

Modeling of Gravity Wave and Instability Processes
in the Middle Atmosphere

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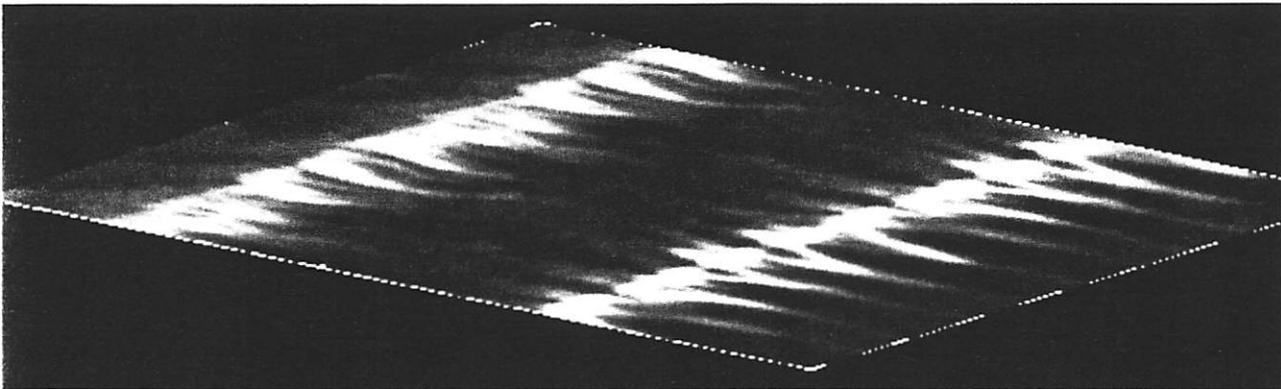
Outline

- 1. Motivations for MA Instability Studies**
- 2. Model Formulation**
- 3. Instability due to Wave Breaking**
- 4. Kelvin-Helmholtz Instability**
- 5. Conclusions**

Modeling Motivations

- **Wave transports of energy and momentum are central to our understanding of middle atmosphere dynamics**
- **Wave interaction and instability processes account for wave saturation, spectral character, and constraints on energy and momentum fluxes**
- **Dynamics of transition from laminar to turbulent flow dictates character of turbulence, efficiency of mixing and transports**

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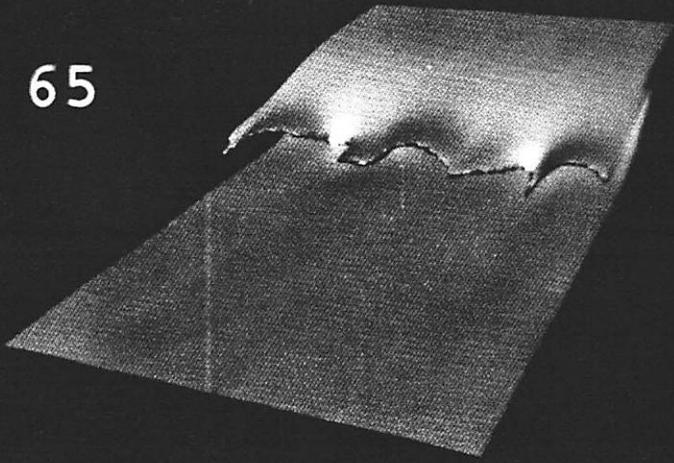
Model Formulation

- **Solves Euler equations with spectral viscosity**
- **Employs spectral collocation techniques**
 - **Fourier in x, y**
 - **Chebyshev in z**
- **Uses domain decomposition for higher resolution, greater efficiency**
 - **wave breaking using two domains**
 - **forcing in low-resolution lower domain (96, 48, 65)**
 - **instability in high-resol. upper domain (192, 96, 129)**
 - **Kelvin-Helmholtz instability using four domains**
 - **Re = 200 to 2000**
- **2D initial evolution, 3D instability evolution following noise insertion at finite amplitude**
- **Boundary and interface conditions**
 - **periodic in x, y**
 - **open in z , using upstream characteristics**

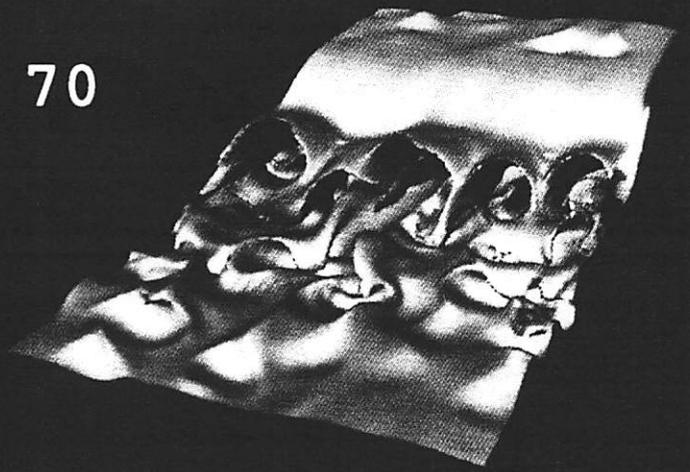
Wave Breaking Simulations

- **high-frequency wave in a shear flow**
 - ~ 30 min period
 - ~ 24 km wavelength
 - ~ 1 km instability depth
- **wave field evolution**
 - **initial instability is convective, streamwise**
 - **secondary instability is dynamical, spanwise and localized (3D KH)**
 - **evolution is rapid and transient, collapse to turbulence $\sim 1 T_b$**

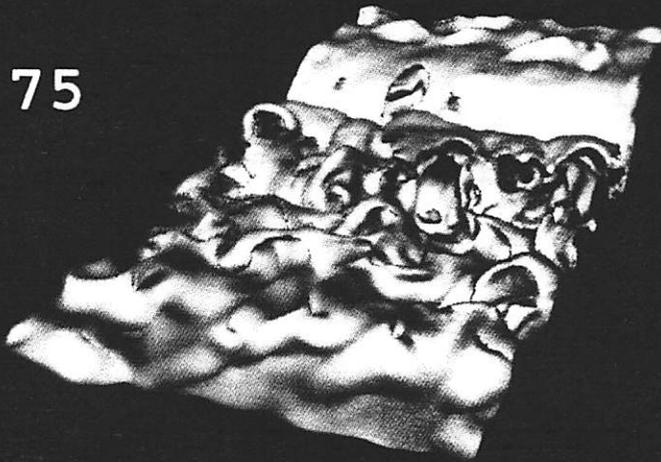
65



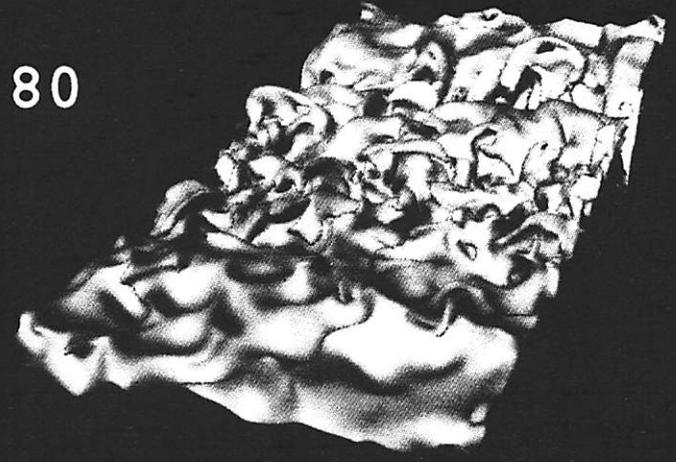
70



75



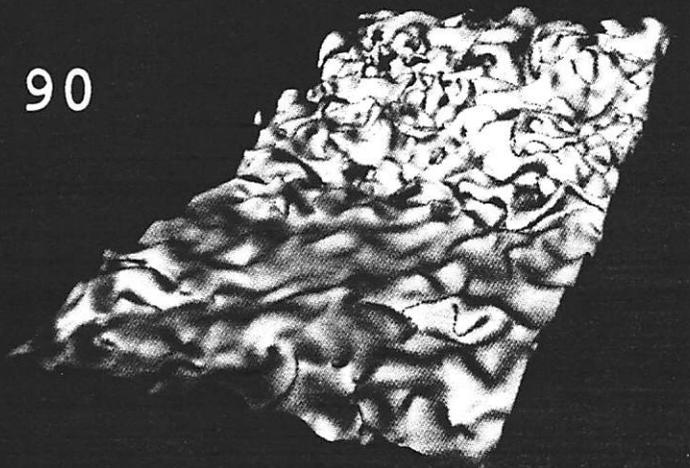
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85

