



# Characterizing High-Resolution Temporal and Spatial Changes of the EIA and ETA

Anton Buynovskiy<sup>1,2</sup>, Jeffrey P. Thayer<sup>1,2</sup>, Marcin D. Pilinski<sup>1,2,3</sup>, Eric K. Sutton<sup>2</sup>, Scott E. Palo<sup>1,2</sup>, Austin Coleman<sup>1</sup>, Dong Lin<sup>4</sup>

<sup>1</sup>CU Boulder/Aerospace Engineering Science, <sup>2</sup>CU Boulder/SWxTREC, <sup>3</sup>CU Boulder/LASP, <sup>4</sup>NCAR/HAO



## Motivation

- Two prominent features in the Ionosphere-Thermosphere (I-T) region are called the Equatorial Ionization Anomaly (EIA) and Equatorial Thermospheric Anomaly (ETA)
- The mechanisms behind the formation of the ETA are not well understood due to the lack of satellite observations with suborbital temporal and small-scale spatial resolutions
- This poster utilizes NCAR's high-resolution TIEGCM to simulate the I-T region while modeling a constellation to simulate instrument sampling
- The constellation is based off the SWARM-EX mission, which consists of 3 identical spacecraft that will take in-situ measurements of the I-T region
- This poster will also explore how ETA formation theories, particularly a theory dealing with the field-aligned ion drag term, can be supported or rejected using the higher temporal and spatial resolutions of a constellation

## Science Background

The EIA is a dayside, equatorial feature with crests at  $\pm 15^\circ$  in magnetic latitude and a trough at the magnetic equator as shown in Fig. 1

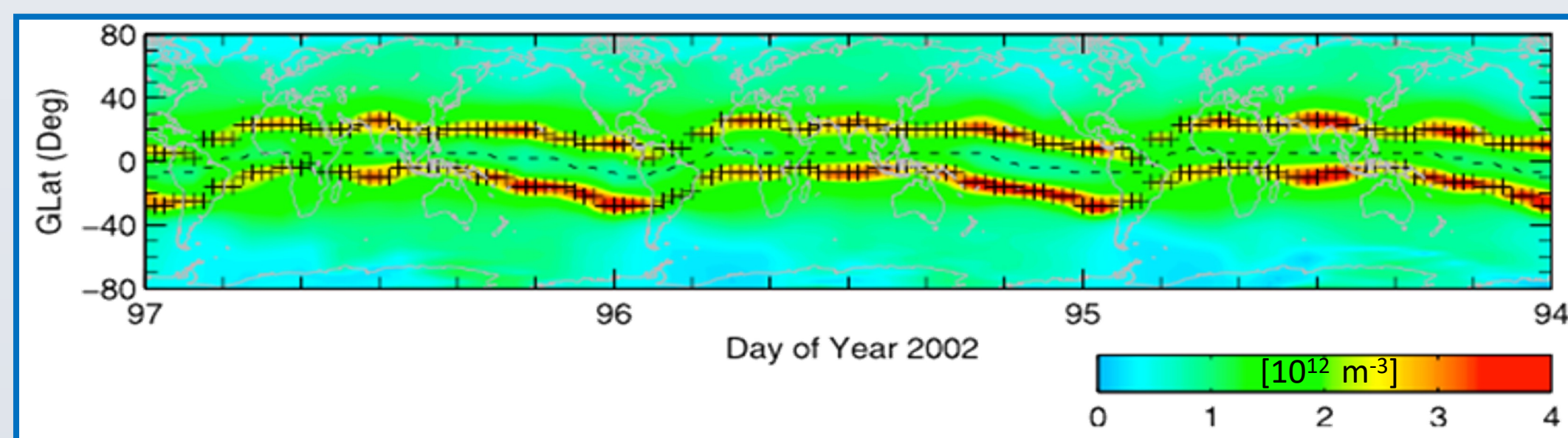


Figure 1: Electron Density at 400 km from the CHAMP mission [1]

The ETA is also a dayside, equatorial feature but has two crests at  $\pm 20-30^\circ$  in magnetic latitude and a trough near the magnetic equator as shown in Fig. 2. Although the ETA is a neutral feature, it is magnetically aligned, showing strong ion-neutral coupling between the EIA and ETA.

