



Equatorial Ionization Anomaly Crest Merging During 19-20 January 2026 Geomagnetic Storm

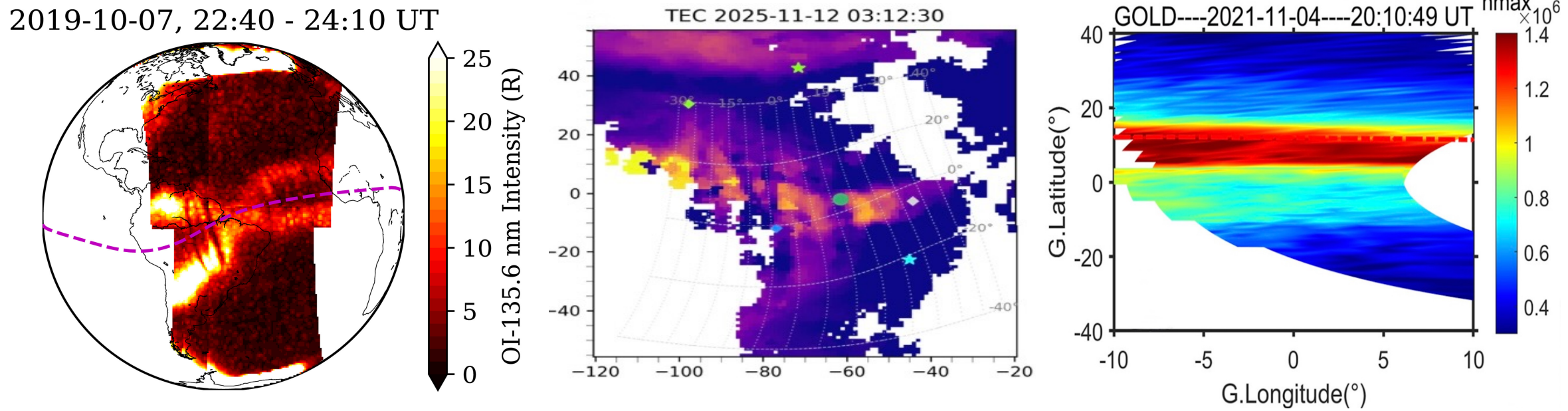
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EQIT-19
Wednesday

1. Background

Equatorial Ionization anomaly (EIA) primarily forms due to eastward electric field. The merged EIA crests have been observed during both geomagnetically quite and active times.



During quiet time. Downward vertical drift + Lower atmospheric forcing. Lasker, et al. (2024)

During MP; Min Dst -231nT at ~04 UT. large equatorward winds + westward PPEF. Zou et al., (2026)

During RP; Min Dst ~ -110 nT at 13:00 UT DDEF + meridional neutral winds. Wu, et al. (2025)

Objective: To investigate underlying physical mechanism driving EIA crest merging during the main phase of 19-20 January 2026 geomagnetic superstorm.

2. Data and Model

Data:

