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# Rossby Waves

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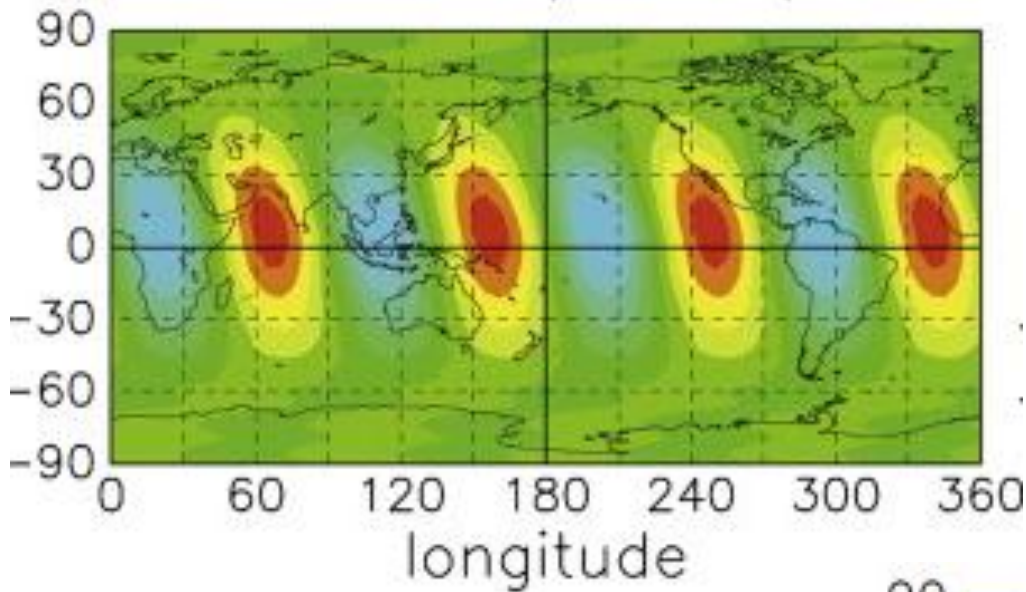
# What is a Rossby Wave?

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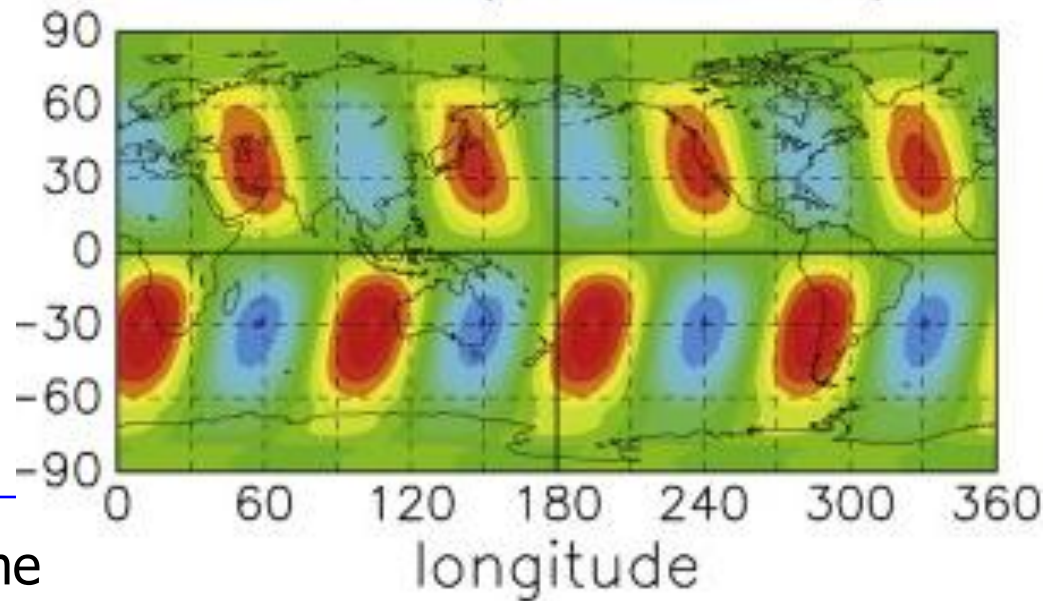
- Rossby waves are global scale atmospheric waves
  - Low zonal wavenumber  $s=1-4$
  - Large zonal wavelength  $\lambda_x = \frac{2\pi a}{s}$  km
    - $a=6370 \cdot \cos(\phi)$  km
    - $s=1, \phi=0$  implies  $\lambda_x = 40023$  km
  - Can be stationary or propagate westward (Don't propagate eastward)
  - Have periods longer than 1 day
    - 2, 5, 10 and 16 days
  - Have low phase speeds  $C_{ph} = \frac{484 \cos(\phi)}{Ts}$  m/s
    - 10 day wave at  $60^\circ\text{N}$  has phase speed of 24 m/s
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# Zonal Wavenumber 4



Equatorially symmetric

Equatorially asymmetric



10day wave movie

These patterns are a snapshot in time



# What is a Rossby Wave?

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- Rossby waves are generally barotropic (no vertical motion)
- Rossby waves are dispersive (group and phase velocities are different)
- Rossby wave are transient
- Generally important in the mid and high latitudes
- Can propagate from the troposphere to the MLT
- Rossby waves owe their existence to the gradient of vorticity - this is the wave restoring force

$$f = 2\Omega \sin(\phi) \quad \text{Coriolis parameter}$$

$$\beta = \frac{\partial f}{\partial y} \quad \text{Latitudinal gradient of } f$$

$$\beta = \frac{2\Omega \cos(\phi)}{a} \quad \beta > 0 \text{ for Rossby wave to exist}$$

