

Shuang Xu, Justin Carstens, Jeff A. France, Cora E. Randall, Jia Yue, James M. Russell III Department of Atmospheric and Planetary Sciences, Hampton University, Hampton, Virginia, USA E-mail: shuang.xu@my.hamptonu.edu

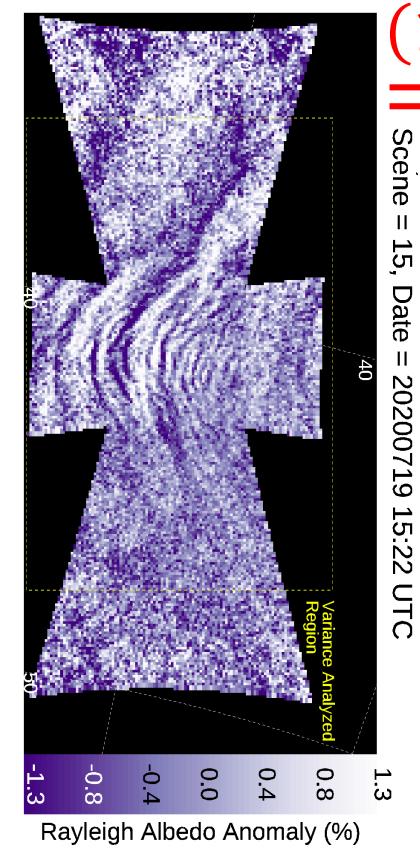
## **Science Questions:**

- What is the morphology of gravity waves (GWs) in the upper stratosphere?
- How do the GWs vary and what are the sources?
- What is the role of stratospheric GWs in coupling of atmospheric regions?

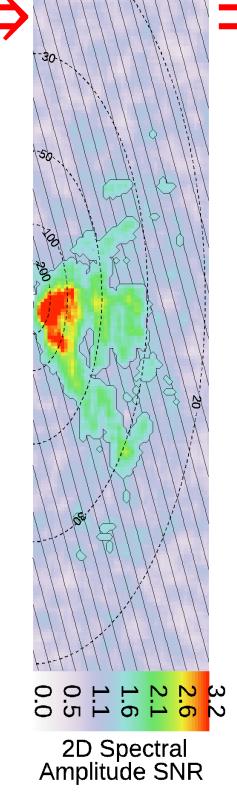
## **1.** Introduction

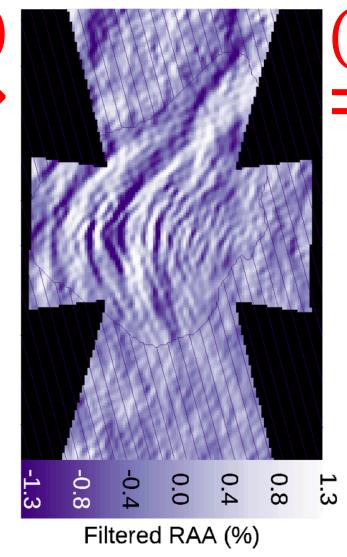
- First near-global maps of GW activity near **50–55 km altitude** from NASA Aeronomy of Ice in the Mesosphere (AIM) Cloud Imaging and Particle Size (CIPS) instrument
- GWs inferred from variances of Rayleigh Albedo Anomaly (RAA) variances of Rayleigh scattering @ 265 nm.
- Comparisons to GW hotspots near **30-40 km altitude** inferred from Atmospheric Infrared Sounder (AIRS) brightness temperature perturbation (BTP) @ 4.3 µm are presented
- **2.** Methodology

Calculate "Peak event frequency" (PEF)



