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The Whole Atmosphere Community Climate Model (WACCM)

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OUTLINE



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- Scientific motivation
- Model development
 - Physical processes
 - Chemistry
- Some initial results
 - Mesospheric thermal inversions
 - Tides in the mesosphere and lower thermosphere
 - The 2-day wave
 - Water vapor in the lower stratosphere
 - The winter stratosphere and the "annular mode"
- The WACCM Workshop



WACCM Motivation

Roble, Geophysical Monographs, 123, 53, 2000



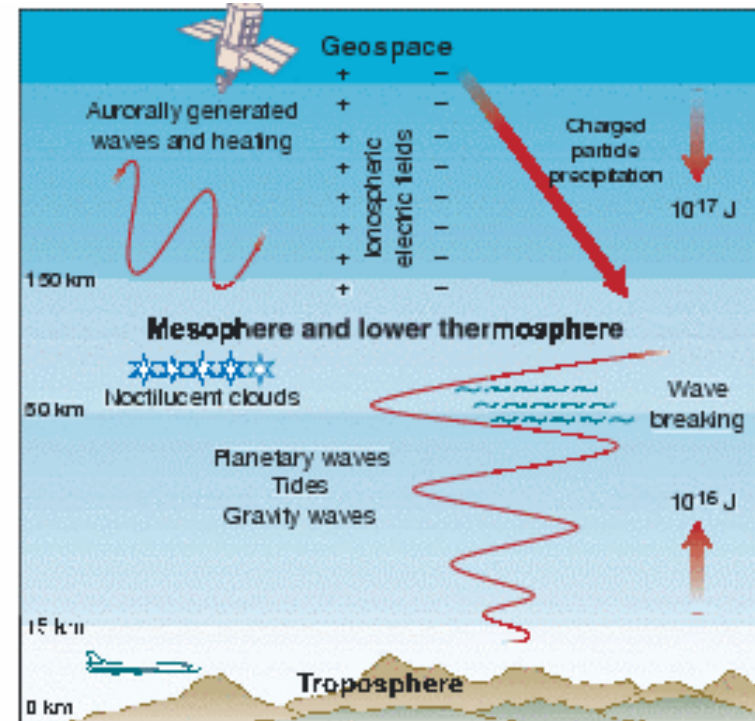
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- Coupling between atmospheric layers:

- Waves transport energy and momentum from the lower atmosphere to drive the QBO, SAO, sudden warmings, mean meridional circulation
- Solar inputs, e.g., auroral production of NO in the mesosphere and downward transport to the stratosphere
- Stratosphere-troposphere exchange

- Climate Variability and Climate Change:

- What is the impact of the stratosphere on tropospheric variability, e.g., the Arctic oscillation or “annular mode”?
- How important is coupling among radiation, chemistry, and circulation? (e.g., in the response to O₃ depletion or CO₂ increase)



Energy transfer in the mesosphere and lower thermosphere. About 10^{16} J of energy propagates up daily from the atmosphere below in the form of waves and tides. During a geomagnetic storm (which occurs about every 5 days), about 10^{17} J is injected per day from space through auroral processes.

Jarvis, “Bridging the Atmospheric Divide”
Science, **293**, 2218, 2001

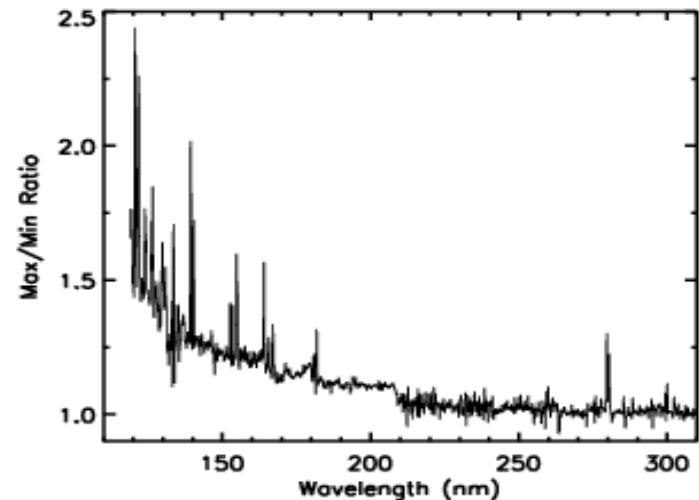
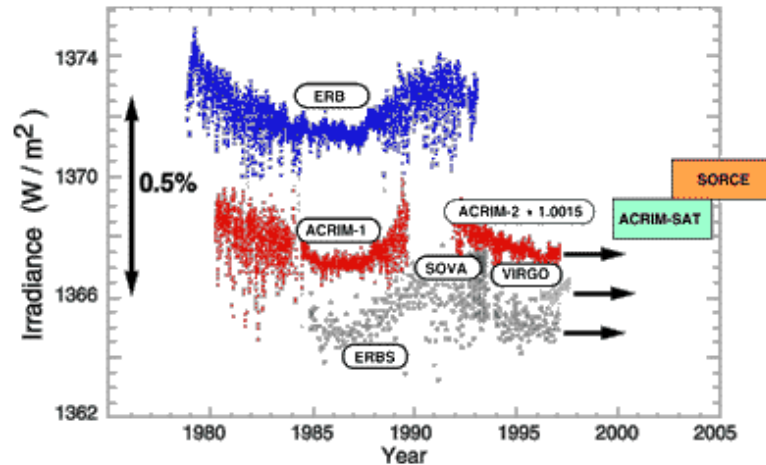


WACCM Motivation



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- Response to Solar Variability:
 - Recent satellite observations have shown that solar cycle variation is:
 - 0.1% for total Solar Irradiance
 - 5-10% at $\approx 200\text{nm}$
 - Radiation at wavelengths near 200 nm is absorbed in the stratosphere
 - => Impacts on global climate may be mediated by stratospheric chemistry and dynamics
- Satellite observations:
 - There are several satellite programs that can benefit from a comprehensive model to help interpret observations
 - e.g., UARS, TIMED, EOS Aura





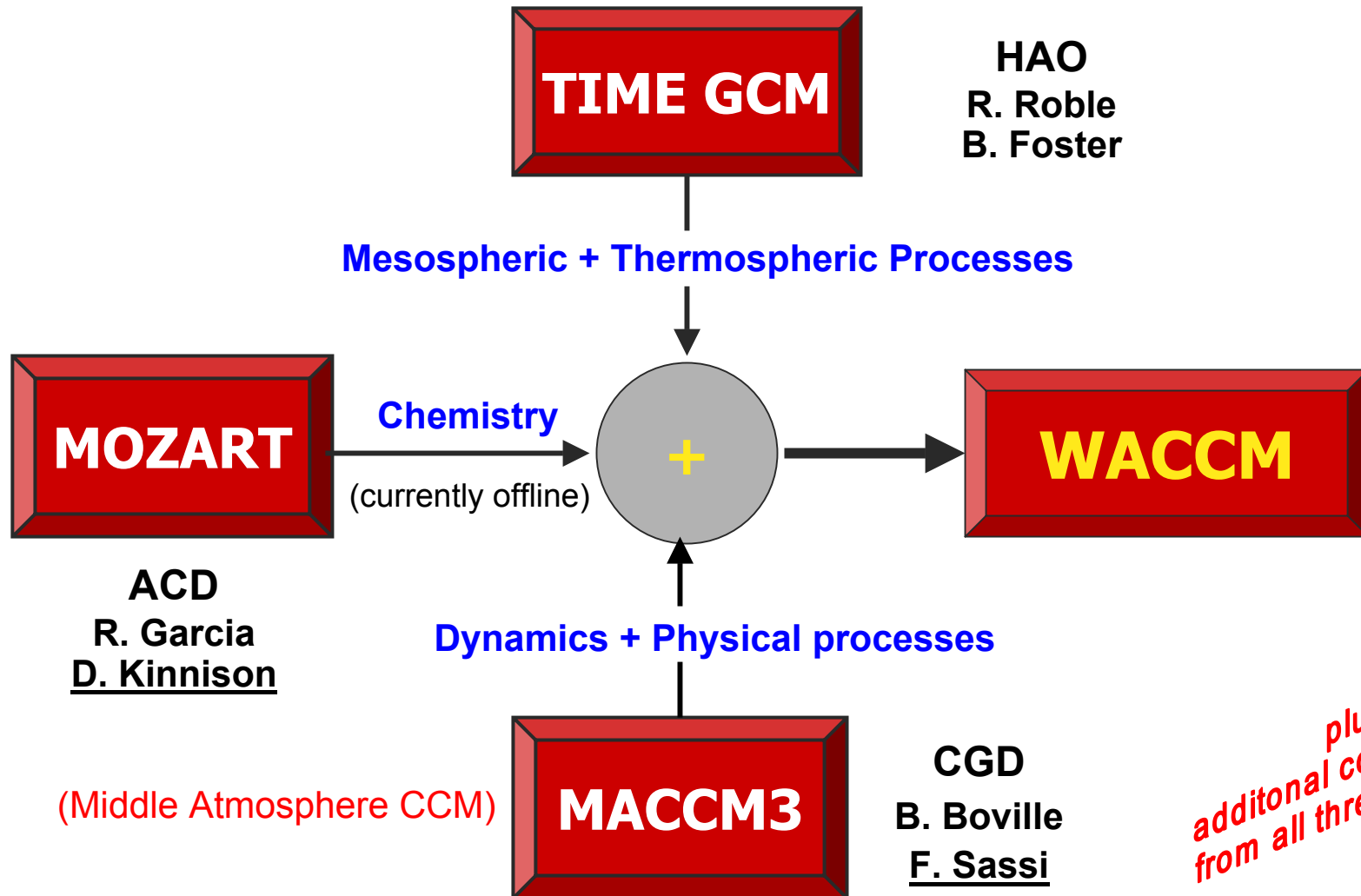
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WACCM Components: A collaboration among 3 NCAR Divisions



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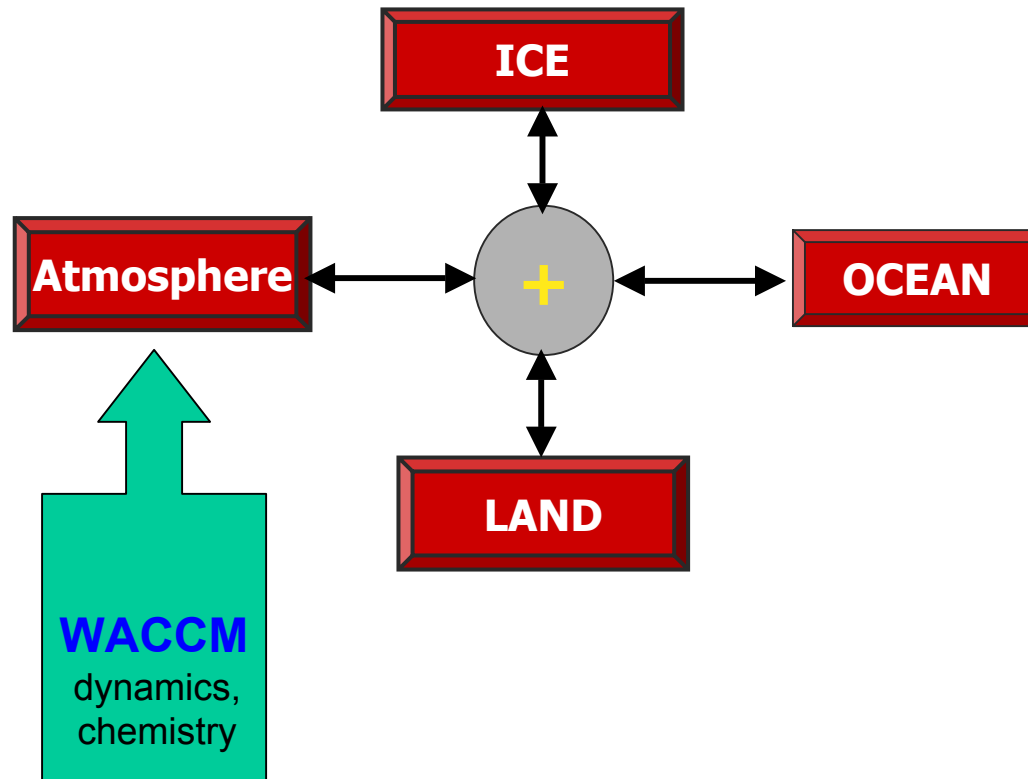




WACCM and the NCAR Community Climate System Model



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WACCM uses the *software framework of the NCAR CCSM*. May be run in place of the standard CAM (Community Atmospheric Model)



Additions to the original MACCM3 code:

- A parameterization of non-LTE IR (15 μm band of CO_2 above 70 km) merged with CCSM IR parameterization (below 70 km)
- Short wave heating rates (above 70 km) due to absorption of radiation shortward of 200 nm and chemical potential heating
- Gravity Wave parameterization extended upward, includes dissipation by molecular viscosity
- Effects of dissipation of momentum and heat by molecular viscosity (dominant above 100 km)
- Diffusive separation of atmospheric constituents above about 90 km
- Simplified parameterization of ion drag

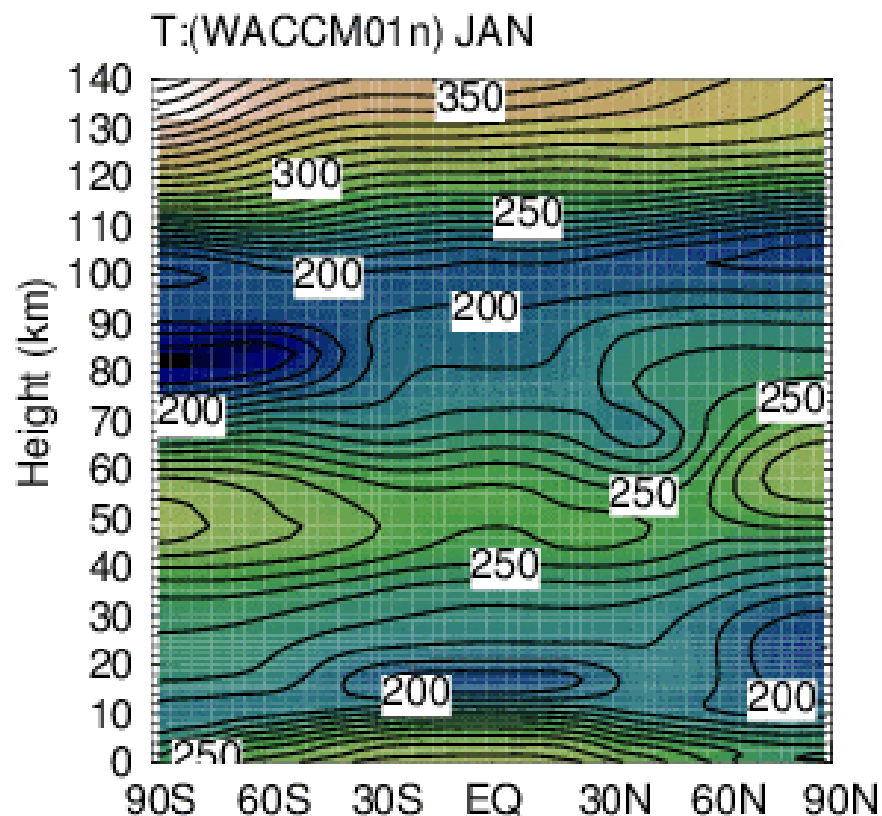
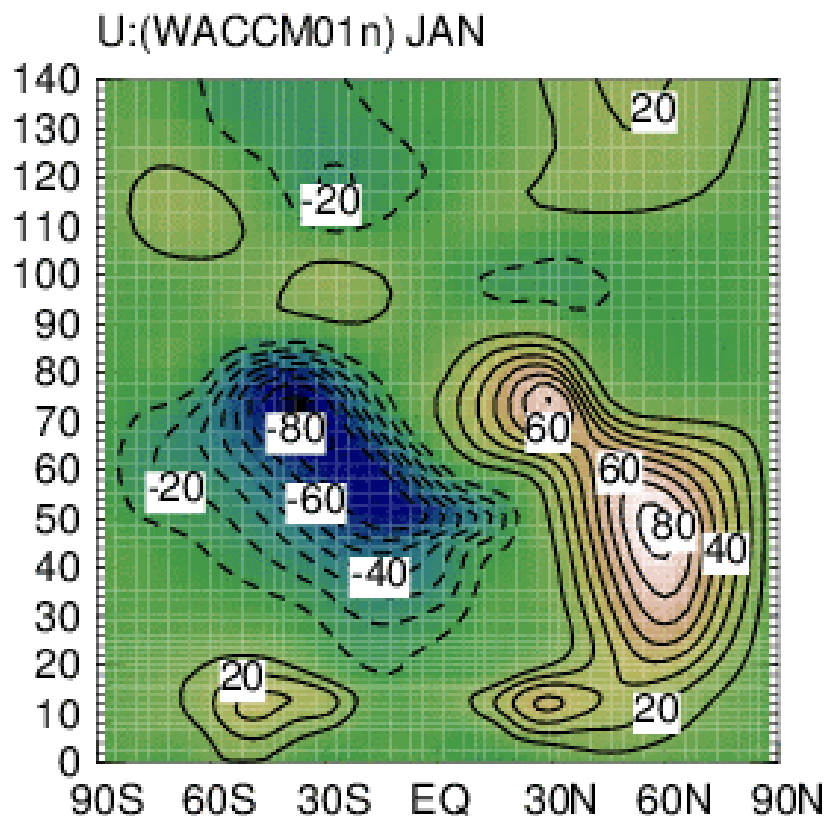


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Zonal Mean Wind and Temperature



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Gross diagnostics (zonal mean behavior)
Complete climatological analysis is planned