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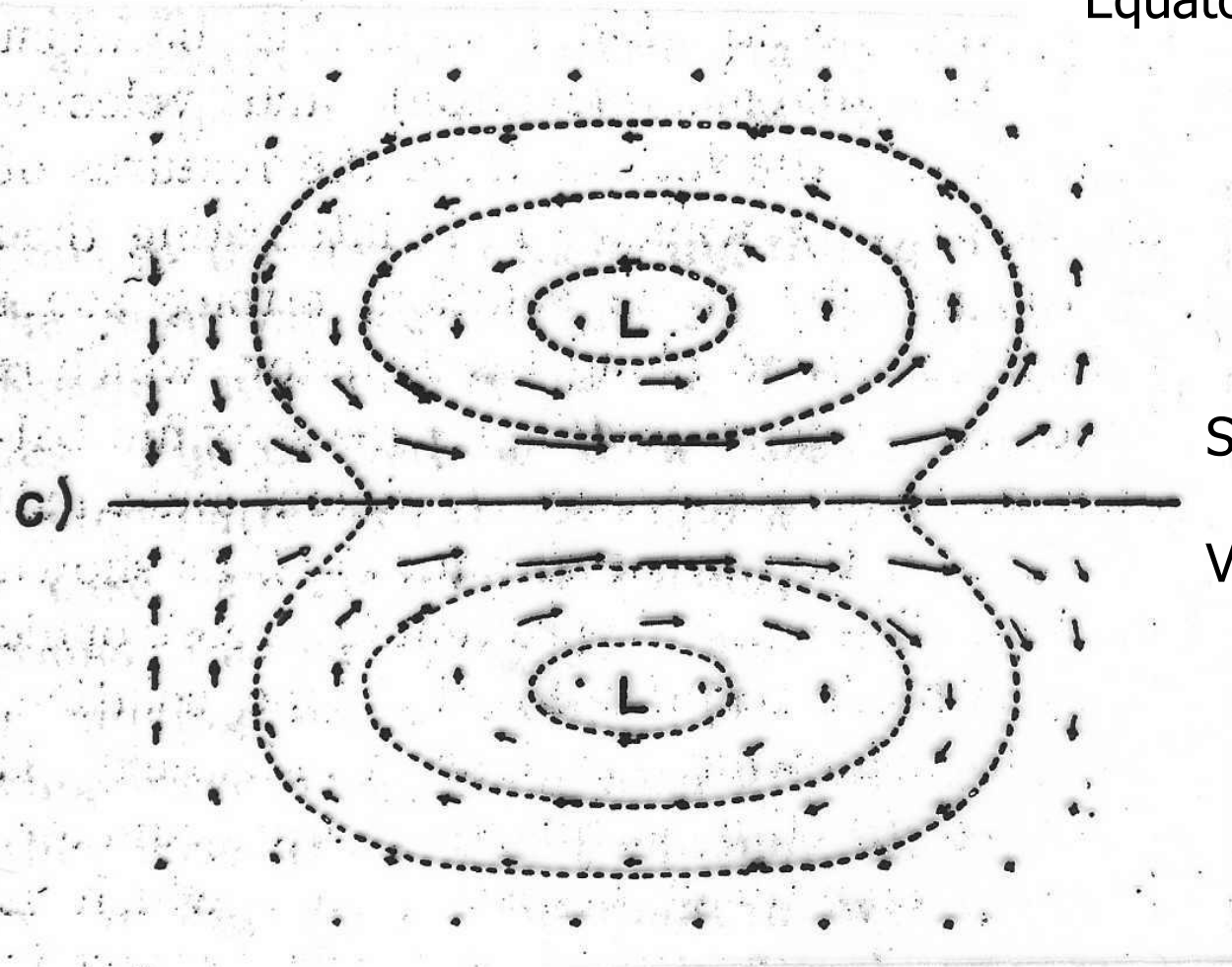
# Symmetric Rossby Wave



