Migrating and Non-migrating Tides in the MLT region

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 - •Migrating diurnal and semidiurnal tides
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 - •Effects of background mean zonal winds
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Definition

Migrating tides

The phase propagates westward synchronized with the diurnal motion of the sun.

$$u \propto e^{i(n\sigma t + s\lambda)}$$
, $\sigma = \frac{2\pi}{1 \text{ solar day}}$,

n=s=1 (diurnal), and n=s=2 (semi-diurnal)

Non-migrating tides

The phase propagates westward or eastward, but not synchronized with the diurnal motion of the sun, or zonally homogeneous oscillation.

$$u \propto e^{i(n\sigma t + s\lambda)}$$
, $\sigma = \frac{2\pi}{1 \text{ solar day}}$,

n=1 (diurnal), and n=2 (semi-diurnal) s = 0, ± 1 , ± 2 , ± 3 , \cdots s $\neq 1$ (diurnal), and s $\neq 2$ (semi-diurnal)

Migrating diurnal tide, October



Diurnal westward moving S=1, u(m/s) z = 98 km

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Nonmigrating diurnal tide, October



Diurnal westward moving S=2, u(m/s) z = 98 km

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Migrating and nonmigrating diurnal tides, October

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Diurnal westward moving S=1+2, u(m/s) z = 98 km

Migrating semidiurnal tide, October



Semidiurnal westward moving S=2, u(m/s) z = 98 km

Non-migrating semidiurnal tide, October

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Semidiurnal westward moving S=1, u(m/s) z = 98 km

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Migrating and non-migrating semidiurnal tides



Semidiurnal westward moving S=2+1, u(m/s) z = 98 km

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λ: N. P.

Figure 176.





Fig. 1. Top Hough functions for diurnal modes normalized to a maximum value of unity. Keys and normalization factors for each Hough mode are as follows: (1,1) solid line, 0.606; (1, 1) dashed line, 1.034; (1, 2) datted line, 1.054; (1, 4) dashed-dotted line, 0.513; (1,2), dashed-double dotted line, 0.641 Bottom Northerly velocity expansion functions for diurnal modes normalized to a maximum value of unity. Normalization factors are 0.026, 0.126, 0.100, 0.024, and 0.015, respectively. Center Westerly velocity expansion functions for diurnal modes normalized to a maximum value of unity. Normalization factors are 0.026, 0.126, 0.100, 0.024, and 0.015, respectively. Normalization factors are 0.038, 0.130, 0.110, 0.024, and 0.018, respectively.

Vertical wavelengths of some tidal modes Isothermal Atmosphere: T=250 K

Diurnal (km) Migrating

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Mode	(1,-2)	(1,-4)	(1,1)	(1,3)	(1,5)
s=1, W	evanescent	evanescent	<u>28</u>	<u>11</u>	7
	X				

Non-mograting

Mode	(1,1)	(1,2)	(1,3)	(1,4)	(1,5)
s=0	<u>102</u>	<u>_26</u>	15	<u>11</u>	9
Mode	(1,1)	(1,2)	(1,3)	(1,4)	(1,5)
s=-1, E	external	_53_		<u>15</u>	<u>11</u>
Mode	(1,1)	(1,2)	(1,3)	(1,4)	(1,5)
s=2,W	<u> </u>	<u>16</u>	<u>11</u>	9	<u>7</u>
Mode	(1,1)	(1,2)	(1,3)	(1,4)	(1,5)
s=-2, E	<u>108</u>	<u>38</u>	<u>21</u>	<u>14</u>	<u>10</u>
Mode	(1,1)	(1,2)	(1,3)	(1,4)	(1,5)
s=5, W	<u>19</u>	<u>13</u>	<u>10</u>	8	_7
Mode	(1,1)	(1,2)	(1,3)	(1,4)	(1,5)
s=-5, E	<u>31</u>	<u>21</u>	<u>15</u>	<u>12</u>	9



Fig.1 Meridional cross section of tide-generating heating for solstice.

