

1994 CEDAR Workshop
Boulder, Colorado
June 20-25, 1994

CEDAR Prize Lecture

by Raymond Roble
HAO/NCAR

**Modelling the Circulation, Temperature and
Compositional Structure of the Upper
Atmosphere (30-500 km)**

TGCM TEAM

HAO

E. Cicely Ridley	-	Mathematical analysis and development
Ben Foster	-	Processors, display and diagnostics
Art Richmond	-	Electrodynamics, AMIE
Barabara Emery	-	Auroral Parameterizations, AMIE
Bill Roberts	-	AMIE displays
Gang Lu	-	AMIE/TGCM runs and analysis
Maura Hagan	-	TGCM/DATA analysis and tidal boundaries

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COLLABORATORS

CASANDRA FESEN

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JEFF FORBES

DAVE FRITTS

VICTOR FOMICHEV

TIM KILLEEN, ALAN BURNS, PAUL HAYS

FRED REES

GONZALO HERNANDEZ

GEOFF CROWLEY

FRANK MARCOS

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ETC., ETC., ETC.

Tom Holzer/HAO/NSF/NASA/SPACE PHYSICS THEORY/
AIR FORCE

LONG RANGE MODEL DEVELOPMENT

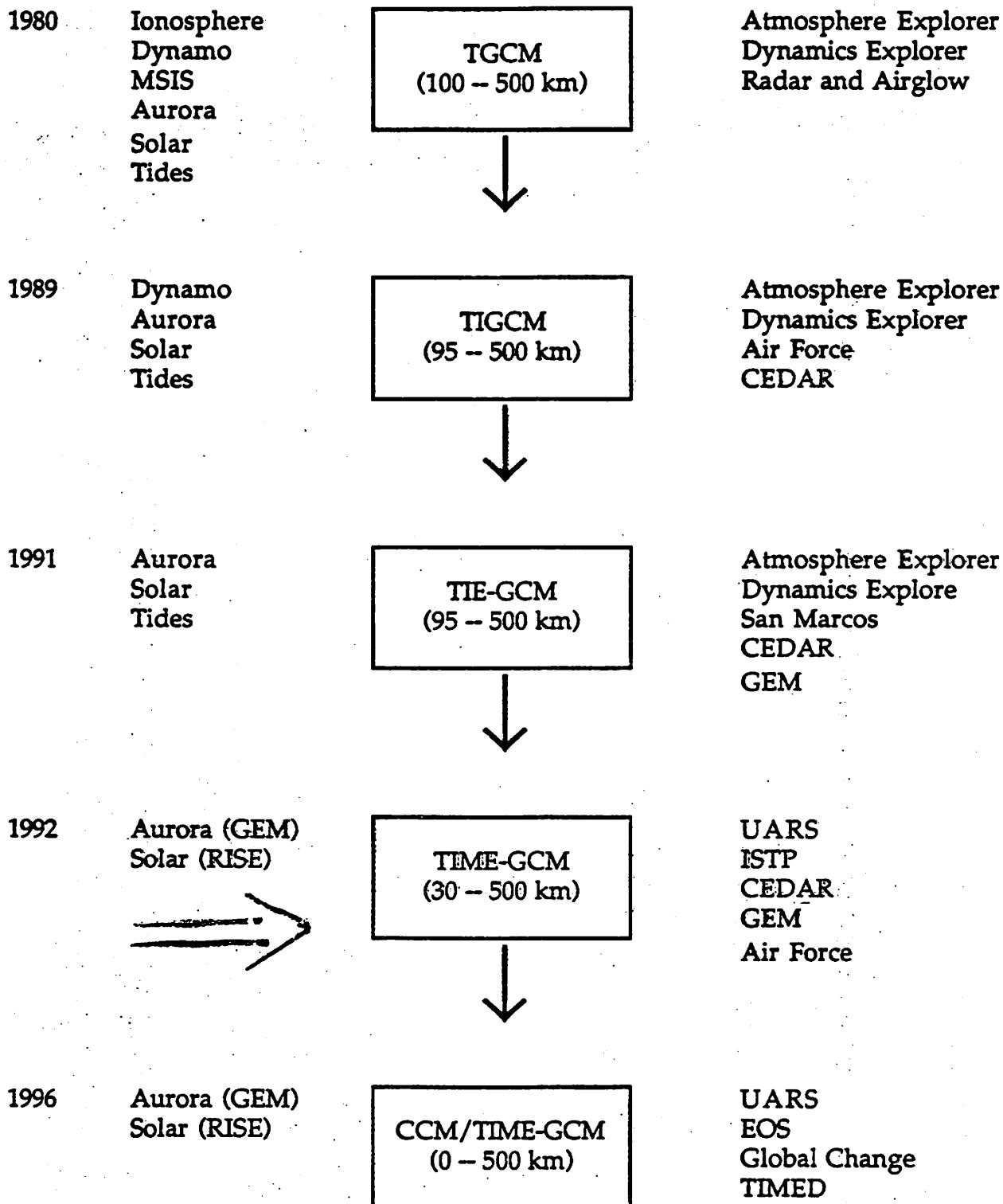


Fig. 1 Schematic illustrating past and future TGCM development. The year of model development and diminishing dependence on empirical specification is given on the left of the boxes and the programs and data sources used for GCM validation and scientific studies is given on the right. CCM refers to the NCAR community climate model.