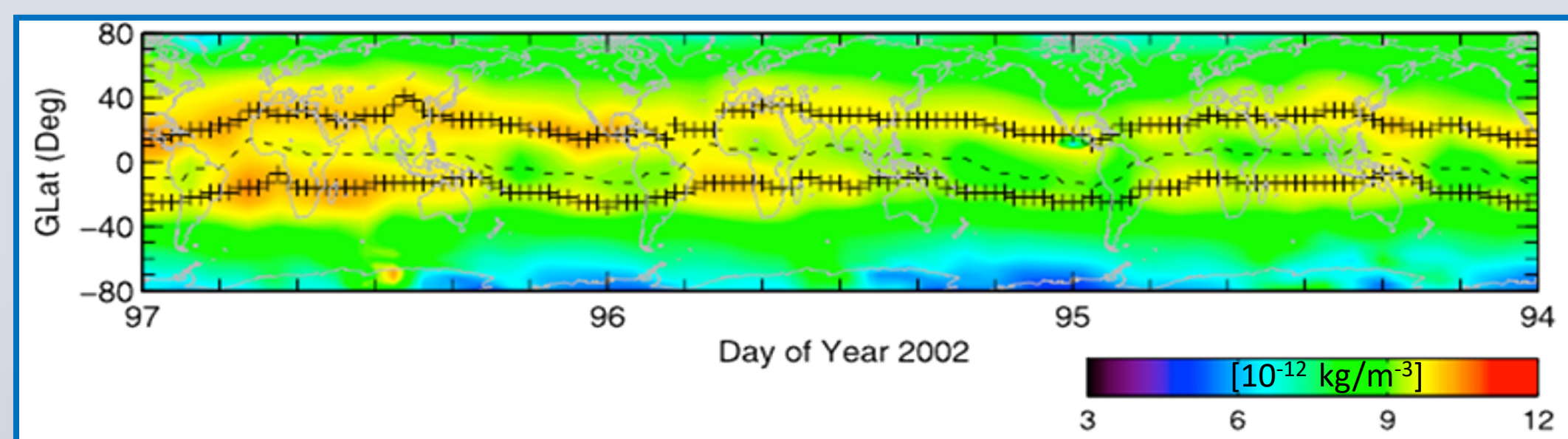


Figure 2: Neutral Density at 400 km from the CHAMP mission [1]

### ETA Formation Theory: Field-Aligned Ion Drag [2]

One of the latest theories dealing with the formation of the ETA is a paper by Hsu et al (2014). This theory states that the field-aligned ion drag force,  $F_{ni,||}$ , is a non-negligible term that creates a trough in the thermosphere, forming the ETA. This theory is broken down into a 4-step process:

- Field-Aligned Ion Drag Force
- Divergence of meridional winds
- Change in vertical winds, adiabatic heating/cooling, and neutral temperature
- Reactive Pressure Gradient to rebalance forces