**Definition of PEF** 

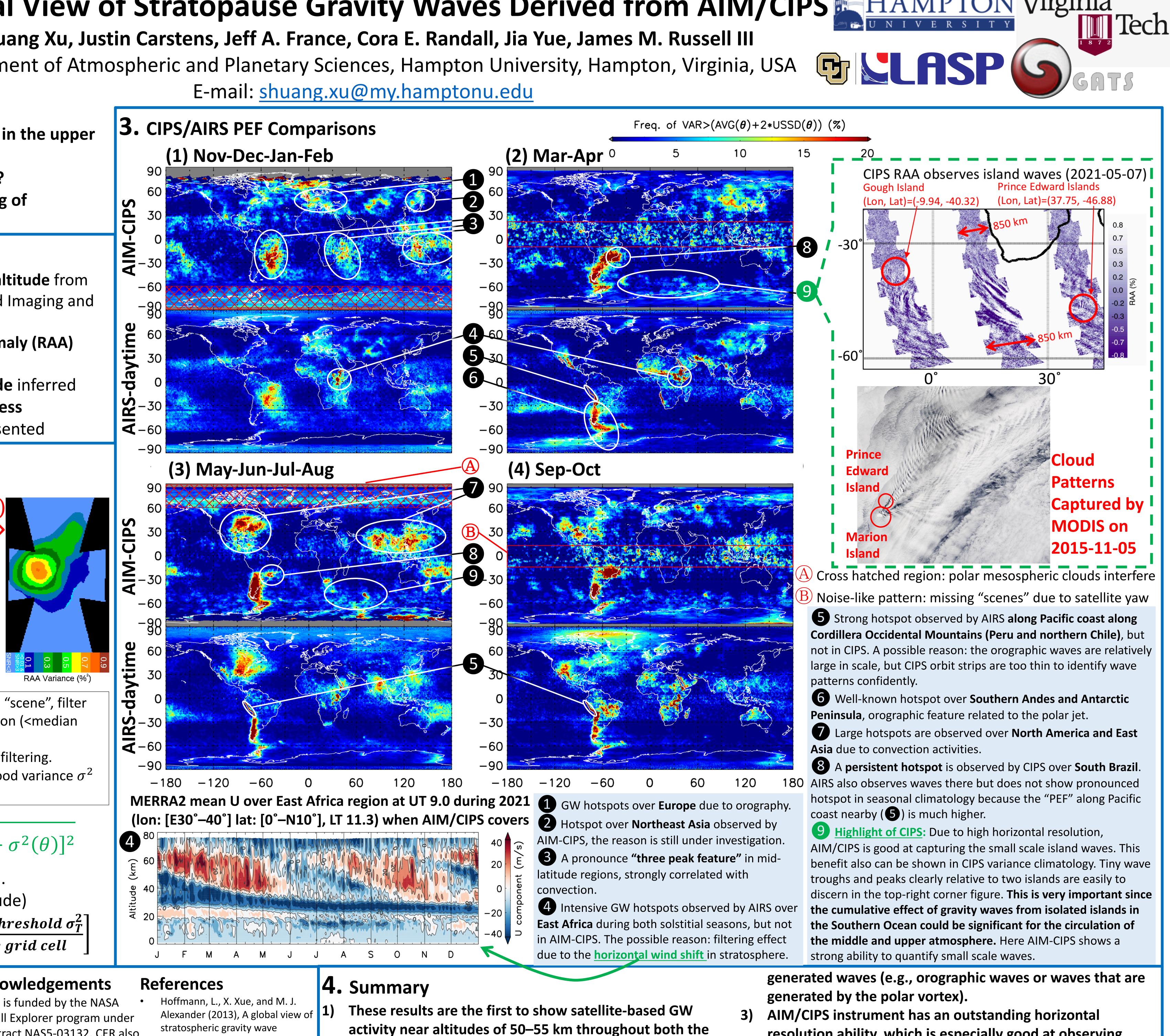




(1) Apply 2D FFT to each "scene", filter out the insignificant region (<median noise, line filled part). **2** Reconstruction after filtering. (3) Calculate neighborhood variance  $\sigma^2$ over each pixel.

| $\sigma_{\mathrm{T}}^{2}(\theta) = \sigma^{2}(\theta) + n \times \sqrt{\frac{1}{N-1} \sum_{i=1}^{N} [\sigma^{2}(\theta) + n \times \sqrt{\frac{1}{N-1}$ | $(\theta_i)$ –   |
|---|--|
| i.e., Threshold(θ)=AVG(θ)+ n×U  |  |
| (factor $oldsymbol{n}=2$ in this study, $oldsymbol{	heta}$ means  | s latitud  |
| $\begin{bmatrix} The PEF in each \end{bmatrix} = \begin{bmatrix} # pixel that ex \\ pixel that ex \end{bmatrix}$  | ceed th  |
| The PEF in each<br>map grid cell (unit: %)=[# pixel that exceed the<br>total # pixel in the grid cell (unit: %)   |  |
| <ul> <li>Compare CIPS and AIRS</li> </ul>   |  |
| <ul> <li>Observation time of CIPS: 2019 Nov ~ 2021</li> <li>Oct (2 years); of AIRS: 2019 Aug ~2021 Sep (w/o 2020 Aug–Sep, 2 years)</li> <li>Horizontal Resolution: CIPS 7.5 km; AIRS</li> </ul>   | <ul> <li>Ackno</li> <li>AIM is<br/>Small<br/>contra<br/>ackno</li> </ul> |
| <ul> <li>Inom20mar Resolution. CIPS 7.5 km, Ans</li> <li>14-40 km</li> <li>Local time: CIPS 9:00-11:30; AIRS 13:30</li> </ul>   | 80NSS<br>the A<br>and d  |
| GW altitude: CIPS 50-55 km; AIRS 30-40 km   | partic<br>Craft,   |

## A Global View of Stratopause Gravity Waves Derived from AIM/CIPS <u>AMPTON</u> Virginia



Il Explorer program under ract NAS5-03132. CER also nowledges NASA grant SSC20K0628. We thank AIM mission operations data systems teams, in icular Dave Welch, Jim , William Barrett, and an Gehr.

hotspots located with Atmospheric Infrared Sounder observations, J Geophys. Res. Atmos., 118, 416-434, doi:10.1029/2012JD018658. Randall, C. E., et al. (2017), New AIM/CIPS global observations of gravity waves near 50–55 km, Geophys. Res. Lett., 44, doi:10.1002/2017GL073943.

northern and southern hemispheres in all seasons. The CIPS RAA variance seasonal maps of GW activity near 50-55 km altitude show many of the same hotspots observed near 30-40 km in the AIRS 4.3-µm brightness temperature variances. Both CIPS and AIRS detect convectively-generated waves and non-convectively-

4)

resolution ability, which is especially good at observing small scale waves (e.g., island waves). This ability can be effectively used to quantify the event frequency/physical influence over atmosphere circulation via small scale waves. Further investigation is being applied to better interpret differences between the climatological results obtained by AIRS and CIPS.