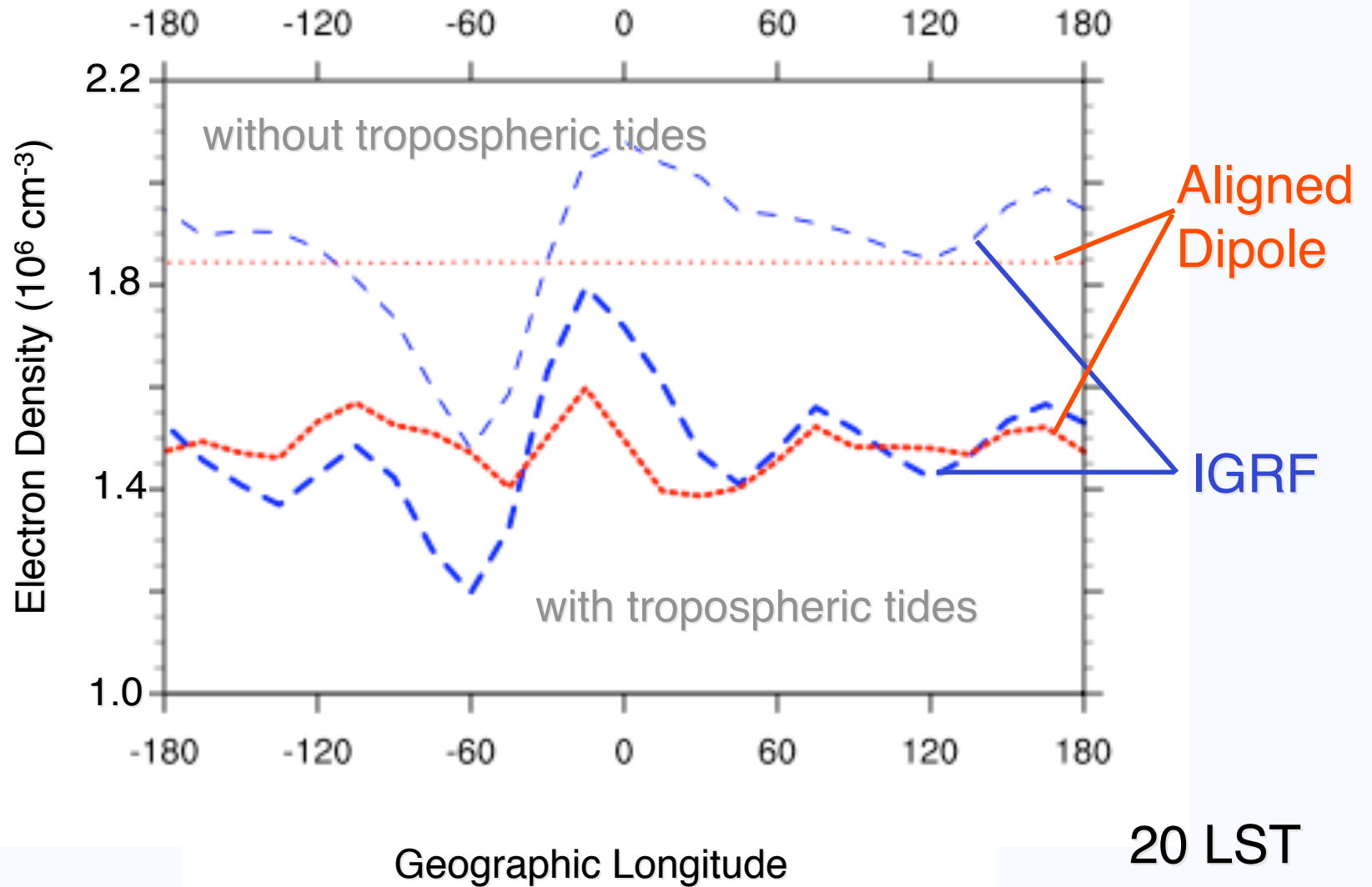
Maura Hagan 29 May 2007

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TIME-GCM Electron Density at 20 LST and 450 km