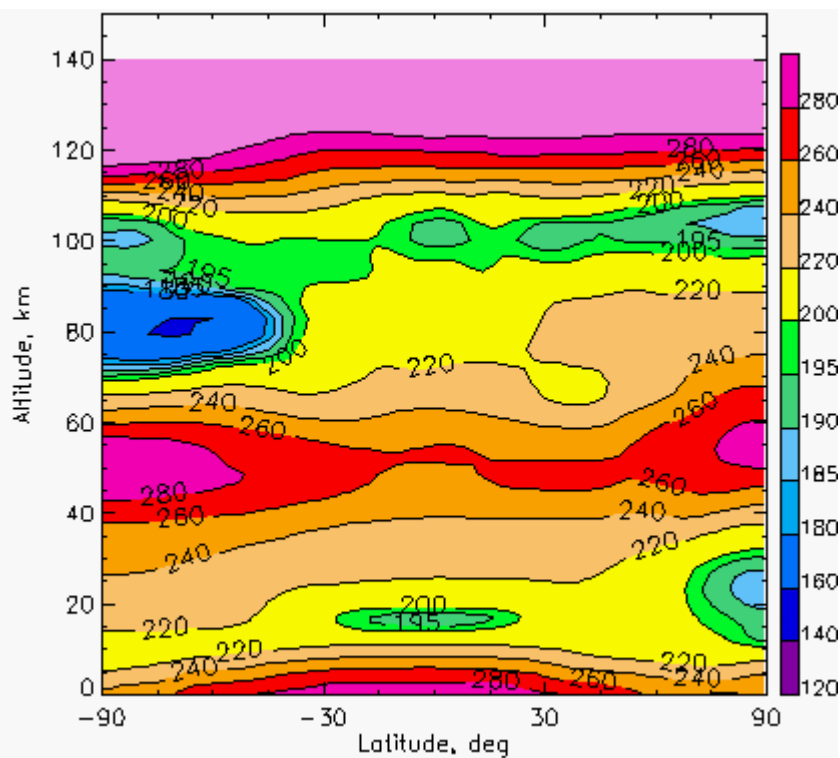


Solstice Temperature Distribution

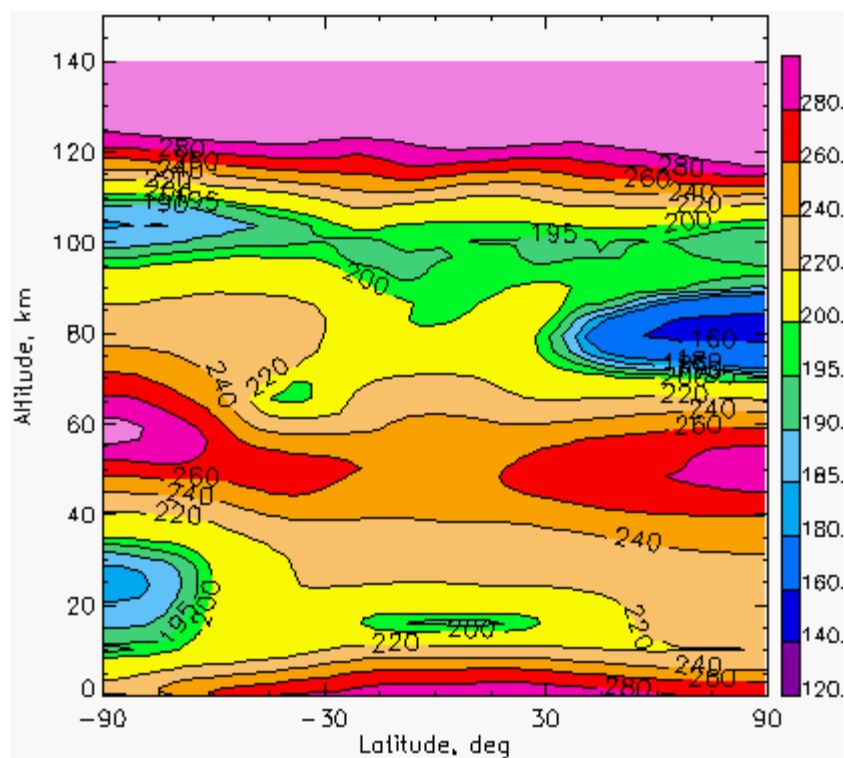


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January



July





Chemistry Module (MOZART-3)

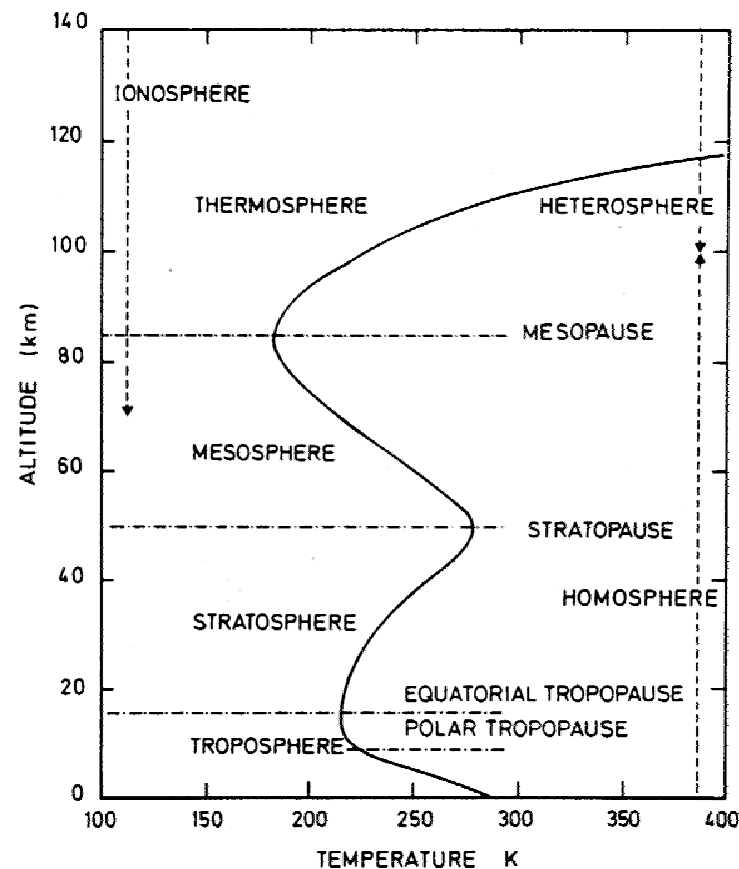
(50 species; 41 Photolysis, 93 Gas Phase, 17 Hetero. Rx)



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Our goal was to represent the chemical processes considered important in the:

- Troposphere, Stratosphere, and Mesosphere:
 - O_x , HO_x , NO_x , ClO_x , and BrO_x
 - Heterogeneous processes on sulfate, nitric acid hydrates, and water-ice aerosols
- Thermosphere (limited):
 - Auroral NO_x production
 - Currently do not include ion-molecule reactions



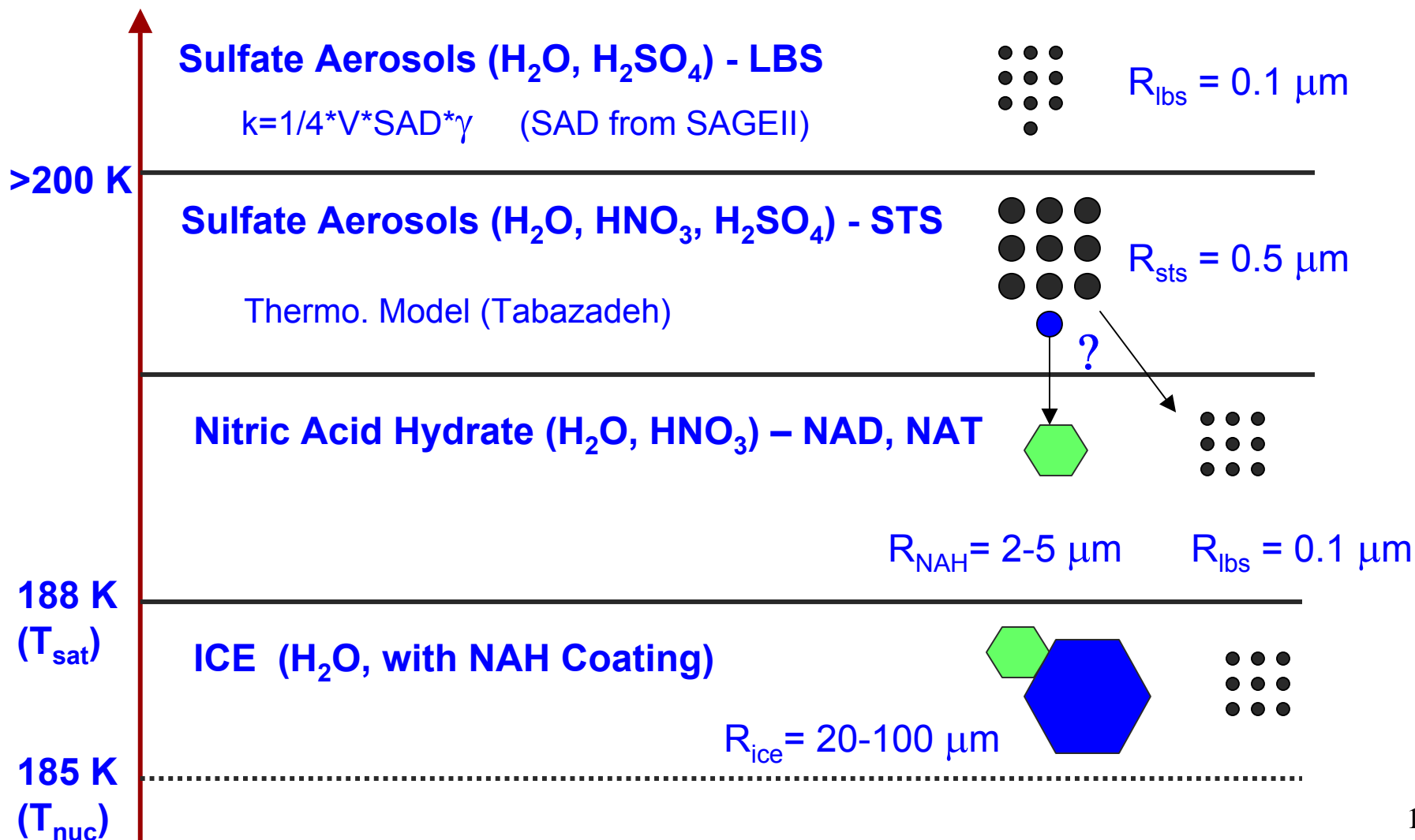


Long-lived Species: (17-species, 1-constant)

- Misc: CO_2 , CO , CH_4 , H_2O , N_2O , H_2 , O_2
- CFCs: CCl_4 , CFC-11, CFC-12, CFC-113
- HCFCs: HCFC-22
- Chlorocarbons: CH_3Cl , CH_3CCl_3 ,
- Bromocarbons: CH_3Br
- Halons: H-1211, H-1301
- Constant Species: N_2

Short-lived Species: (32-species)

- O_x : O_3 , O , $\text{O}(^1\text{D})$
- NO_x : N , $\text{N}(^2\text{D})$, NO , NO_2 , NO_3 , N_2O_5 , HNO_3 , HO_2NO_2
- ClO_x : Cl , ClO , Cl_2O_2 , OCIO , HOCl , HCl , ClONO_2 , Cl_2
- BrO_x : Br , BrO , HOBr , HBr , BrCl , BrONO_2
- HO_x : H , OH , HO_2 , H_2O_2
- HC Species: CH_2O , CH_3O_2 , CH_3OOH



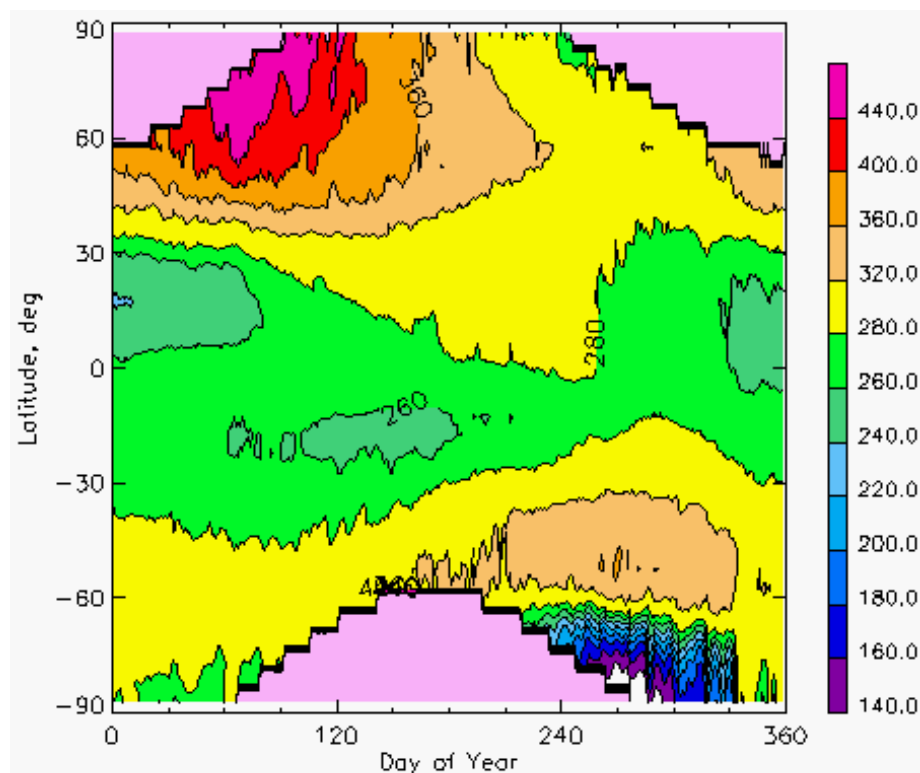


Total Column Ozone (Dobson Units)

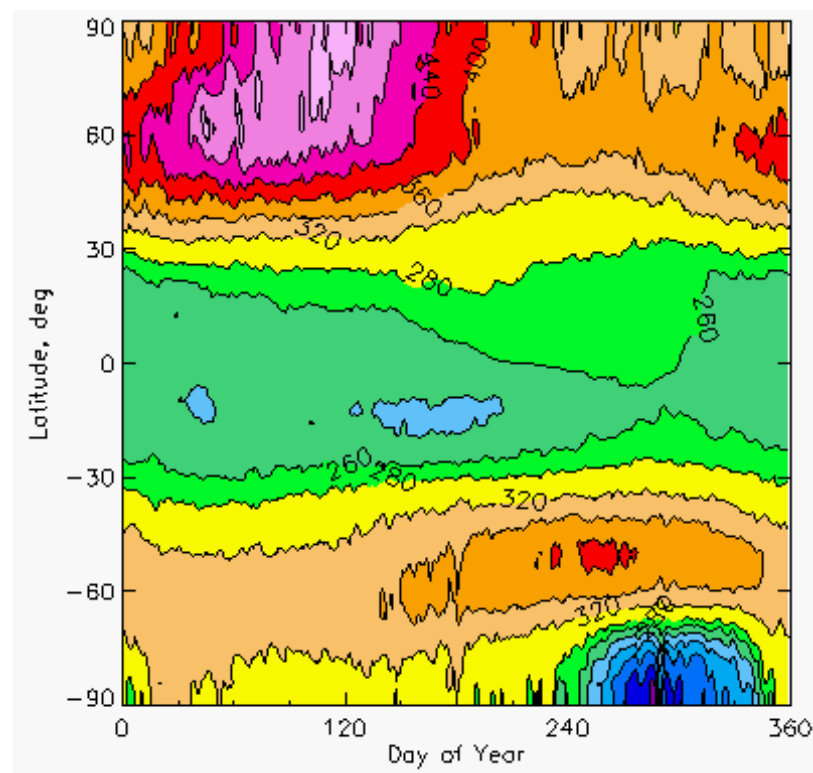


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Earth Probe TOMS, 1999 (daily)



WACCM (daily)





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Vertical T profiles during inversions: Observed and modeled



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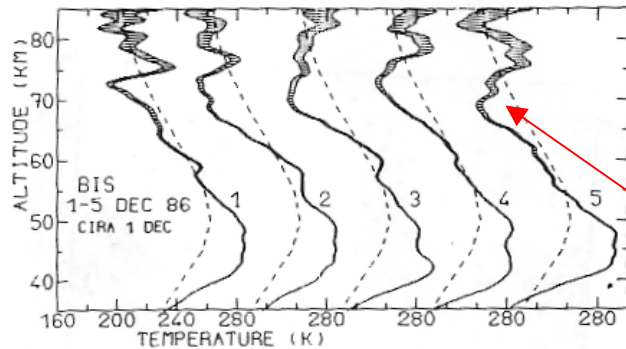
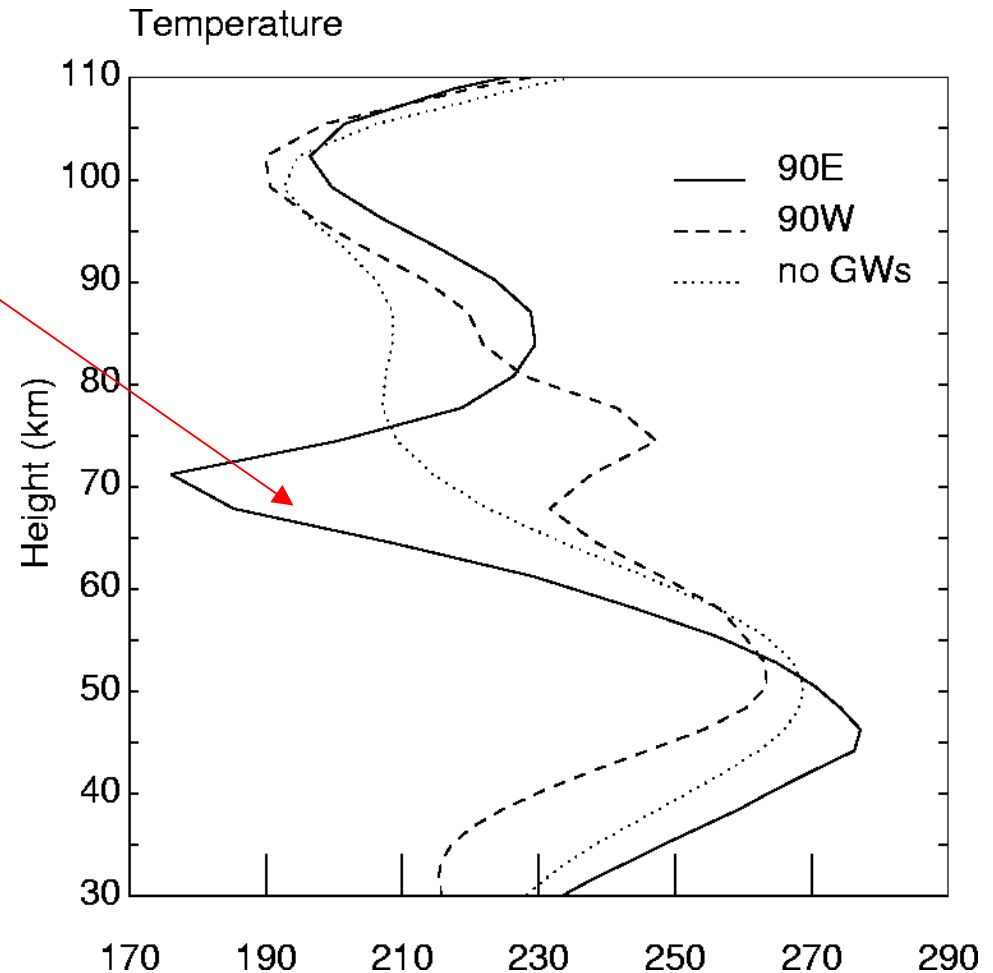


Fig. 1 Nightly mean temperature profiles at Biscarrosse from December 1 to December 5, 1986. The error bars ($\pm 1 \sigma$) are indicated by the shaded area. The CIRA 1972 profile is shown for comparison (dotted line). Perturbations with vertical wavelengths shorter than 1.5 km have been filtered.





Wintertime mesospheric inversions: Observed and modeled



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Hauchecorne et al, JGR, 1987

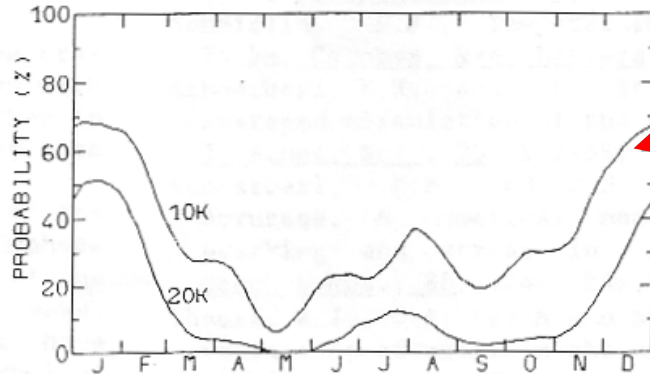


Fig. 6 Annual variation of the probability of occurrence of a mesospheric inversion for the 10 K and 20 K criteria. A ± 30 days triangular filter has been applied to smooth the data.