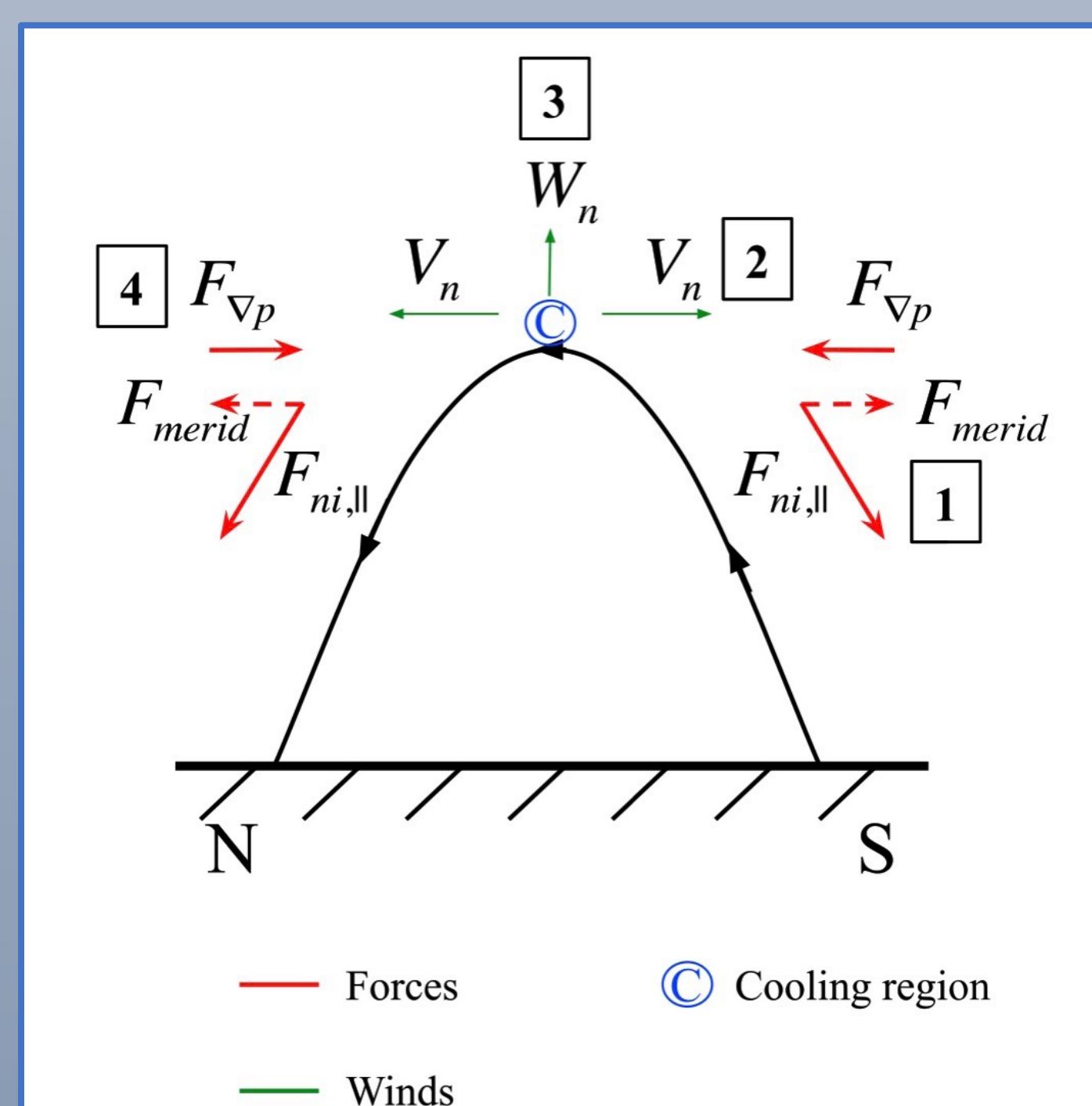


Figure 3: Four Step ETA Formation Diagram

## Model Approach & Constellation Design

### Model Approach

- NCAR's High-Resolution Thermosphere-Ionosphere-Electrodynamic General Circulation Model (HR-TIEGCM) is utilized [3]
- HR-TIEGCM can numerically solve for the coupled I-T system and accommodate the different timescales for plasmas and neutrals
- HR-TIEGCM has a grid resolution of  $0.625^\circ \times 0.625^\circ$  with a 1-minute sampling timestep
- Two simulations are run in a quiet environment ( $K_p = 2$ ,  $F_{10.7} = 120$ ) during solar max at vernal equinox with run 1 having  $F_{ni,||}$  turned on and run 2 having  $F_{ni,||}$  turned off

### Constellation Design

- A satellite fly-through scheme is implemented to simulate instrument sampling
- Constellation parameters are based off SWARM-EX's mission design with a nominal altitude of 400km, an inclination of  $60^\circ$ , and 5 minute separation between each of the 3 satellites
- This design allows the I-T region to be observed on suborbital timescales and spatial scales not previously observed
- Figure 4 shows the ground track for the SWARM-EX constellation and Fig. 5 shows a visual of the spatial and temporal resolutions