ROSSBY WAVES

Equatorially symmetric

Northward ↑



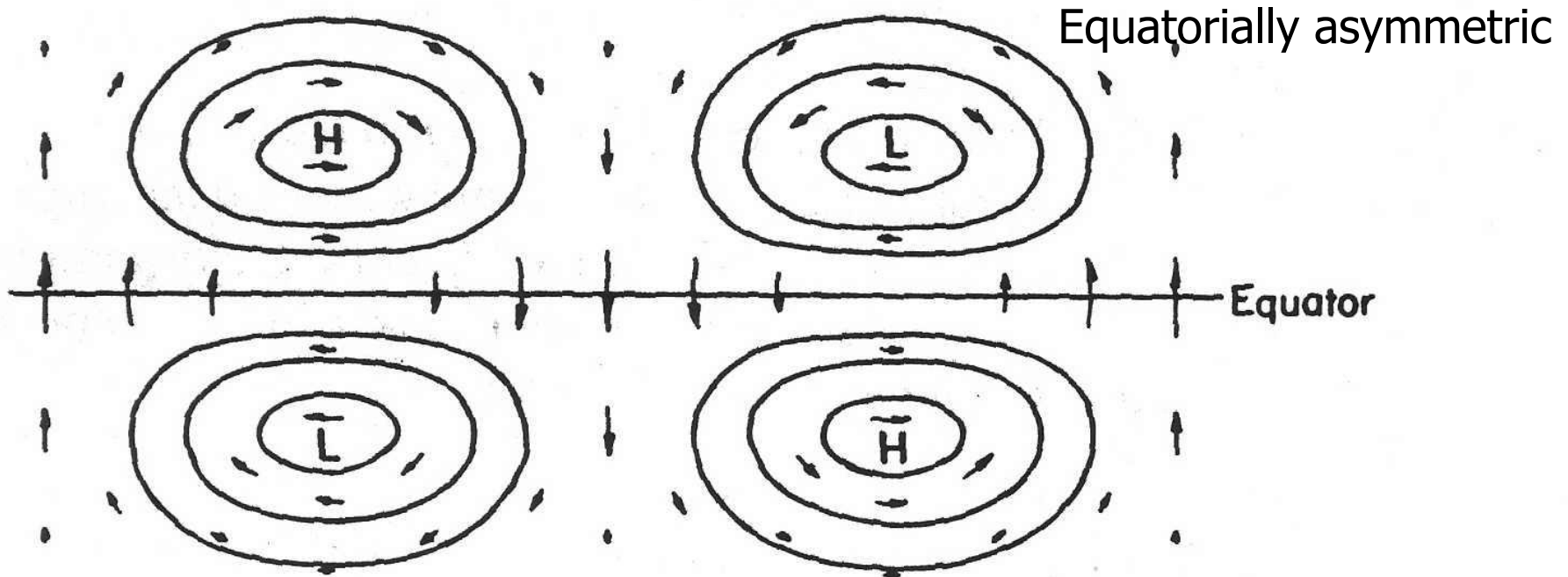
Scalar: Temperature

Vector: Winds

Eastward →

# Mixed Rossby Gravity Mode

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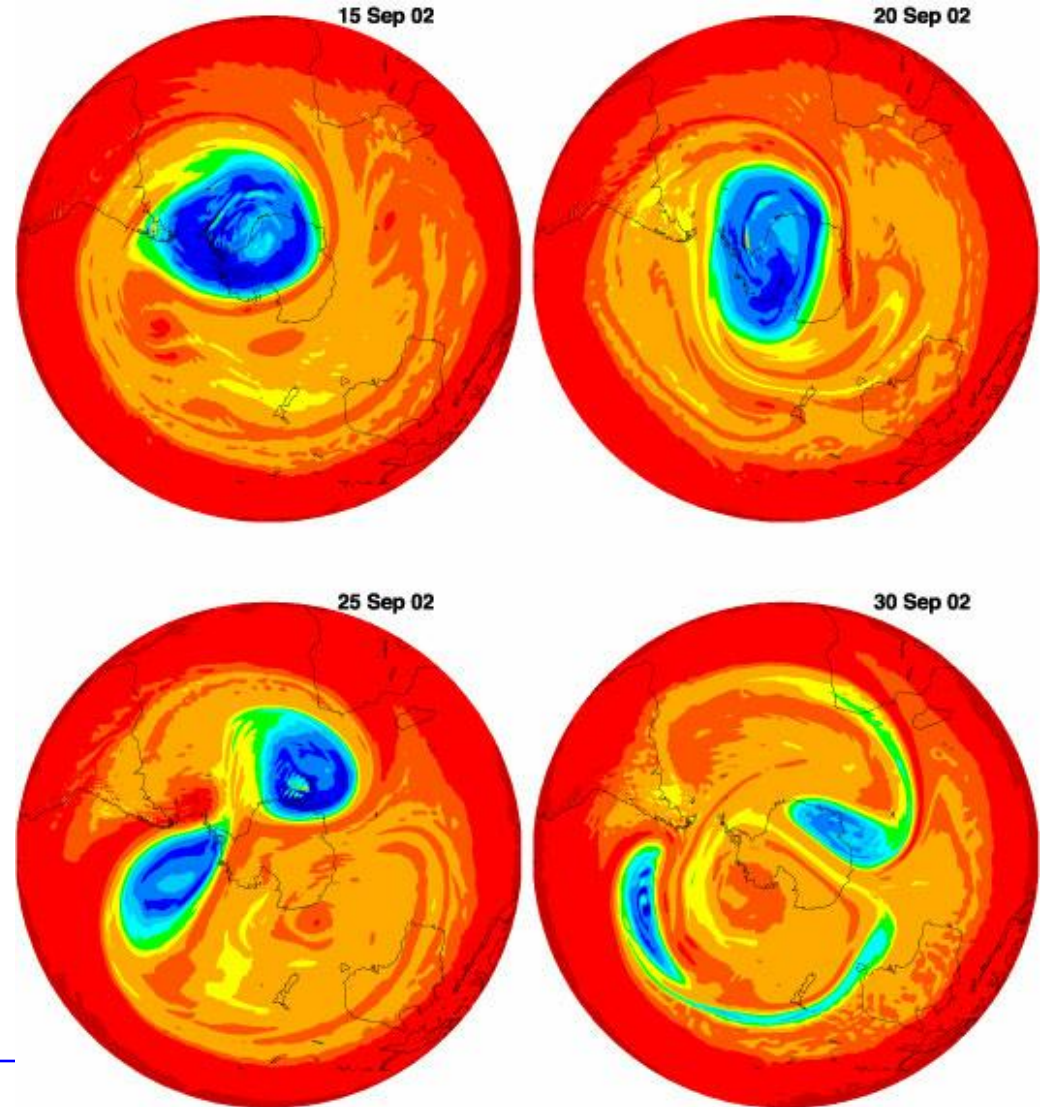
Plan view of horizontal velocity and height perturbations associated with an equatorial Rossby-gravity wave. (Adapted from Matsuno, 1966.)

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# Stationary Planetary Waves

- Wave 1 & 2 stationary planetary waves (SPW1, SPW2)
- Play a key role in sudden stratospheric warmings (SSWs)



SPW Antarctic movie

SPW global movie

Baldwin et al.