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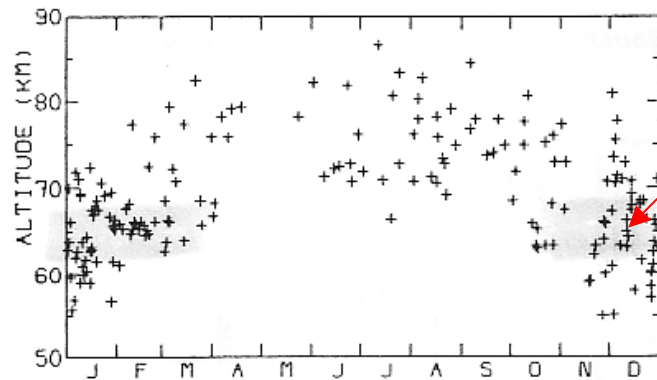
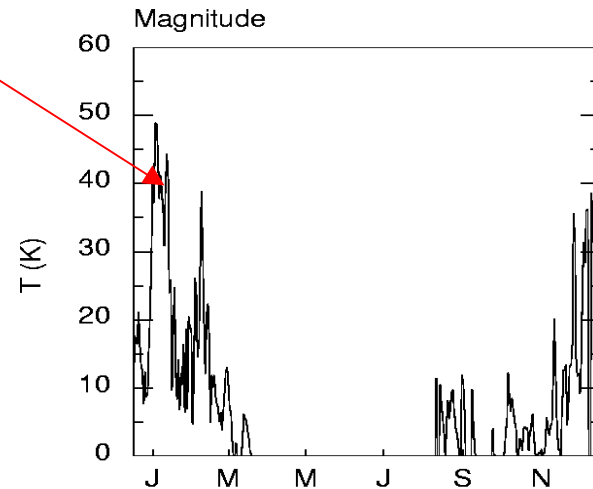
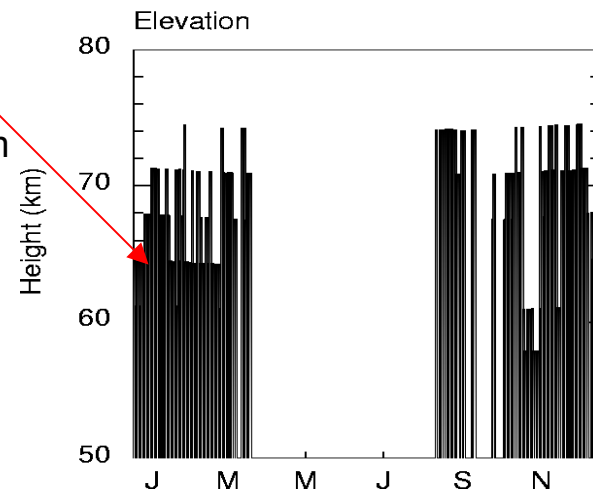


Fig. 5 Altitude of the secondary minimum of temperature versus the day of the year for all OHP profiles from June 1981 to September 1986 for which a temperature inversion has been detected (10 K criterion).





Mean zonal wind and planetary wave breaking: Standard run with GW parameterization

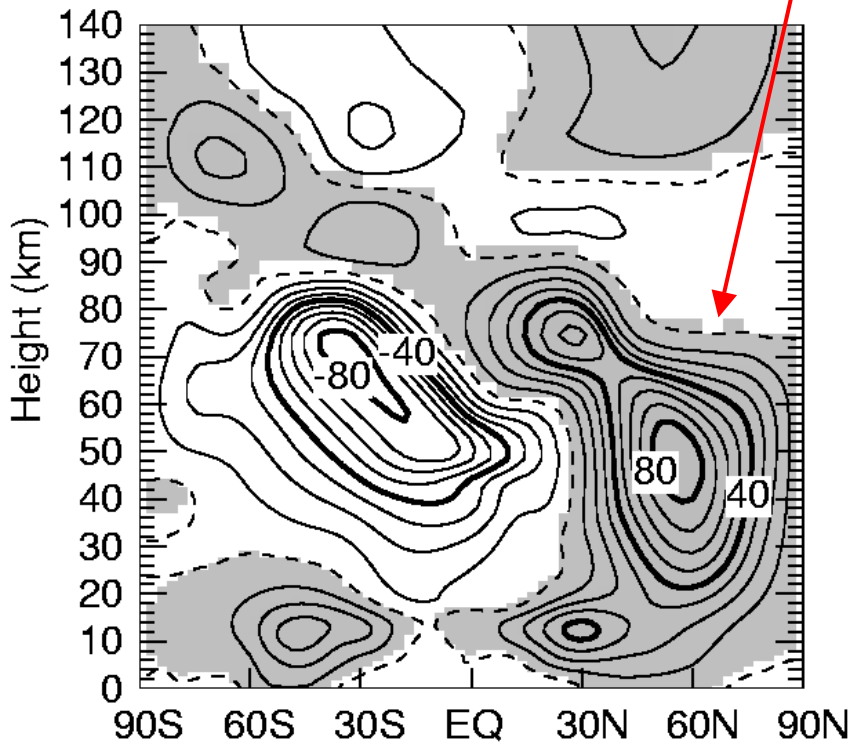


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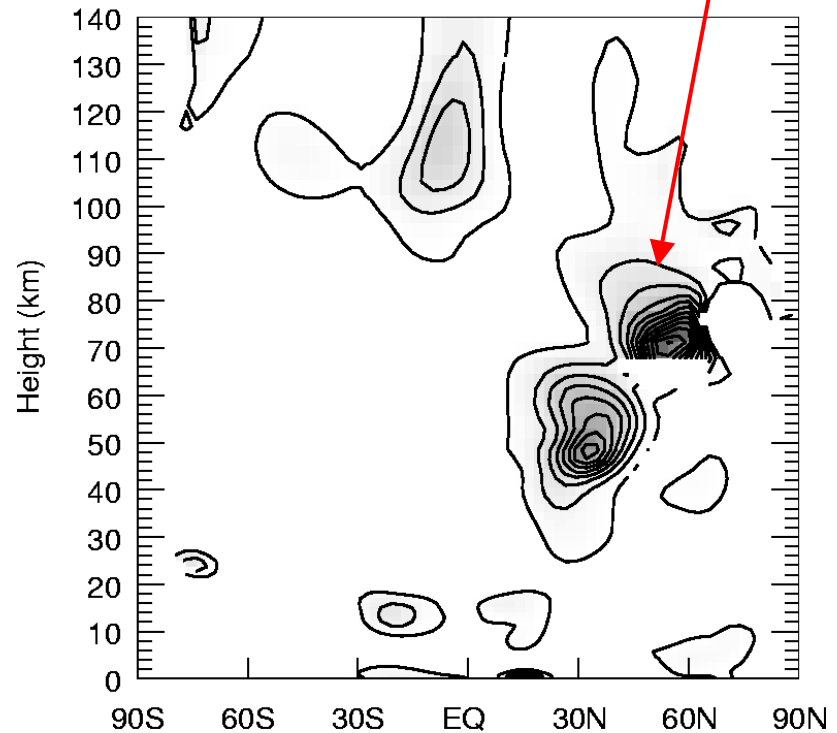
wind reversal,
zero-wind line
near 75 km

note "compact"
mesospheric surf-zone

a) Zonal Wind



Kyy : with parameterized GWs

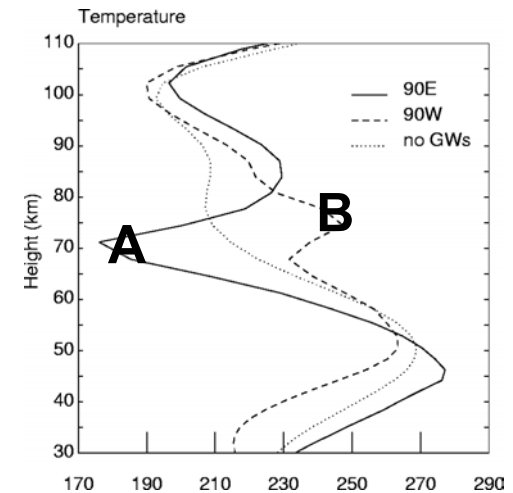
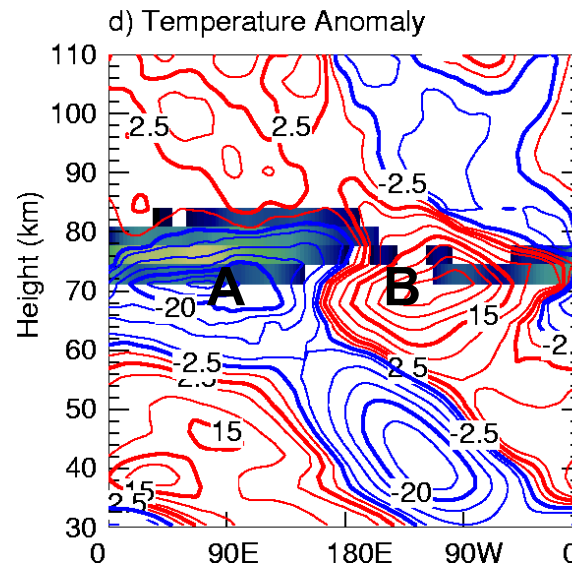
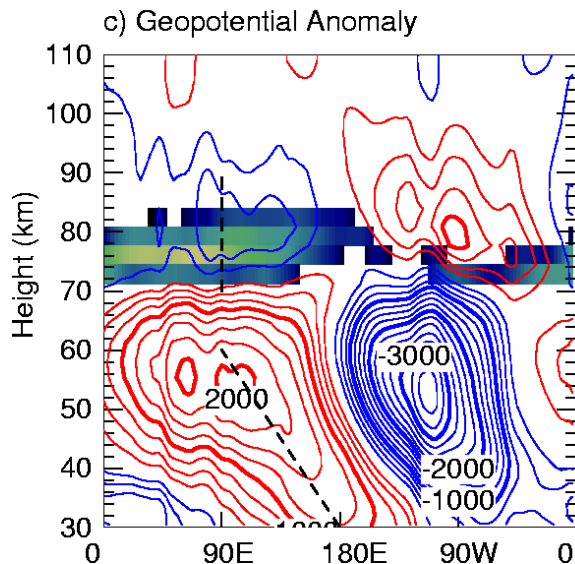
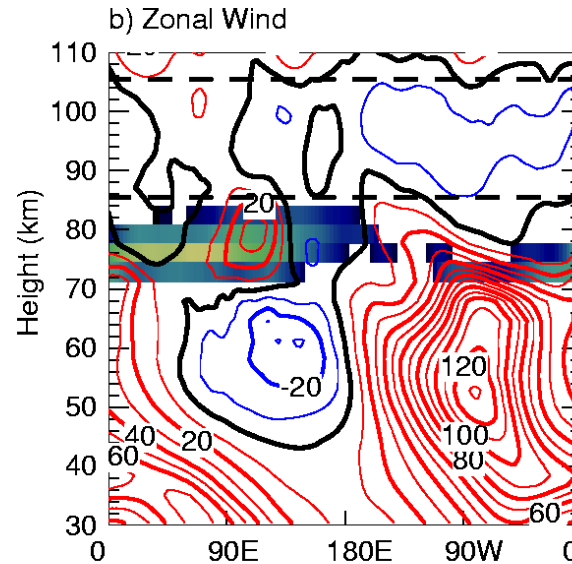
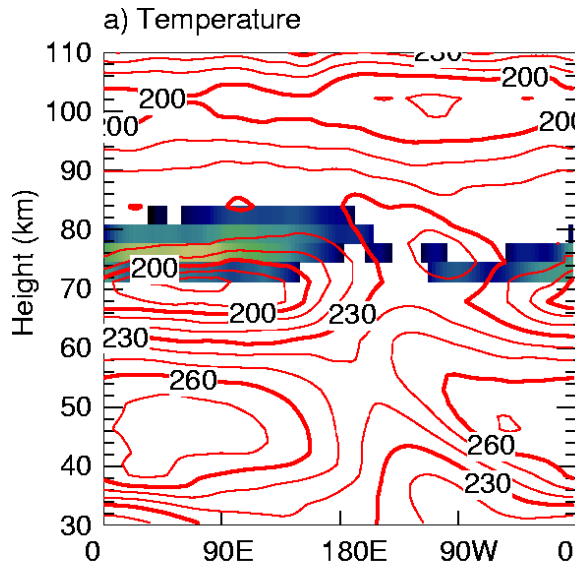




Structure of wintertime T inversions simulated with WACCM



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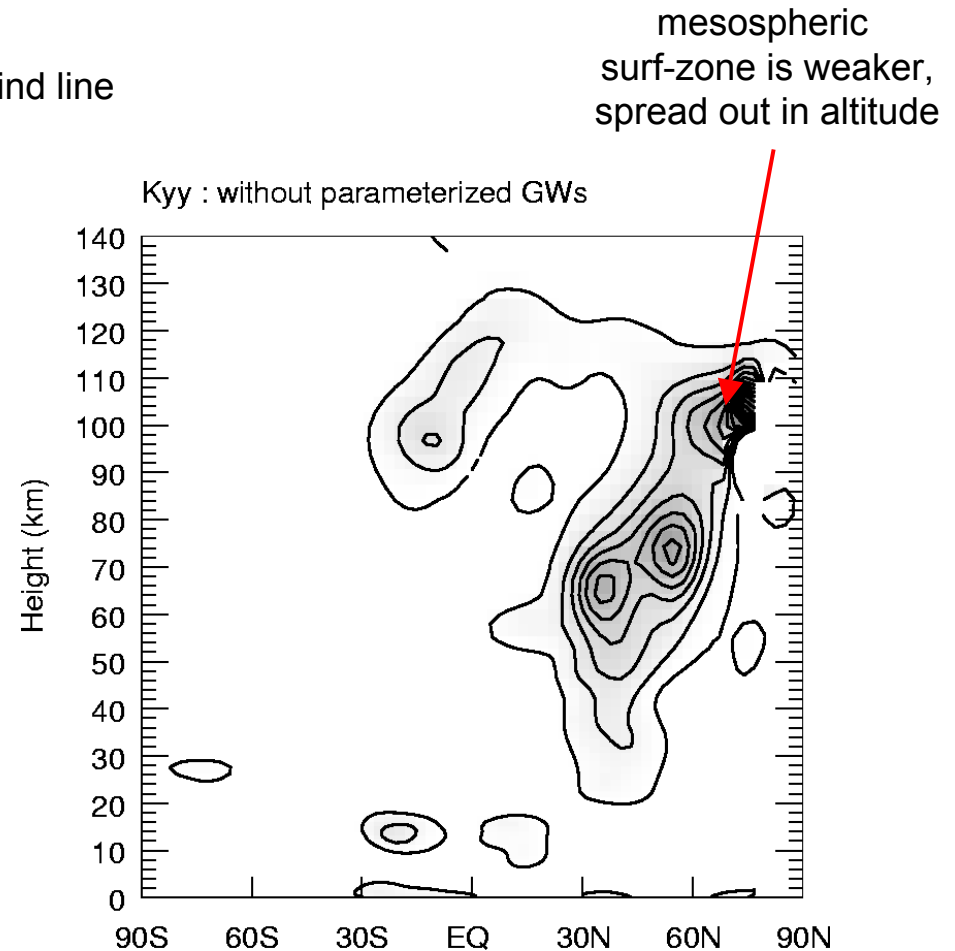
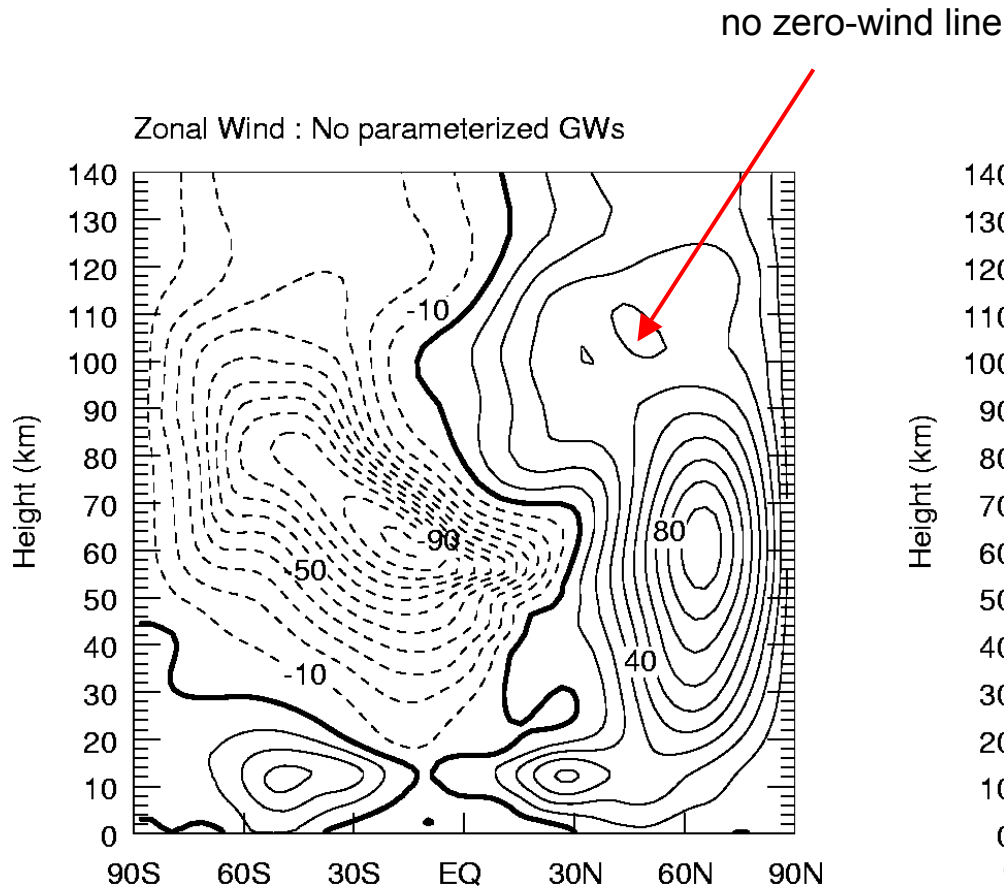


Effects of removing gravity wave drag: Zonal wind and planetary wave breaking



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**Model run with Rayleigh friction
Replacing gravity wave drag**