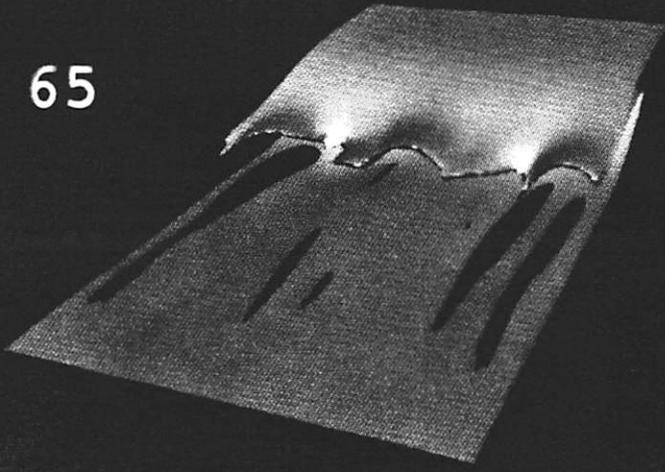


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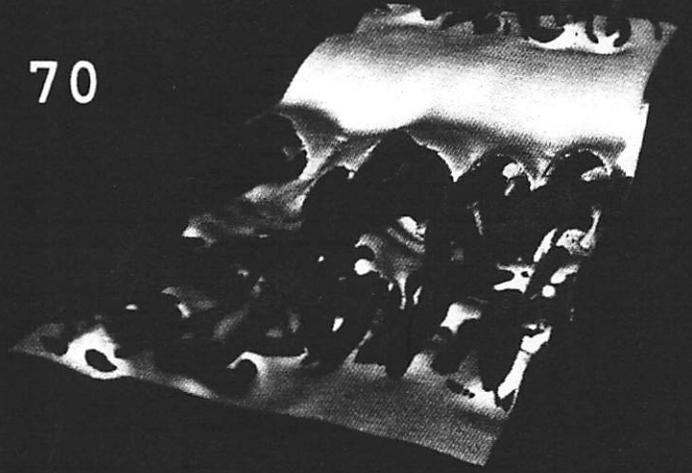


Wave breaking shown with isosurface of θ

65



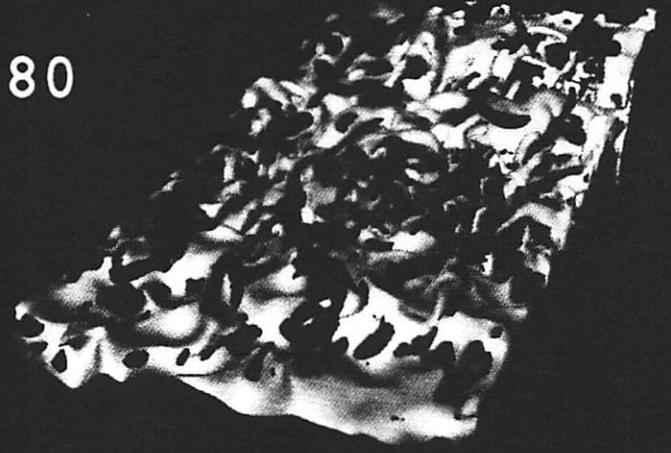
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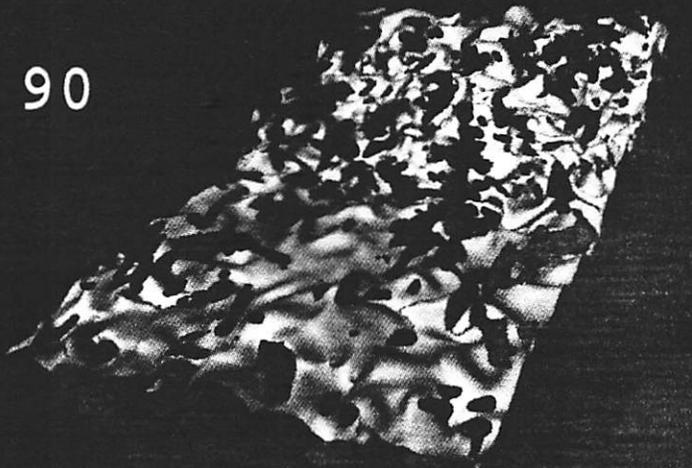
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85



90



Wave breaking with isosurface of θ
 and of positive (red) and negative (blue)
 streamwise vorticity

Eddy Kinetic Energy Equation

$$\left(\frac{\partial}{\partial t} + \hat{\mathbf{u}} \cdot \nabla\right) K_e + \frac{\partial}{\partial x} \langle p' u' \rangle + \frac{\partial}{\partial z} \langle p' w' \rangle$$
$$\approx -\hat{\rho} \langle u' u'_i \rangle \frac{\partial}{\partial x} \tilde{u}_i - \hat{\rho} \langle u'_i w' \rangle \frac{\partial}{\partial z} \hat{u}_i + \frac{\hat{\rho} g}{\hat{\theta}} \langle \theta' w' \rangle$$

Vorticity Equation

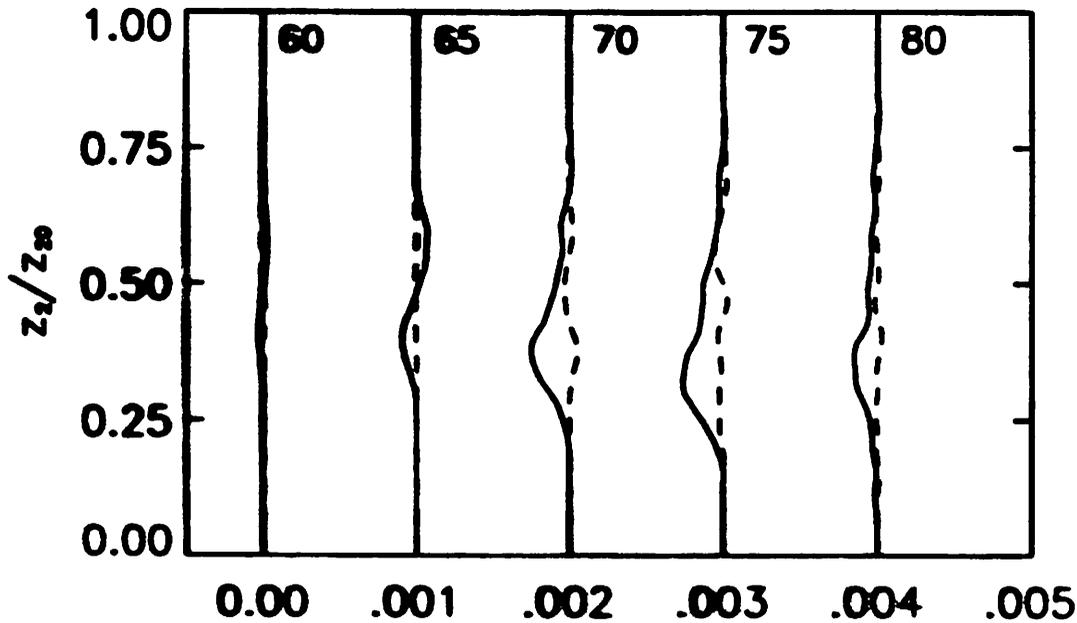
$$\frac{d\omega_i}{dt} \approx \omega_j S_{ij} + \left\{ \frac{\nabla \rho}{\rho} \times \frac{\nabla p}{\rho} \right\}_i$$