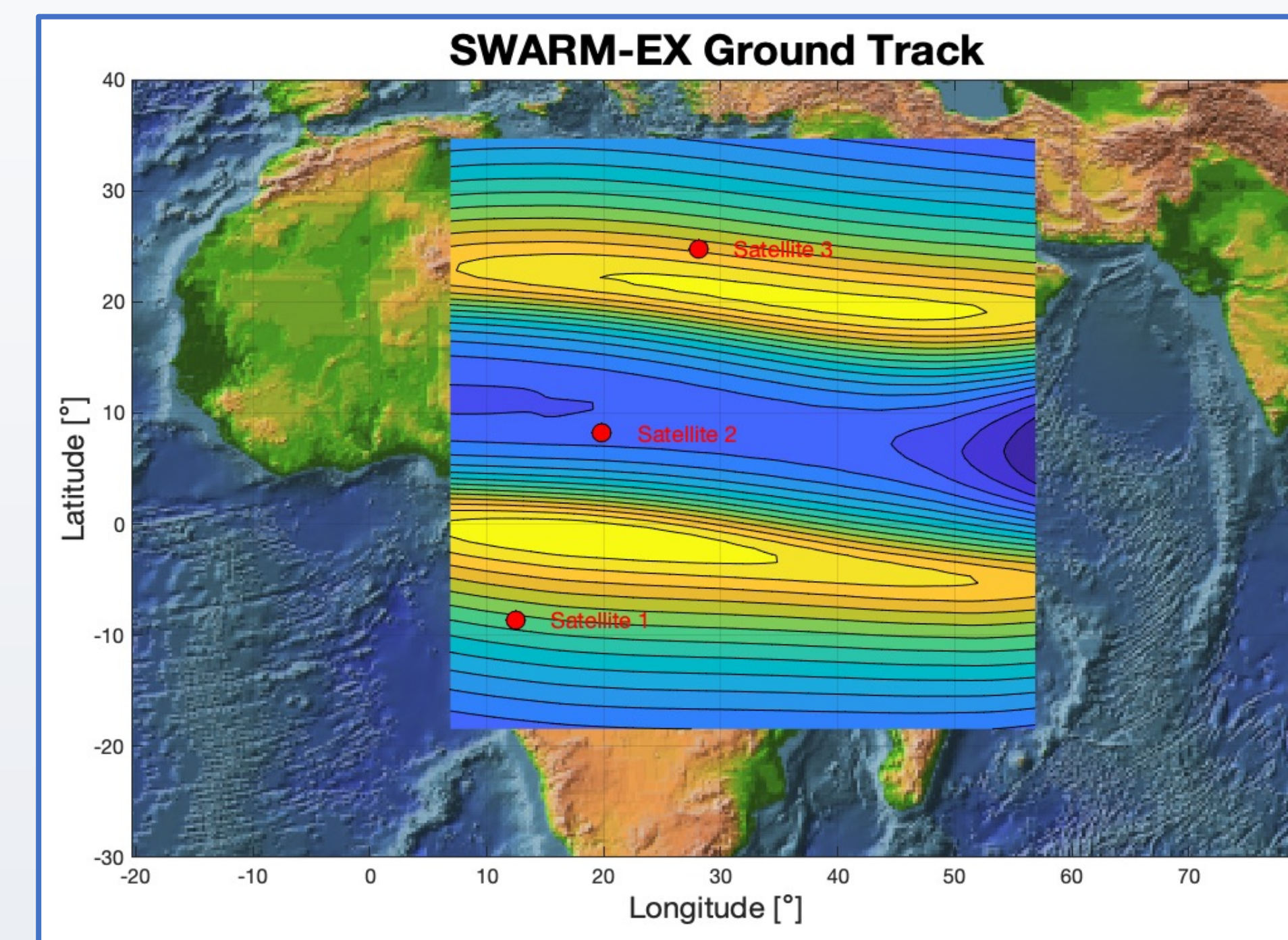


Figure 4: SWARM-EX Ground Track Visual

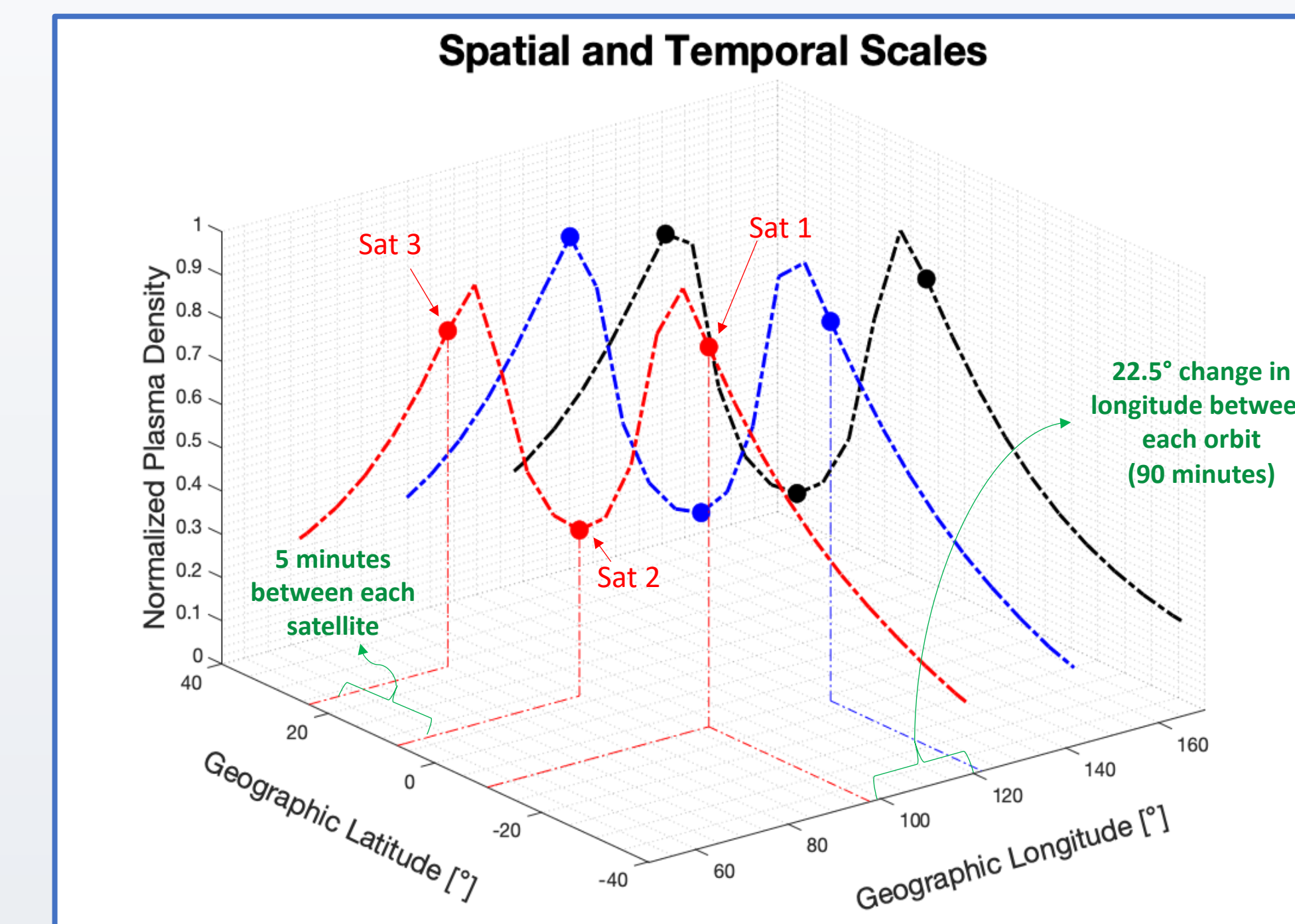


Figure 5: SWARM-EX Spatial and Temporal Scales

## Results

### Constellation Observations

- Figure 6 shows the simulation output from the HR-TIEGCM
- The magnetic dip in the North American (NA) sector shown in Fig. 7 is an example where significant spatial changes in magnetic topology can be mistaken for temporal changes in plasma and neutrals if not sampled adequately
- To differentiate between spatial and temporal change, the EIA and ETA must meet the persistence criteria
- Persistence can be defined as a feature that is spatially coherent with little change in phase and whose uncorrelated geophysical noise is below some percent change
- Table 1 shows the persistence criteria for the EIA and ETA inside and outside the NA sector based on the 3-satellite constellation