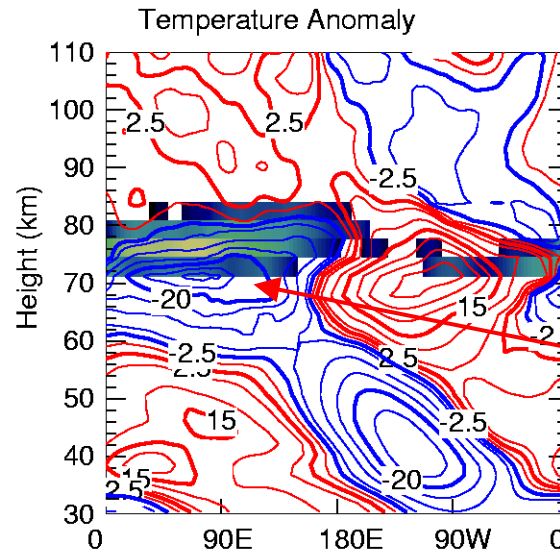
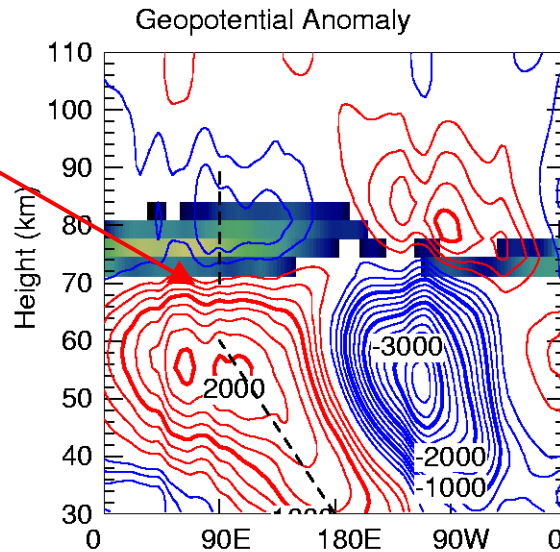


Effects of removing gravity wave drag: Planetary wave structure



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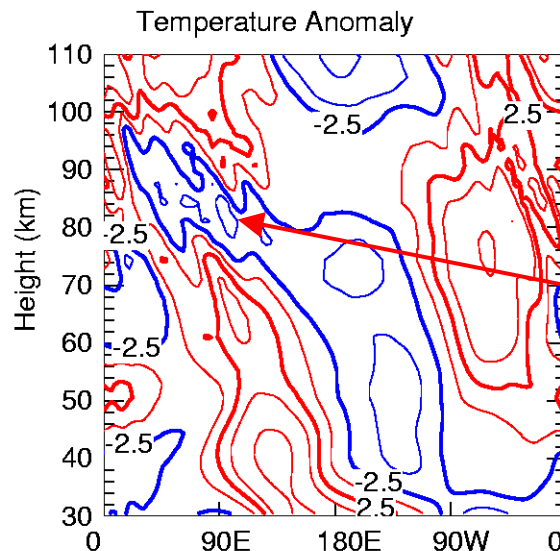
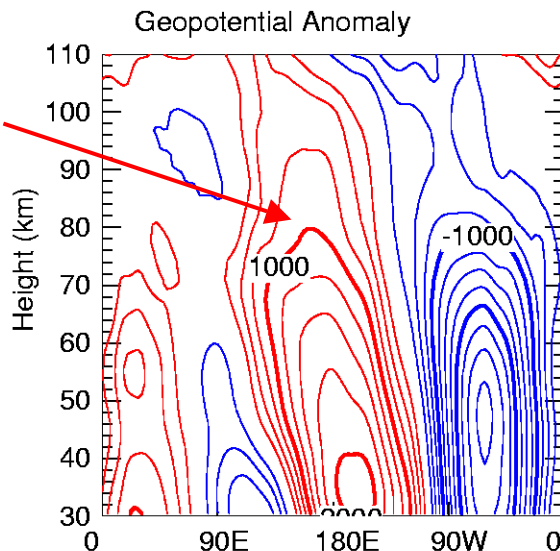
rapid decrease
of geopotential
with height



Standard model
run with GWs

large T'

gradual decrease
of geopotential
with height



Model run with
Rayleigh friction

smaller T'



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Tides: standard case



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Model:

- Forcing via shortwave heating (O_3 , H_2O , O_2) and convective heating
- Gravity wave parameterization
- Comparisons with observations

Analysis method:

- 3-hourly output saved for successive 5-day periods throughout year
- Extract wavenumber, frequency components via Fourier analysis ($k = 0, 6$ and $f = \pm 3$ cpd)

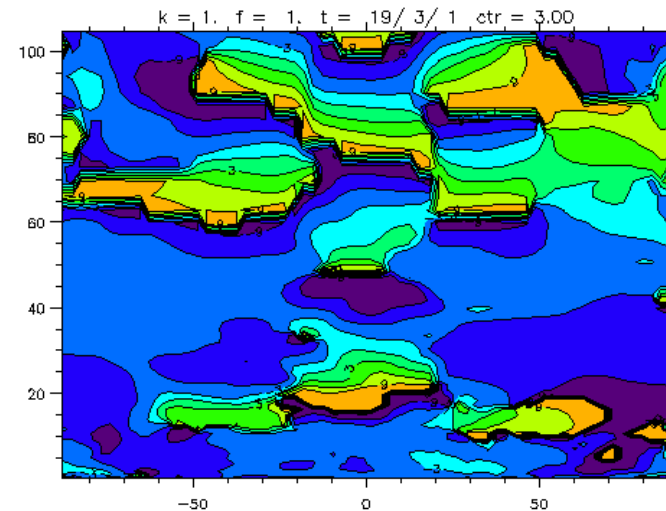
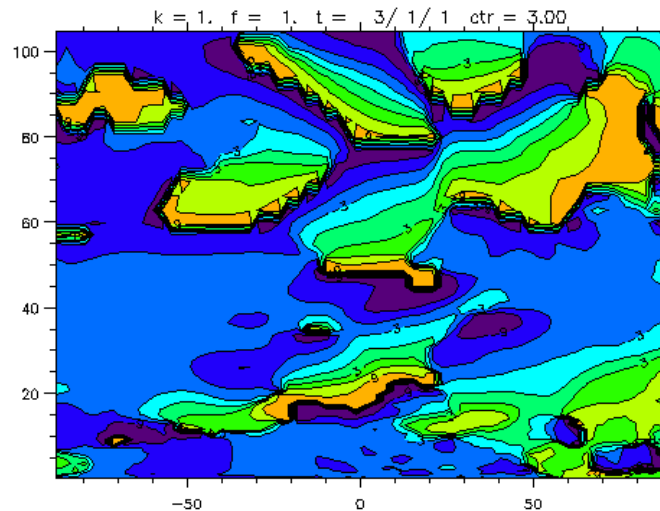
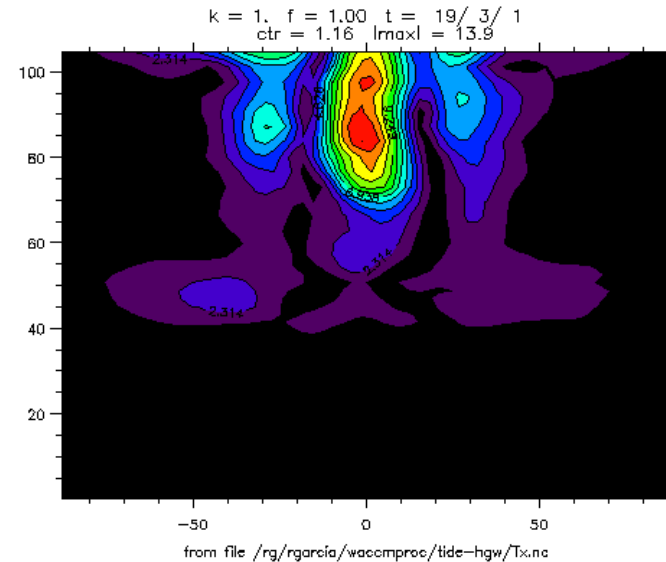
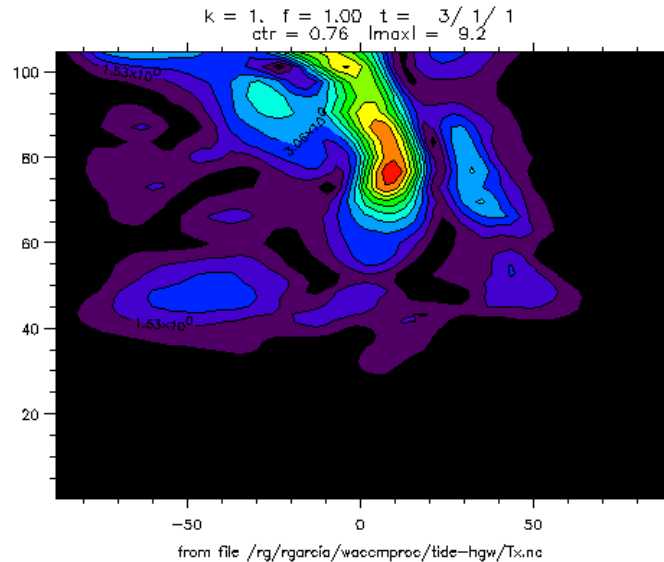


Diurnal T tide (k=1 westward): January and March



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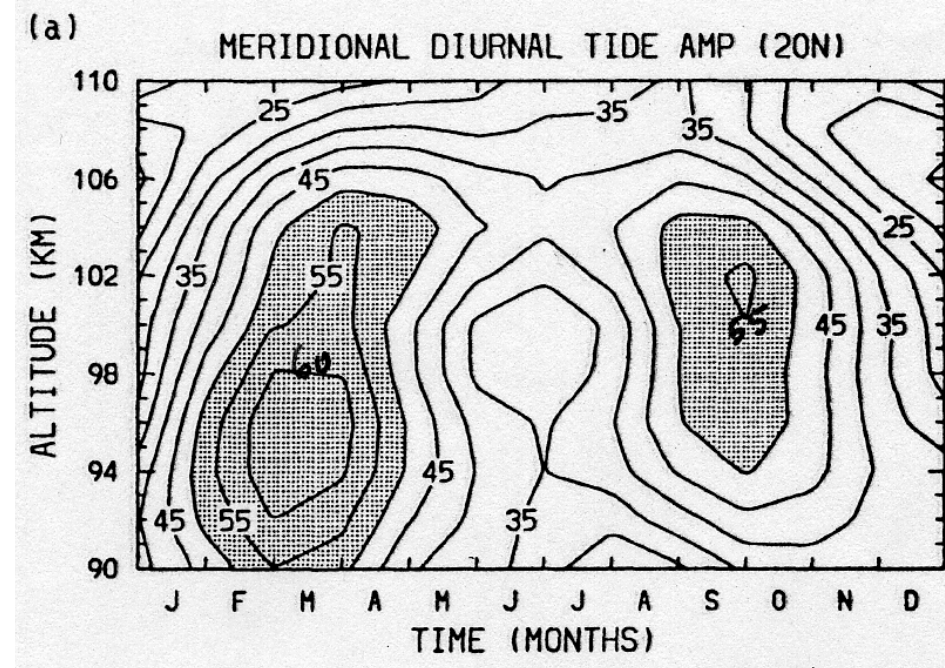
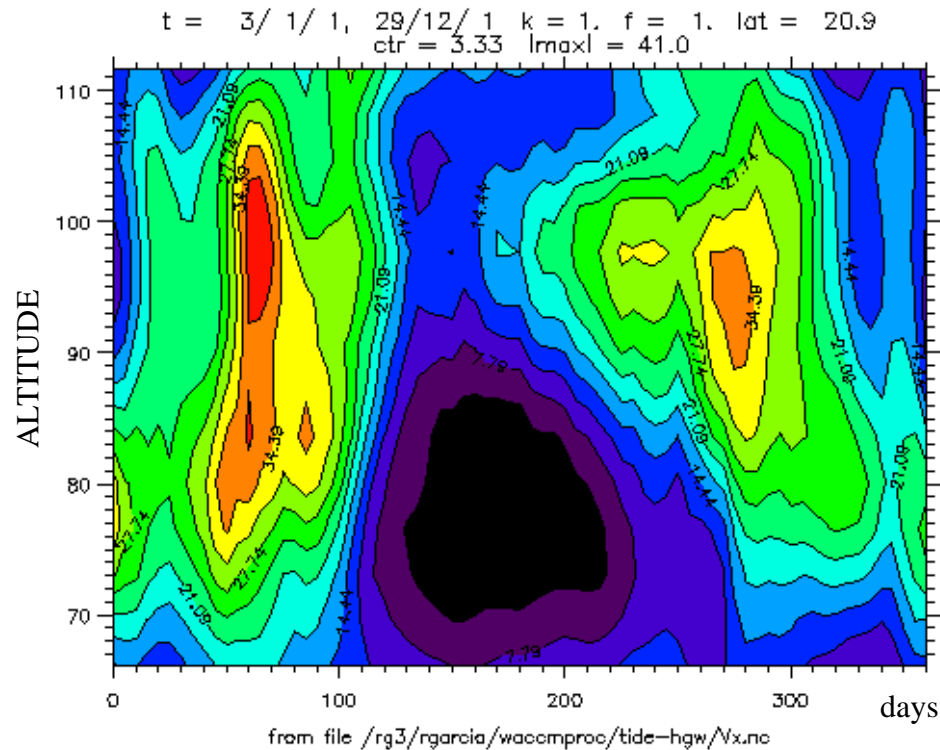
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Diurnal V tide - annual cycle at 20 N



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V from UARS/WINDI
(McLandress et al, 1996)

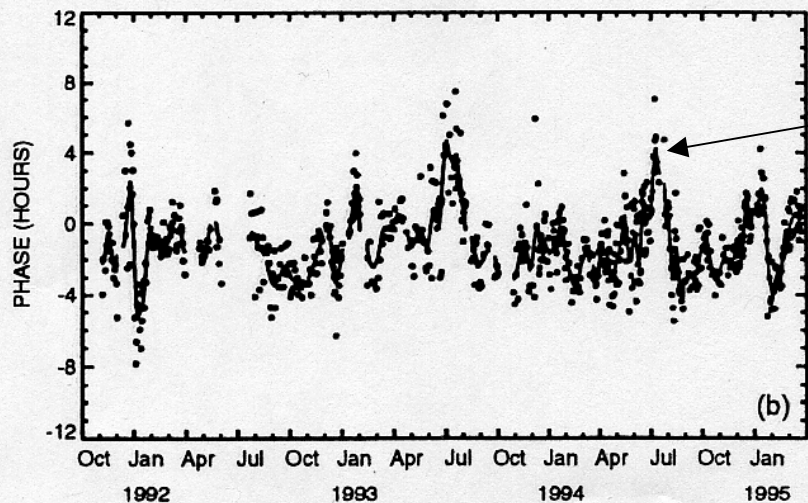
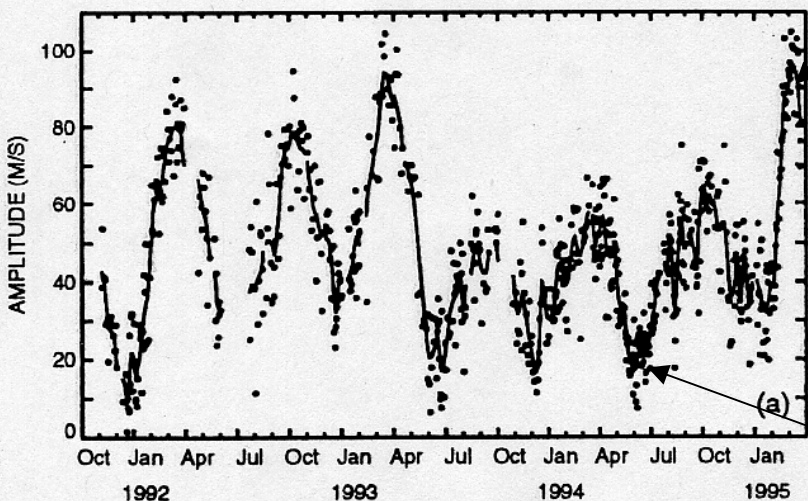


Diurnal V tide – annual cycle at 20N, 95 km



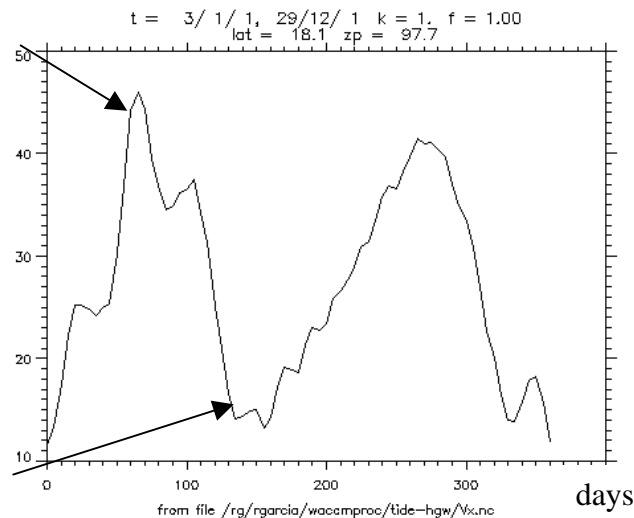
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UARS/HRDI V
At 95 km, 20 N
(Burrage et al., 1995)



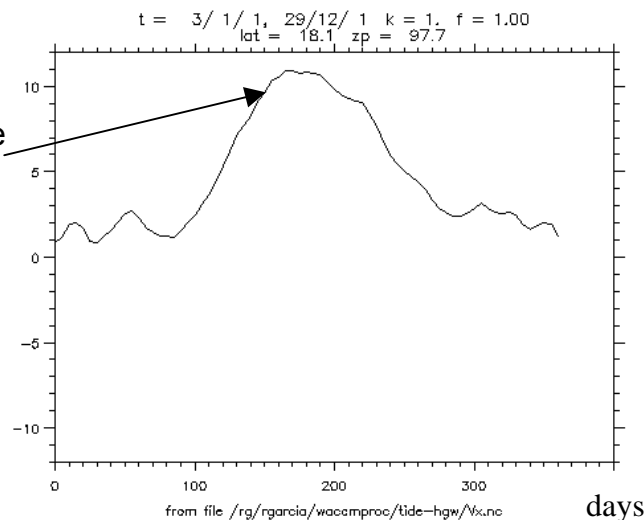
WACCM V
At 97 km, 18 N

maxima at equinoxes



deep minima at solstices

phase change at summer solstice



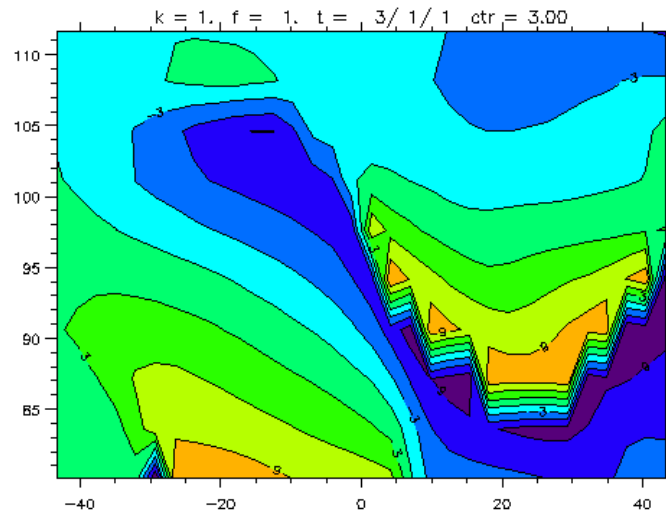
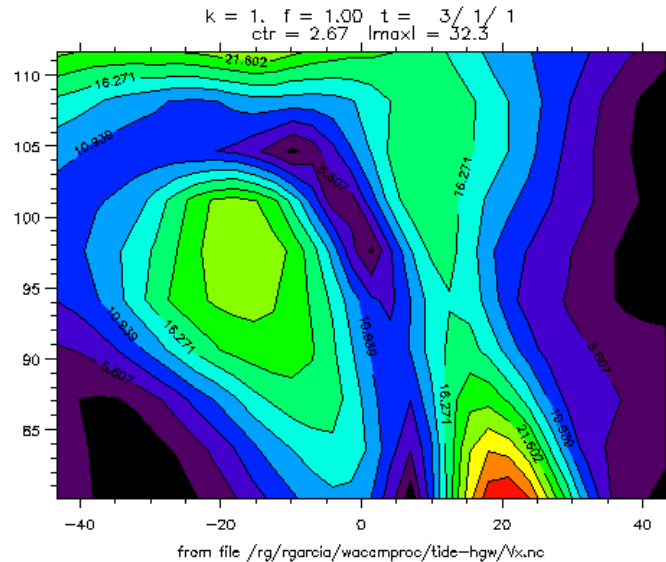