where

$$S_{ij} = \frac{1}{2} (\partial_i v_j + \partial_j v_i)$$

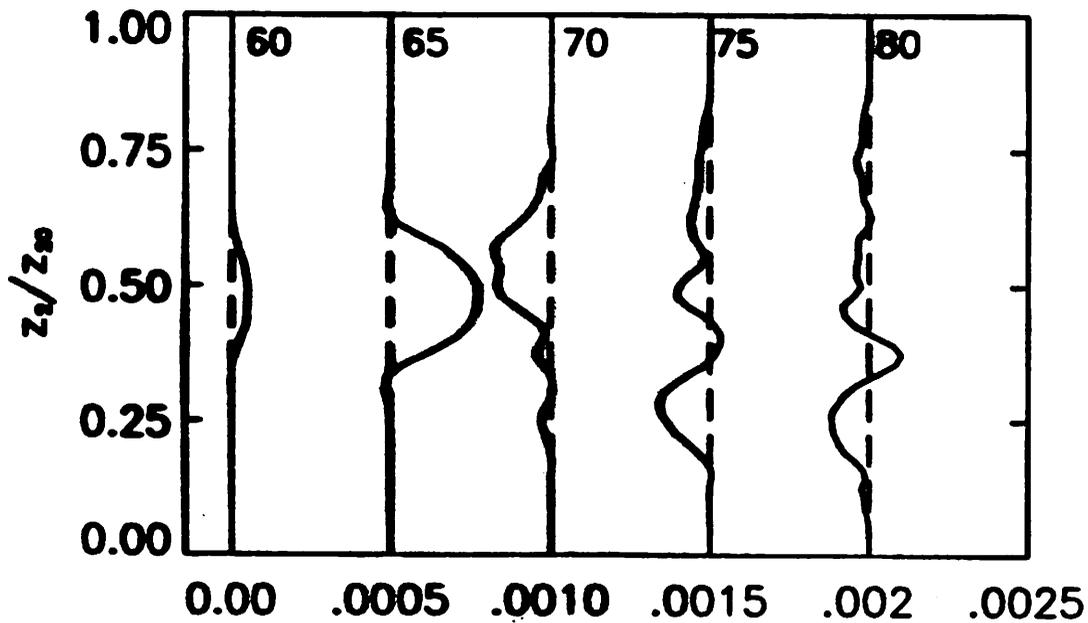
eddy momentum fluxes (domain averaged)

$$\overline{u'w'} \quad \text{---} \quad \overline{v'w'} \quad \text{---}$$



eddy heat fluxes (domain averaged)

$$\overline{\theta'w'}$$





Modeling of Breaking Gravity Wave

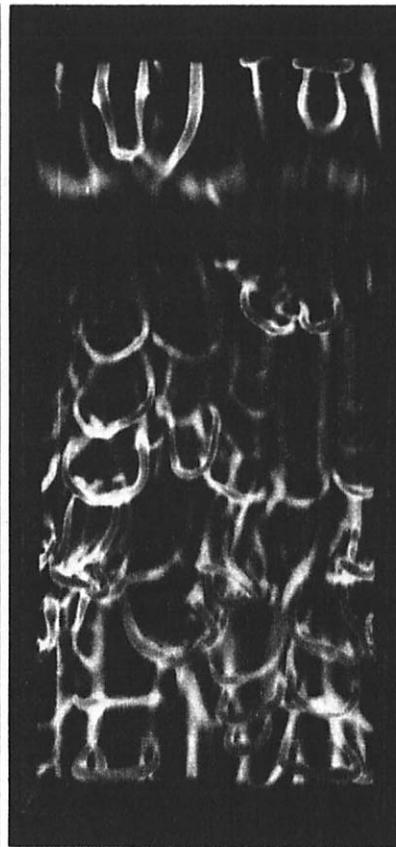
- Vortices rendered by $\lambda_2 < 0$ of $S^2 + R^2$, viewed from below