TIE-GCM OUTPUT

- Neutral gas temperature, T_n
- Neutral winds, U, V, W
- Height of constant pressure surface, h
- Neutral composition and density

Major - O, O_2 , and N_2

Minor - $N(^2D), N(^4S), NO, He$, and Ar

- Ion composition

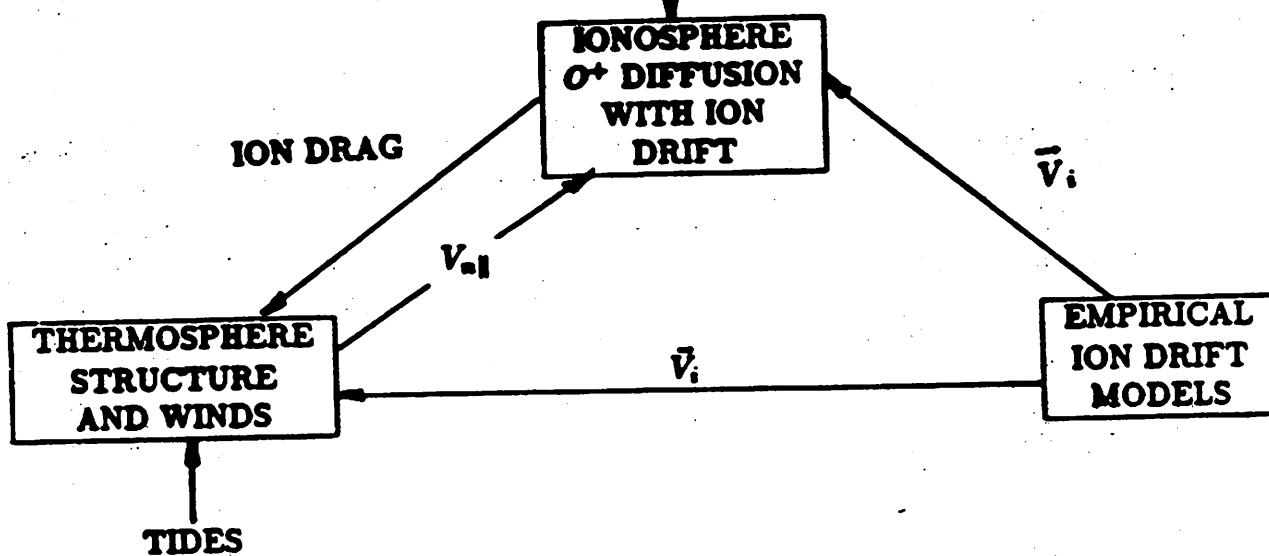
$O^+(^2P), O^+(^2D), O^+(^4S), NO^+, O_2^+, N_2^+$,
and $N^+ \sum n_i = n_e$

Ion temperature, T_i

Electron temperature, T_e

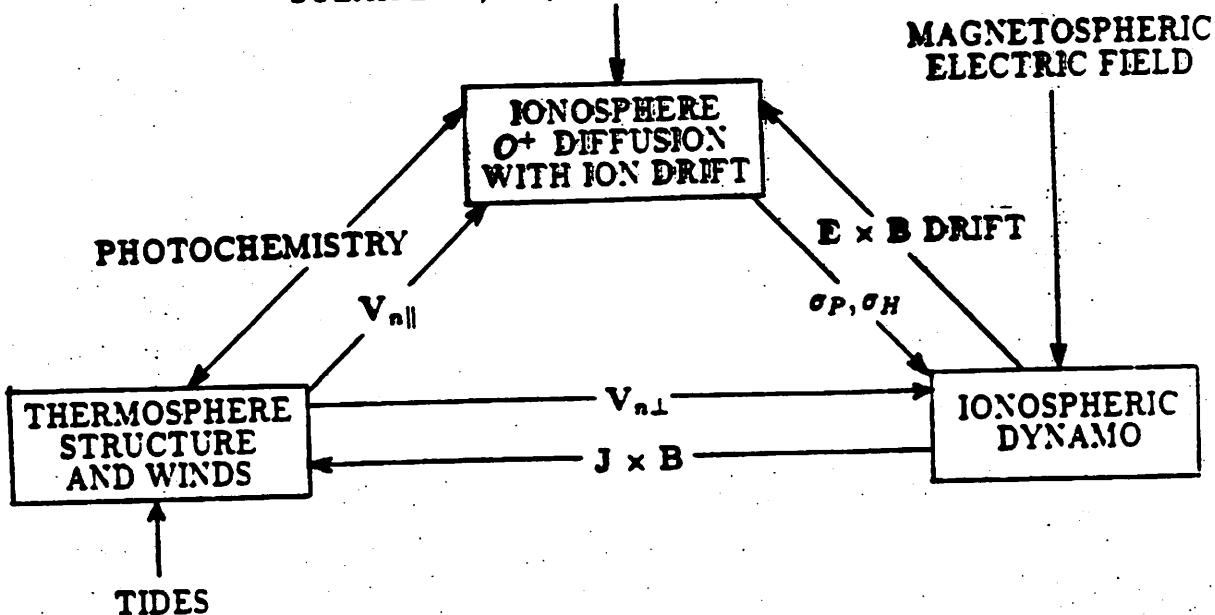
- Global distribution of electric fields \vec{E} and currents \vec{J}

SOLAR EUV, UV, AND AURORAL INPUTS



TIE-GCM

SOLAR EUV, UV, AND AURORAL INPUTS



STANDARD MODEL.

