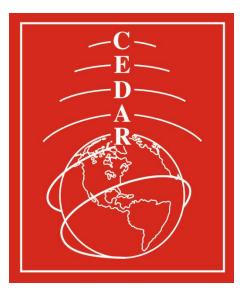
Mesospheric bore evolution characteristics

Brian Laughman NorthWest Research Assoicates CEDAR Postdoc Report #2 June 27th, 2012





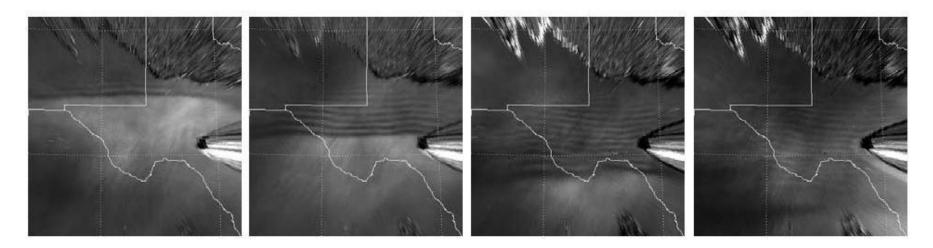


Mesospheric Bore Evolution Characteristics

- What are Mesospheric Bores?
- Numerical Models
- BDO theory
- Results
 - Forcing dependence
 - Ducting environment dependence
- Conclusions

What are mesospheric bores?

- Roughly 2D, weakly non-linear, ducted wave phenomenon
- Characterized by crest creation
- Observed at roughly 80 100 km
- Propagate at speed of ~ 75 m/s
- Length scales of ~ 20 km



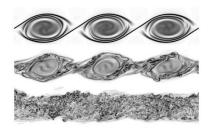
4 time ordered images of a propagating bore courtesy of Steve Smith. CSU lidar data shows a strong thermal inversion of 50 K between 85 and 90 km.

Other Bores

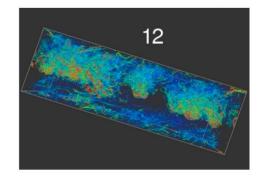


Navier-Stokes model

We numerically solve the Navier – Stokes equations



Momentum Continuity Heat Equation Ideal Gas Law



- It is a DNS model Direct Numerical Simulation
- It uses the Boussinesq Approximation incompressible
- 2-D simulations evolve *potential temperature* and *vertical velocity*

BDO model : A useful approximation

The BDO equation makes two approximations.

1. Fairly-long wavelength \rightarrow weak dispersion

$$c = c_0 (1 - \delta |k|)$$

2. Small-but-finite amplitude \rightarrow weakly nonlinear effects:

• Steepening • No breaking • No recirculation

Evolves displacement:

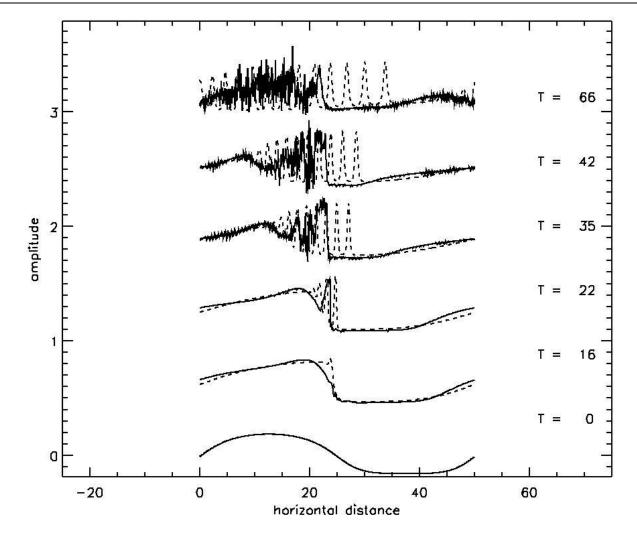
 $\eta(x, z, t) = A(x, t) \phi(z) \longrightarrow \bullet$ Modal equation

Evolution equation

Steady-state solution:

$$A(x,t) = \frac{a\lambda^2}{\left(x - c_b t\right)^2 + \lambda^2}, \quad c_b = c_0 + \frac{\delta}{\lambda}, \quad a\lambda = \frac{4\delta}{\alpha}$$

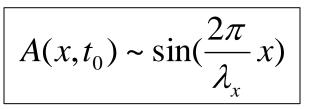
The BDO equation : large amplitude limit