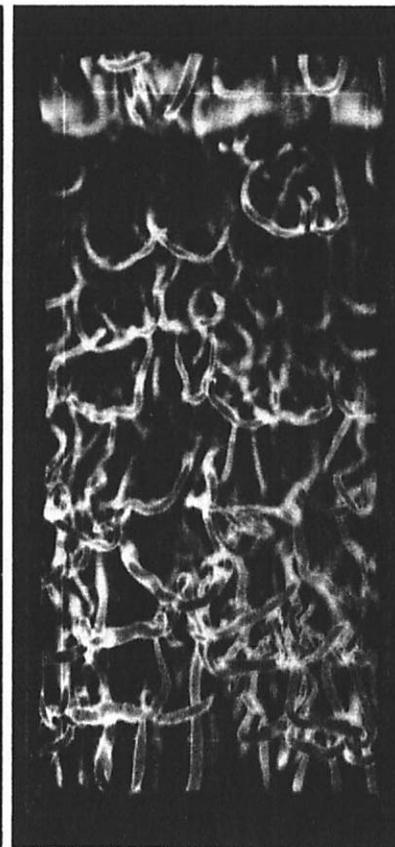
t = 65.1



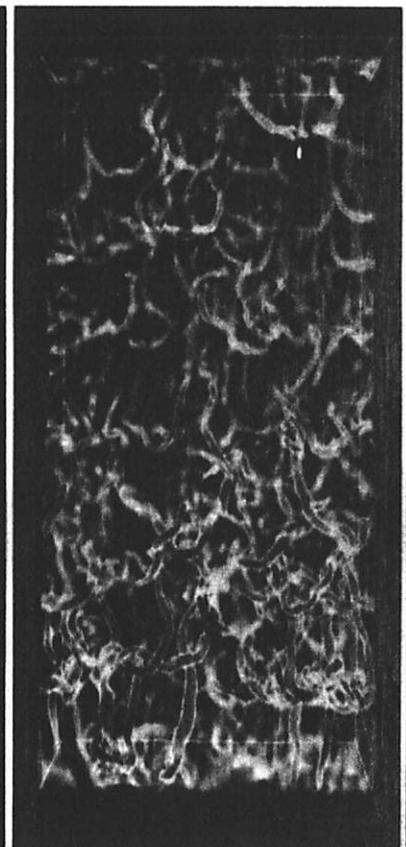
67.5



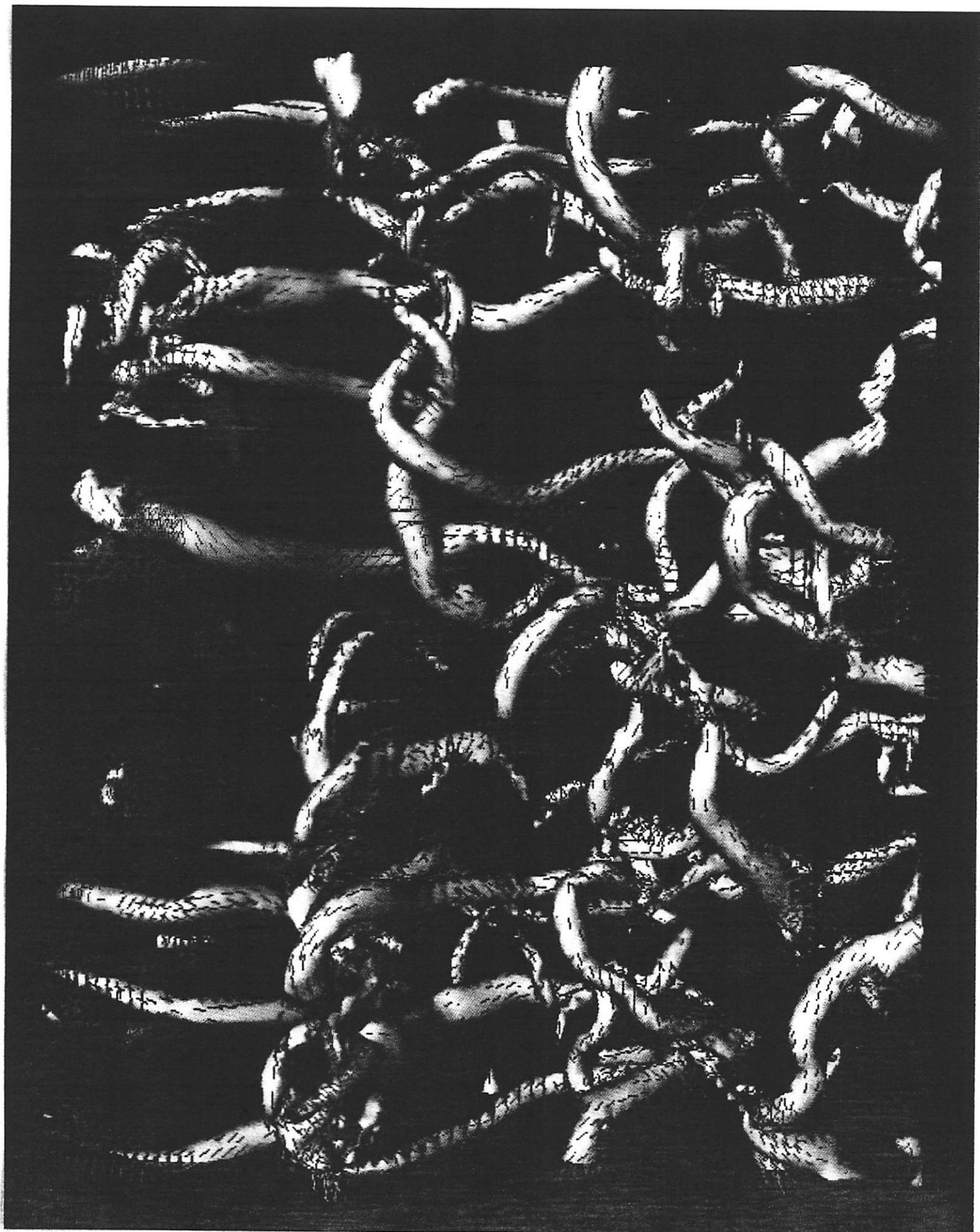
70.0



72.5



75.0



$\tau = 75$





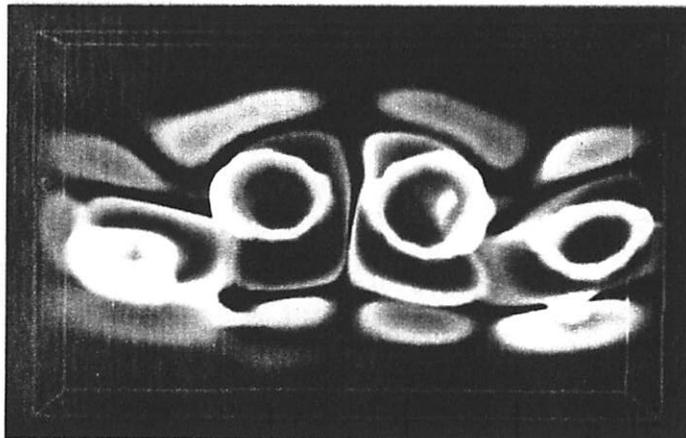
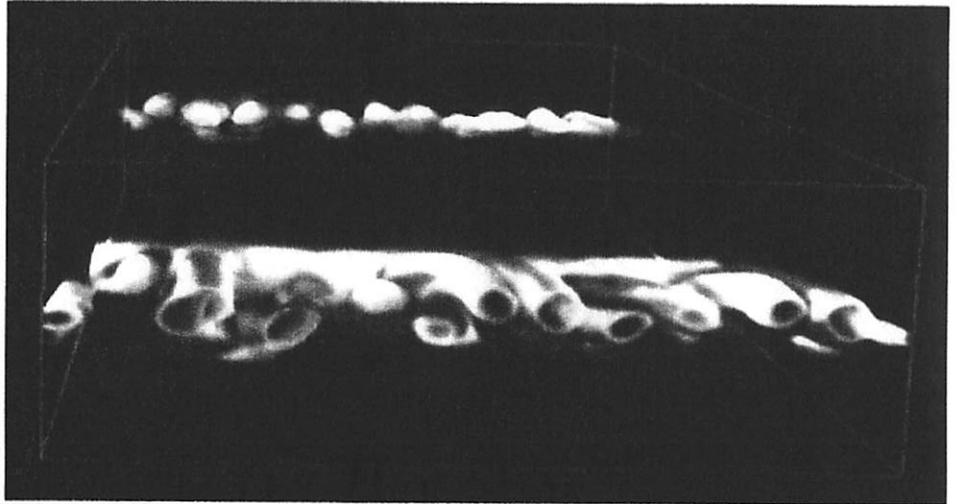
NORWEGIAN DEFENCE RESEARCH ESTABLISHMENT

Baroclinic generation of vortices at $t=62.5$

Vortices



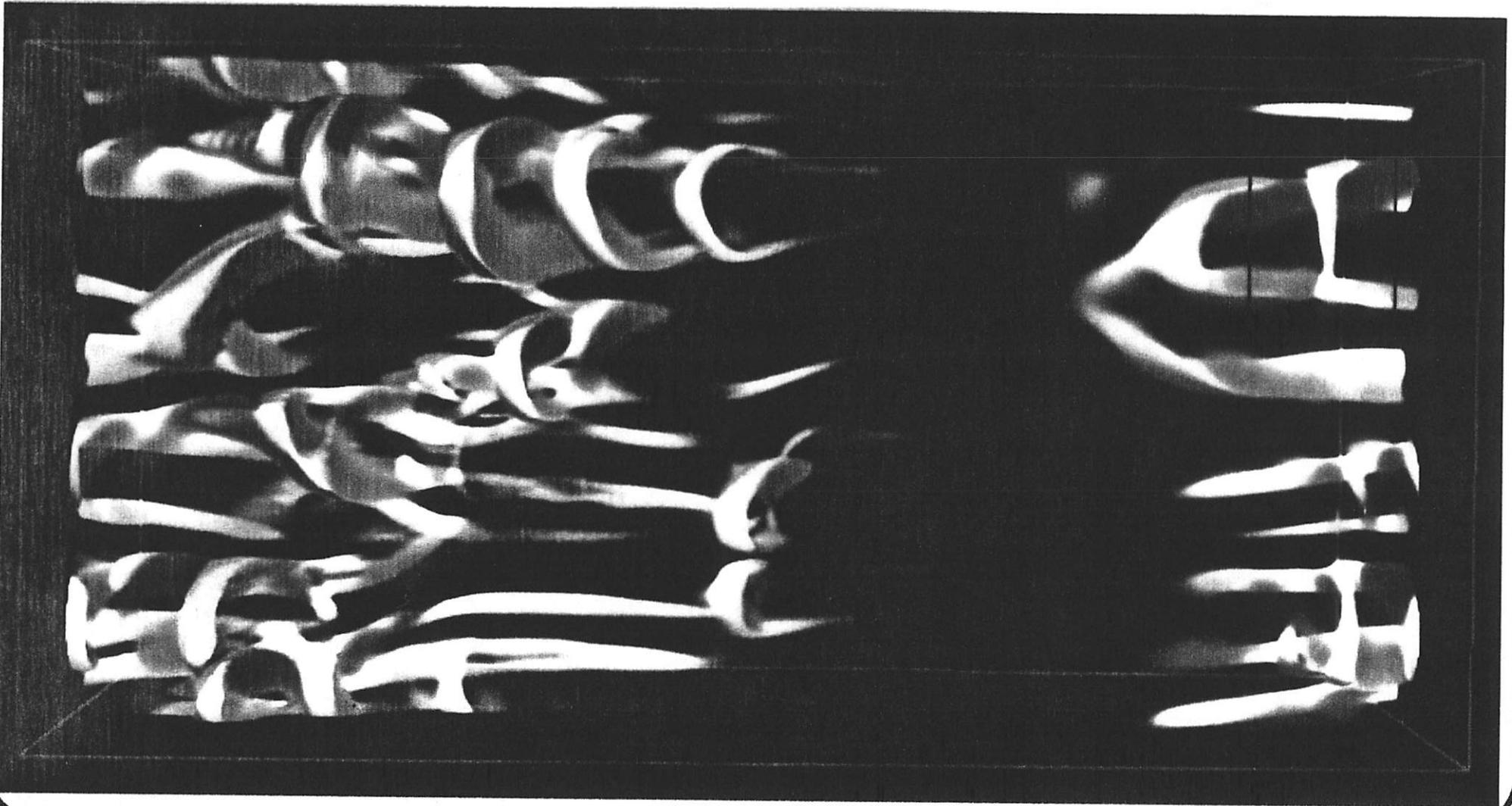
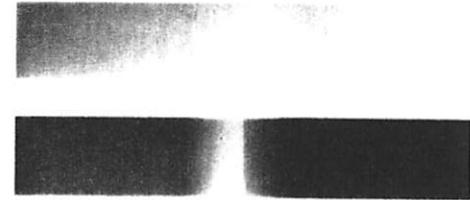
Solenoidal sources