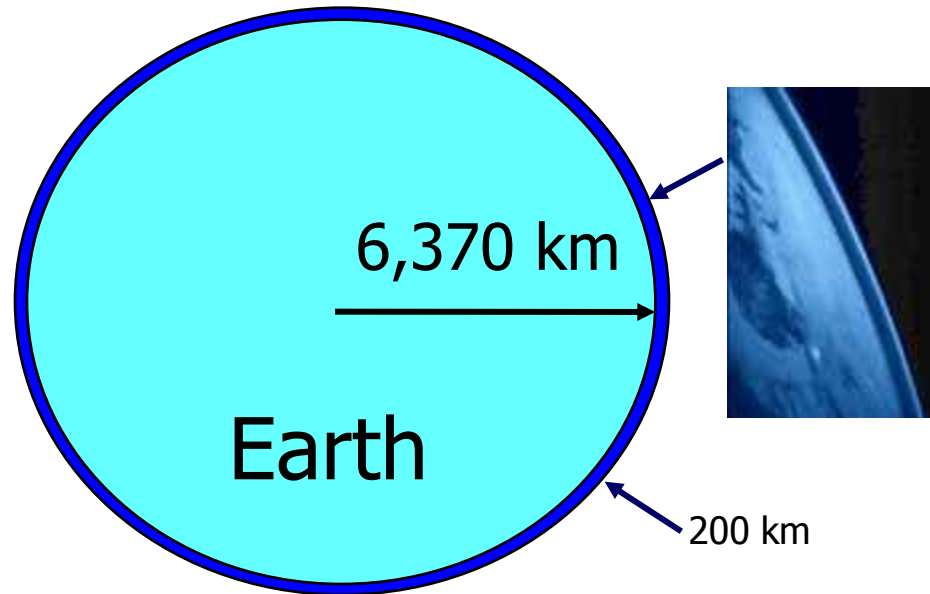
[http://www.atmosp.physics.utoronto.ca/SPARC/News20/20\\_Baldwin.html](http://www.atmosp.physics.utoronto.ca/SPARC/News20/20_Baldwin.html)

# What is the theory?

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Navier-Stokes equation on  
a thin fluid surrounding  
a rotating sphere.





# The Primitive Equations on geometric height coordinates

- Horizontal Momentum  $\frac{D\mathbf{U}}{Dt} = -\frac{1}{\rho} \nabla p - f\hat{\mathbf{k}} \times \mathbf{U} + \mathbf{F}_{visc}$
- Vertical Momentum (hydrostatic balance)  $\frac{dp}{dz} = -\rho g$
- Thermodynamic (conservation of energy)  $Q = c_p \frac{DT}{Dt} + \frac{1}{\rho} \frac{Dp}{Dt}$
- Continuity (conservation of mass)  $\frac{D\rho}{Dt} + \rho \nabla \cdot \mathbf{U} = 0$
- Ideal gas law for a dry atmosphere  $p = \rho RT$

$$\frac{D}{Dt} = \frac{\partial}{\partial t} + u \frac{\partial}{\partial x} + v \frac{\partial}{\partial y} + w \frac{\partial}{\partial z} \quad u \equiv \frac{Dx}{Dt} \quad v \equiv \frac{Dy}{Dt} \quad w \equiv \frac{Dz}{Dt}$$

# How are planetary waves defined?



Primitive Equations

$$f(t, \lambda, \phi, z)$$

Zonal Mean Equations

$$\bar{f}(t, \phi, z) = \int_{\lambda} f(t, \xi, \phi, z) d\xi$$

Nonlinear Perturbation Equations

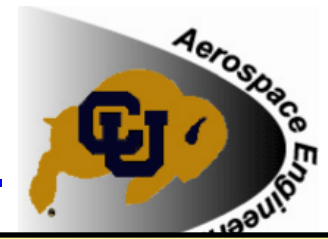
$$f'(t, \lambda, \phi, z) = f(t, \lambda, \phi, z) - \bar{f}(t, \phi, z)$$

Solutions are Periodic in Time ( $\sigma$ -frequency) and Longitude (s-wavenumber)

$$f'(t, \lambda, \phi, z) = \sum_{\sigma} \sum_{s} A^{\sigma, s}(\phi, z) \cos(\Omega \sigma t - s \lambda + \Psi^{\sigma, s}(\phi, z))$$

$\Omega = 2\pi/24$  radians/hour,  $\sigma$  = normalized freq (cycles/day),  $T = 24/\sigma$  period (hours)

# Analytical Solutions



Assume:

No zonal mean winds  $\bar{U} = 0$

Isothermal atmosphere  $T = T_0$

No dissipation

Solve Laplace's Tidal Equation

$$f'(t, \lambda, \phi, z) = \sum_{\sigma} \sum_s \sum_n F_n^{\sigma,s}(\phi) G_n^{\sigma,s}(z) \cos(\Omega\sigma t - s\lambda + (\lambda_z)_n^{\sigma,s} z + \Psi_n^{\sigma,s}(\phi))$$

$F_n^{\sigma,s}(\phi) = \Phi_n^{\sigma,s}(\phi)$  – related to Hough Functions

$$\Phi_n^{\sigma,s}(\phi) = \Theta_n^{\sigma,s}(\phi)$$

$$F_n^{\sigma,s}(\phi) = V_n^{\sigma,s}(\phi) = \frac{-j}{2a\Omega(\sigma^2 - \sin^2 \phi)} \left( s \tan \theta + \sigma \frac{\partial}{\partial \phi} \right) \Phi_n^{\sigma,s}(\phi)$$

$G_n^{\sigma,s}(\phi)$  – solution to the vertical structure equation

# Solutions for a Linearized, Isothermal Atmosphere



$$f'(t, \lambda, \phi, z) = \sum_{\sigma} \sum_s \sum_n F_n^{\sigma,s}(\phi) G_n^{\sigma,s}(z) \cos(\Omega\sigma t - s\lambda + (\lambda_z)_n^{\sigma,s} z + \Psi_n^{\sigma,s}(\phi))$$

Solutions are periodic in

Time (t) –  $\Omega\sigma$  is frequency

Longitude ( $\lambda$ ) – s is wavenumber

Altitude (z) –  $\lambda_z$  is vertical wavelength

Solar tides are harmonics of  $\Omega$  ( $\sigma \geq 1$ )