TIE-GCM
350 km, 0 UT, Equinox, Solar Maximum

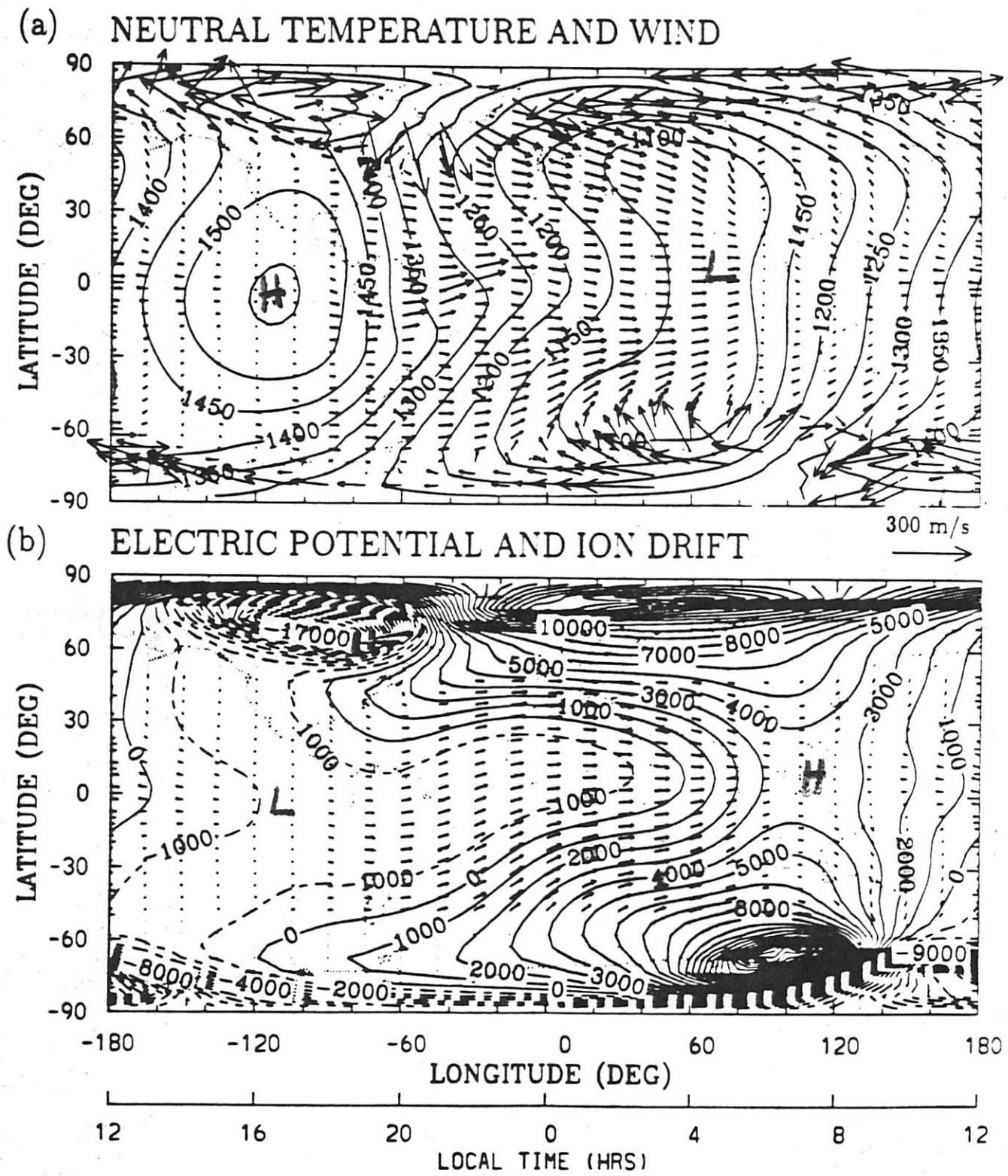


Figure 2

NCAR
THERMOSPHERE-IONOSPHERE-ELECTRODYNAMICS
GENERAL CIRCULATION MODEL
(TIE-GCM)

- Primitive equations of dynamic meteorology adapted to thermospheric heights
- Eulerian model of the ionosphere
- Horizontal grid 5° latitude $\times 5^\circ$ longitude, geographic
- Vertical grid – 25 constant pressure surfaces, 2 grid points per scale height, 95 to 500 km
- Time step 240 or 300 S
- Global dynamo with realistic magnetic field

INPUT

- Solar EUV and UV radiation 5 to 250 nm
- Empirical ionospheric convection and auroral particle precipitation models
- Structure of upward propagating tides and other features from middle atmosphere

THERMOSPHERE-IONOSPHERE-MESOSPHERE-ELECTRODYNAMICS GENERAL CIRCULATION MODEL (TIME-GCM)

- TIE-GCM is extended downward to 30 km (10 mb) altitude
- Major species transport for $\underline{O_2}$, $\underline{N_2}$, and $O_X(\underline{O} + \underline{O_3})$
- Minor species transport for H_2O , H_2 , CH_4 , CO , CO_2 , $NO_X(\underline{NO} + \underline{NO_2})$,
 $\underline{N(^4S)}$, $HO_X(H + HO_2 + OH)$
- Minor species in photochemical equilibrium for $\underline{O(^1D)}$, $\underline{N(^2D)}$, H_2O_2 ,
 $O_2(^1\Delta_g)$, and $O_2(^1\Sigma_g)$
- NLTE- CO_2 radiational cooling (Fomichev parameterization); $O_3(9.6\mu m)$
- Fritts gravity wave parameterization
- Boundary conditions
 - (a) Lower: specified mixing ratios for O_2 , N_2 , H_2O , H_2 , CH_4 ,
 CO_2 , CO , O_X , and NO_X G-S model - 10 mb
 - (b) Upper: diffusive equilibrium for all species except H where
exospheric escape is specified
 - (c) Dynamics based on tidal theory at various frequencies (10 mb)
• NMC 10mb GEOPOTENTIAL HEIGHT DATA
- D-region ion chemistry
- Energetic electron and proton particle degradation