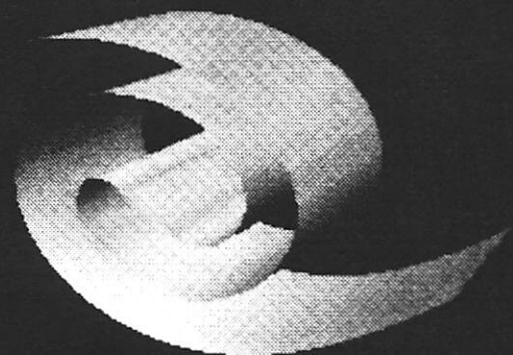
NORWEGIAN DEFENCE RESEARCH ESTABLISHMENT

**Vortices at $t=67.5$,
strain source of streamwise
vorticity $(\omega_j S_{ij})_1$**

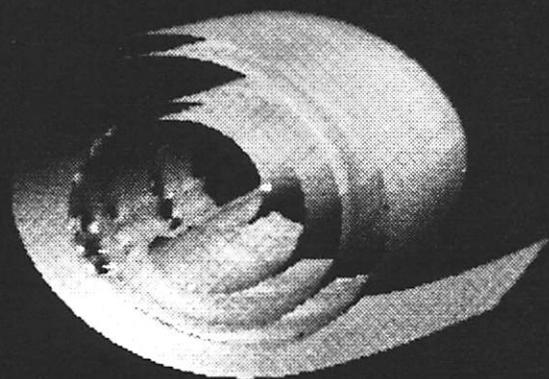


Kelvin-Helmholtz Instability

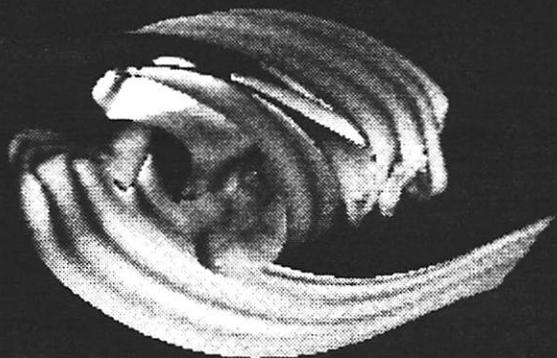
- **unstable shear flow in uniform stratification**
 - $U(z) = U_0 \tanh(z/h)$, $U_0 = 28$ m/s, $h = 300$ m
 - wavelength ~ 4 km
 - $Ri = N^2/Uz^2 = 0.05$
 - $Re = 200$ to 2000
- **KH evolutions**
 - remain 2D, $Re < 200$
 - secondary convective instability, $Re > 250$
 - secondary dynamical instability, $Re > 1000$
 - secondary instabilities
 - accelerate KH breakdown, restratification
 - mixing and transports are very different in 2D and 3D



16



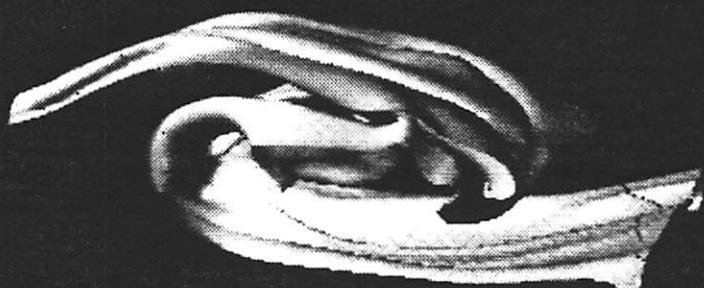
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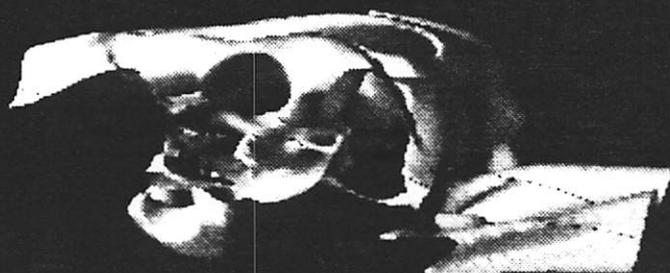
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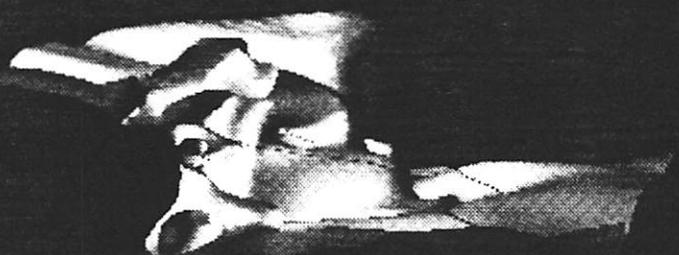
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48