$\sigma=1$  Diurnal Tide (T=24 hours)

$\sigma=2$  Semidiurnal Tide (T=12 hours)

$\sigma=3$  Terdiurnal Tide (T=8 hours)

Rossby Waves are subharmonics of  $\Omega$  ( $\sigma \geq 1$ )

$\sigma=0.5$  2 day wave

$\sigma=0.2$  5 day wave

$\sigma=.01$  10 day wave

# Latitudinal Structure



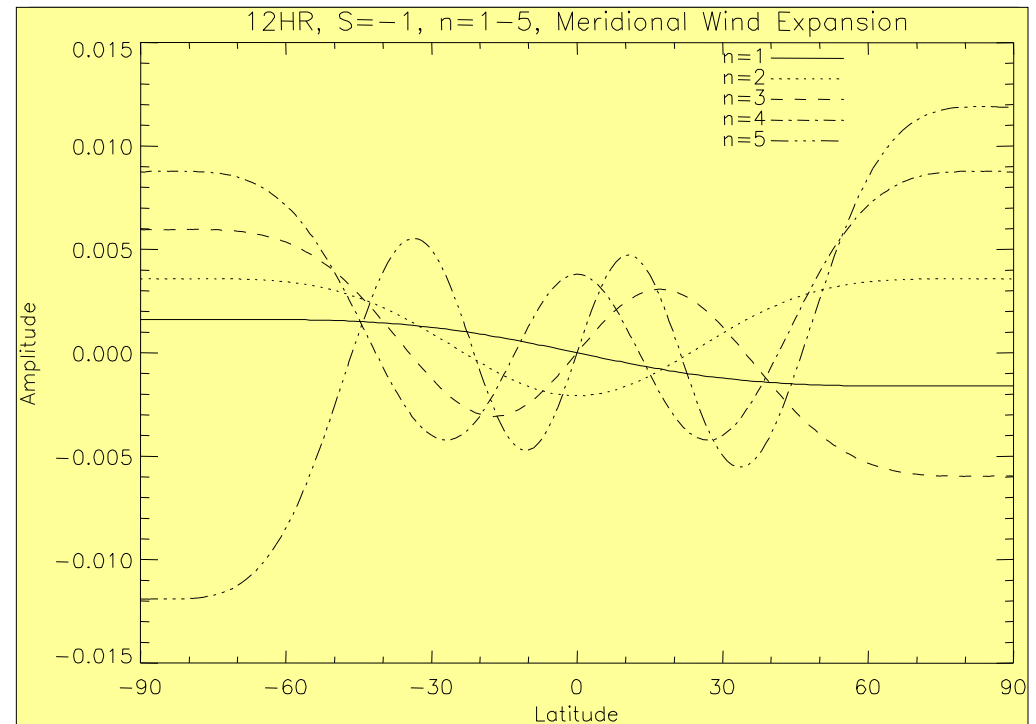
$$f'(t, \lambda, \phi, z) = \sum_{\sigma} \sum_s \sum_n F_n^{\sigma, s}(\phi) G_n^{\sigma, s}(z) \cos(\Omega \sigma t - s \lambda + (\lambda_z)_n^{\sigma, s} z + \Psi_n^{\sigma, s}(\phi))$$

Latitudinal amplitude structure

Depends upon:

- Frequency ( $\sigma$ )
- Wavenumber ( $s$ )
- Meridional index ( $n$ )

Non Migrating Semidiurnal Tide ( $\sigma=2, s=-1, n=1-5$ )