Miyahara (1884)

UARS (WINDII + HRDI)







Fig. 16. Meridional cross section of the induced mean zonal winds. Dashed lines denote easterlies. The contour interval is 20 ms⁻¹. Wu et al. (1993)

Numerical simulation (non-linear)



Fig. 1. Contours of meridional wind amplitude (*solid lines*) and phase (*dashed lines*) of the diurnal tide observed by WINDII for March/ April 1992 and 1993 at a 45°E to 135°E, b 135°E to 135°W, c 135°W

to 45° W, and d 45° W to 45° E longitude sectors. Amplitude contours > 50 m/s are shaded. The phase (local time of maximum) is contoured using a 3-h interval

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Fig. 2a-d. Contours of meridional wind amplitude (solid lines) and phase (dashed lines) of the diurnal tide observed by WINDII for December 1992, 1993 and January 1993, 1994 at: a 45°E to 135°E, b

135°E to 135°W, c 135°W to 45°W, and d 45°W to 45°E longitude sectors. Amplitude contours >40 m/s are shaded. The phase (local time of maximum) is contoured using a 3-h interval



Fig. 1. Average spectrogram of the hourly meridional winds over the four directions of measurement. The spectrogram is formed by sliding 10-day spectra in increments of 3 days from 19 January 1995 to 26 January 1996



Fig. 3.13. Altitude distribution of the amplitude of the solar diurnal component of *u* at 15° intervals of latitude; isothermal basic state assumed. After Lindzen (1967a).

Wu, Miyahara, and Miyashi (1885)

A nonlinear simulation of the thermal diurnal tide



Fig. 4. Vertical profiles of (a) amplitudes and (b) phases, of zonal velocities of the diurnal tide at various latitudes.



FIGURE 3. Amplitude of the diurnal temperature escillation over the equator for different distributions of the cooling rate coefficient: ----, standard; ---, without photochemical acceleration; ----, zero.



FIGURE 7. Time-height cross-section of the Richardson number over the equator.

Lindzen (1967) 19



Fig. 1. The radiative damping rate (solid) and the vertical eddy diffusion coefficient (dashed).

Akmaev et al. (1992)

Monthly mean amplitudes of diurnal tides (UARS - HRDI)



Vertical eddy diffusion coefficients derived from HRDI data



Khattatov et al., (1997)

Middle <u>Atmosphere Circulation Model</u> at Kyushu University <u>MACMKU</u>

•T21L55, General Circulation Model (GCM) •Height Range: Ground through about 150 km •Solar Radiation: having diurnal cycle, H₂O, O₃, and O₂ heating, H₂O: Predicted in the model (Troposphere: non-zonal) O3, and O2: zonally symmetric distribution •Infrared radiation: Fomichev's parameterization Troposphere: Chou's parameterization •Land Temperature: Predicted in the model •SST: Prescribed monthly mean SST Tropospheric physical processes Latent heat, Topography, etc. •Dissipation processes in the MLT Molecular viscosity and conductivity Ion drag (Local time dependent) Dry convective adjustment Eddy diffusion Rayleigh Friction (Only for the zonal mean zonal winds in the MLT) No gravity wave drag parameterization



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Jan. D. W. S=5 140 É 130 110 HE 100 HE 100 0 STANDARD 90 80 70 60 60 30 0 -30 Ν LATITUDE

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Chapter 4: Sources of Forcing



Fig.(4.1.2) Latitude dependence of total forcing which includes latent heating, heating due to dry convection, eddy thermal conduction and insolation absorption by water vapor in the troposphere and ozone absorption heating in the middle atmosphere for s = +1, (a), and s = -1, (b).

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Ebanayake Dr. Thesis



Fig.(4.1.3) As in Fig.(4.1.1) except for s = +2, (a), and s = -2, (b).

Linear model result by the GCM heating.

(Ekanayake et al. 1997)



GCM result



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CONTOUR INTERVAL = 3.000E+00



Fig, 1a

number s=1 becomes large in the summer season at the north and south polar MLT regions. Two different possible excitation mechanisms of the non-migrating semidiurnal tide are suggested (Forbes et al., 1995; Portnyagin et al., 1998). One is non-migrating heating in the troposphere associated with latent heat release. The other is nonlinear interaction between the migrating semidiurnal tide and a stationary planetary wave with zonal wavenumber s=1. In this section the later mechanism is investigated using the output data of the MACMKU.