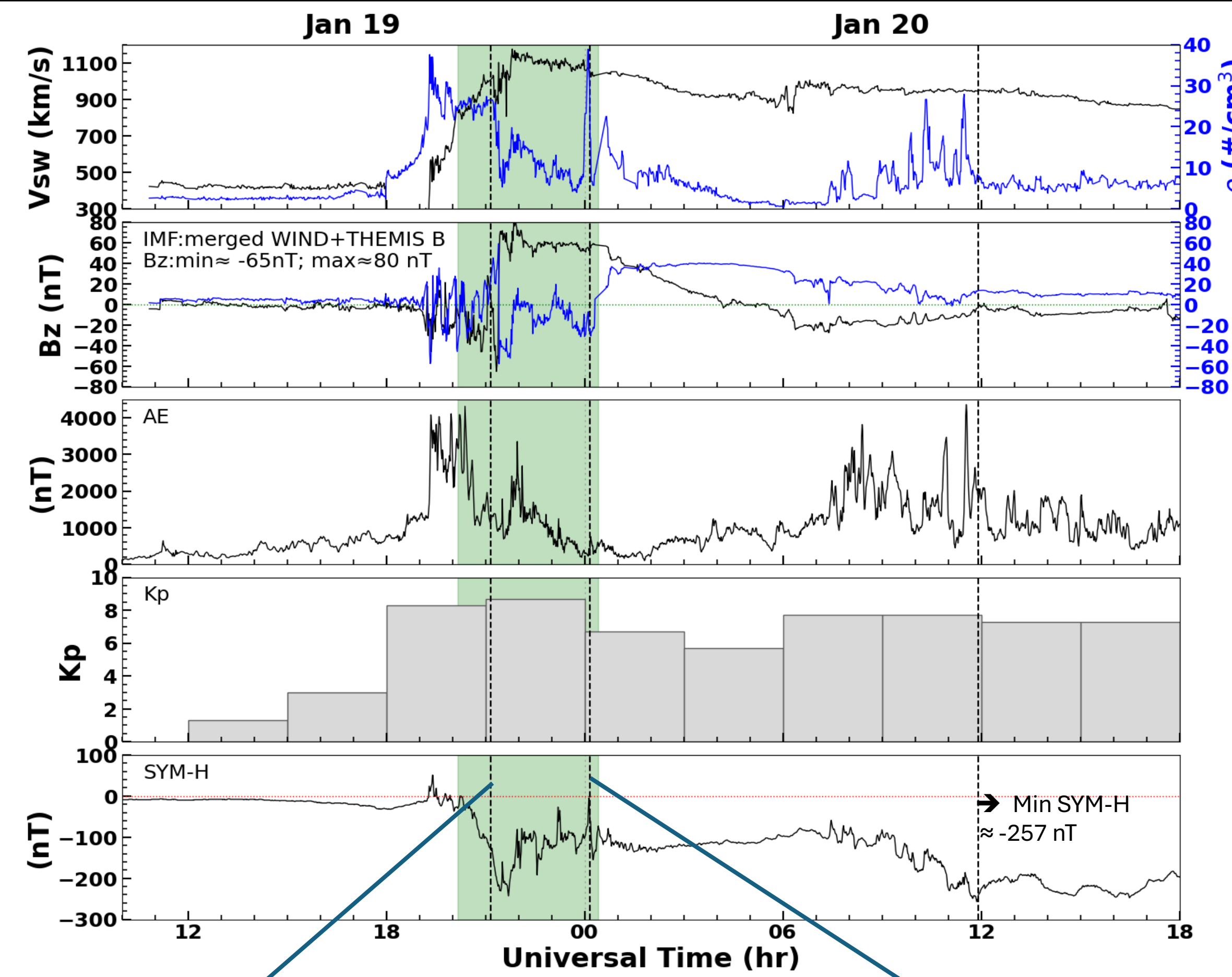
- GOLD level 2 NMAX + Madrigal TEC
- neutral winds at Millstone Hill (42° N, 72° W)

Model setup:

Global Ionosphere Thermosphere Model (GITM) with self-consistent 3D dynamo

- High Latitude forcing and auroral precipitation: SWMF/ Ridely Ionosphere Model based AMIE files