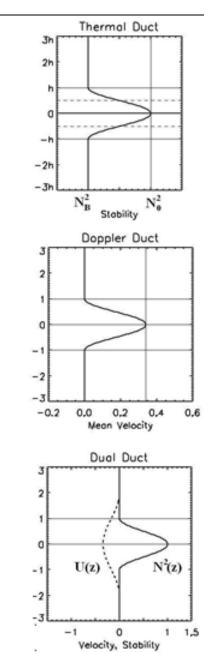
- BDO handles nonlinearity through crest creation
- NS model predicts breaking

Results

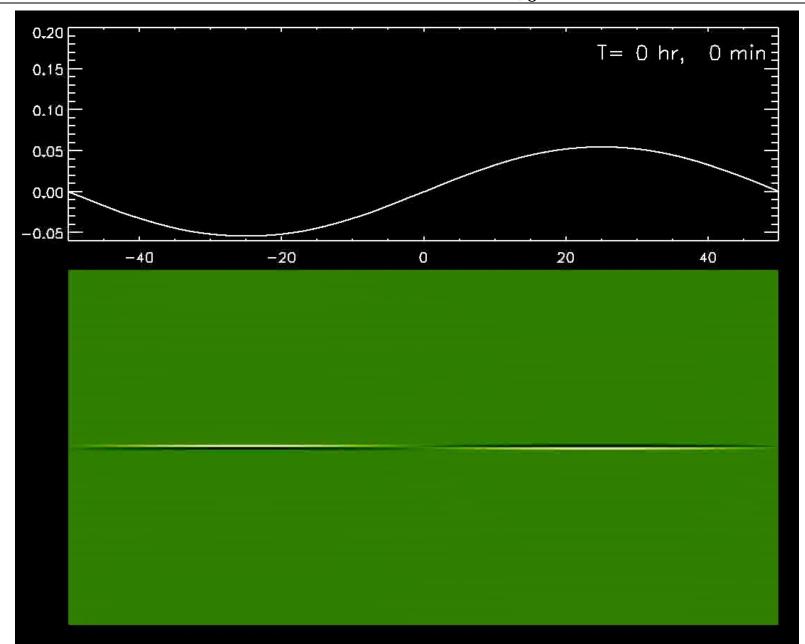
- Forcing geometry effects (simple thermal duct)
- 1. Wavelength effects
- 2. Amplitude effects



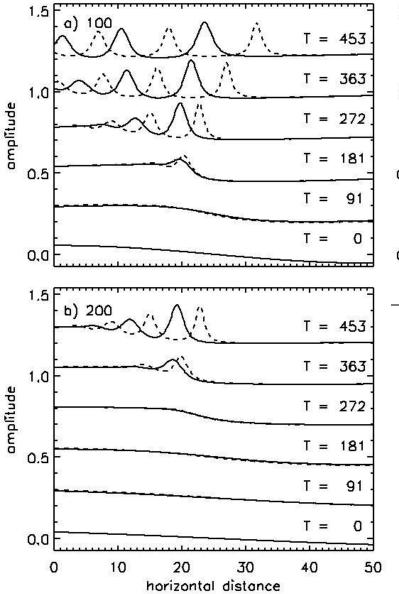
- Ducting environment effects
- 1. Duct shape (cosine v. sech-squared)
- 2. Duct thickness (*increased h*)
- 3. Viscous effects
- 4. Non-zero background stability
- 5. Isolated Doppler duct
- 6. Co-located ducts (Headwinds/Tailwinds)

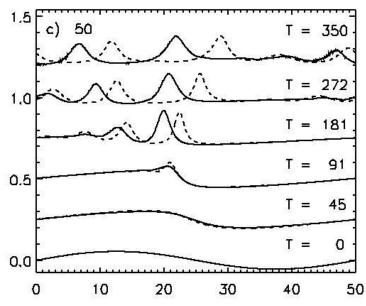


Results : Thermal duct, $\lambda = 100$, $\eta_0 = 0.1$, $\nu = 0$



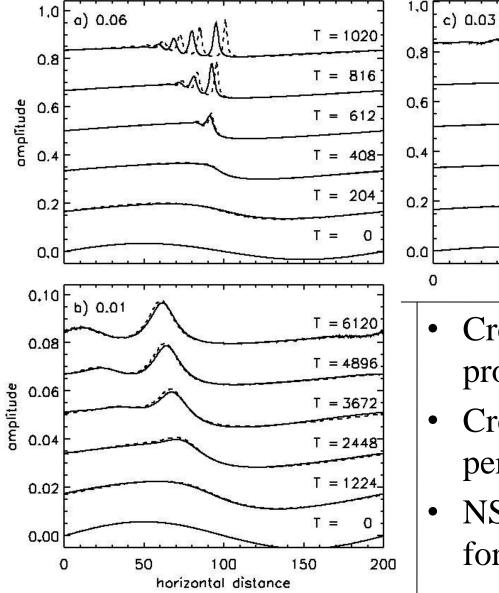
Results : Sine wave, wavelength variation