56

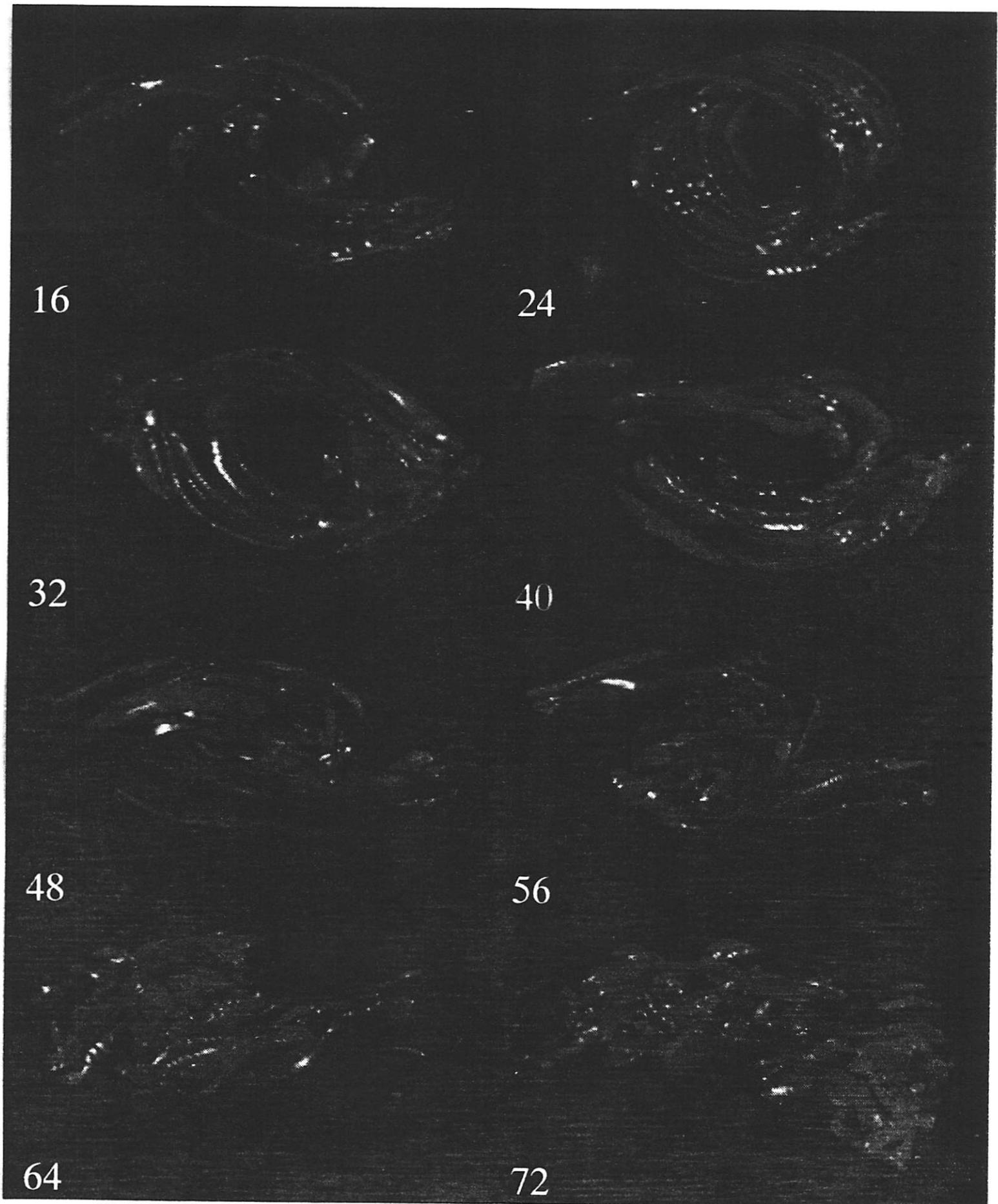


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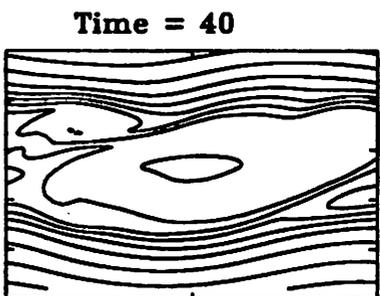
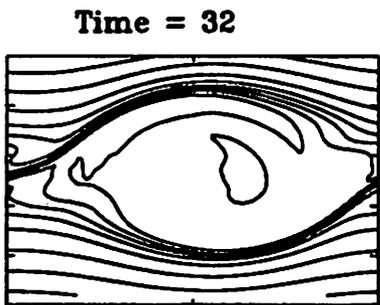
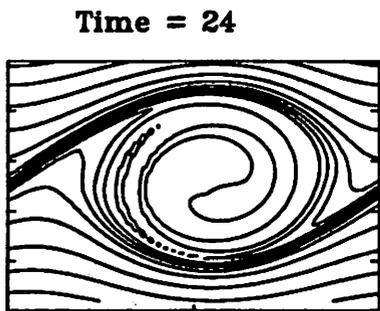
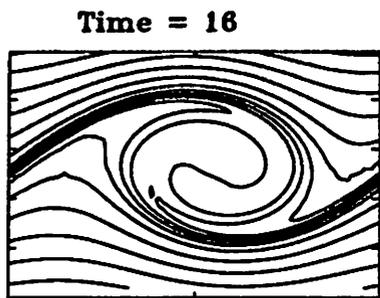
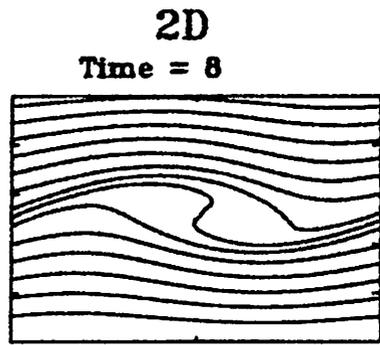
72

Contour of $\theta = 1.035$ for $Re = 500$

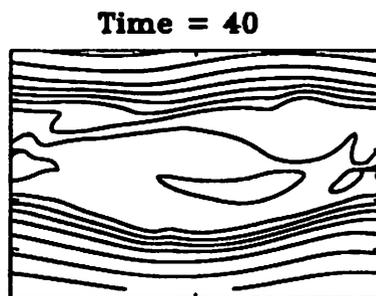
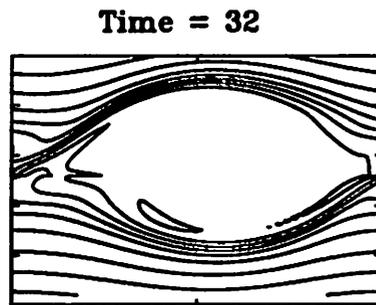
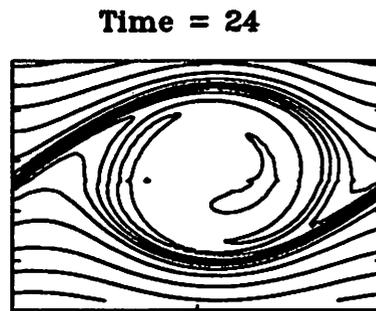
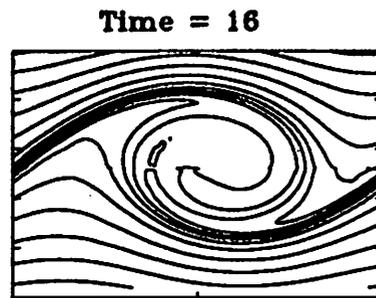
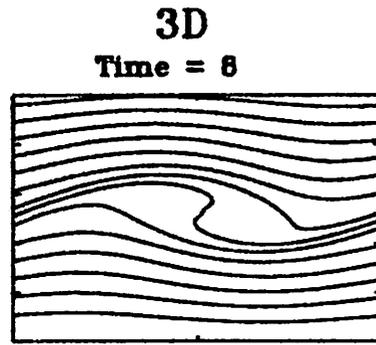


Contours of positive (red) and negative streamwise vorticity for Reynolds number = 500

