Vz Terdiurnal Atmospheric Tides seen with ICON at E-region Altitudes Rythm Agarwal, Scott L. England, Guiping Liu

atmospheric tides and planetary waves are commonly observed at altitudes just below the ionosphere. Until recently, routine observations of the winds across a wide range of altitudes was scarce at ionospheric altitudes. Recent wind observations from ICON that span the Eand F-regions offer an opportunity to examine the global scale waves that are present, and ultimately determine their impacts on the ionosphere. This work utilizes these wind observations at E-region ionospheric altitudes to identify which tides are present. Specifically, this work will examine the terdiurnal tides that are present and compare these to the betterstudied diurnal and semidiurnal tides. Changes with season, altitude will be presented.



Atmospheric tides and planetary waves represent the dominant dynamical component of 80-120 km height; MLT (mesosphere-lower (thermosphere) Categories of tides: Diurnal: wavenumber, n = 1; phase: 0-360° Semidiurnal: wavenumber, n = 2; phase:

Terdiurnal: wavenumber, n = 3; phase:



Verifies the MIGHTI data collection process – nighttime has lesser tides because of decreased photoionization Fig. 5: Zonal wind on Local time vs This is the plot of zonal wind data which spans across 45 days and is plotted Latitude is limited to 0-5 degrees north When plotted against Local time and Longitude, an upward propagating winc Again, the daytime values are higher Verifies the MIGHTI data collection process – nighttime has lesser tides because of decreased photoionization. **Data Processing** A wave model was designed to fit the

Data Model

Fig. 4: Migrating terdiurnal tide on Local

which spans across 45 days and is

✤ Latitude is limited to 0-5 degrees nort

When plotted against Local time and

Longitude, the daytime amplitudes are

denser as compared to the nighttime.

This is the plot of zonal wind data

time vs Longitude; s = 3

plotted over 24 hours.

✤ Altitude is 91.33 km.





5. Conclusions

We have examined a year's worth of data from MIGHTI in the form of migrating and nonmigrating tides. With our focus on the lesserknown terdiurnal tides, these still seems to have significant amplitudes in the lower thermosphere.

 \checkmark Tides with n, s = 3 show seasonal variations

Fig. 3: Geographic coverage of MIGHTI observations, after Englert et al. 2017a; Harding et al. 2017a.

✤Nearly circular orbit; Altitude ~ 600 km; 27° inclination; Latitudes 12° South - 42° North MIGHTI A and B's phase of observed interference fringes, factored by oxygen red $(\lambda = 630 \text{ nm})$ and green line $(\lambda = 557.7 \text{ nm})$ emission.

Vertical profiles of horizontal thermospheric winds are measured

Red line phase change: 1.8 µrad for every 1 m/s wind speed

Green line phase change: 2.10 µrad for every 1 m/s wind speed

Tides with n = 3, s = 1, show a lot of variations in shorter time scales. Tides with n = 3, s = 0, also show a lot of variations in shorter time scales, but the peaks are at lower altitudes than s = 1. Tides in the meridional winds have peaks in the regions where there are no peaks in the corresponding zonal wind tides.

6. Future Work

Study more years' data to interpret annual patterns in the tides. Data from 2021 and 2022 can be used.

Incorporate red line data in the analysis to examine tides at higher altitude. Analyze phase and temperature values for more