Diurnal V tide structure - solstice

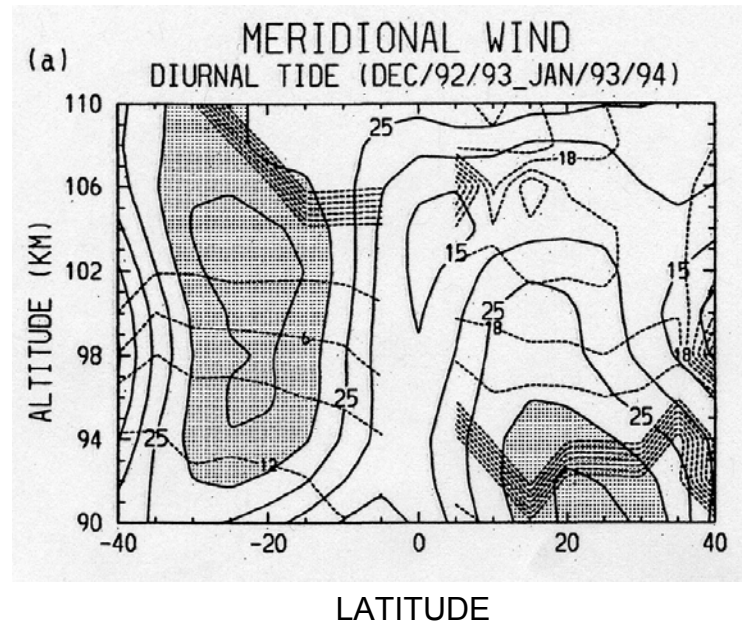


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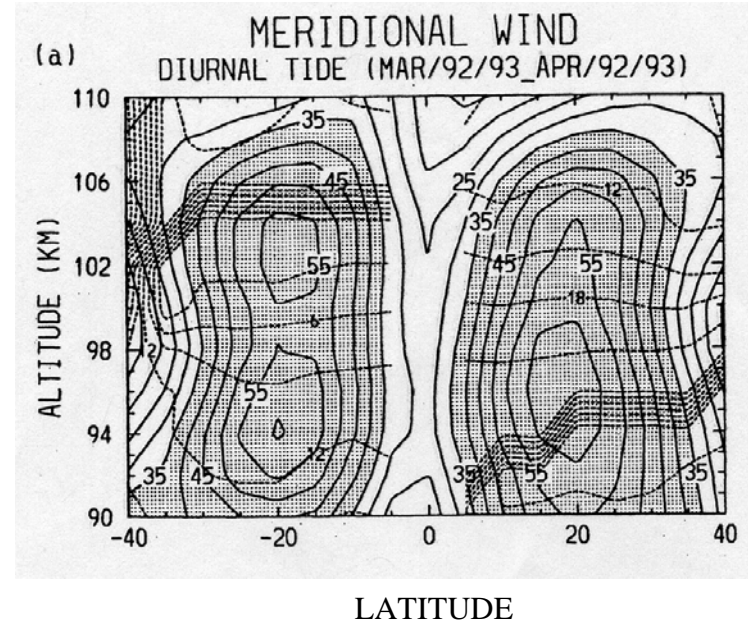
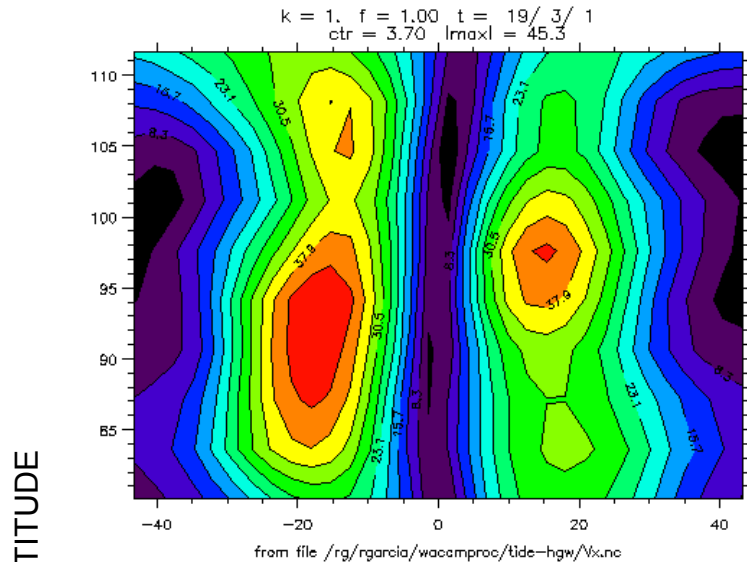
V from UARS/WINDI
(McLandress et al, 1996)



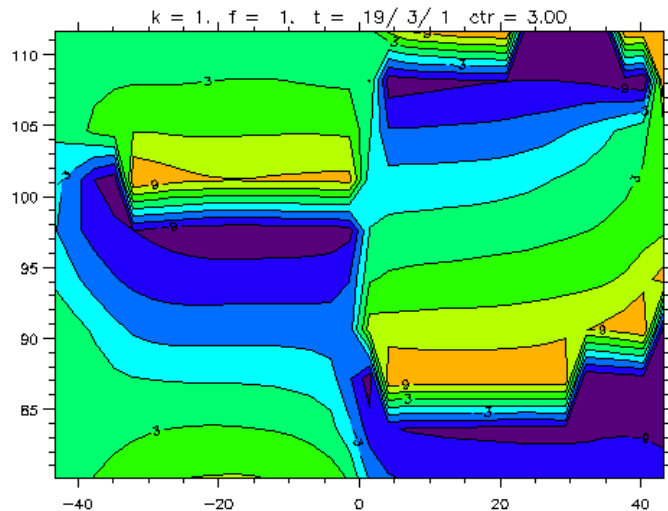
Diurnal V tide structure - equinox



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V from UARS/WINDI
(McLandress et al, 1996)

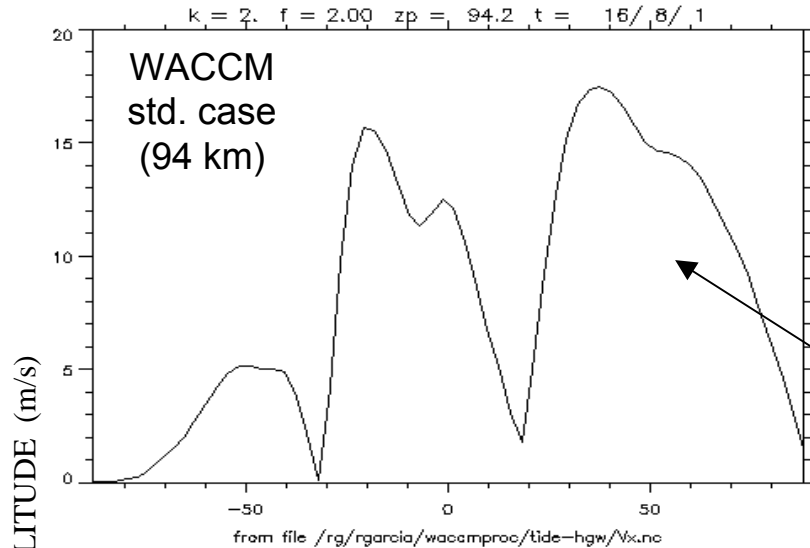




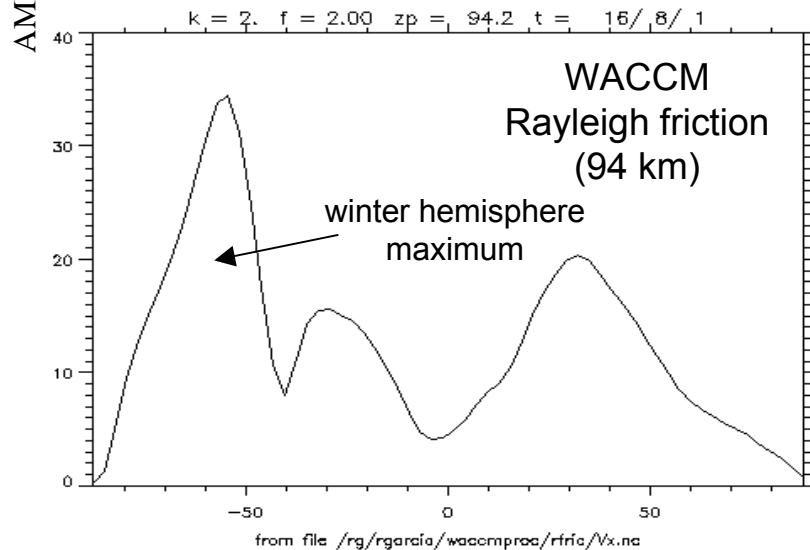
Semidiurnal V Tide: Effects of dissipation mechanisms (GW drag vs. Rayleigh friction)



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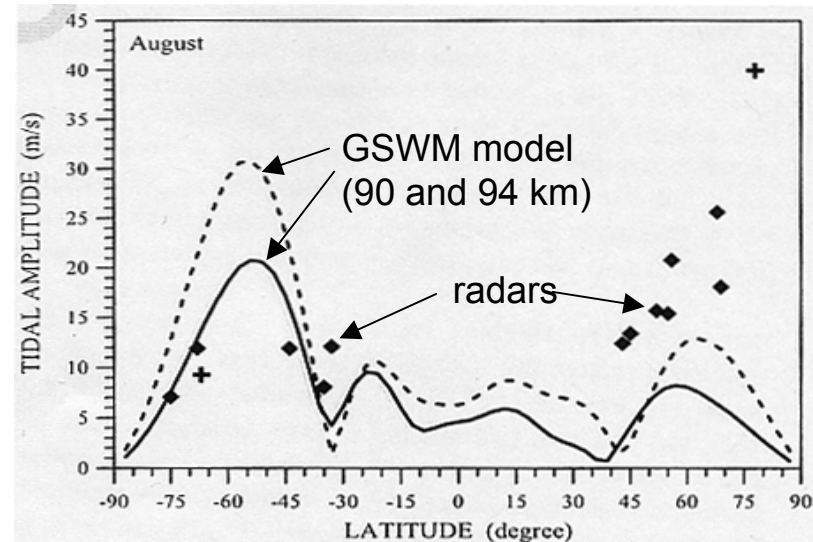
summer hemisphere
maximum



winter hemisphere
maximum

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Semidiurnal V Tide
Latitudinal Structure
WACCM vs. Pancheva et al's
(JASTP, 2002) radar data (August)





Tidal contributions due to shortwave and convective heating

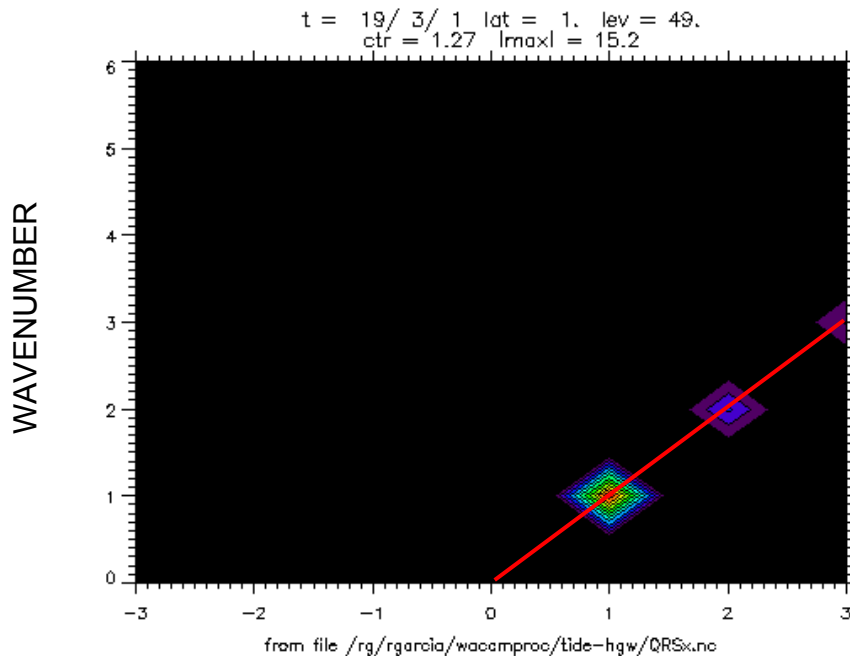


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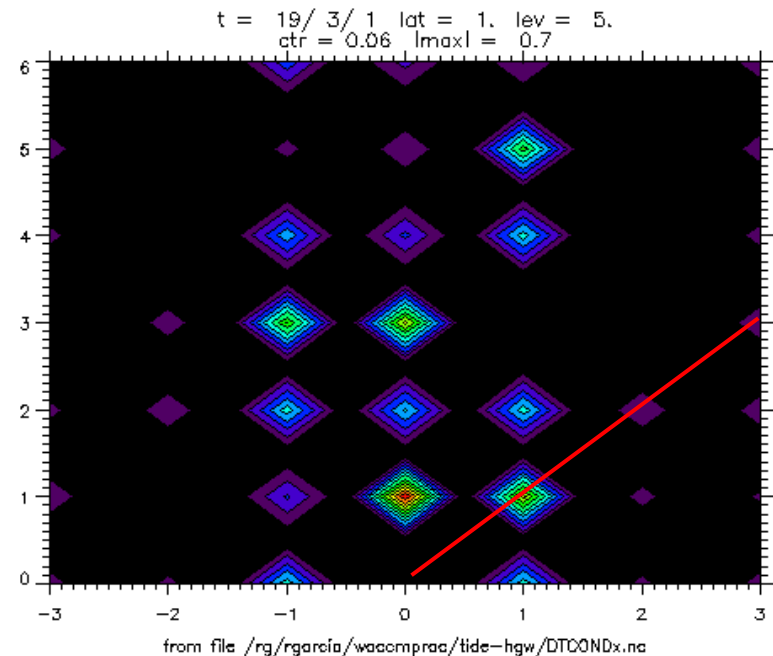
In the model,

- **Shortwave** heating excites mainly **migrating** components
- **Convective** heating excites **migrating and non-migrating** components

Short-wave (O_3) heating at 50 km



Convective heating at 5 km



FREQUENCY (cpd)



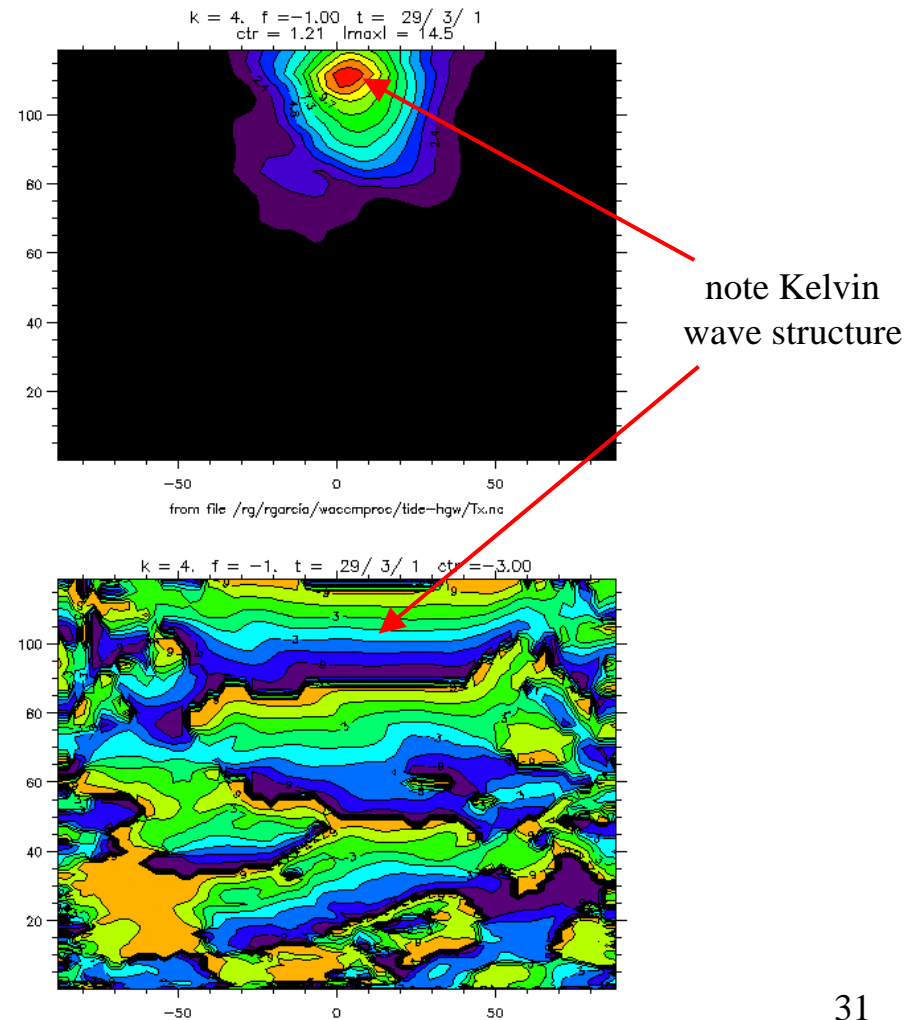
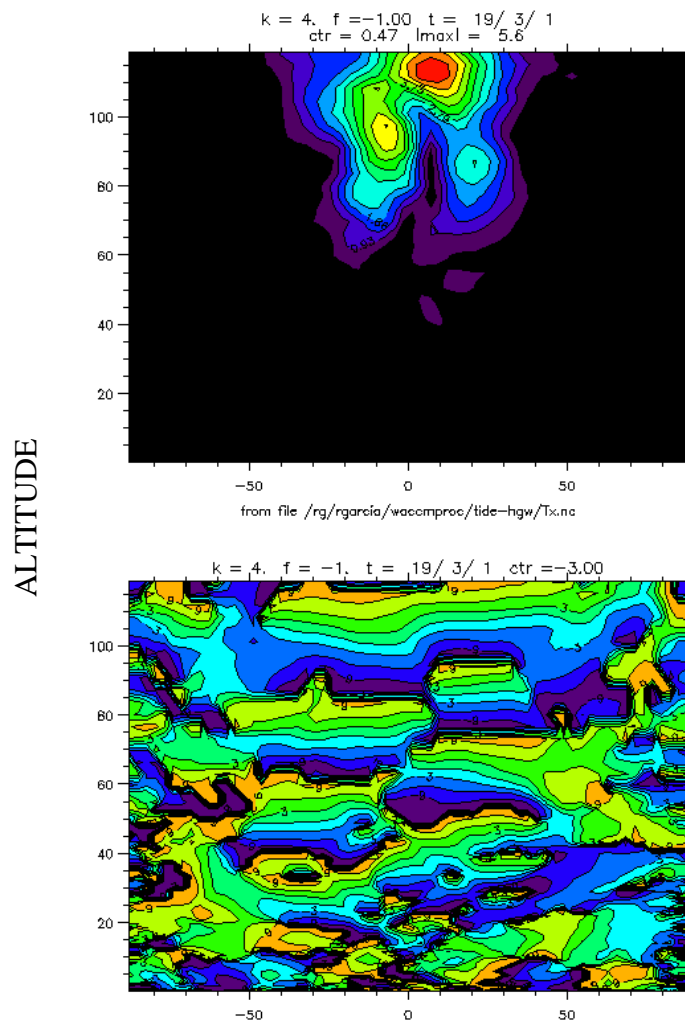
Diurnal "non-migrating" tides