Re = 500 Potential Temperature

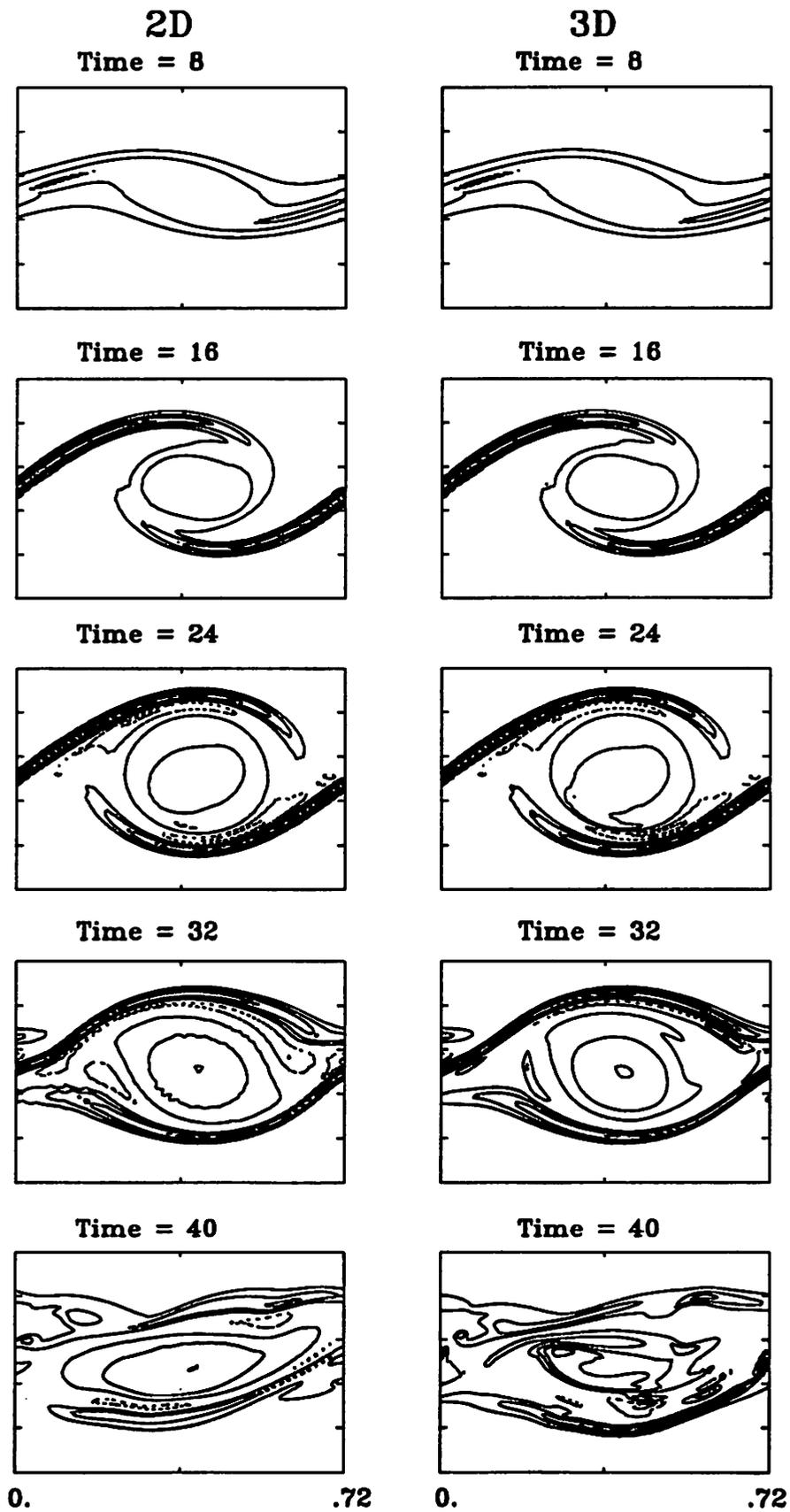


0. .72



0. .72

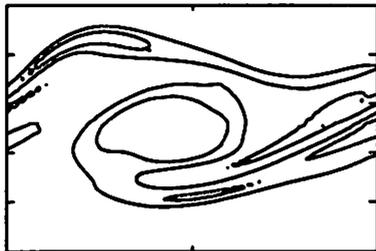
Re = 500 Spanwise Vorticity



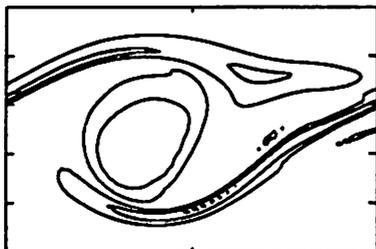
Re = 500 Spanwise Vorticity

2D

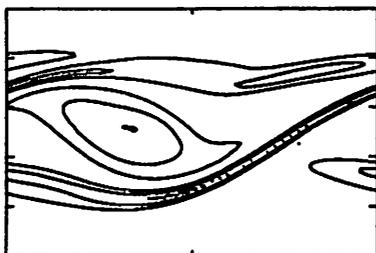
Time = 48



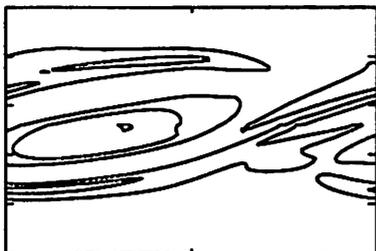
Time = 56



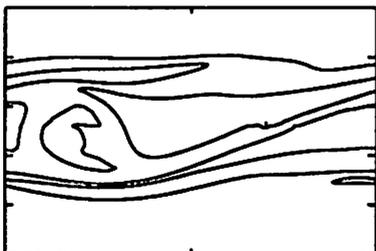
Time = 64



Time = 72



Time = 80

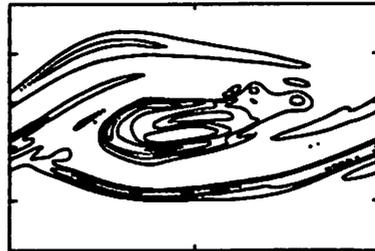


0.

.72

3D

Time = 48



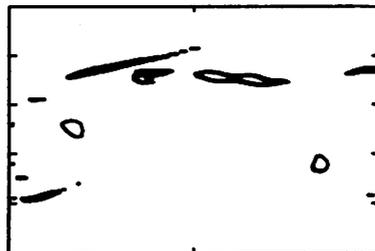
Time = 56



Time = 64



Time = 72



0.