# Dispersion relation for $s=1$

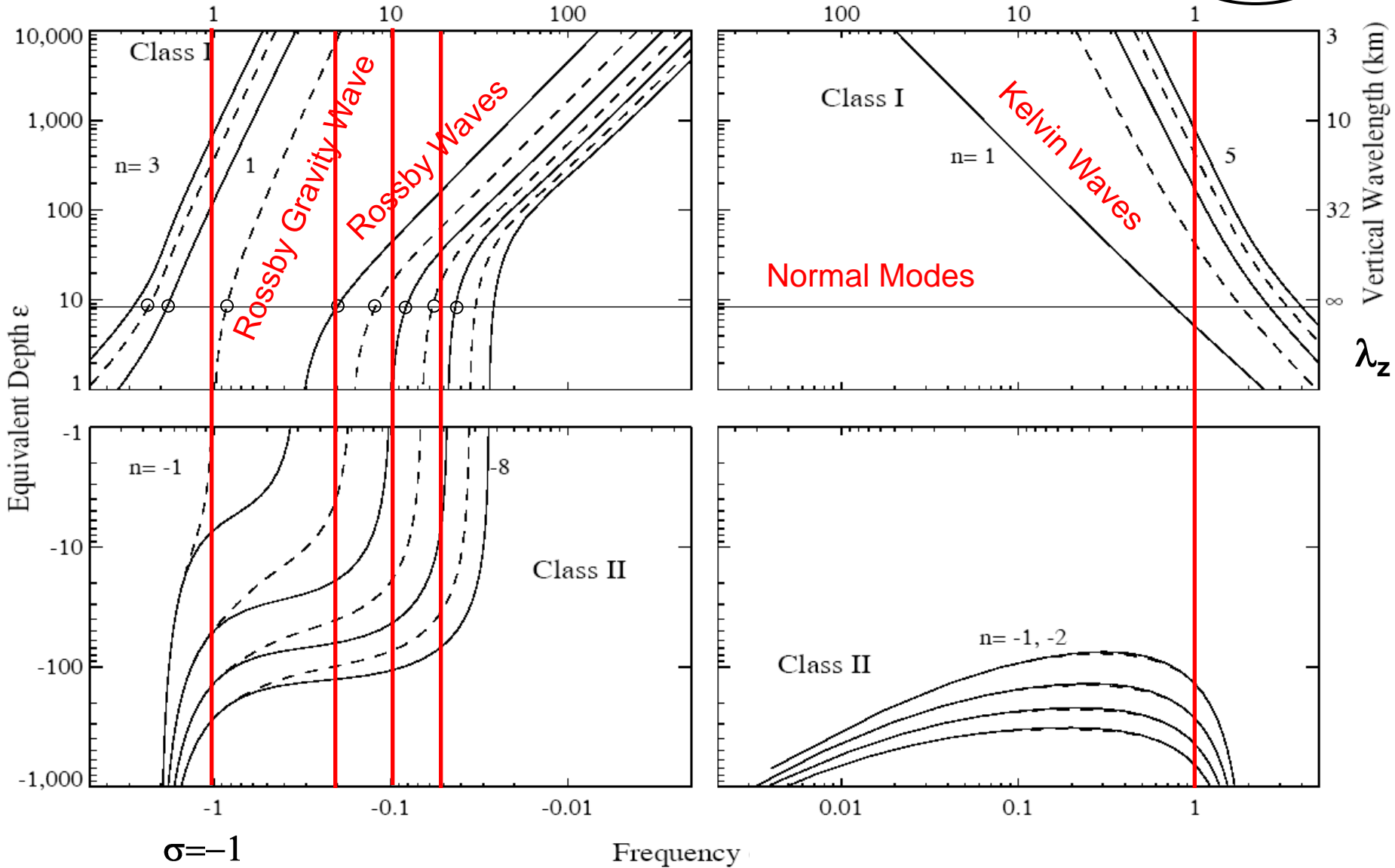


Westward propagating

Period (days)

Eastward propagating

$S = 1$

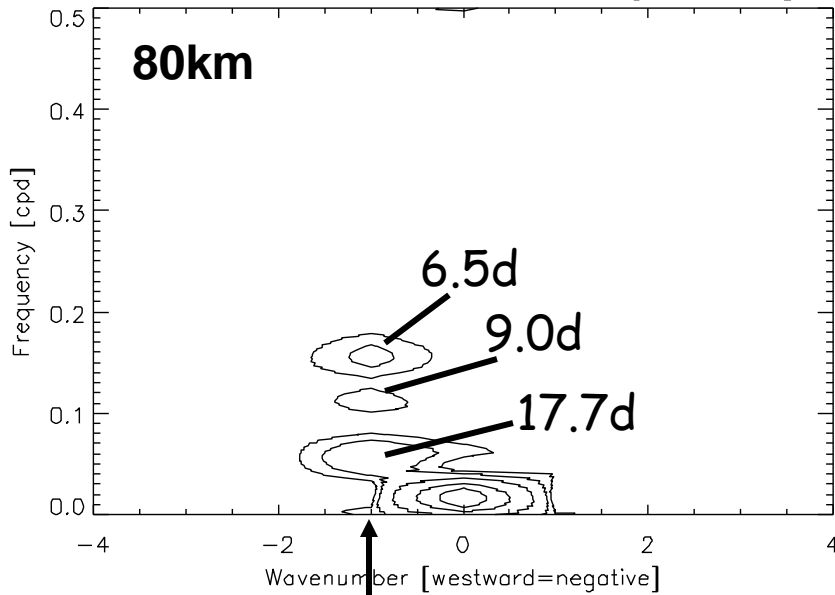




# SABER Power Spectra, 32N-37N



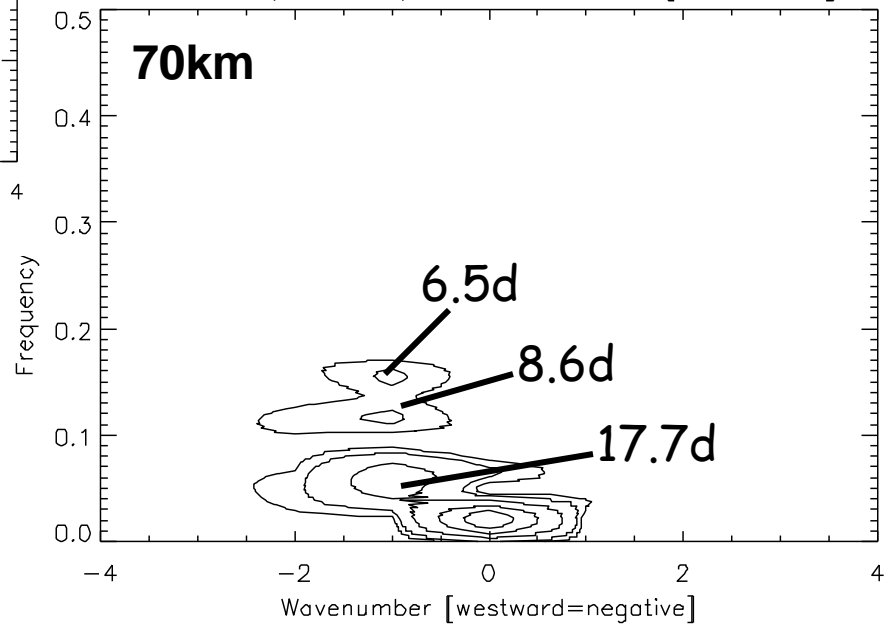
SABER Spectra April 2002, 80km [32N-37N]