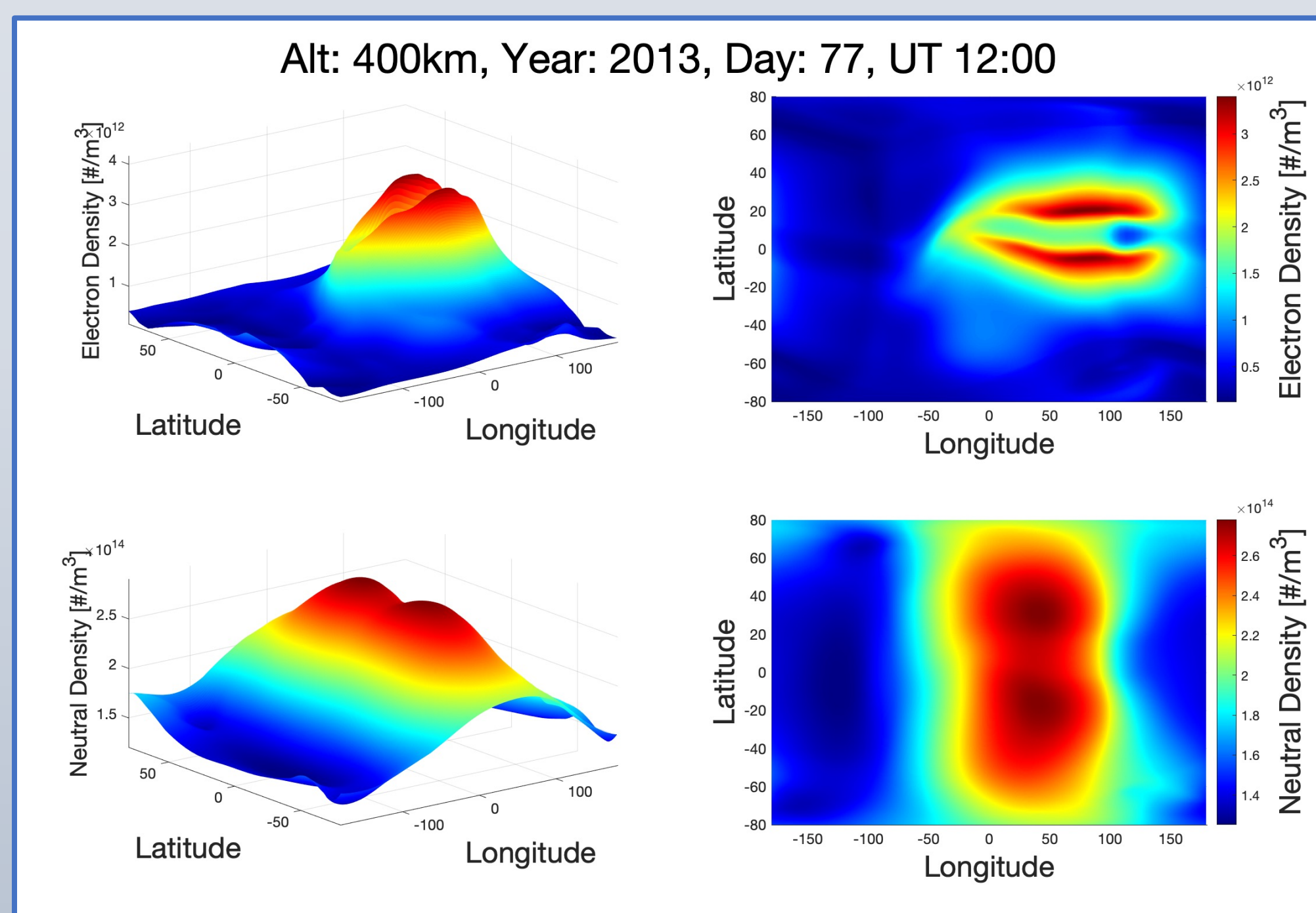


Figure 6: Global Representation of the I-T Region

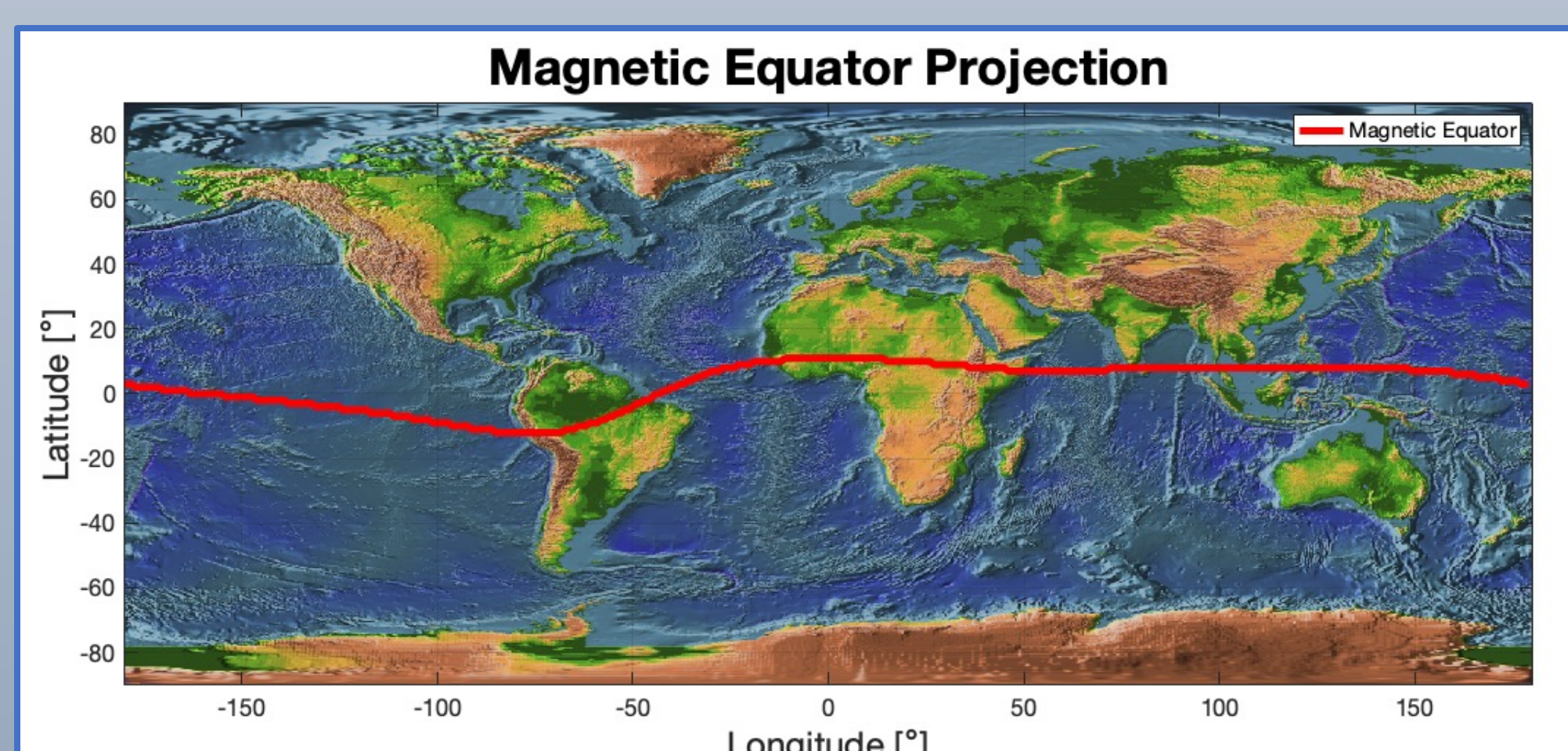


Figure 7: Magnetic Dip in the North American Sector

Table 1: Persistence Criteria

Feature	Value
ETA	$< \pm 0.5\%$ change over 5 min inside the NA sector $< \pm 0.2\%$ change over 5 min outside the NA sector
EIA	$< \pm 6.0\%$ change over 5 min inside the NA sector $< \pm 2.0\%$ change over 5 min outside the NA sector

### Theory Observations

- Figure 8 observes key parameters mentioned in the field-aligned ion drag ETA theory between runs 1 and 2 for one satellite
- The theory indicates a decrease in neutral temperature and neutral density at the magnetic equator, which is evident in Fig. 8
- Figure 9 shows how a satellite constellation can isolate persistent features to avoid disturbances unrelated to the ETA formation using the persistence criteria from Table 1