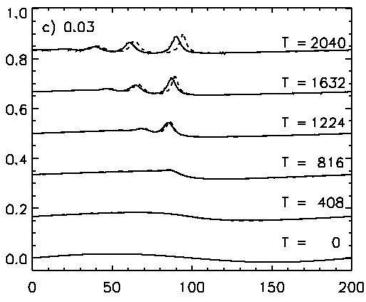




- Crest creation time proportional to wavelength.
- Crest geometry is roughly independent of perturbation wavelength.

Results : Sine wave, amplitude variation



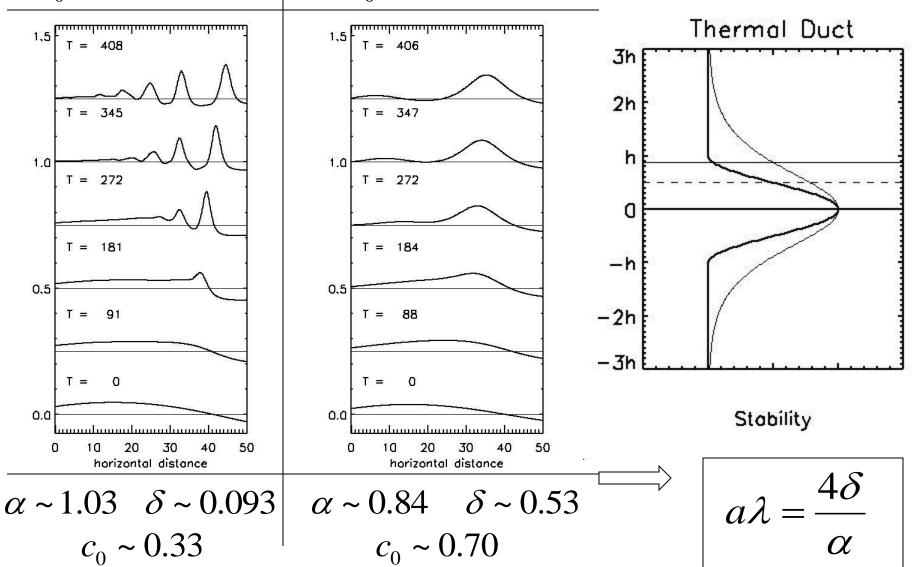


- Crest creation time inversely proportional to amplitude.
- Crest geometry is a function of perturbation amplitude.
- NS/BDO agreement improves for smaller amplitudes.

Results : Duct shape

 $\sim N_0^2 \cos(\pi z/h)$

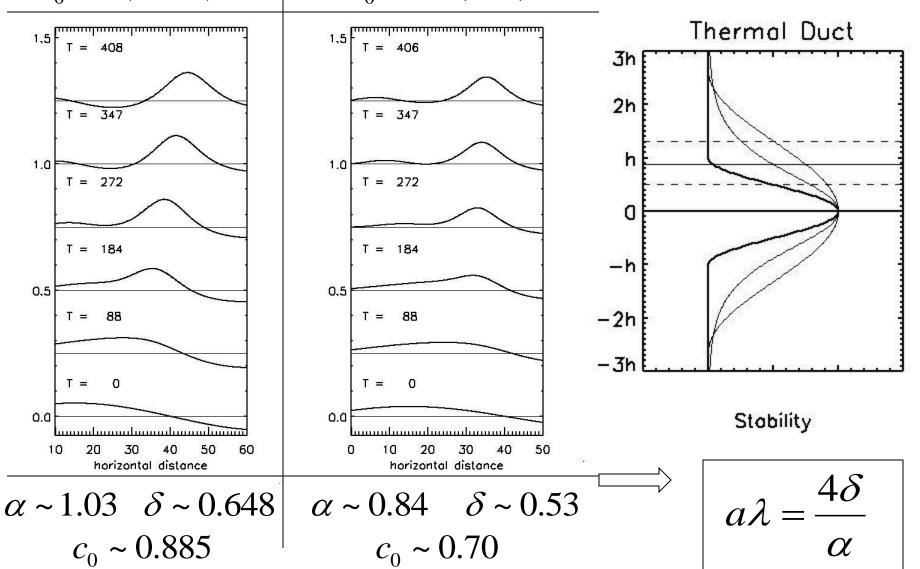
~ N_0^2 sech²(z/h)



