

## Motivation

- Meteor head echoes abundant in high-power large-aperture (HPLA) radar data: up to thousands/hour
- Want to identify head echoes without searching through clutter
- Want method that works on **any** radar facility or pulse code
  - Existing method only works at Jicamarca with short pulse<sup>1</sup>
  - Head echoes can be observed at non-HPLA facilities, albeit more infrequently<sup>2,3</sup>, and technique would be valuable there

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## Meteor Experiment

Simultaneous head echo observations at three high-power radar facilities were taken before dawn on October 10<sup>th</sup> & 11<sup>th</sup>, 2019.



### **Resolute Bay Incoherent Scatter Radar** ~120 head echoes/hour

Frequency: 442.5 MHz Pulse code: Min. Sidelobe 51, 51 µs with 1 µs baud Pulse period: 1.4 ms Beam angle: 26° azimuth, 86° elevation

### Millstone Hill Observatory (MHO)

### ~600 head echoes/hour

442.9 MHz
Barker-7, 42 µs with 6
2 ms
270° azimuth, 45° elev

### Jicamarca Radio Observatory (JRO) ~5000 head echoes/hour

Frequency: 50 MHz Pulse code: Min. Sidelobe 51, 51 µs with 1 µs baud Pulse period: 1.25 ms Beam angle: 90° elevation

## General Approach

- Split raw data into chunks of size  $2 \times 150 \times 150$  (or similar) with some overlap
- Real and imaginary components form separate channels
- Classify chunks as head echo (label = 1) or non head echo (label = 0)
- Group together adjacent positive chunks for further analysis



## Meteor Head Echo Detection at Multiple High-Power Radar Facilities via a **Convolutional Neural Network Trained on Synthetic Data**

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(RISR-N)

 $6 \mu s baud$ 

vation



## Training via Synthetic Data

- **Problem:** Need O(10000) head echoes to train the network for a specific meteor radar experiment!
- Solution: generating synthetic head echoes
- Head echoes reflect pulse code + Doppler shift  $\rightarrow$  easy to synthesize
- We assume meteoroid range is exponential, and received power curve is a sinc function, with some random parameters...

$$r(t) = \left(r_0 - \frac{v_0^2}{a_0}\right) + \frac{v_0^2}{a_0}e^{(a_0/v_0)t}$$

 $r_0 \in [75, 130] \text{ km}$ **Observed head echo range:**  $v_0 \in [12,73] \text{ km/s}$  $a_0 \in [0,70] \text{ km/s}^2$  $SNR \in \sim [10, 40] \, dB$ **Received voltage signal (complex):**  $c(t) \in \mathbb{C}$  (pulse code)  $\phi(t) \in \mathbb{R}$  (Doppler phase)  $nf \in \mathbb{R}$  (noise floor)  $a, b, \phi_0 \in \mathbb{R}$  (constants) Matched filter Training example Real clutter

$$b(t) = \frac{4\pi f}{c} \int_0^t v(t)dt + \phi_o$$

 $rx(t) = 10^{\frac{1}{20}(SNR+nf)}c(t)e^{i\phi(t)}\operatorname{sinc}(a+bt)$ Radar clutter is hard to synthesize but is more abundant than head echoes. Can identify clutter in data and add to synthetic head echoes:



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Total number of parameters

	<b>RISR-N</b>	MHO	JRO (non-EEJ)	JRO (EEJ)
Total test set size	11868	7140	4836	4836
<pre># positive (truth)</pre>	157	148	290	122
<pre># negative (truth)</pre>	11631	6848	4373	4574
<pre># inconclusive (truth)</pre>	80	144	173	140

# truth data. Results at each facility:

Classification accuracy chunks correctly classified Sensitivity (% of head ecl chunks that are identified) Sensitivity to  $\geq 15$  dB ec Sensitivity to  $\leq 15$  dB ec Precision (% of head echo classifications that are corre

Overall sensitivities greater than 70% demonstrate the technique's capability to discriminate between head echoes and clutter. Sensitivities greater than 50% for head echoes less than 15 dB demonstrate detection of a non-exhaustive selection of the weakest head echoes.

### **Class activation maps for some head echoes in raw data...**



## **Conclusions and Future Work**

- population analyses

<b>References:</b>	Ac
[1] Li et al., 2022	Gr
[2] Janches et al., 2014	Co
[3] Panka et al., 2021	Tre
[4] Li et al., 2016	Sig
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### **Results: Performance Metrics**

We manually labelled "truth" data at each facility and excluded inconclusive streak-like objects from analysis. Training set statistics:

Each neural network corresponding to each facility is tested against

	<b>RISR-N</b>	MHO	JRO	JRO
			(non-EEJ)	(EEJ)
(% of )	99.6	99.4	96.2	97.4
ho	75.8	80.8	77.4	71.3
choes	93.4	97.0	91.7	77.8
choes	51.5	65.3	59.4	58.5
o ect)	96.7	88.1	66.4	50.3



• Convolutional neural networks (CNNs) are highly capable of distinguishing head echoes from other radar signatures

• The CNN technique is sensitive to the weakest head echoes, which may be missed by the human eye when manually searching

• Can apply technique at any facility and pulse code to greatly speed up

• Future meteor experiments can benefit from improved meteor shower identification and estimation of neutral densities<sup>4</sup> with the technique • Future work will more accurately quantify sensitivity biases of the method, and refine the synthetic data model and CNN architecture to improve sensitivity to the weakest head echoes

> cknowledgements: Thanks to the NSF Grant AGS-1920383 and NSF ant AGS-2048349 for supporting this work.

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