MAJOR TIME-GCM OBJECTIVES

- How deep into the atmosphere do effects of solar variability penetrate?
- How do disturbances from lower atmosphere affect middle (30 - 100 KM) and upper atmosphere (100 - 500 KM)?
 - + Couplings
 - * Solar Radiative
 - * Auroral
 - * Chemical
 - * Radiative
 - * Electrical
 - * Dynamic
- How will the upper atmosphere and ionosphere respond to global change?

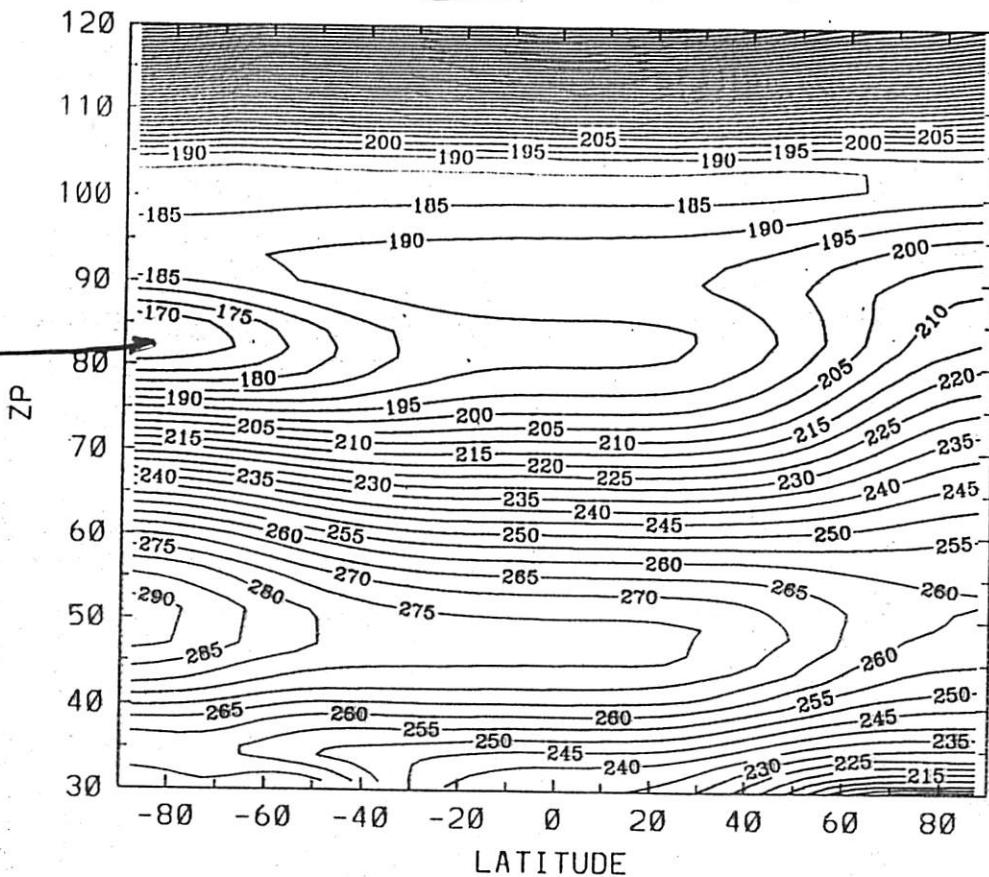
DEC. SOLSTICE

TN ZONAL MEANS UT= 0.00

FR

RAYLEIGH
FRICTION

165K



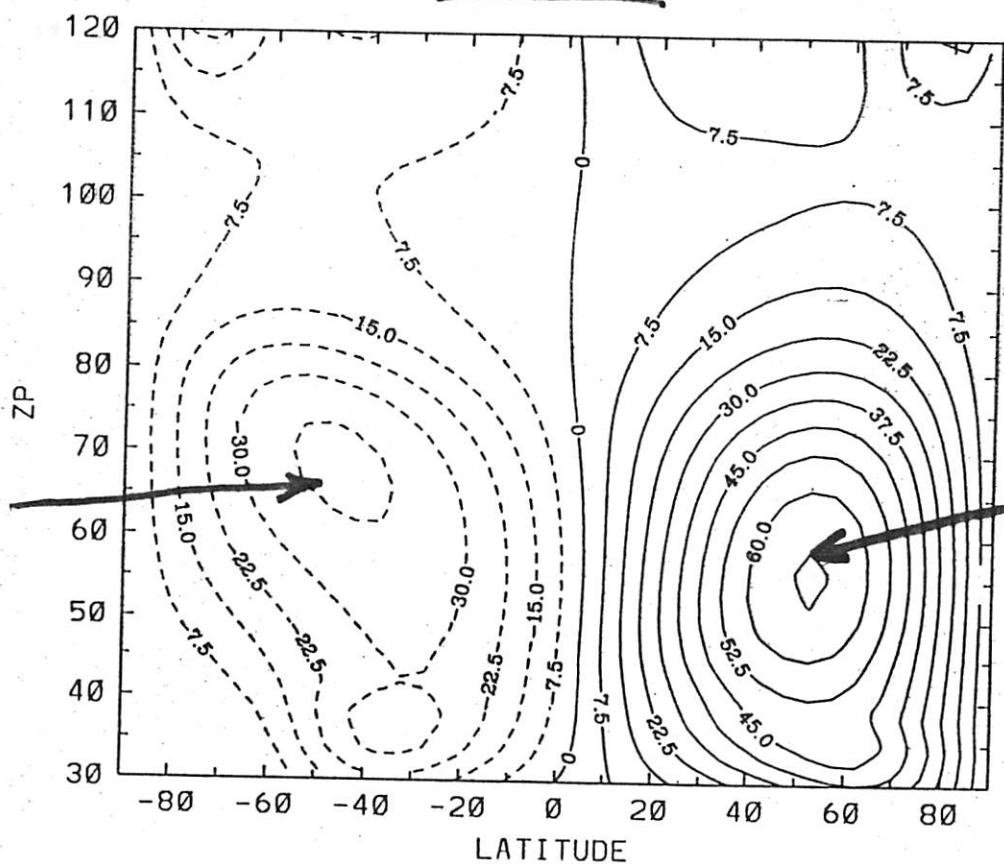
K

UN ZONAL MEANS UT= 0.00

m/s

-40

70 m/s

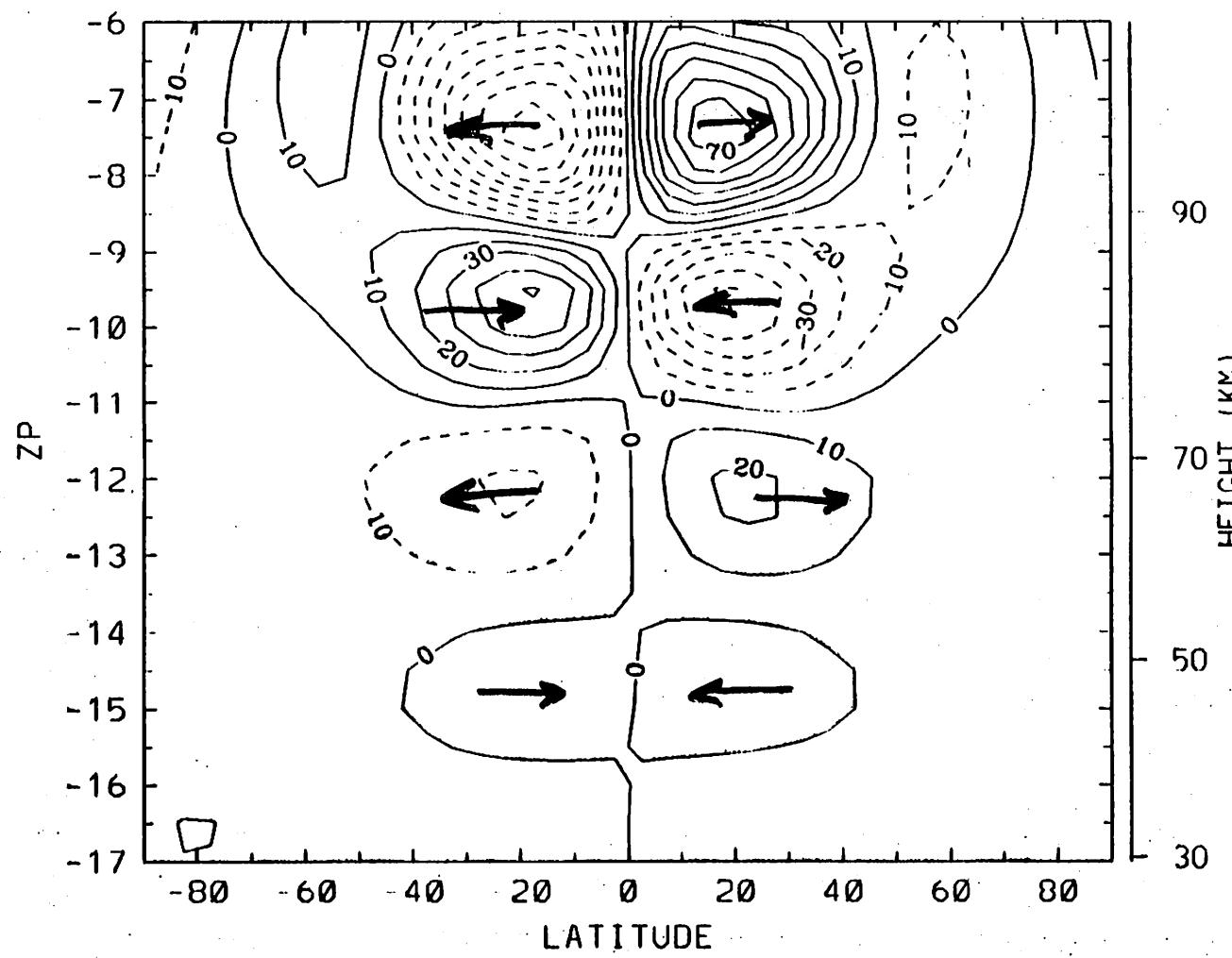


V (POSITIVE NORTHWARD)
(m/s)