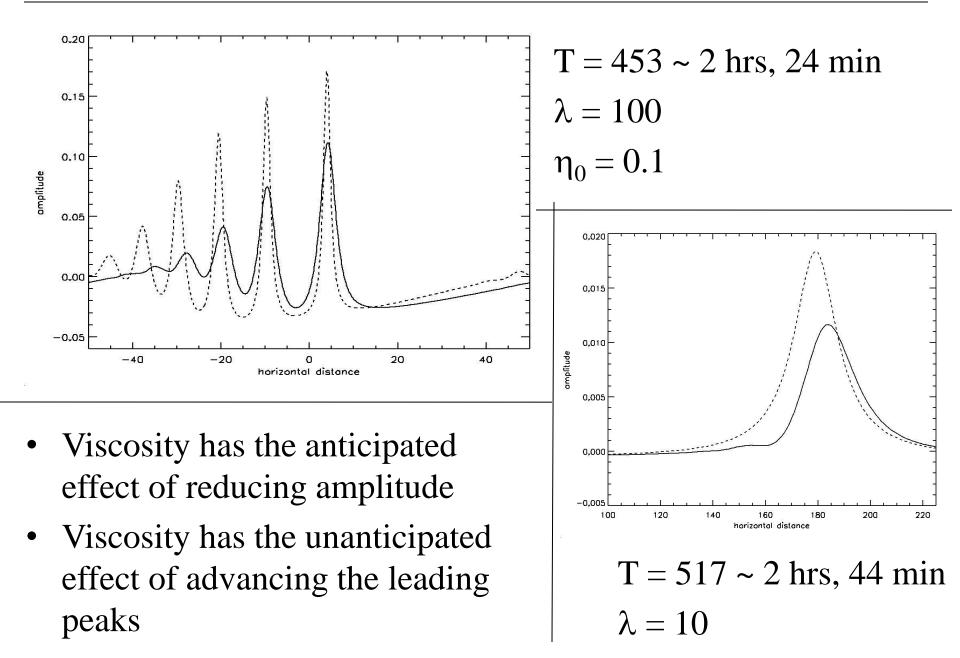
Results : Duct thickness

 $\sim N_0^2 \cos(\pi z/h)$

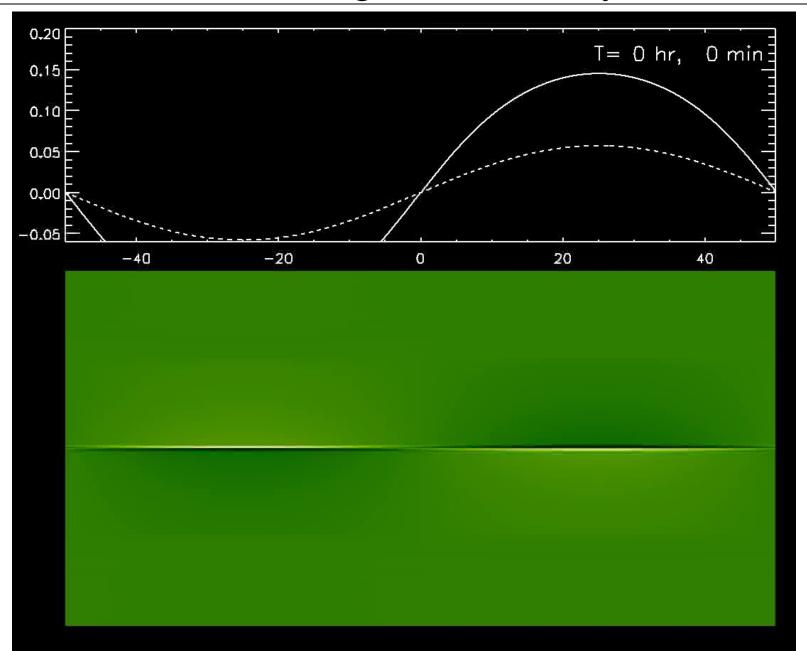
~ N_0^2 sech²(z/h)