3. Geomagnetic Conditions

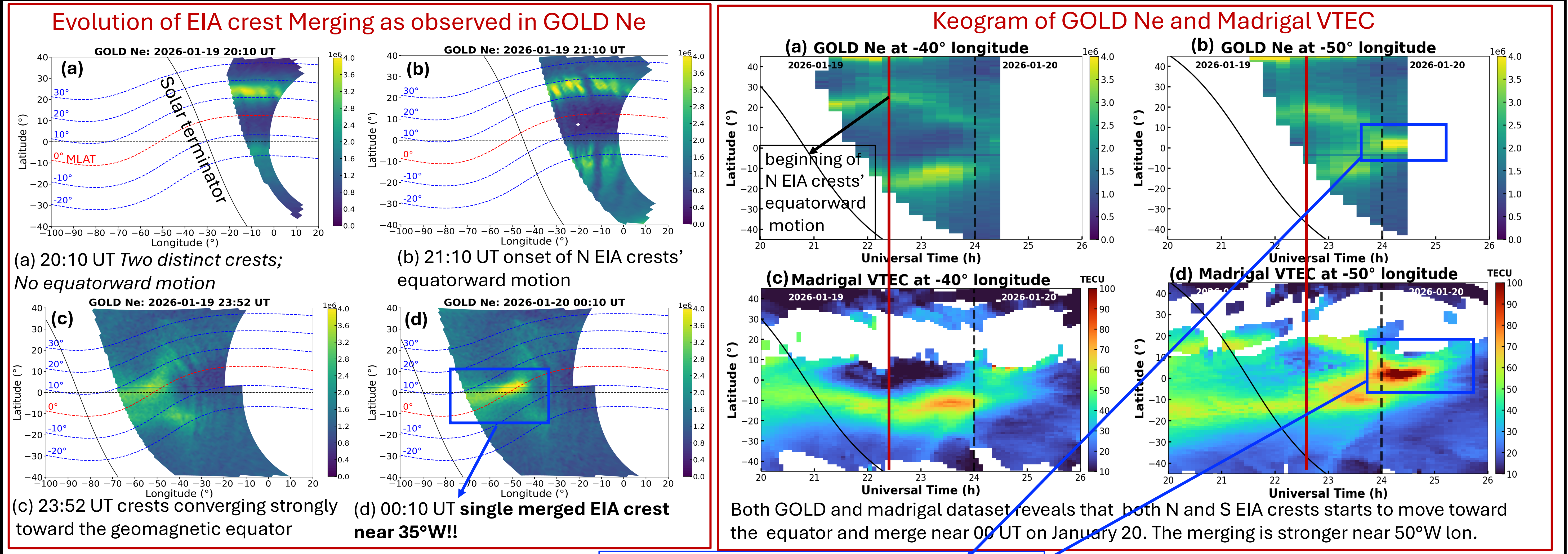


21: 10 UT onset of Equatorward motions of crests

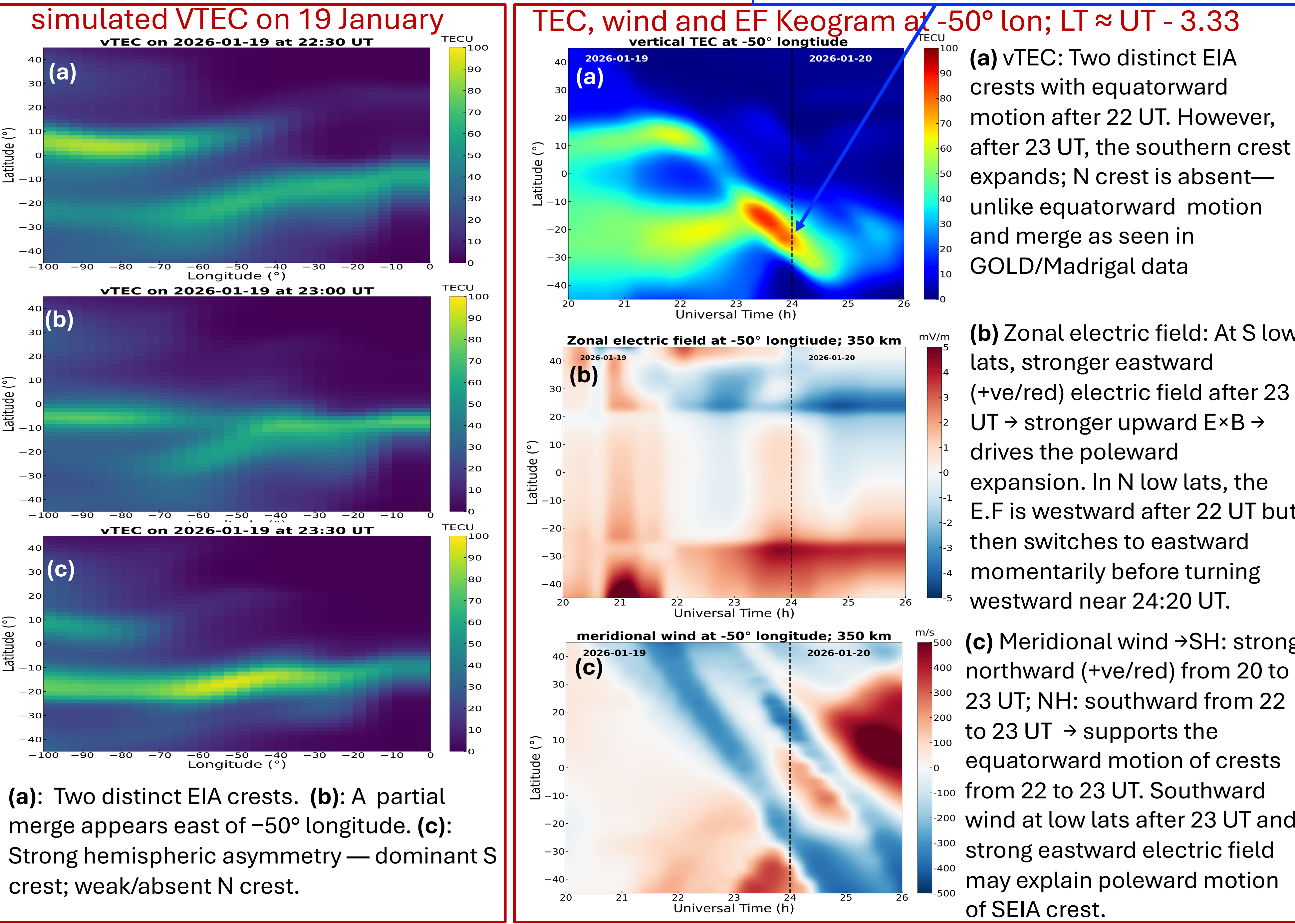
00:10 UT single EIA crest near 35° W longitude

Green shaded region: GOLD observation period on 19 January.