VN

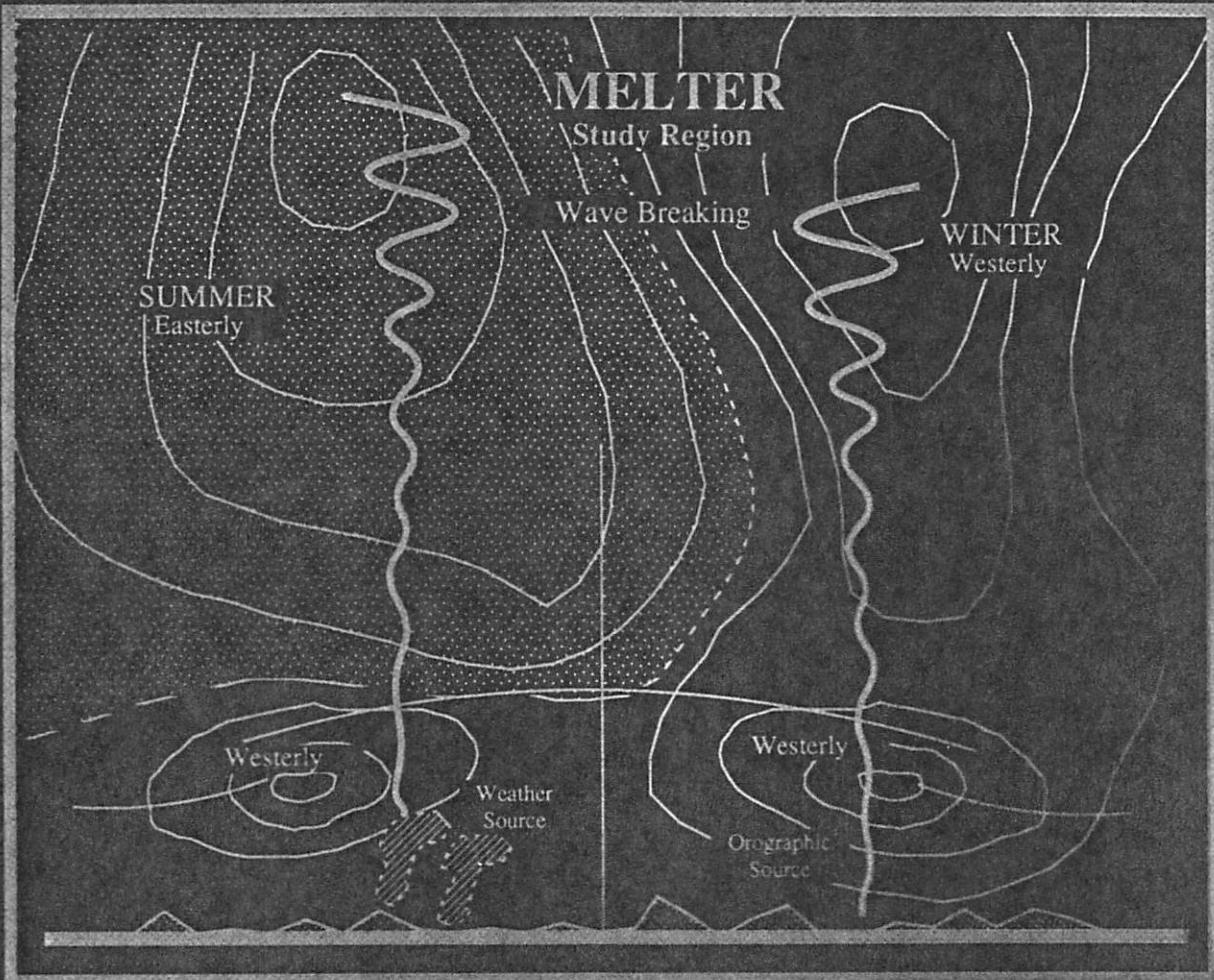
LON = -180.00 UT = 0.00

NOON



MIN -76.24, MAX 77.43, INTERVAL 10

HISTORY = /ROBLE/RGR92/TISEZ3

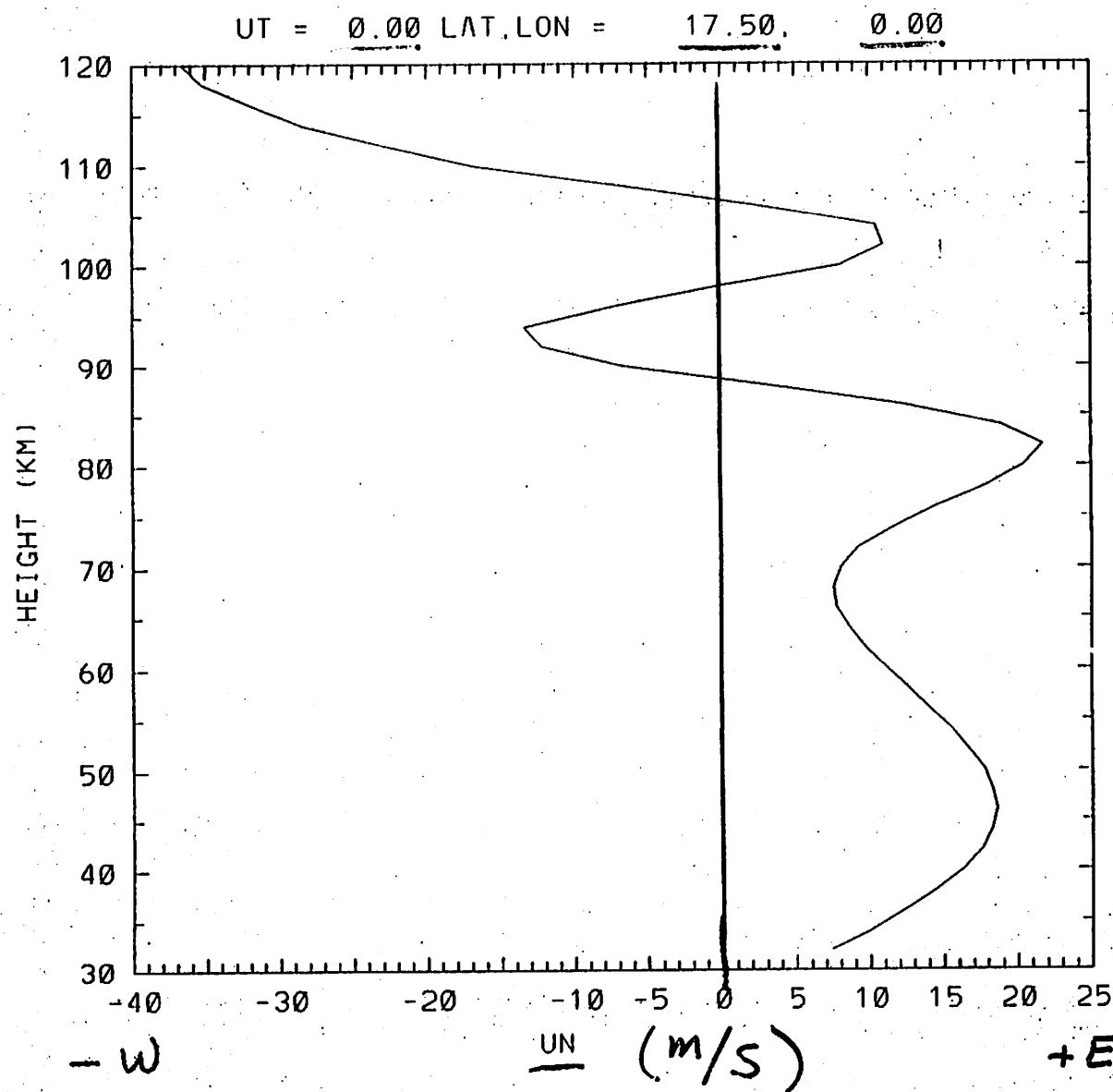


Gravity Wave Disturbances

Gravity Waves Propagate
If They Have The Correct Phase Speed.
Thus They Are Present Both In The Winter and Summer
Stratosphere And Mesosphere.

