

Motivation

- Results from CHAMP accelerometry revealed mesoscale (few thousand kilometers) mass density structures that coincided with signatures of wind influence at the high-latitude thermosphere [1]
- Co-located density and wind structures can indicate a potential relationship between mass density and winds that produce perturbations detectable by highly sensitive accelerometers
- High-latitude features (e.g., mass density cusp enhancement) are difficult to model solely considering direct energy inputs into the thermosphere [2, 3]
- This poster focuses on how mechanical mechanisms (e.g., ion drag) can distribute energy and produce mesoscale perturbations in the high-latitude mass density and wind structures

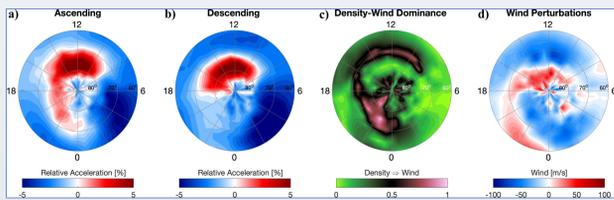


Figure 1: Results from Buynovskiy et al. (2025) demonstrating relative accelerations in the ascending and descending satellite trajectories with an analysis of density-wind dominance and detected wind perturbations

Results

Demonstrating First Law of Thermodynamics Compliance

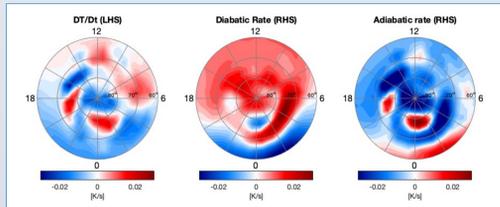


Figure 2: Validation of the first law of thermodynamics using TIEGCM energy equation terms (Eq. 4a)

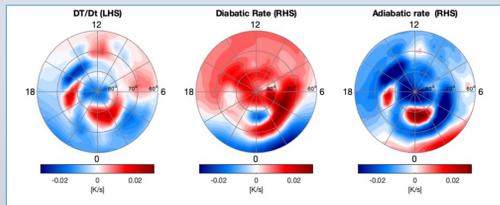


Figure 3: Validation of the first law of thermodynamics using observable-based diabatic and adiabatic heating terms (Eq. 4b)

- Figure 2 demonstrates that **TIEGCM satisfies the first law of thermodynamics**
 - Left-hand side (total temperature rate) equals the right-hand side (sum of the diabatic and adiabatic terms)
- Figure 3 shows that the **diabatic and adiabatic heating rates can be directly computed from observables**, reproducing the balance in Fig. 2
- Note: The first law of thermodynamics computes the total (material) derivatives, composed of the local time rate of change and advective terms
 - Including advective terms can make a qualitative analysis with observables difficult to interpret
 - In this analysis, the advection terms approximately cancel, which allows a reduced form of the first law as shown in Eq. 5

$$\frac{\partial T}{\partial t} \approx \left(\frac{P}{P_0}\right) \frac{R}{c_p} \frac{\partial \theta}{\partial t} + \left(\frac{RT}{c_p P}\right) \frac{R}{c_p} \frac{\partial p}{\partial t} \dots \text{(Eq. 5)}$$

Where T is the neutral temperature, P is the in-situ pressure, P₀ is a reference pressure (1e-5 Pa for this case), θ is the potential temperature, R is the specific gas constant, and c_p is the specific heat at constant pressure

Tracking Diabatic vs Adiabatic Influence

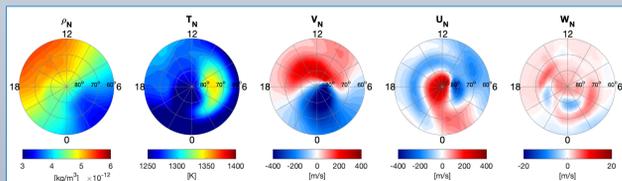


Figure 4: Thermospheric state for run 1 (baseline) at UT = 01:00

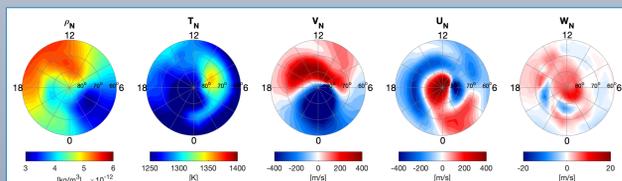


Figure 7: Thermospheric state for run 2 (enhanced ion drag) at UT = 01:00

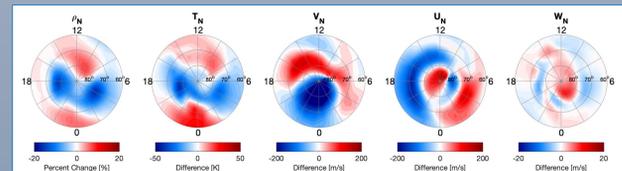


Figure 10: Thermosphere state comparison between run 1 and run 2

Theoretical Framework

- Thermosphere-Ionosphere-Electrodynamics General Circulation Model (TIEGCM)
 - Physical model that self-consistently solves the 3D momentum (Eq. 1), energy (Eq. 2), and continuity equation of the entangled ionosphere-thermosphere system and coupled with the magnetosphere [4]
 - Outputs thermospheric state properties and individual heat terms, enabling a detailed energy analysis
- Linking the TIEGCM energy equation to thermodynamic principles
 - TIEGCM's energy equation is directly mapped to the first law of thermodynamics shown (Eq. 4a)
 - At 400 km, negligible eddy diffusion simplifies the energy balance to **adiabatic** (mechanical work) and **diabatic** (external heating/cooling) processes
 - Adiabatic** and **diabatic** heat terms can also be expressed using temperature, potential temperature, and vertical winds (Eq. 4b) [5]
- Imposing mechanical forcing in TIEGCM
 - To investigate how mechanical forcing can modify the mass density and wind structures, two model scenarios are developed at analyzed constant altitude:
 - Run 1: Baseline
 - K_p = 2, F_{10.7} = 150 solar flux units, solar maximum, June solstice
 - Run 2: Increased mechanical forcing
 - Identical solar inputs as Run 1
 - The momentum equation (Eq. 1) is modified by doubling the ion drag force at high-latitudes