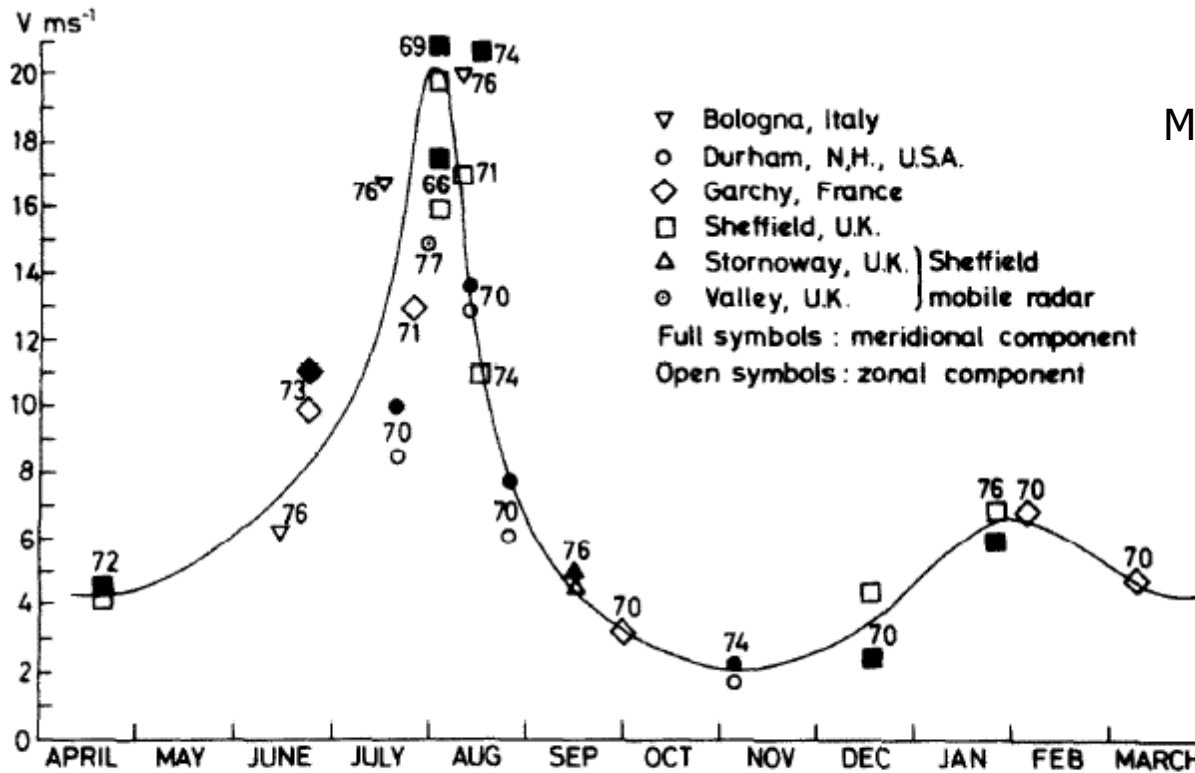
Zonal wavenumber 1  
Westward propagating

April 2002

SABER Spectra April 2002, 70km [32N-37N]