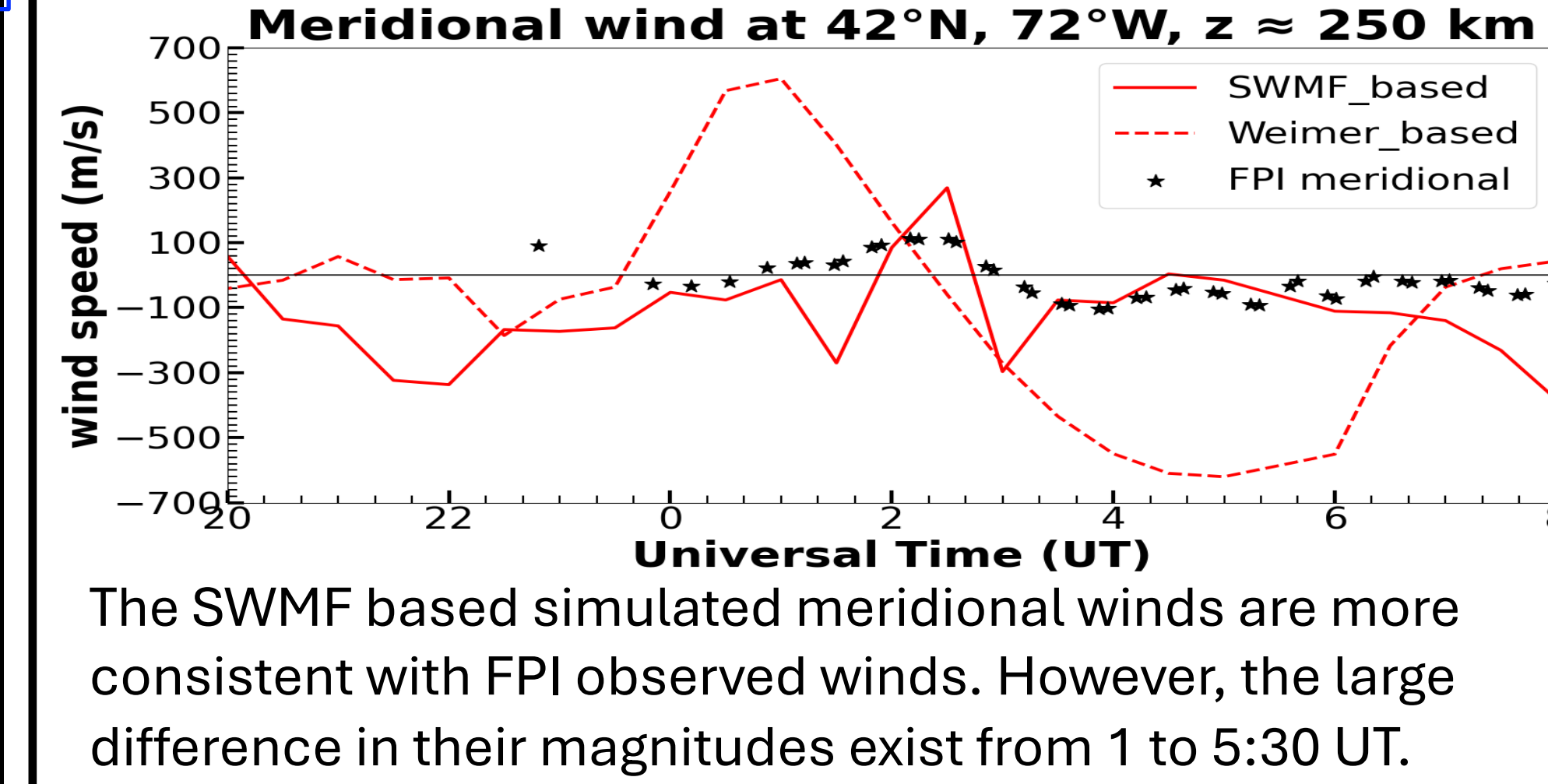
4. Results: Observations



5. Results: GITM simulations (Preliminary)



6. Data-model comparison



7. Main Take ways and Future works

- Within 3 hours of equatorward movement, both N and S EIA crests merged into single EIA crest near 35°W longitude. GOLD Ne and Madrigal TEC are consistent on timing.
 - At -50°, unlike observation, the GITM simulated SH TEC near crest expands poleward instead of merging.
- Future work:**
- Match model's neutral winds with FPI data.
 - Investigate why the simulated electric field is +ve in SH.
 - Separate the E-field (PPEF/DDEF) vs wind roles using the continuity equation (E×B vs wind vs chemistry).

8. Acknowledgements and References

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Lasker et al. (2024). 10.1029/2023JA032224

Wu et al., (2025). 10.1029/2024JA032896

Zou et al., (2026). 10.1029/2025GL121182