

Abstract

The E-region ionosphere is the region where field-aligned currents terminate because of enhanced ionneutral collisions enabling perpendicular closure currents. The closure of these currents enables energy to be dissipated through Joule heating. Thayer, 1998; Aikio et al., 2012; Cai et al., 2013; and Zahn et al., 2021; have shown that the neutral wind can play a significant role in enhancing or reducing energy dissipation. The observations presented were primarily derived from incoherent scatter radar observations that self-consistently calculated the neutral wind. Sangalli et al., 2009 presented one of the only case studies to quantify the impacts of the neutral winds on Joule heating using rocket-based observations. Since 2009, several other campaigns have derived neutral winds using the chemical tracer technique, and combined with observations from the Poker Flat Incoherent Scatter Radar (PFISR), there is a unique opportunity to quantify the impact of the neutral winds on the Joule heating over a range of magnetic local time sectors and geomagnetic activity levels. Direct TMA observations provide high-resolution (~1 km) neutral wind data that is notably independent of other measurement techniques.

In this investigation, we present results that combine PFISR data with in-situ data taken during five separate sounding rocket campaigns: Joule II, MIST, JETS, and INCAA. For each campaign, altituderesolved neutral wind profiles have been determined via TMA observations which are triangulated using the line-of-sight projection method between 80-130 km. Where applicable, the impact of in-situ measured electric fields are compared with the impact PFISR-derived results have on energy exchange parameters. The Joule heating rate, including the effects of the neutral winds, are quantified during these experiments and the associated geomagnetic conditions. Our results assess the impact independently derived neutral winds have on Joule heating in the E-region spanning over multiple magnetic local time sectors and geomagnetic activity levels.

Motivation & Methods

- **Science Question: How does independently derived neutral winds impact energy transfer in the E-region during different MLT sectors and geomagnetic conditions?**
- To address the science question above, a total of four campaigns were considered. For each campaign in question, three selection criterion was met: The dispersal of TMA at E-region altitudes that was triangulated and processed into a neutral profile, the campaign must be at high latitudes with either low, medium, or high geomagnetic activity levels, and the campaigns must have coincident groundbased measurements for data verification. Two of the campaigns investigated can be viewed in Table

| Campaign | Date | MLT Sector | KP |
|----------|---------------------|------------|-----|
| Jets | 3/2/2017 - 5:41 UT | Dusk | 4.3 |
| INCAA | 4/7/2022 - 12:47 UT | Morning | 2.0 |

The Pedersen conductivity (conductance) was calculated using the equation below which is taken from Section 2.2 in Kelley, 2009. The ion-neutral collision frequency contained in this equation is calculated for both the resonant and non-resonant cases using section 4 in Shunk & Nagy, 2009. NRLMSIS-2.1 provides the neutral temperatures while PFISR provides both the plasma density and ion temperatures required. The ratio of the ion cyclotron frequency to the ion-neutral collision frequency (κ_i) is calculated with magnetic fields from IGRF and the aforementioned ion-neutral collision frequency.

$$\sigma_P = ne^2 / [M\nu_{in}(1+\kappa_i^2)]$$

The perpendicular ion drifts are estimated in the E-region with the methodology described in Heinselman and Nicolls, RS, 2008. The electric field estimated in the F-region used long pulse ISR data and is assumed to map vertically into the E-region. PFISR electric fields are compared against insitu measurements. Both sources for electric fields are used to calculate passive energy deposition rate which is also integrated with respect to height.

$$q_j^E = \sigma_P \mathbf{E}_{\perp}^2 \qquad \qquad Q_j^E = \Sigma_P \mathbf{E}_{\perp}^2$$

The vapor trail technique was used to provide measurements of the altitude resolved E-region neutral winds. The vapor trails of trimethylaluminum (TMA) were photographed and triangulated to obtain neutral velocity profiles utilizing the line-of-sight projection method providing profiles between 100-140 km. These neutral profiles are included in the calculation of the Joule heating rate which is also integrated with respect to height.

$$q_j = \sigma_P (\mathbf{E}_{\perp} + \mathbf{U}_n \times \mathbf{B})^2 \quad Q_j = \int \sigma_P (\mathbf{E}_{\perp} + \mathbf{U}_n \times \mathbf{B})^2$$

◆ In order to accurately attribute the contribution of neutral wind to Joule heating, the percent difference between the passive energy deposition rate and the Joule heating rate is calculated for both campaigns using the equation below. For negative percent differences, the neutral wind effectively increases the rate of Joule heating and for positive percent differences, the Joule heating rate is decreased when neutral winds are included in the calculation.

$$\% \text{ diff} = \frac{q_j^E - q_j}{q_j} \times 100\%$$

The Role of Neutral Winds in MIT Energy Transfer A. M. Pepper¹ and S. R. Kaeppler¹ 1. Department of Physics and Astronomy, Clemson University, Clemson, SC, 29634

 $\mathbf{B})^2 dz$





was found to have an average reduction of 26% between 110-140 km during the enhanced geomagnetic activity prior to launch.

* Using independently derived neutral wind profiles provides a high degree of precision to the calculations presented. The agreement between the PFISR based and in-situ based calculations allow for high resolution investigations both temporally and spatially. * The differences in the ion and neutral forcing for these MLT sectors and geomagnetic activity levels directly impacts the dissipation of energy in the E-region ionosphere. Acknowledgements: We thank R. Pfaff (PI of Jets) for providing neutral wind profiles and guidance essential in this study. We also acknowledge NASA grant #80NSSC19K0810 for funding.