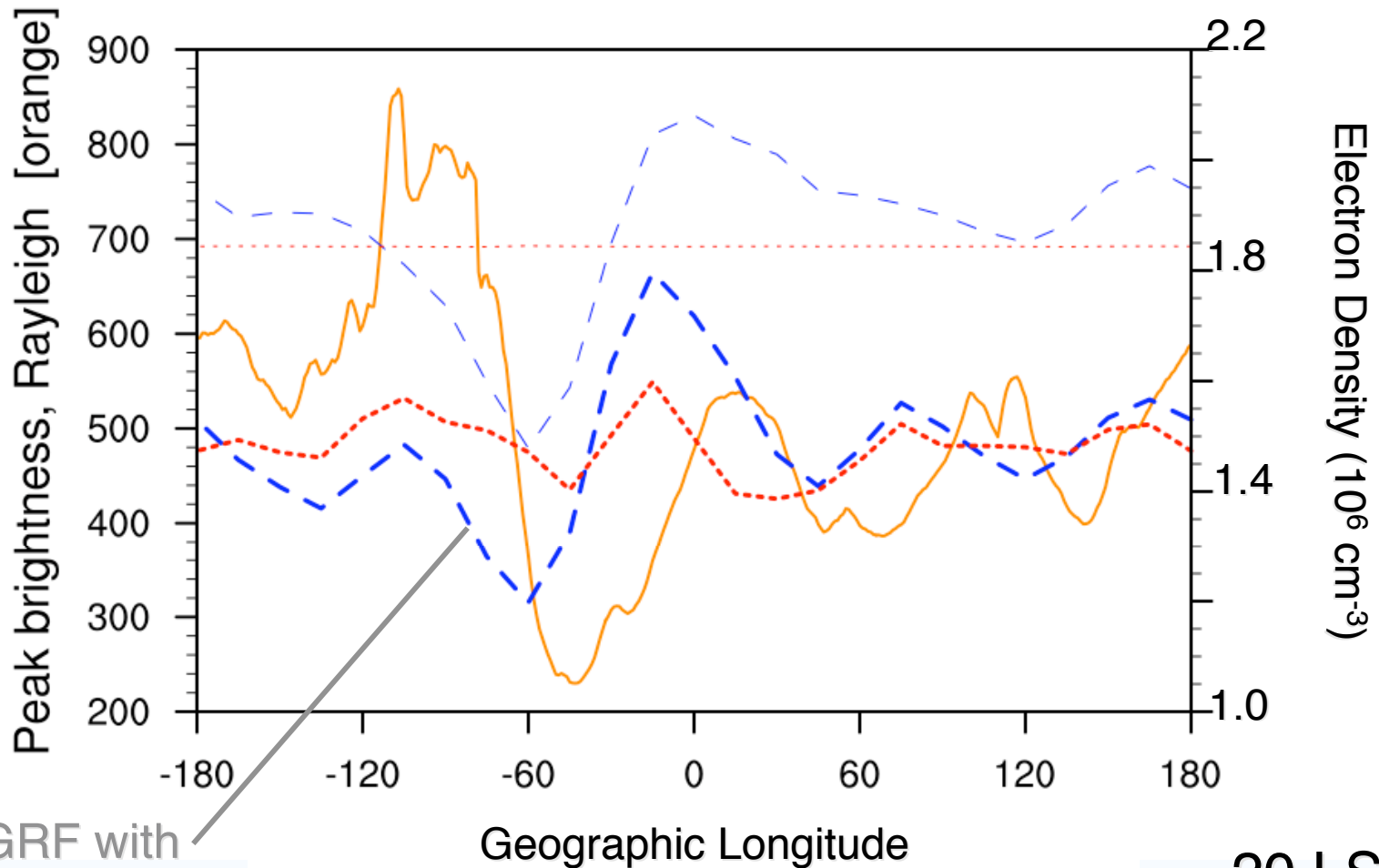


TIME-GCM Northern Hemisphere Peak Electron Density



TIME-GCM Northern Hemisphere Peak Electron Density

IMAGE FUV Northern Hemisphere Peak Brightness



IGRF with tropospheric tides

20 LST

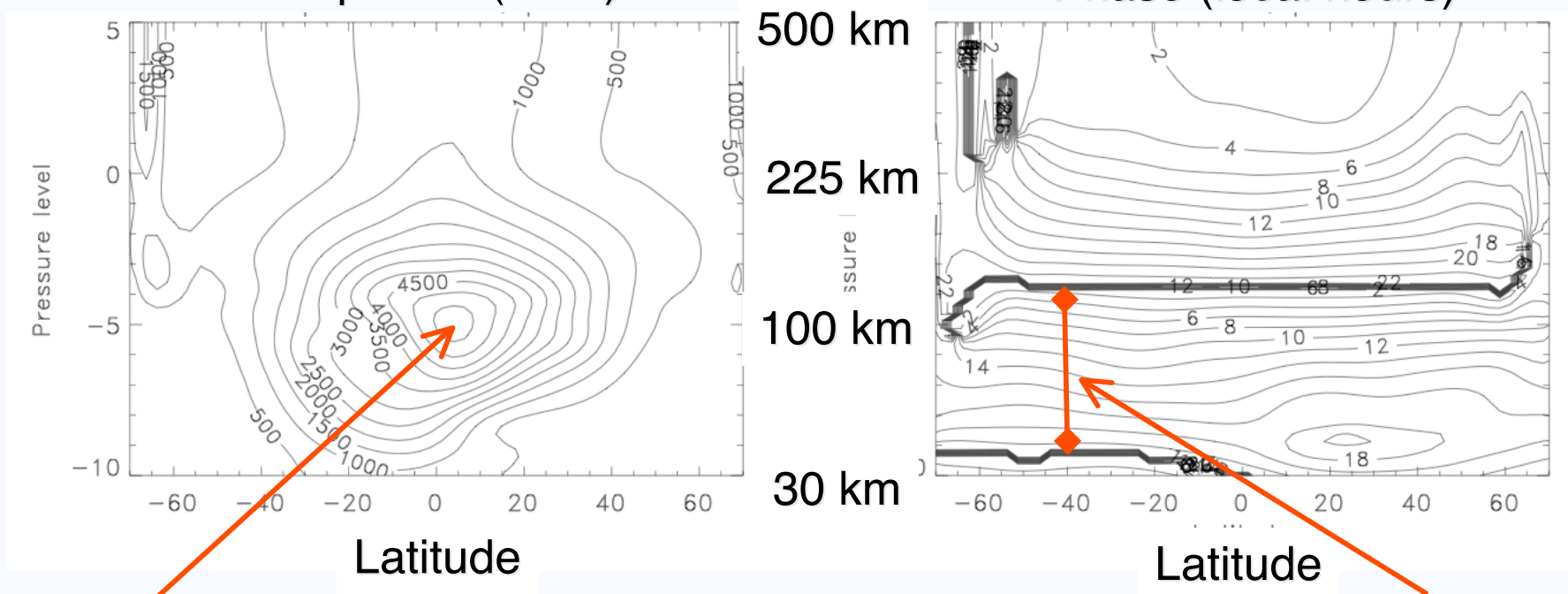


TIME-GCM Zonal Wind DE3 Perturbations

Standard IGRF Run

Amplitude (cm/s)

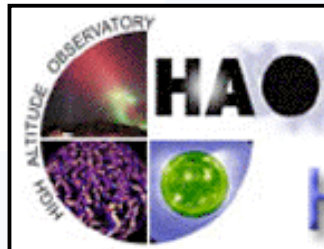
Phase (local hours)



55 m/s

Peak amplitude is **stronger** than
 GSWM prediction (Hagan and Forbes, 2002)
 UARS/HRDI analysis result (Talaat and Lieberman, 1999)
 TIMED/TIDI analysis result (Oberheide et al., 2006)
 λ_z is **longer** than
 GSWM prediction (Hagan and Forbes, 2002)

$\lambda_z \sim 75-85$ km



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Conclusions - *Hagan et al., 2007*

The DE3 is excited by parameterized latent heat release associated with raindrop formation in deep tropical convective clouds in the GSWM.

NCAR TIME-GCM simulations can replicate observed 4-peaked ionospheric longitudinal variations only when GSWM DE3 forcing is included at the lower boundary.

These results provide strong evidence of a connection between persistent global meteorological weather patterns and quiescent space weather.



NCAR

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DASI Science Challenges

▪ **Select Quiescent Variability Science Questions**

- ☛ what are the magnitudes & sources of **global** quiescent thermosphere-ionosphere (TI) variations?
- ☛ does the quiescent **global** variability “pre”condition the TI response to space weather events?

▪ **Diagnostic Wish List**

- ☛ global T_n , U_n , V_n , ρ , N_e , T_i , U_i , V_i , W_i , ΔH , ΔD , ΔZ , (?)
- ☛ leverage existing ground-based & space-borne resources
- ☛ correlative analyses to quantify wave sources and signatures

▪ **Modeling Component**

- ☛ develop/assess data assimilation (DA) models (e.g., GAIM); use DA to quantify optimum data requirements (quality-location-cadence) for fielding new instruments
- ☛ correlative/interpretive numerical modeling efforts