3-1. Source of excitation

The idea of the nonlinear interaction forcing is as follows. The nonlinear interaction terms in the governing equation system expressed by the pressure coordinate system in the MACMKU are given by the advection terms. The nonlinear forcing terms due to a planetary wave and the semidiurnal tide are given by

$$F_{\lambda} = \frac{u_{pl}}{a\cos\theta} \frac{\partial u_{sd}}{\partial \lambda} + \frac{v_{pl}}{a\cos\theta} \frac{\partial}{\partial \theta} (u_{sd}\cos\theta) + \omega_{pl} \frac{\partial u_{sd}}{\partial p} + \frac{u_{sd}}{a\cos\theta} \frac{\partial u_{pl}}{\partial \lambda} + \frac{v_{sd}}{a\cos\theta} \frac{\partial}{\partial \theta} (u_{pl}\cos\theta) + \omega_{sd} \frac{\partial u_{pl}}{\partial p} , \qquad (3.1)$$

$$F_{\theta} = \frac{u_{pl}}{a\cos\theta} \frac{\partial v_{sd}}{\partial \lambda} + \frac{v_{pl}}{a} \frac{\partial v_{sd}}{\partial \theta} + \omega_{pl} \frac{\partial v_{sd}}{\partial p}$$

$$+\frac{u_{sd}}{a\cos\theta}\frac{\partial v_{pl}}{\partial \lambda} + \frac{v_{sd}}{a}\frac{\partial v_{pl}}{\partial \theta} + \omega_{sd}\frac{\partial v_{pl}}{\partial p} + (2u_{sd}u_{pl})\frac{\tan\theta}{a}, \qquad (3.2)$$

$$F_{T} = \frac{u_{pl}}{a\cos\theta} \frac{\partial T_{sd}}{\partial\lambda} + \frac{v_{pl}}{a} \frac{\partial T_{sd}}{\partial\theta} + \omega_{pl} (\frac{\partial T_{sd}}{\partial p} - \frac{R}{c_{p}p} T_{sd}) + \frac{u_{sd}}{a\cos\theta} \frac{\partial T_{pl}}{\partial\lambda} + \frac{v_{sd}}{a} \frac{\partial T_{pl}}{\partial\theta} + \omega_{sd} (\frac{\partial T_{pl}}{\partial p} - \frac{R}{c_{p}p} T_{pl}), \qquad (3.3)$$

Result of a linear response model with Th of GCM.

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(MACMKU)

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Fig.1b

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Fig. 12a

Effects of background mean zonal winds

 $\overline{u}(\theta, z)$:mean zonal winds

Doppler shift of tidal frequency ω

$$\frac{\partial}{\partial t} + \frac{\overline{u}}{a\cos\theta} \frac{\partial}{\partial\lambda} = i(\omega + s\overline{u})$$

S: zonal wave number

$$\theta$$
: Latitude



Fig. 12. Location where Doppler shifted frequency is equal to Coriolis Frequency for W6 (-----) and E6 () waves. Lower latitude region than the location indicates internal region.

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Miyahara et al. (1983)

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CONTOUR INTERVAL - 5.000E+00

6 UT

Jan. Sem-diurnal S=1 v(m/s)



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3 UT Jan. Sem-diurnal S=1 v(m/s) ٦

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CONTOUR INTERVAL = 5.000E+00

9 UT

Jan. Sem-diurnal S=1 v(m/s)



CONTOUR INTERVAL = 5.000E+00

Semidiurnal westward moving S=1, v(m/s) = 108 km, S. H.

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Semidiurnal tides, Southern Hemispher, January

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Semidiurnal tides S=0 to 7, v(m/s) = 108 km, S. H.



Concluding remarks: Migrating tides Non-migrating tides Both tides exist in the MLT region Confirmed by obsevartions and numerical simulations

Excitation mechanism? Migrating tides: H₂O, O₃, O₂ heating

> Non-migrating tides Moist convective heating Tide-planetary waves interactions

Longitudinal variation of tidal amplitudes Interference between migrating and non-migrating tides

Need more observations TIMED mission