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k=4 eastward T - March 19

k=4 eastward T - March 29



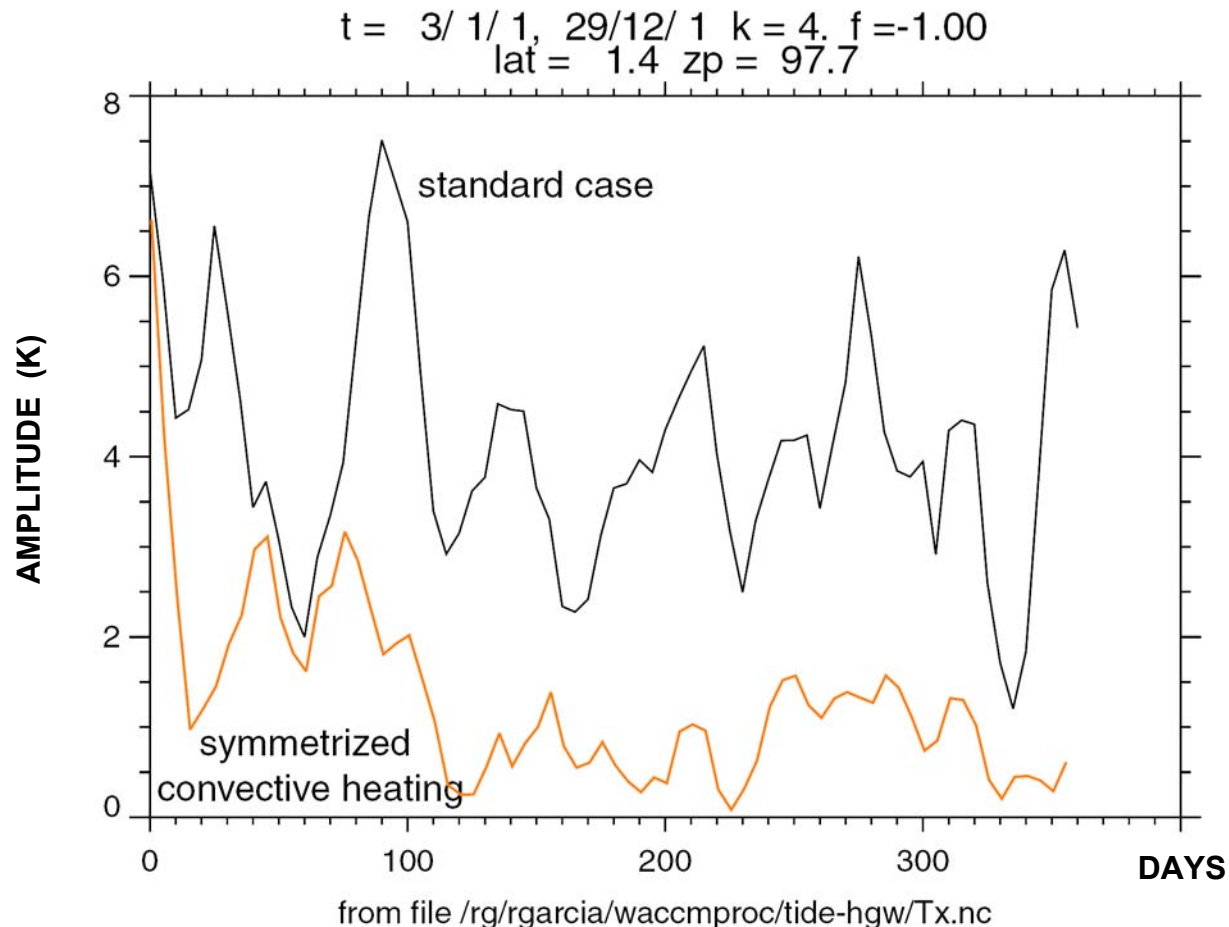


Role of convective Heating in the excitation of "non-migrating" tides



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Annual Cycle: T at 98 km, Equator: $k = 4$ diurnal, eastward



Convective forcing of the tide is eliminated by "symmetrizing" the heating, i.e., zonally-averaging Q'_{conv} and applying the average at every longitude



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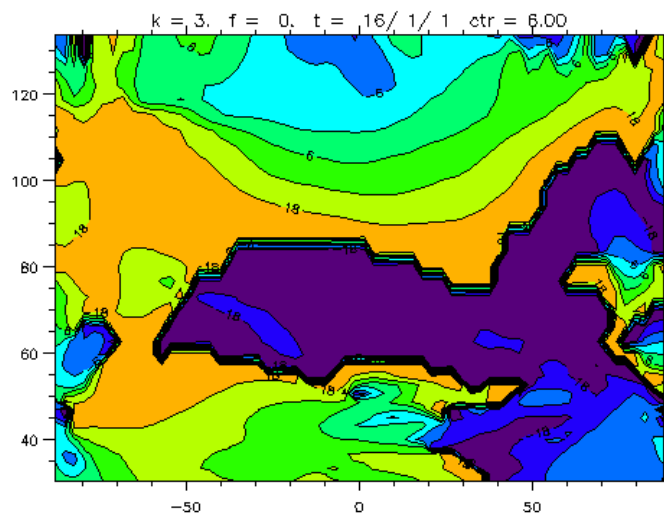
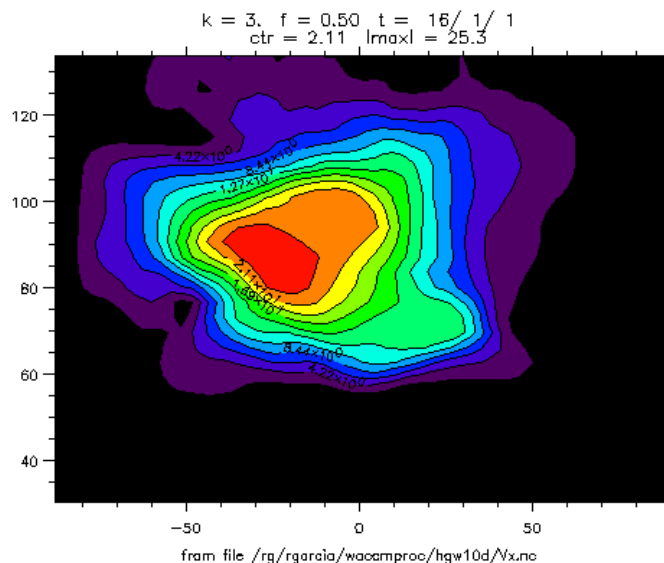


V' structure of 2-day wave (January): Modeled and observed

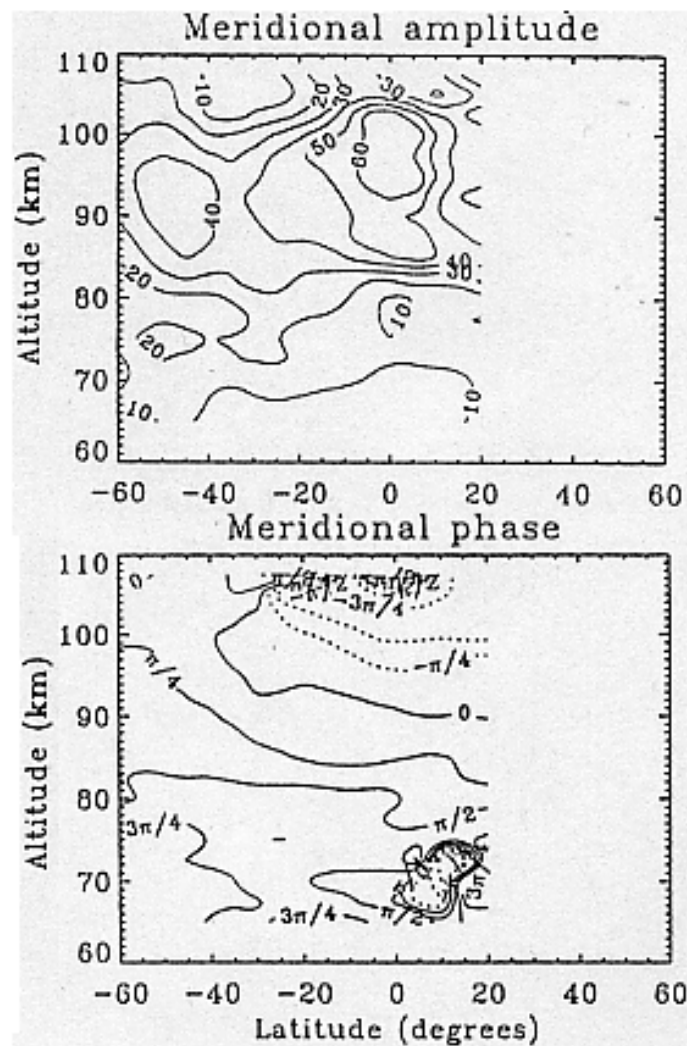


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Wu et al (GRL, 1993) HRDI data

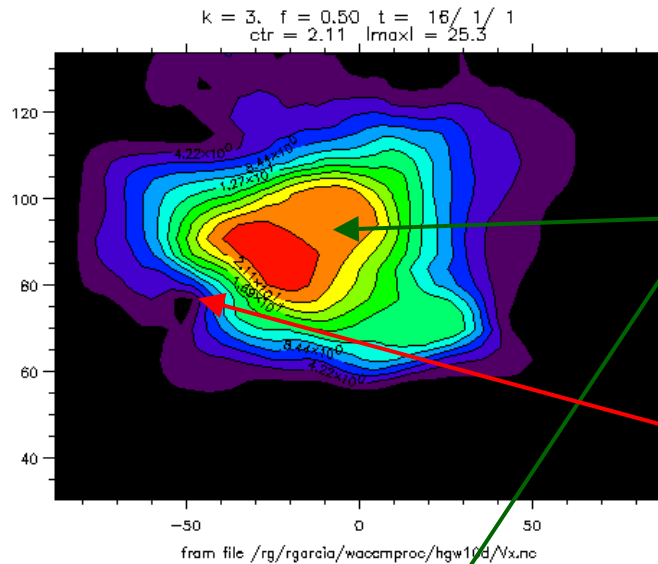


V' structure of 2-day wave ($k = 3$, $t \sim 2$ days) and zonal-mean wind in January



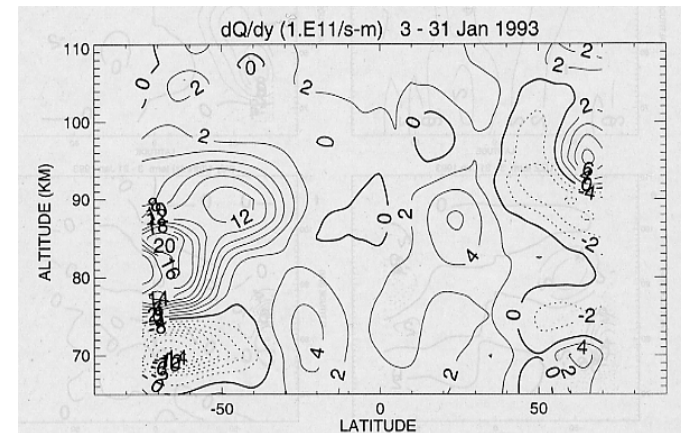
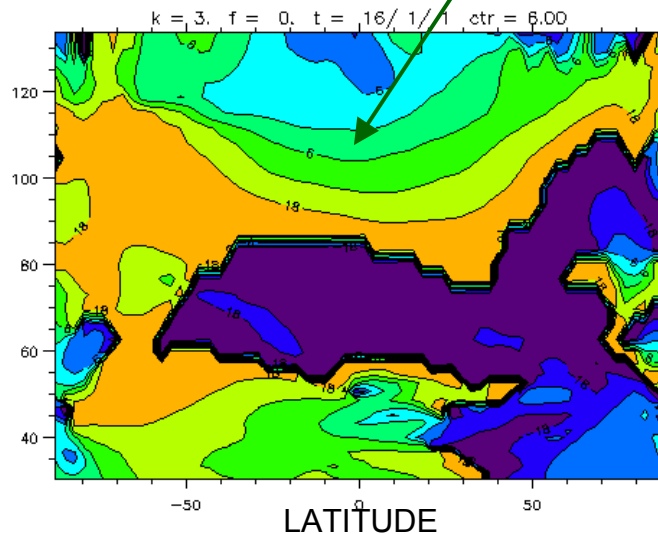
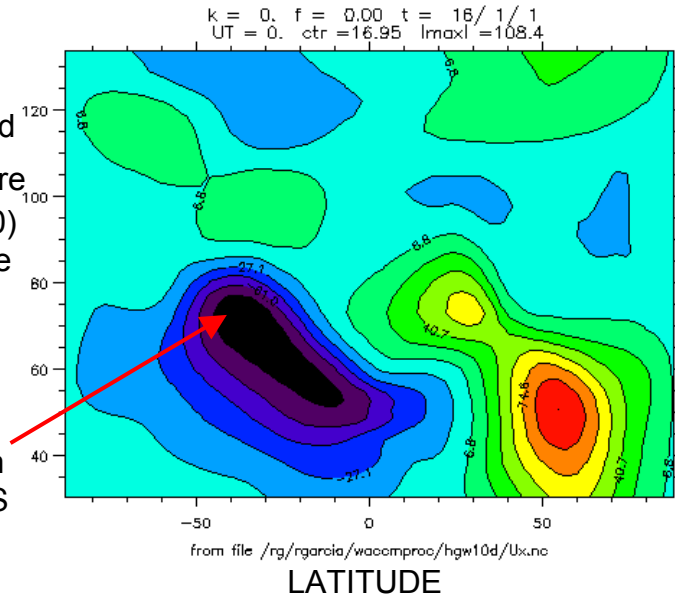
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Amplitude and
phase structure
resemble (3,0)
normal mode

q_y reverses
in this region
~75 km, 45 S



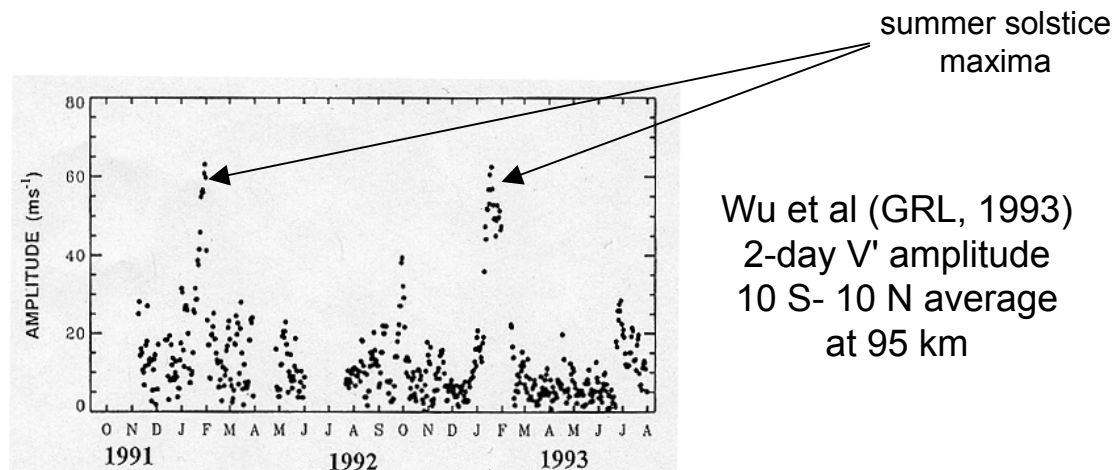
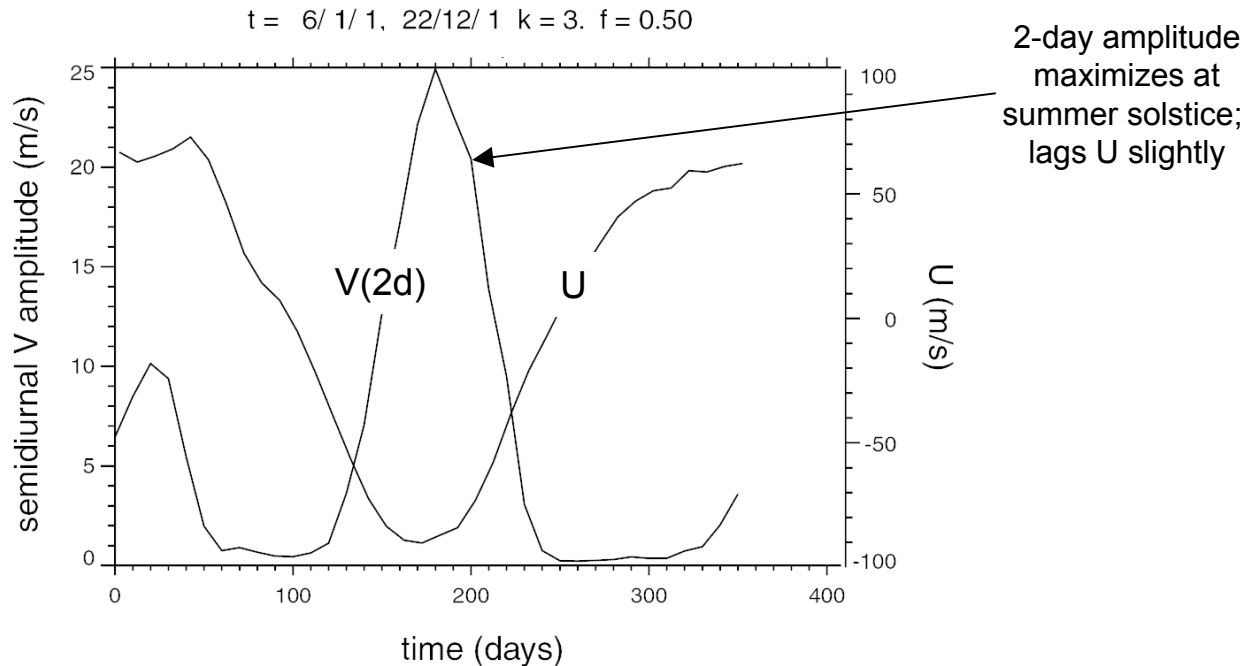
PV gradient from CIRA U (Fritts et al, JGR, 1999)



Seasonal evolution of 2-day wave amplitude (95 km) at Equator and zonal mean wind (75 km) at 30 N