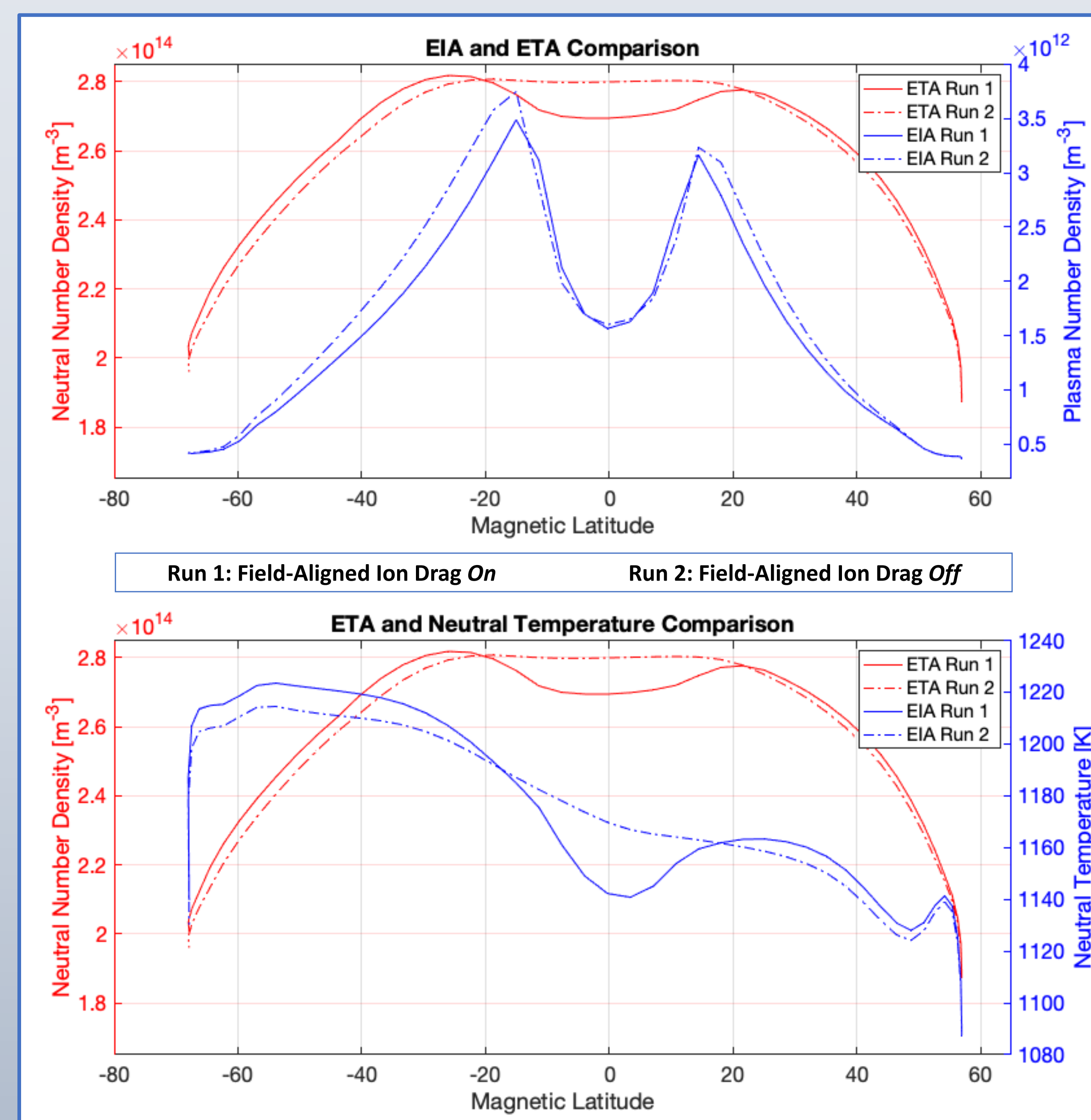


Figure 8: Comparison of Various Parameters Between Runs 1 and 2

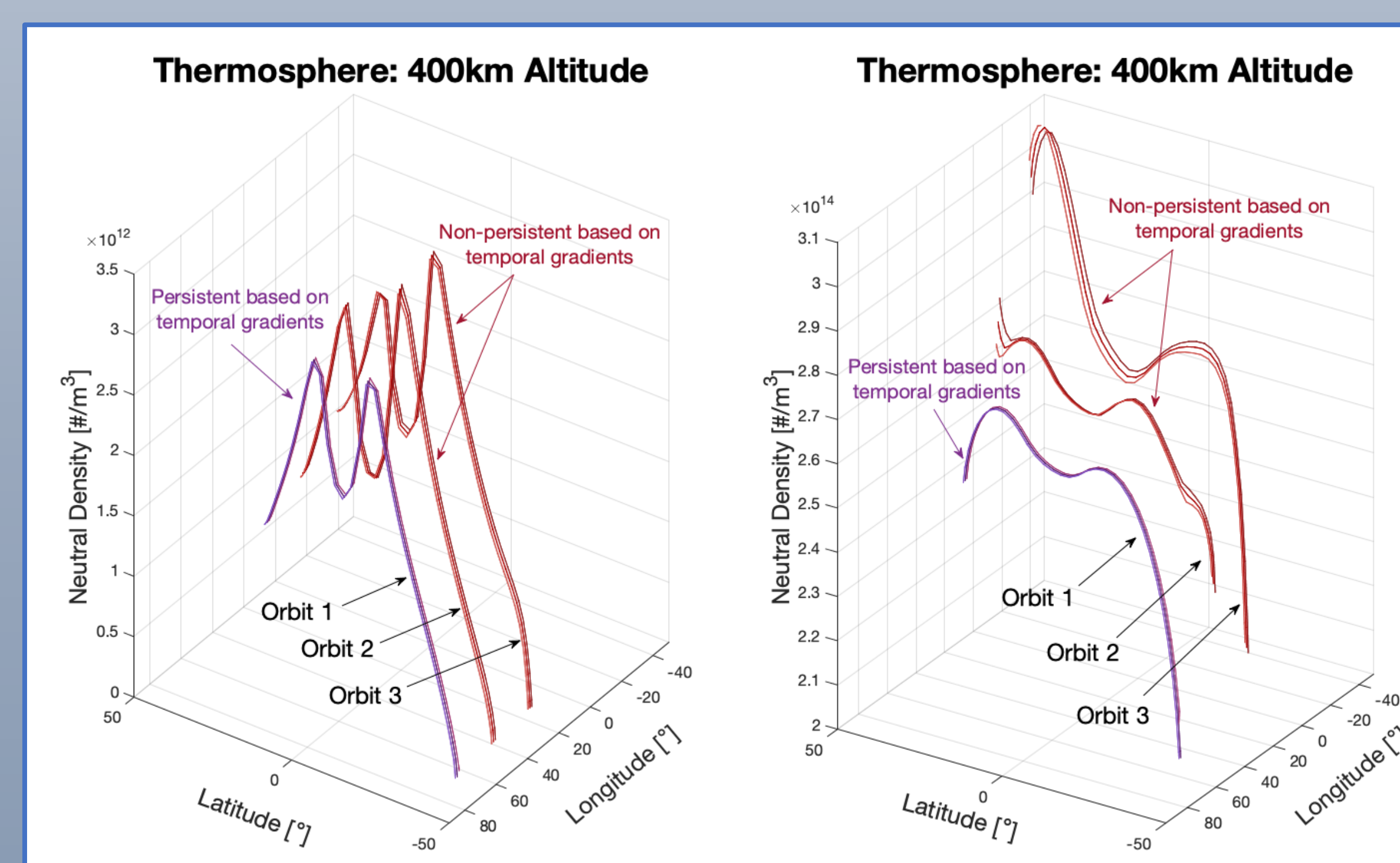


Figure 9: Representation of Persistence During a Geomagnetic Storm

## Discussion

- HR-TIEGCM simulates the ETA comparable to orbital data and with the HR-TIEGCM being capable of resolving mesoscale changes in the I-T region, orbital data from a satellite constellations can be directly compared to explore ETA formation theories
- Instrument sampling of the ETA using HR-TIEGCM showed the ETA density and temperature changes aligned with the field-aligned ion drag theory proposed
- Using a quantitative definition for persistence, a multi-satellite constellation can capture small scale changes to determine whether the satellites are observing ETA formation processes or transients
- In the case where the magnetic kink is present, the persistence definition can establish that the EIA/ETA features are persistent and spatially coherent