EQUINOX

ZONAL WIND

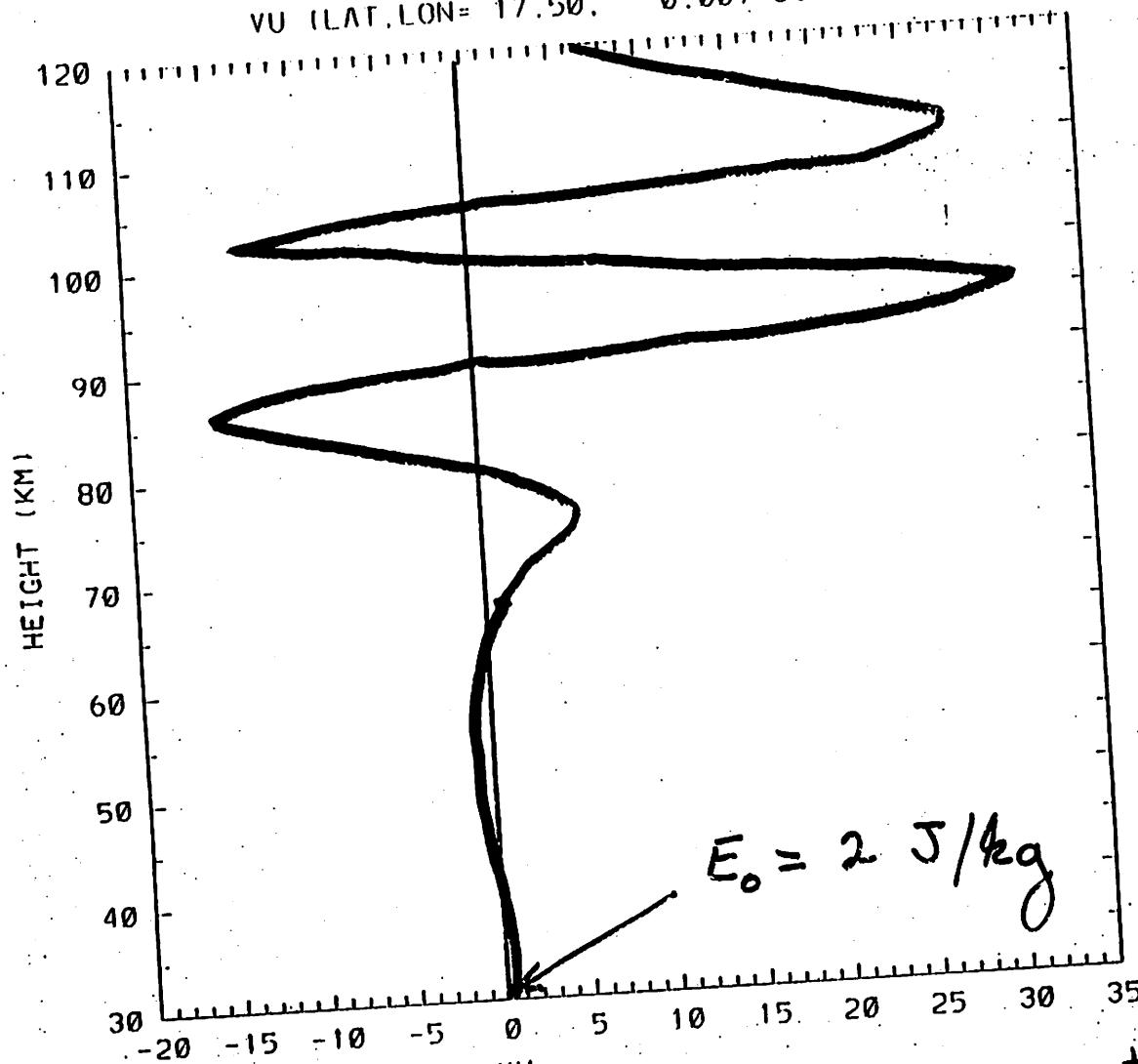


HISTORY=/ROBLE/RGR94/T15GAR3

EQUINOX

ZONAL WIND ACCELERATION