$$\frac{\partial \mathbf{v}_n}{\partial t} = F_{\text{Pressure Gradient}} + F_{\text{Ion Drag}} + F_{\text{Viscous}} + F_{\text{Coriolis}} + F_{\text{Advection}} + F_{\text{Other}} \dots \text{(Eq. 1)}$$

$$\frac{\partial T_n}{\partial t} = H_{\text{Cond}} + H_{\text{Eddy Diff}} + H_{\text{Advec}} + H_{\text{Adiabatic}} + (H_{\text{Solar}} + H_{\text{Chem}} + H_{\text{Joule}}) + H_{\text{Cool}} \dots \text{(Eq. 2)}$$

$$\frac{\partial T_n}{\partial t} = \frac{g e^z}{p_0 c_p} \frac{\partial}{\partial z} \left(\frac{K_T}{H} \frac{\partial T_n}{\partial z} \right) - \mathbf{v}_n \cdot \nabla T_n - W \frac{\partial T_n}{\partial z} - W \frac{R^* T_n}{c_p \bar{m}} + \frac{Q^{\text{exp}} - e^z L^{\text{exp}}}{c_p} - L^{\text{imp}} T_n \dots \text{(Eq. 3)}$$

$$\left(\frac{DT}{Dt} = \frac{1}{c_p} \frac{Dq}{Dt} + \frac{RT}{c_p} \frac{1}{\rho} \frac{Dp}{Dt} \right) = \left(\frac{DT}{Dt} = \left(\frac{p}{p_0} \right)^{\frac{R}{c_p}} \frac{D\theta}{Dt} + \frac{g}{c_p} w_d \right) \text{(Eq. 4b)}$$

Purple = Total temperature rate
Red = Diabatic heating rate
Blue = Adiabatic heating rate

$$\theta \equiv T_n \left(\frac{p_0}{p} \right)^{\frac{R}{c_p}}$$

Discussion

Connecting the First Law of Thermodynamics to TIEGCM

- TIEGCM has been verified to satisfy the first law (see Fig. 2) at the high-latitude thermosphere
- Thermosphere energy exchange can be reduced to:
 - Internal energy** changes (total temperature rate)
 - Diabatic heating** (solar heating, joule heating, etc.)
 - Adiabatic processes** (expansion/compression work)
- Thermosphere state variables (temperature, potential temperature, vertical winds) enable the tracking of diabatic and adiabatic contributions via Eq. 4b
 - This framework provides a diagnostic tool to assess the drivers of energy distribution in the thermosphere without requiring the full energy equation (Eq. 3)**
- Divergent vertical winds captures the adiabatic temperature changes due to expansion/compression without external heating
- Potential temperature variations capture the net diabatic heating contributions

Thermosphere Response to Induced Mechanical Forcing

- Ion drag (a momentum term) was doubled in a TIEGCM simulation, producing the following changes after one hour:
 - Mass density changes up to 20%
 - Temperature changes up to 50 K
 - Horizontal wind changes up to 200 m/s
- The thermosphere response followed a self-consistent feedback:
 - Momentum forcing induced → winds respond → heating terms adjust → further dynamic response
- Adiabatic heating/cooling can match or exceed diabatic contributions**, with a strong effect seen with a momentum-induced case study (see Figure 9 vs Figure 10)
 - Direct heat sources are not needed to significantly modulate the thermosphere mass density, temperature, and winds

Implications Modeling and Observing Mesoscale Structures

- The formation of mesoscale structures (e.g., neutral cusp enhancement) must consider both direct energy sources (e.g., joule heating) in addition to momentum-driven sources responsible for net adiabatic effects**
 - Figure 10 demonstrates how direct and indirect energy mechanisms can produce similar heating/cooling
- In-situ satellite measurements can track diabatic and adiabatic energy transfer using accurate measurements of few observables: pressure, temperature, and wind
 - Can help explain the underlying mechanisms responsible for the observed acceleration perturbation responses from Fig. 1

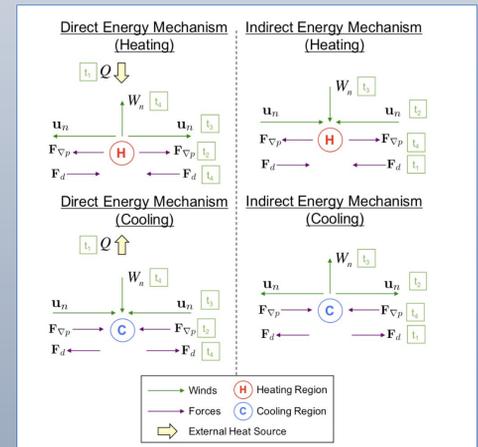


Figure 10: Direct vs indirect energy mechanisms [6]

References

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