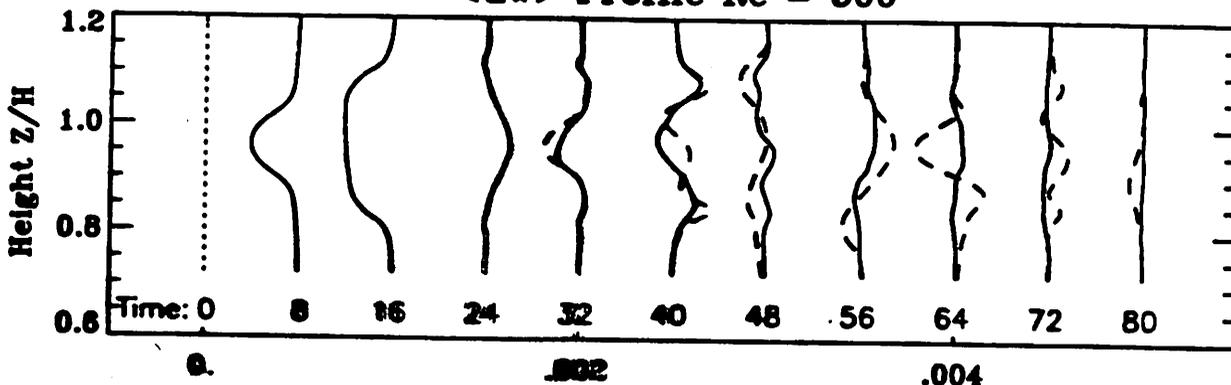
.72

KH Mean Flow Evolutions

3D — 2D ---

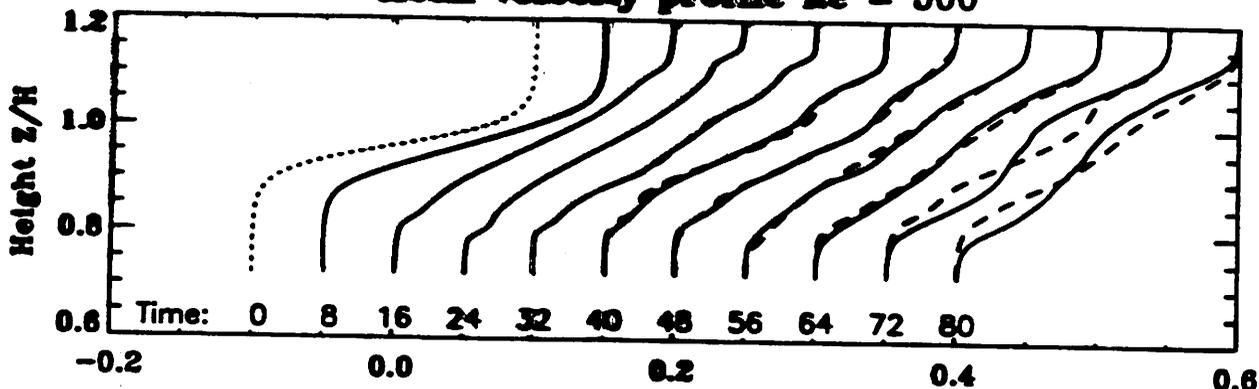
$\langle uw \rangle$ Profile Re = 500

$\overline{u'w'}$



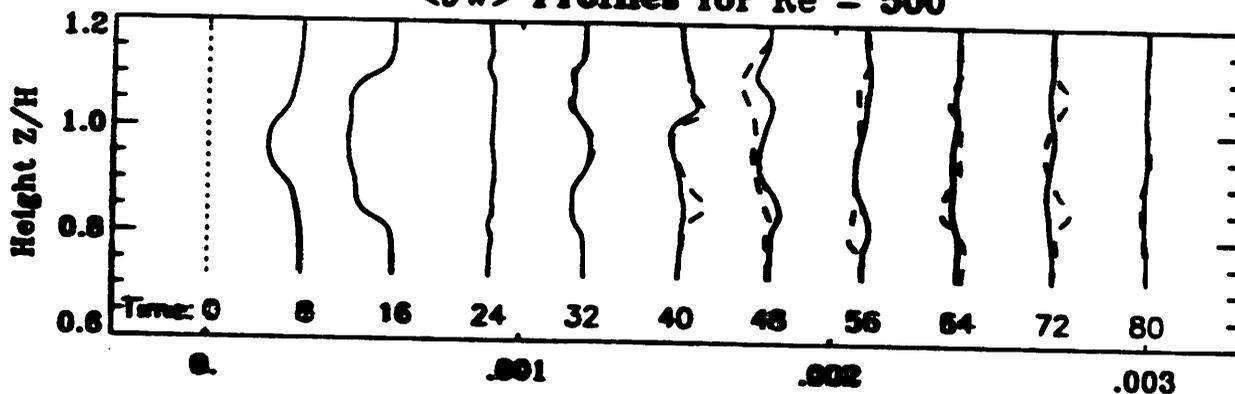
Mean velocity profile Re = 500

$\bar{u}(z)$



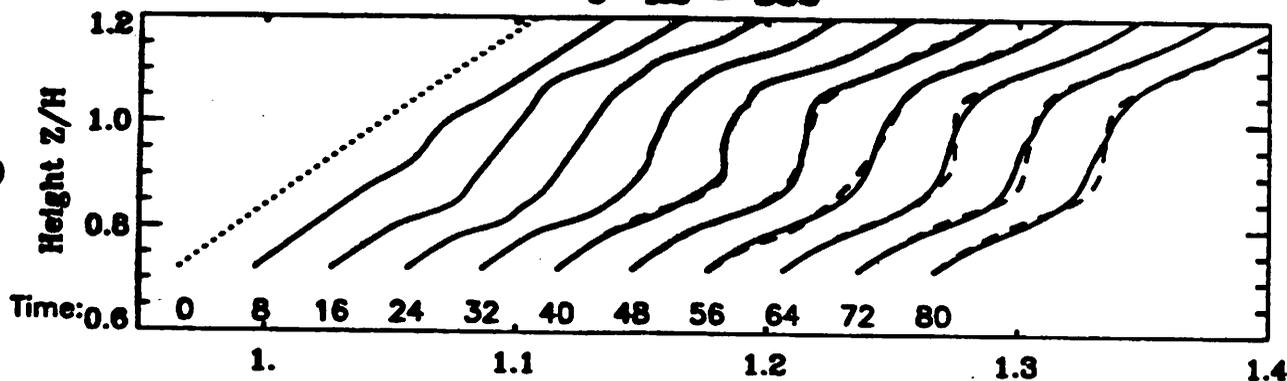
$\langle \theta w \rangle$ Profiles for Re = 500

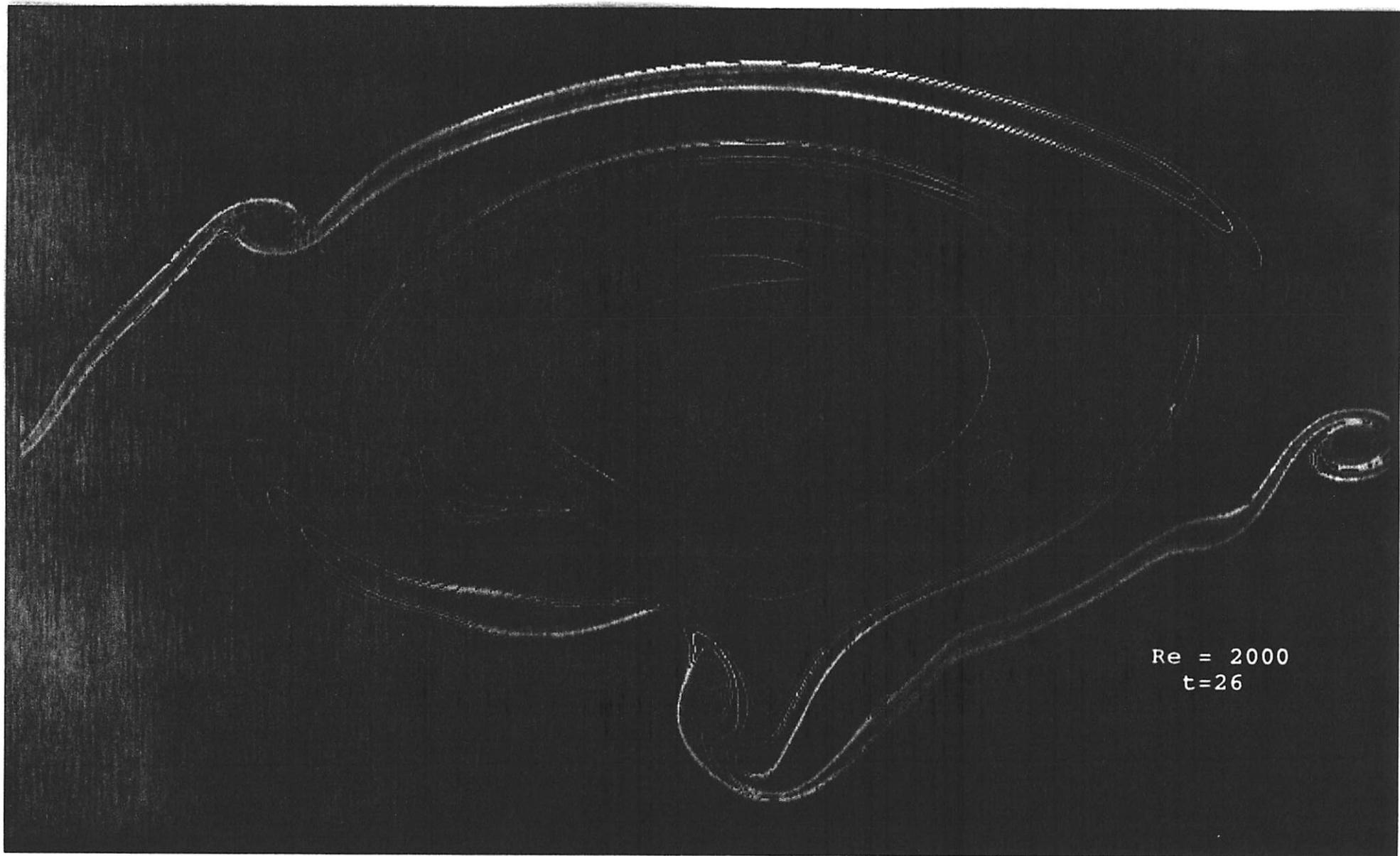
$\overline{\theta'w'}$



$\bar{\theta}(z)$ Re = 500

$\overline{\theta(z)}$





Conclusions

- **Wave breaking is inherently three dimensional**
 - **primary instability is convective in nature over large range of wave frequencies**
 - **secondary dynamical instability (KH in 3D) arises due to stretching of vortex sheets**
 - **vorticity dynamics drives transition to turbulence**
 - **intertwined vortex tubes**
 - **intense vortex interactions**
 - **vortex fraying, fragmentation => cascade of energy and enstrophy to smaller scales**
- **Kelvin-Helmholtz instability exhibits secondary instability**
 - **convective, streamwise instability, $Re > 250$**
 - **dynamical, spanwise aligned inst., $Re > 1000$**
 - **2D and 3D evolutions have very different**
 - **vorticity dynamics**
 - **implications for mixing and transports**