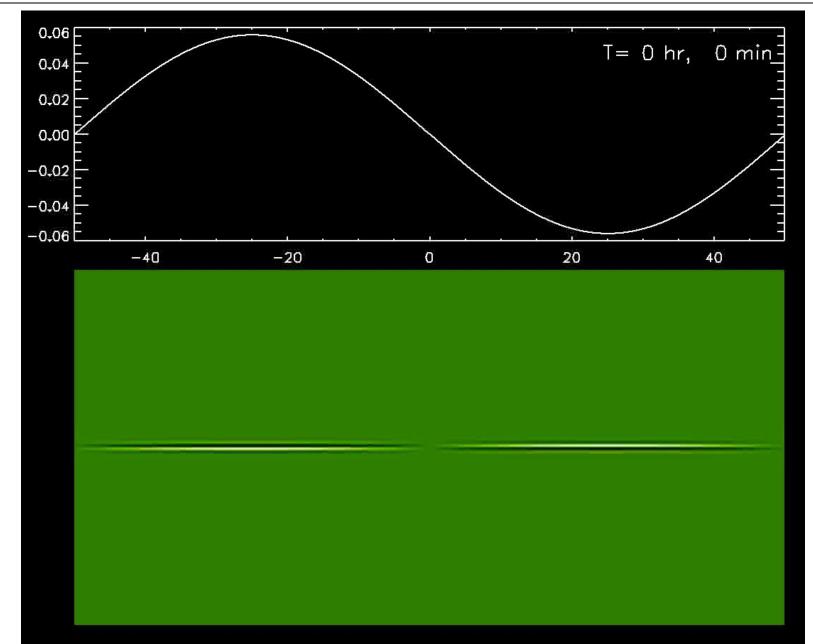
Results: Viscosity = $100 \text{ m}^2/\text{s}$



Results : Non-zero background stability, $v = 100 \text{ m}^2/\text{s}$



Results : Velocity Duct (high viscosity)

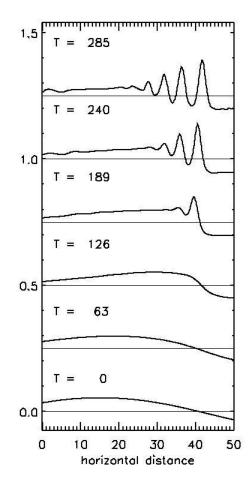


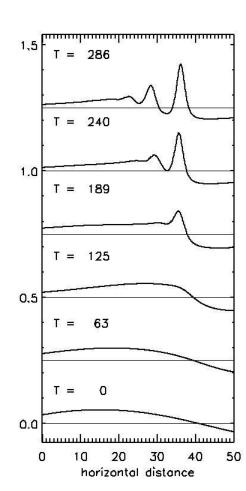
Results : Headwind

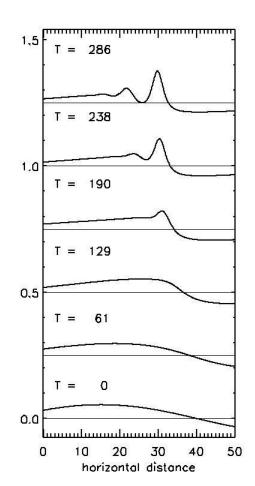
Co-located thermal and doppler duct

 $\sim -U_0 \cos(\pi z/2) \sim -U_0 \cos(\pi z/10)$

 $l_{s}^{2} = \frac{N^{2}(z)}{(U(z) - c)^{2}} - \frac{\partial_{zz}U(z)}{(u_{0} - c)}$



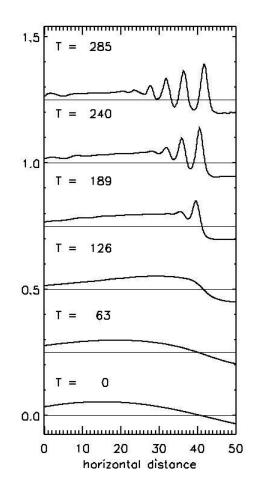


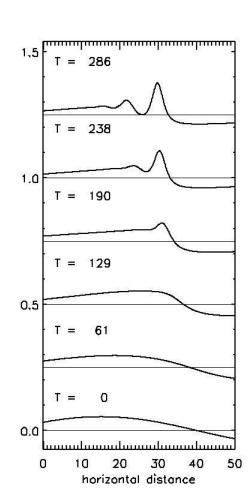


Results : Tailwind

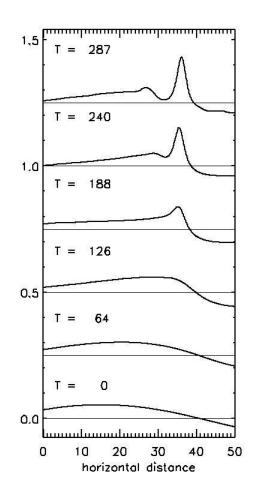
Co-located thermal and doppler duct

 $\sim -U_0 \cos(\pi z/2)$





$$\sim U_0 \cos(\pi z/2)$$



Conclusions

- Bore evolution dependence on imposed forcing
- Bore evolution dependence on ducting environment

Future Work

- Ultimately, extension to 3D, breaking bores, turbulence
- Temporally varying ducts
- Add complexity to BDO
- Pose more realistic forcings