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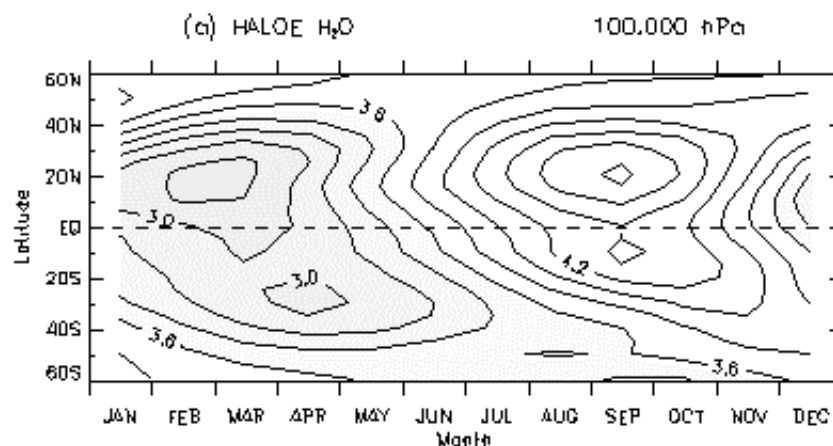
Water seasonal cycle at 100 mb (courtesy of Mijeong Park and Bill Randel)



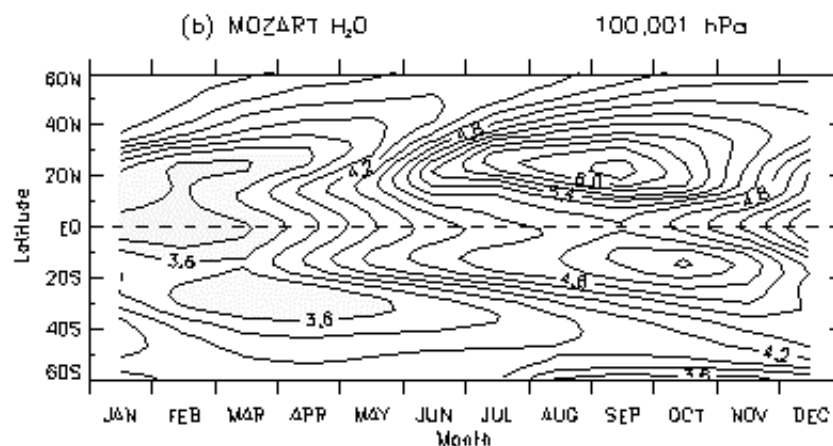
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Water Vapor Distributions

HALOE



MOZART

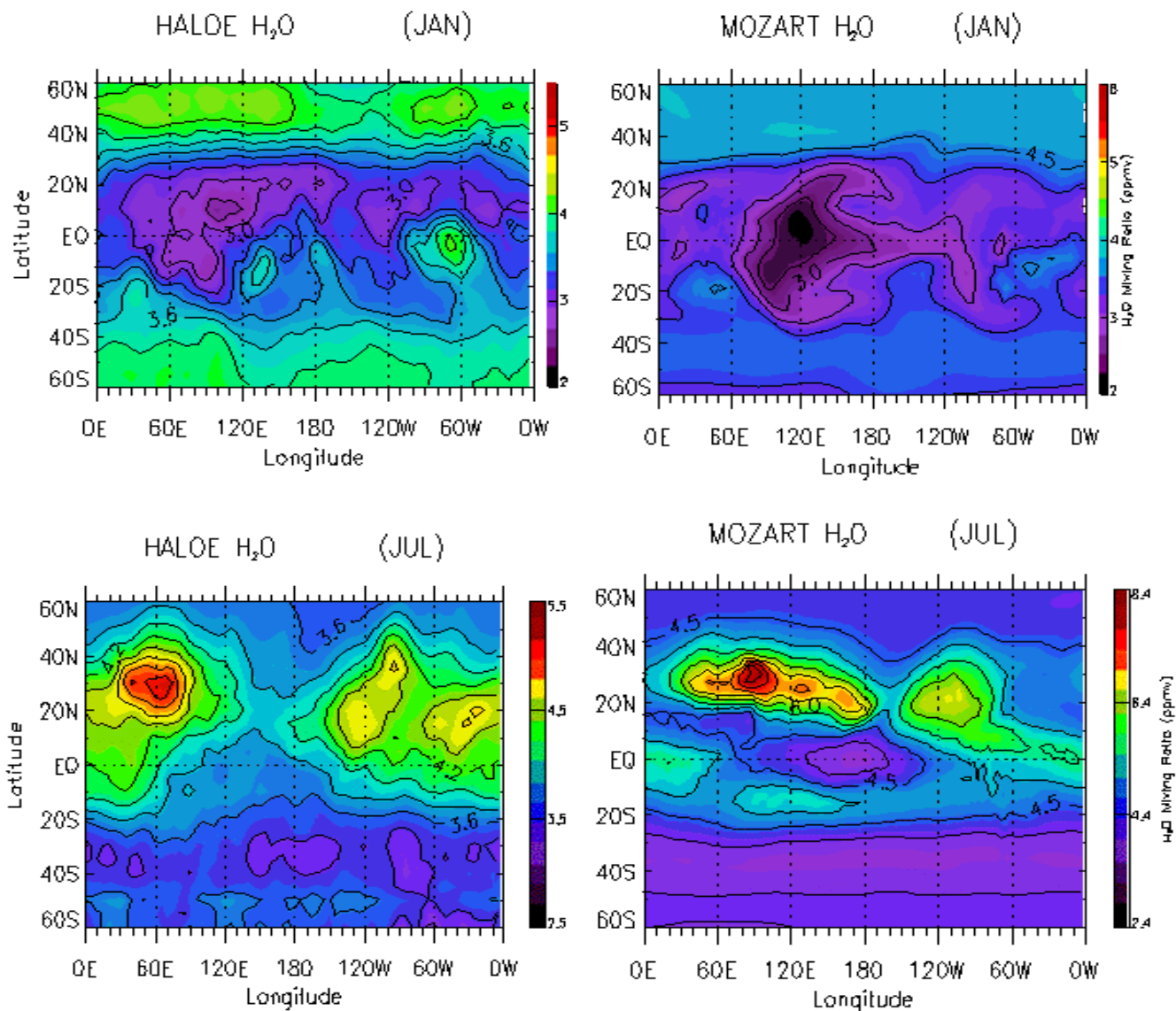




Water vapor at 100 mb: January and July (courtesy of Mijeong Park and Bill Randel)



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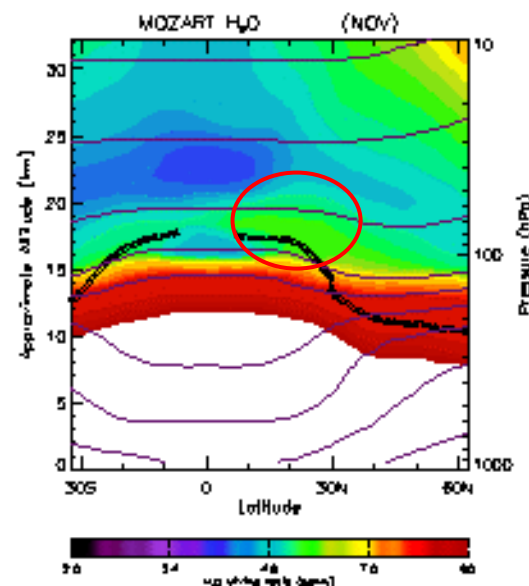
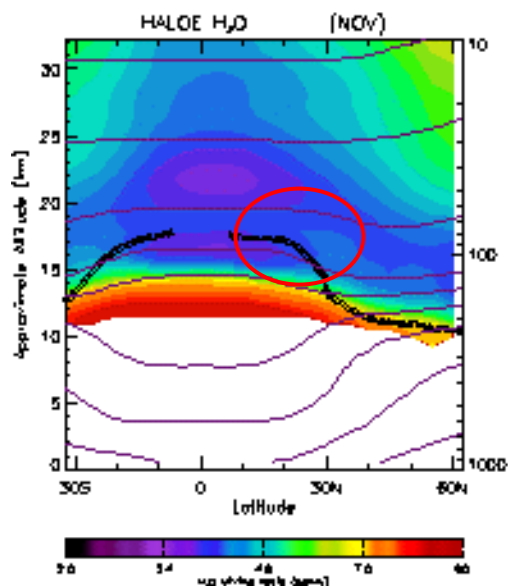
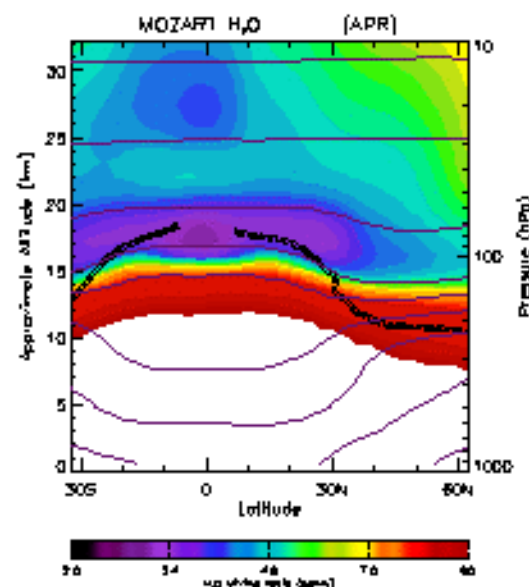
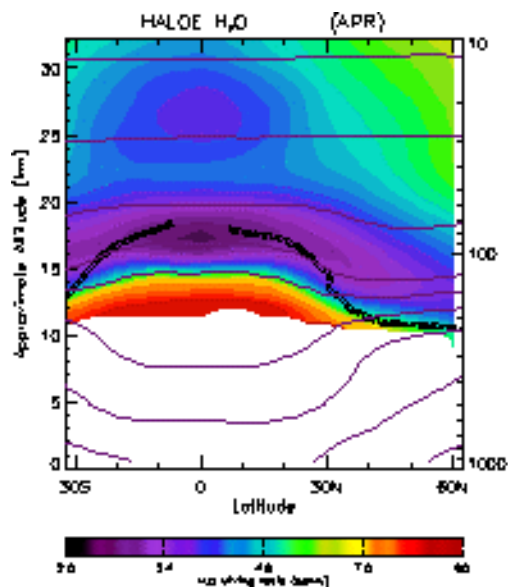




Water vapor in April, November: Meridional section through monsoon region (courtesy of Mijeong Park and Bill Randel)



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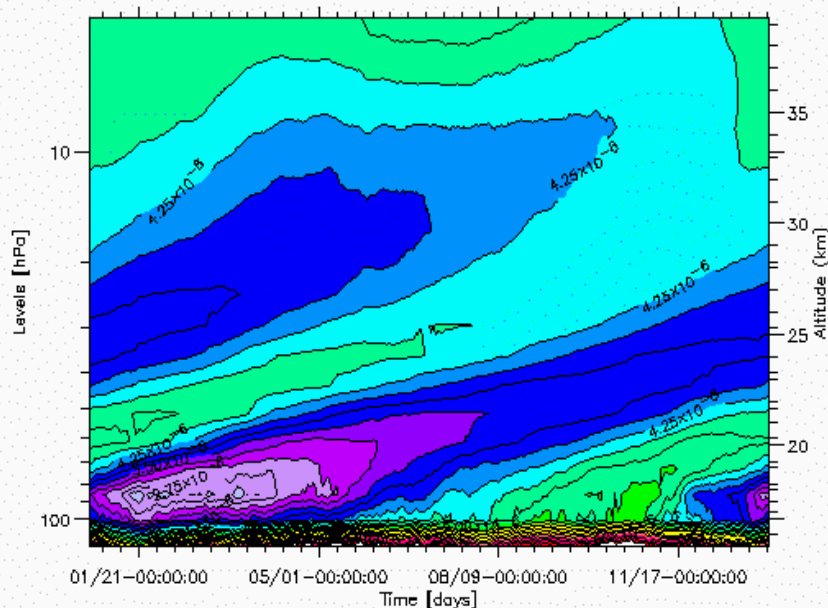
Water vapor "Tape Recorder"

Seasonal evolution at 1.4 and 10 N

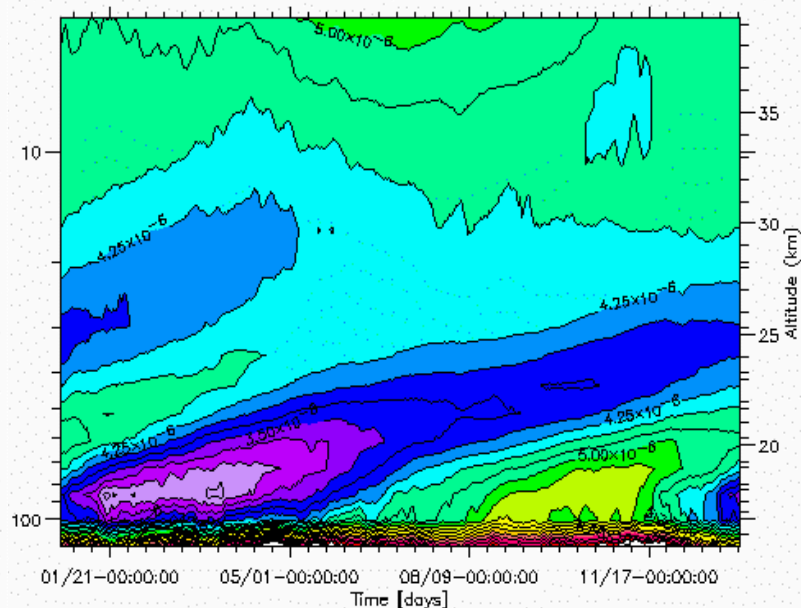


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H2O_VMR_avrg [VMR], lon average, lat 1.42857



H2O_VMR_avrg [VMR], lon average, lat 10.0000





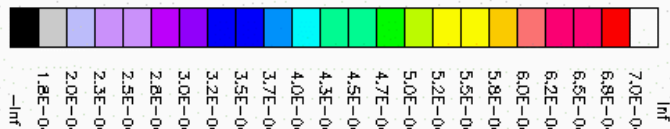
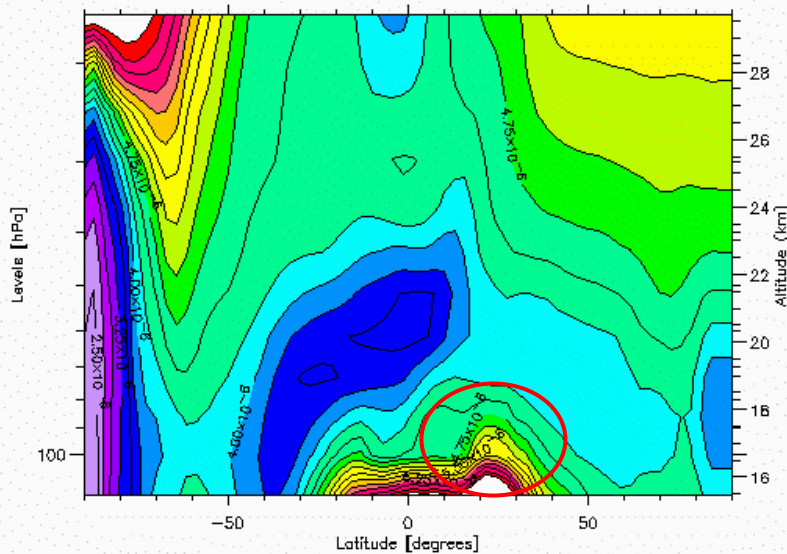
Computed annual cycle of water vapor in the lower stratosphere



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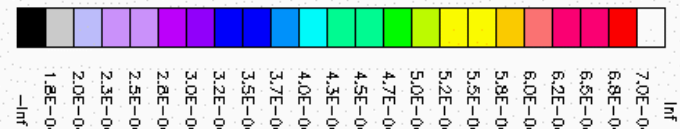
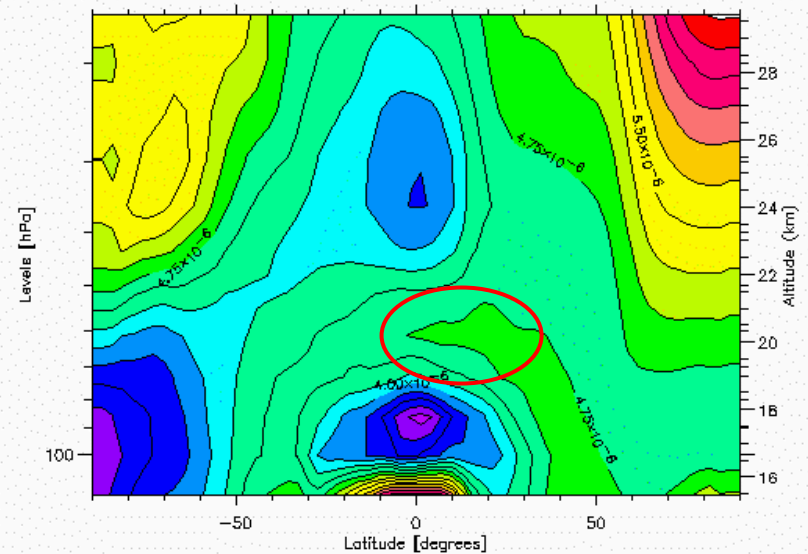
JULY

H2O_VMR_avg [VMR], 07/11-00:00:00 UT, lon average



JANUARY

H2O_VMR_avg [VMR], 01/04-00:00:00 UT, lon average





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The winter stratosphere and the tropospheric annular mode: Observations



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Northern annular mode in geopotential height: EOF index composites (Baldwin and Dunkerton, *Science*, 2001)

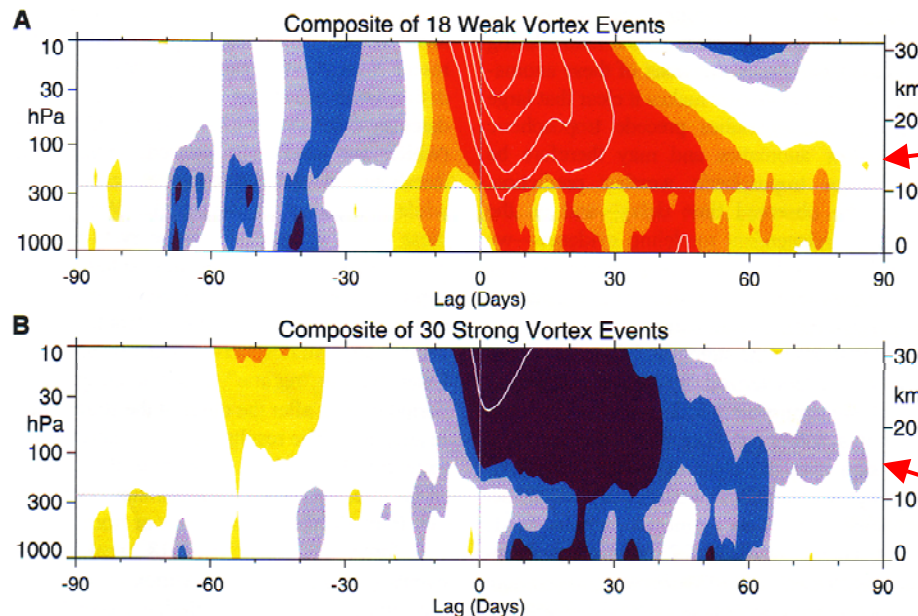


Fig. 2. Composites of time-height development of the northern annular mode for (A) 18 weak vortex events and (B) 30 strong vortex events. The events are determined by the dates on which the 10-hPa annular mode values cross -3.0 and $+1.5$, respectively. The indices are nondimensional; the contour interval for the color shading is 0.25 , and 0.5 for the white contours. Values between -0.25 and 0.25 are unshaded. The thin horizontal lines indicate the approximate boundary between the troposphere and the stratosphere.

annular mode in P_s

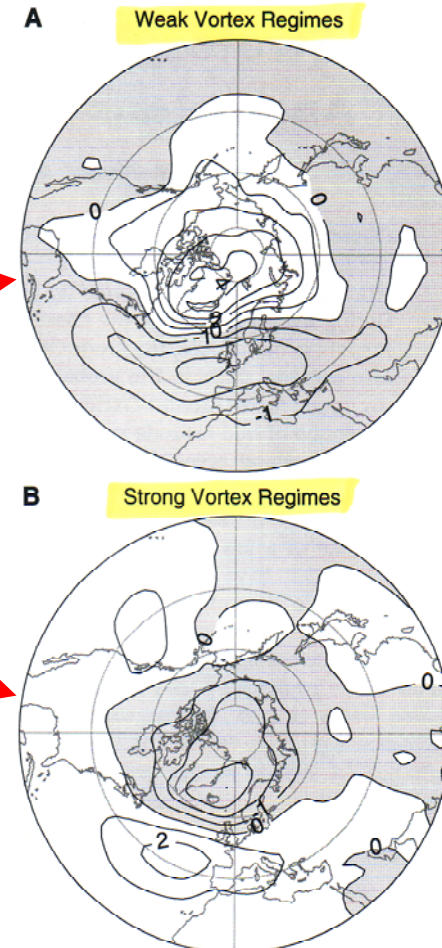


Fig. 3. Average sea-level pressure anomalies (hPa) for (A) the 1080 days during weak vortex regimes and (B) the 1800 days during strong vortex regimes.