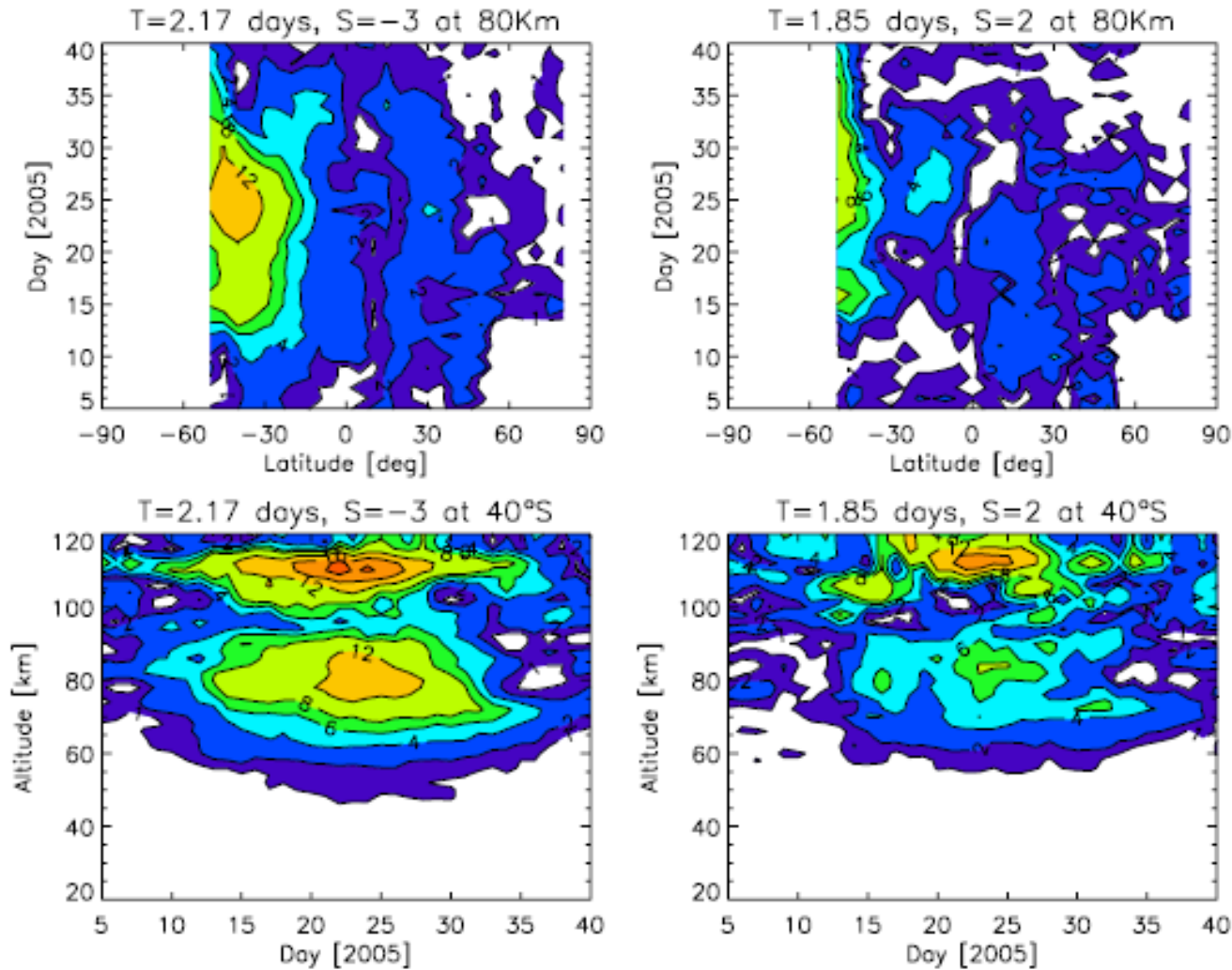
# 2-day wave



Mueller and Nelson, 1978

Figure 2. The amplitude of the 51 h meteor wind oscillation at various times of the year based on mid latitude records for the northern hemisphere during the period 1966 - 1977. Values relate to an average altitude of 95 km.

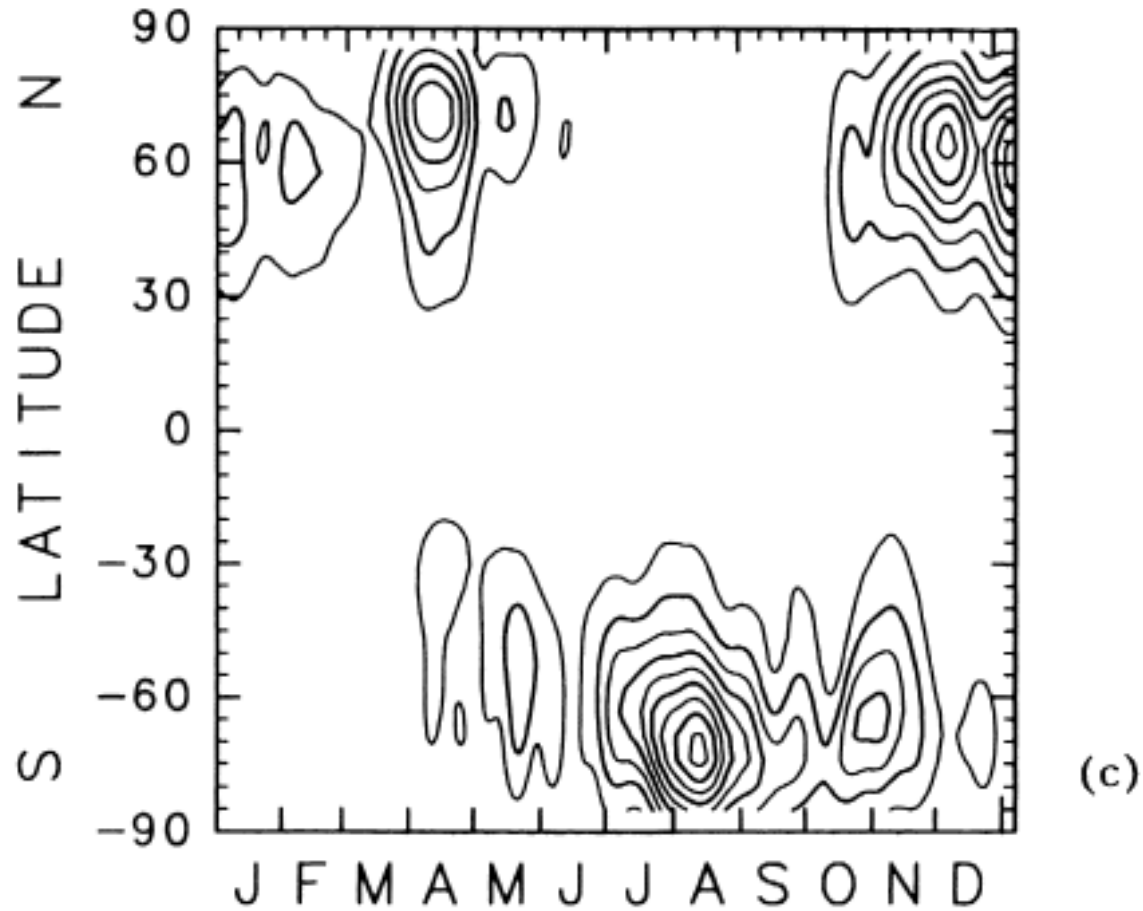
# 2 Day Wave in SABER



2day wave movie

Palo et al, 2007

# 16-day wave



# Rossby Normal Modes

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Wave	(s,n)	(s,  n  -s)	$h_n$ (km)	Cph@eq m/s	
5-day	(1,-2)	(1,1)	10.5	97	First symmetric
10-day	(1,-3)	(1,2)	10.5	48	First asymmetric
16-day	(1,-4)	(1,3)	10.5	30	Second symmetric
4-day	(2,-3)	(2,1)	10.5	60	First symmetric
2-day	(3,-3)	(3,0)	10.5	81	Mixed Rossby-Gravity, asymmetric

Table adapted from Forbes, 1995

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# Summary

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- Rossby waves are global scale disturbances ( $s=1-4$ )
- $\beta$  is the restoring force for Rossby waves
- Stationary planetary waves (SPW1, SPW2) are Rossby waves and are important in northern hemisphere winter dynamics and play a major role in sudden stratospheric warmings.
- Many of the large scale disturbances observed in the MLT are Rossby Waves. These waves propagate westward and are related to atmospheric normal modes
  - Quasi 2-day wave
  - Quasi 4-day wave
  - Quasi 10-day wave
  - Quasi 16-day wave



# Rererences: Dissertations

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Note that dissertations often provide seminal references on a subject and tutorial information

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