- To capture all aspects of the ion-neutral coupling in the I-T region, a larger constellation is required to optimize the different timescales of plasmas and neutrals (NASA GDC multi-satellite mission will enable such investigations)

## Future Work

- Explore methods of extracting spatial and temporal gradients from a satellite constellation to determine limiting resolution factors
- Explore constellation designs that can utilize the least number of spacecraft to capture I-T processes
- Continue characterizing correlations between the EIA and ETA trough locations as well as their associated densities
- Explore other parameters that can be observed in a constellation to validate/reject ETA formation theories
- Follow up on the SWARM-EX mission to compare orbital data to HR-TIEGCM simulations

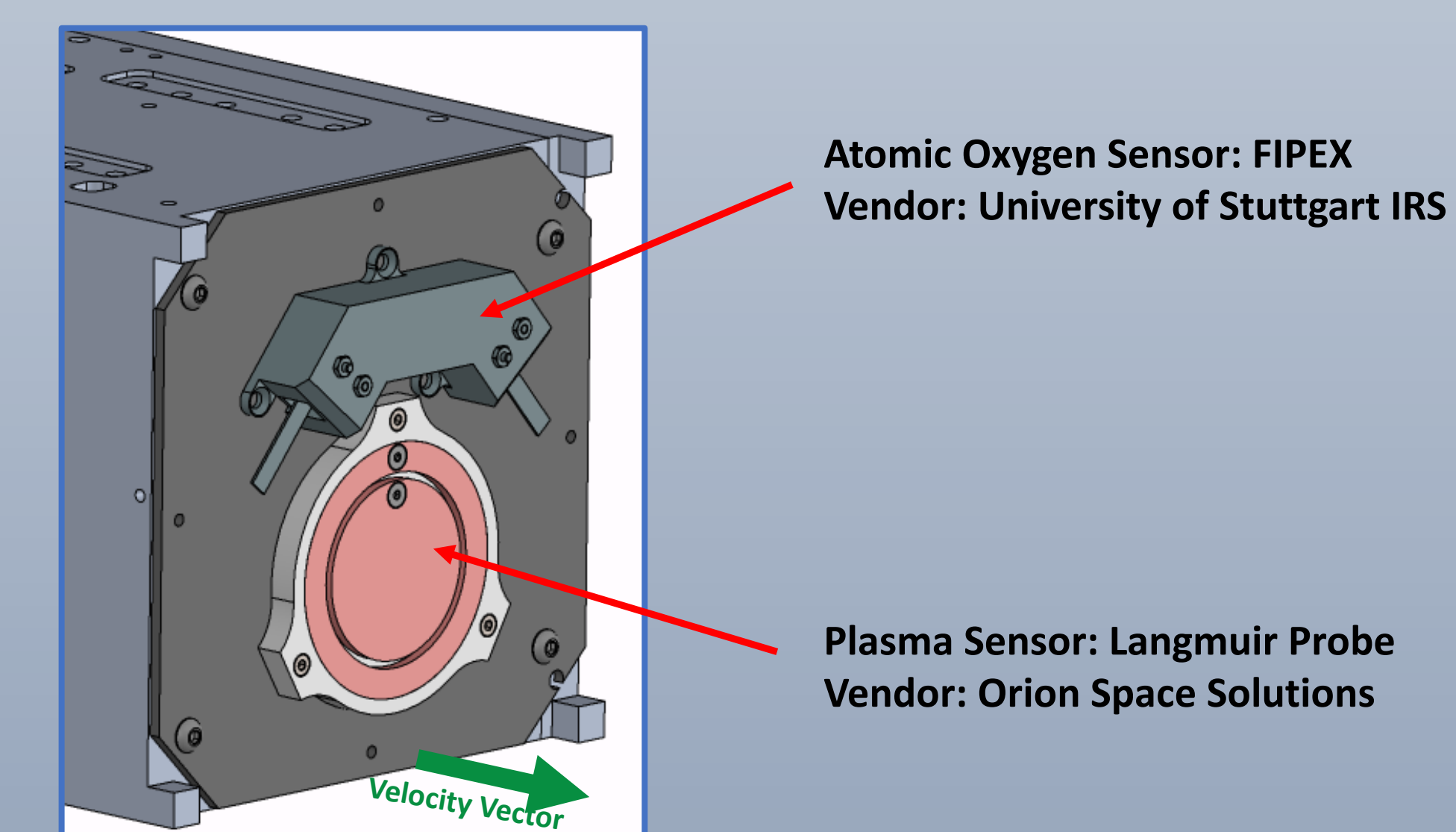


Figure 10: SWARM-EX Instrument Deck

## References

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