Leading EOFs of 10 mb Geopotential and P_s from 46 yrs of WACCM monthly-mean results

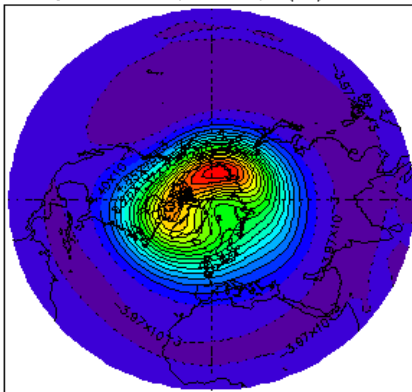


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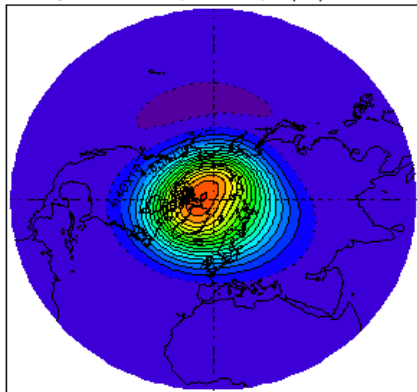
EOF1 of Geopotential at 10 mb, JAN-MAR
(variance explained 61-76%)

EOF1 of Surface Pressure, JAN-MAR
(variance explained 36-47%)

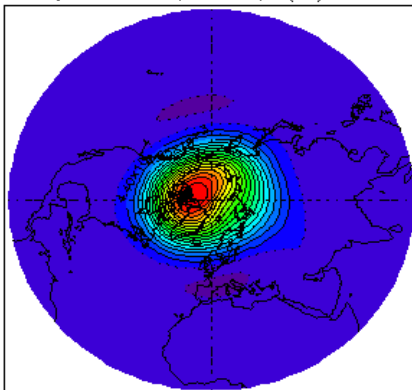
PHI 1, var=61.25%, mo=01, lv(36)=1.1e+01



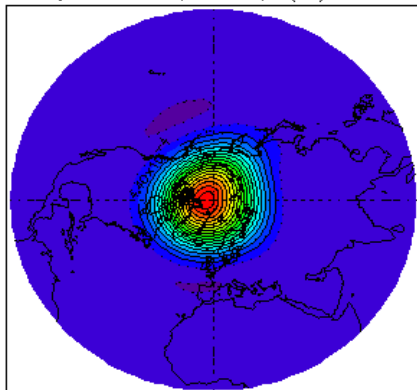
PHI 1, var=70.99%, mo=02, lv(36)=1.1e+01



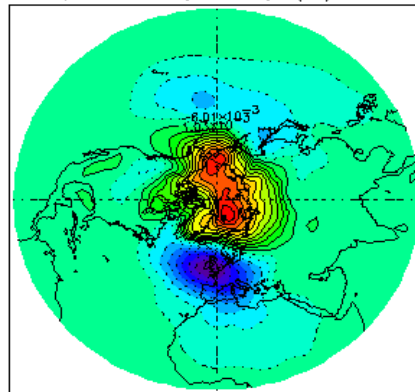
PHI 1, var=74.72%, mo=03, lv(36)=1.1e+01



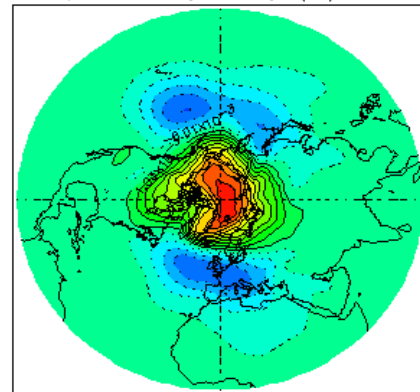
PHI 1, var=76.11%, mo=04, lv(36)=1.1e+01



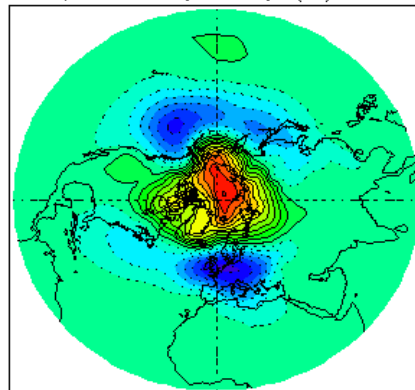
PSF 1, var=36.06%, mo=01, lv(66)=1.0e+03



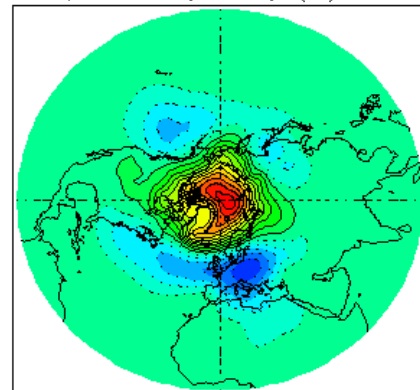
PSF 1, var=43.68%, mo=02, lv(66)=1.0e+03



PSF 1, var=46.92%, mo=03, lv(66)=1.0e+03



PSF 1, var=42.74%, mo=04, lv(66)=1.0e+03





Correlations between monthly-mean EOF 1 of PHI (10 mb) and P_s



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	P_s Jan	P_s Feb	P_s Mar	P_s Apr
Φ Jan	0.11	0.35	0.44	0.28
Φ Feb	0.27	0.45	0.65	0.56
Φ Mar	0.26	0.48	0.50	0.64
Φ Apr	0.10	0.35	0.29	0.54

Correlation coefficients > 0.4 are significant (95%)

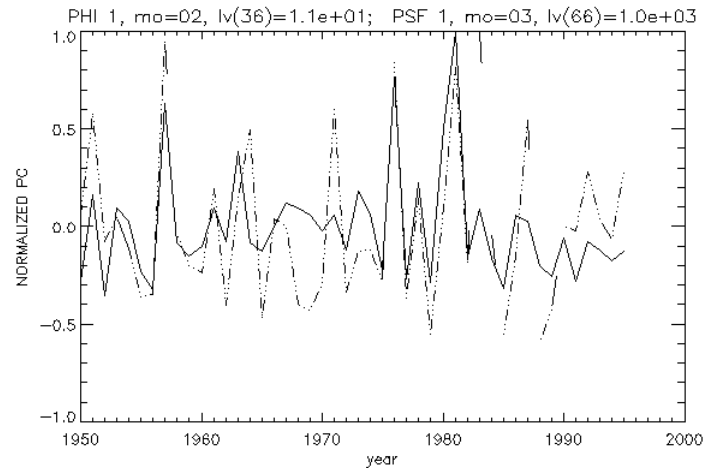


Correlations between PHI (10 mb) EOF 1 and P_s EOF 1: two examples

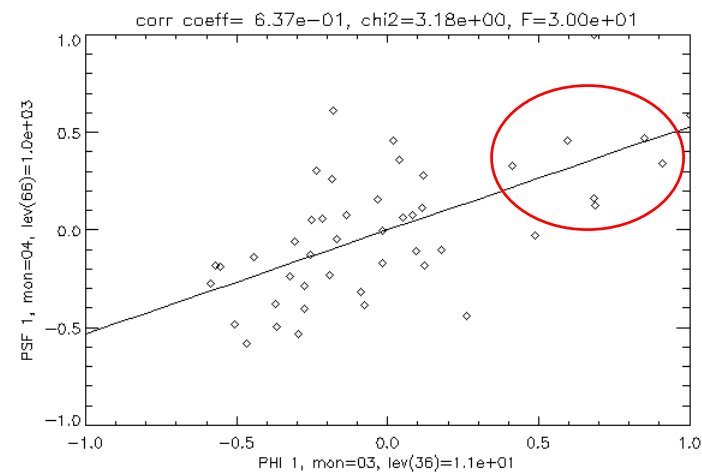
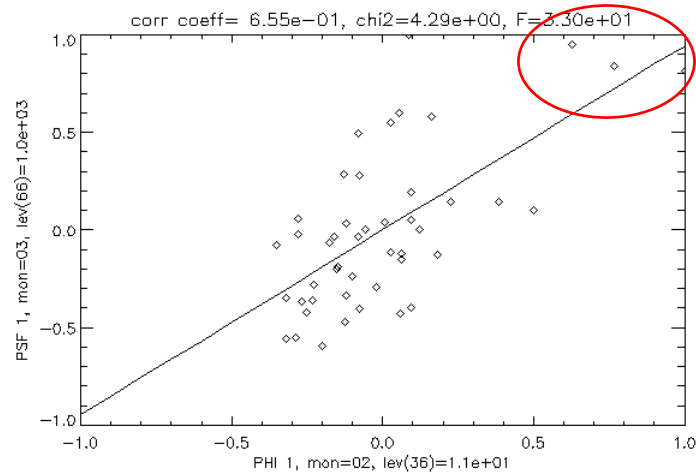
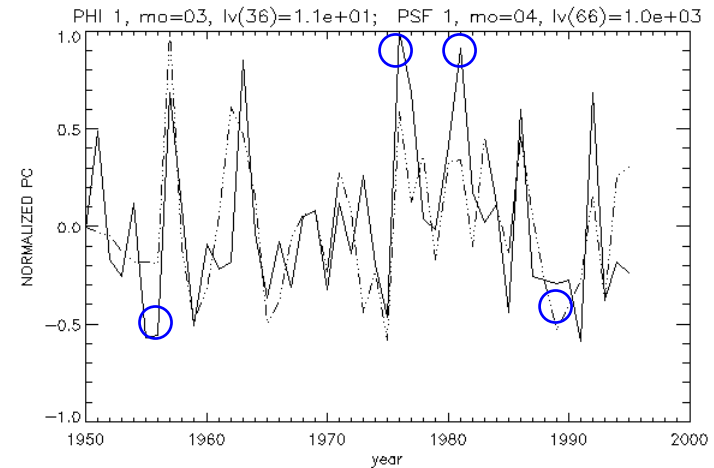


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PHI (10 mb) FEB vs. P_s MAR



PHI (10 mb) MAR vs. P_s APR





Reconstruction of PHI(10 mb) and P_s (4 EOFs) for years of high and low PHI (10 mb) EOF 1

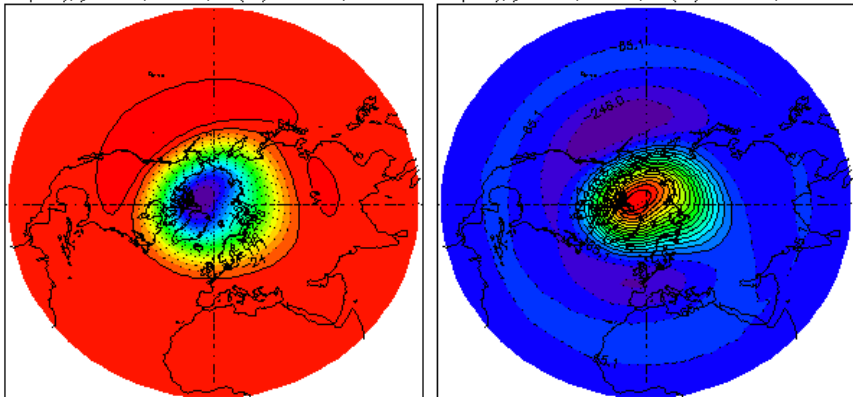


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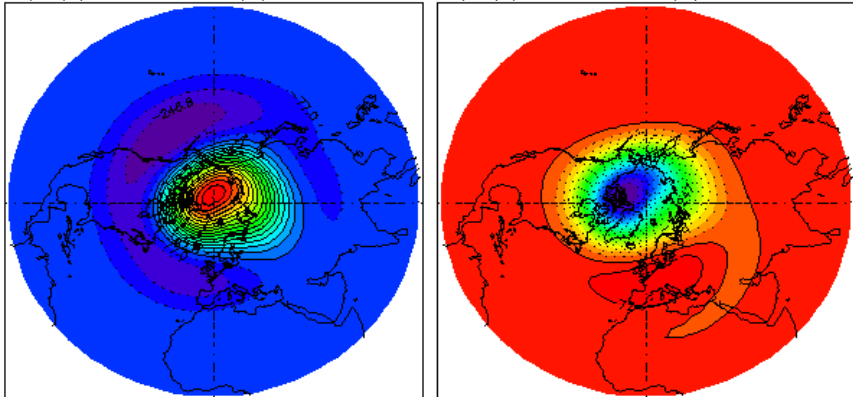
PHI (10 mb) MAR

P_s APR

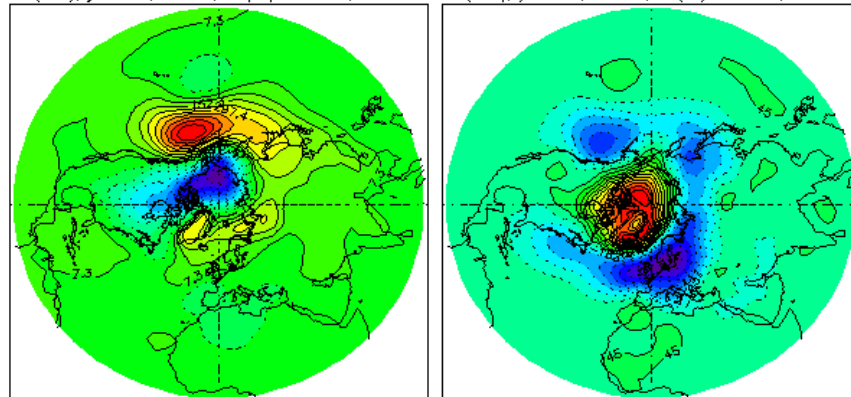
PHI(1-4), yr=1956, mo=03, lev(36)=1.1e+01, ctr= 43.7 PHI(1-4), yr=1976, mo=03, lev(36)=1.1e+01, ctr= 90.4



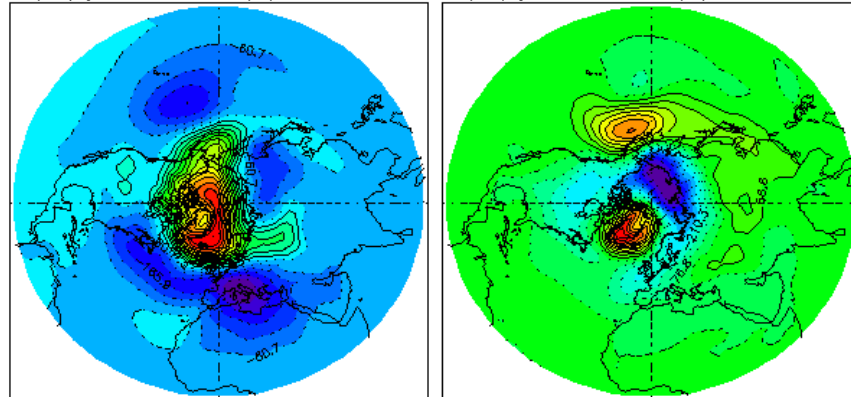
PHI(1-4), yr=1981, mo=03, lev(36)=1.1e+01, ctr= 54.9 PHI(1-4), yr=1985, mo=03, lev(36)=1.1e+01, ctr= 39.9



PSF(1-4), yr=1956, mo=04, lev(66)=1.0e+03, ctr= 72.5 PSF(1-4), yr=1976, mo=04, lev(66)=1.0e+03, ctr=114.0



PSF(1-4), yr=1981, mo=04, lev(66)=1.0e+03, ctr= 52.6 PSF(1-4), yr=1985, mo=04, lev(66)=1.0e+03, ctr= 66.7





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WACCM Workshop: June 20-21 (1:00 PM – 5:30 PM)



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Purpose:

- To provide a forum for scientists of diverse backgrounds to discuss topics related to coupling of atmospheric layers, from the troposphere to the lower thermosphere
- To establish guidelines for community involvement in WACCM

Format:

- A small number of invited talks (40-45 minutes), with an extended period of discussion (20 minutes) following each presentation
- No contributed talks. Workshop participants are encouraged to come prepared to contribute to the general discussion



WACCM Workshop Agenda



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THURSDAY, JUNE 20 (1:00 PM)

Front Range Theater, Radisson Hotel Conference Center

- 1:00 pm: Water vapor in the middle atmosphere
(A. Dessler, University of Maryland, USA)
- 2:00 pm: Effects of solar variability on the middle atmosphere
(K. Kodera, Meteorological Research Institute, Tsukuba, JAPAN)
- 3:00 pm: BREAK
- 3:30 pm: Dynamics of the MLTI Region
(J. Forbes, University of Colorado, Boulder, CO, USA)
- 4:30 pm: Modeling and Observations of the Atmospheric Thermal Structure, Chemical Composition, and Radiation Balance using WACCM and Satellite Data
(M. Mlynczak, NASA L.R.C., Hampton, VA, USA)



WACCM Workshop Agenda



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FRIDAY, JUNE 21 (1:00 PM)

Front Range Theater, Radisson Hotel Conference Center

- 1:00 pm: Convective parameterizations and tropical dynamics
(L. Ricciardulli, Remote Sensing Systems, Santa Rosa, CA, USA)
- 2:00 pm: Modeling the Quasibiennial Oscillation
(M. Giorgetta / E. Manzini, Max Planck Institute, Hamburg, GERMANY)
- 3:00 pm: BREAK
- 3:30 pm: The winter stratosphere and the "annular mode"
(T. Dunkerton, NorthWest Research Associates, Bellevue, WA, USA)
- 4:30 pm: General Discussion: Model Status and Guidelines for Use
(led by B. Boville, NCAR/CGD)