VU (LAT, LON = 17.50, 0.00) $u_l = 0.00$



-W

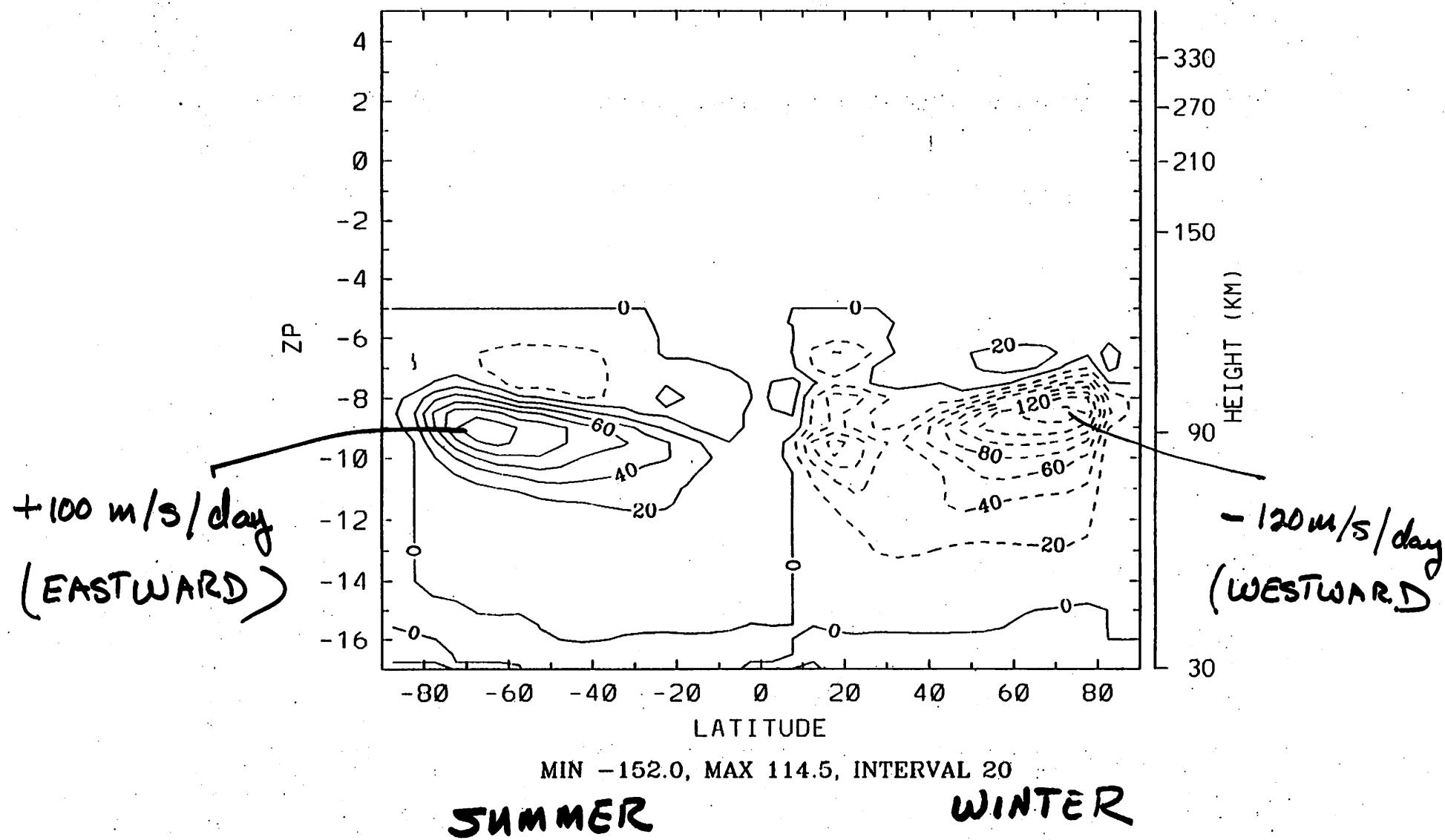
$A(u)$ m/s/day +E

GRAVITY WAVE DRAG (m/s/day)

DEC.
SOLSTICE

F_4

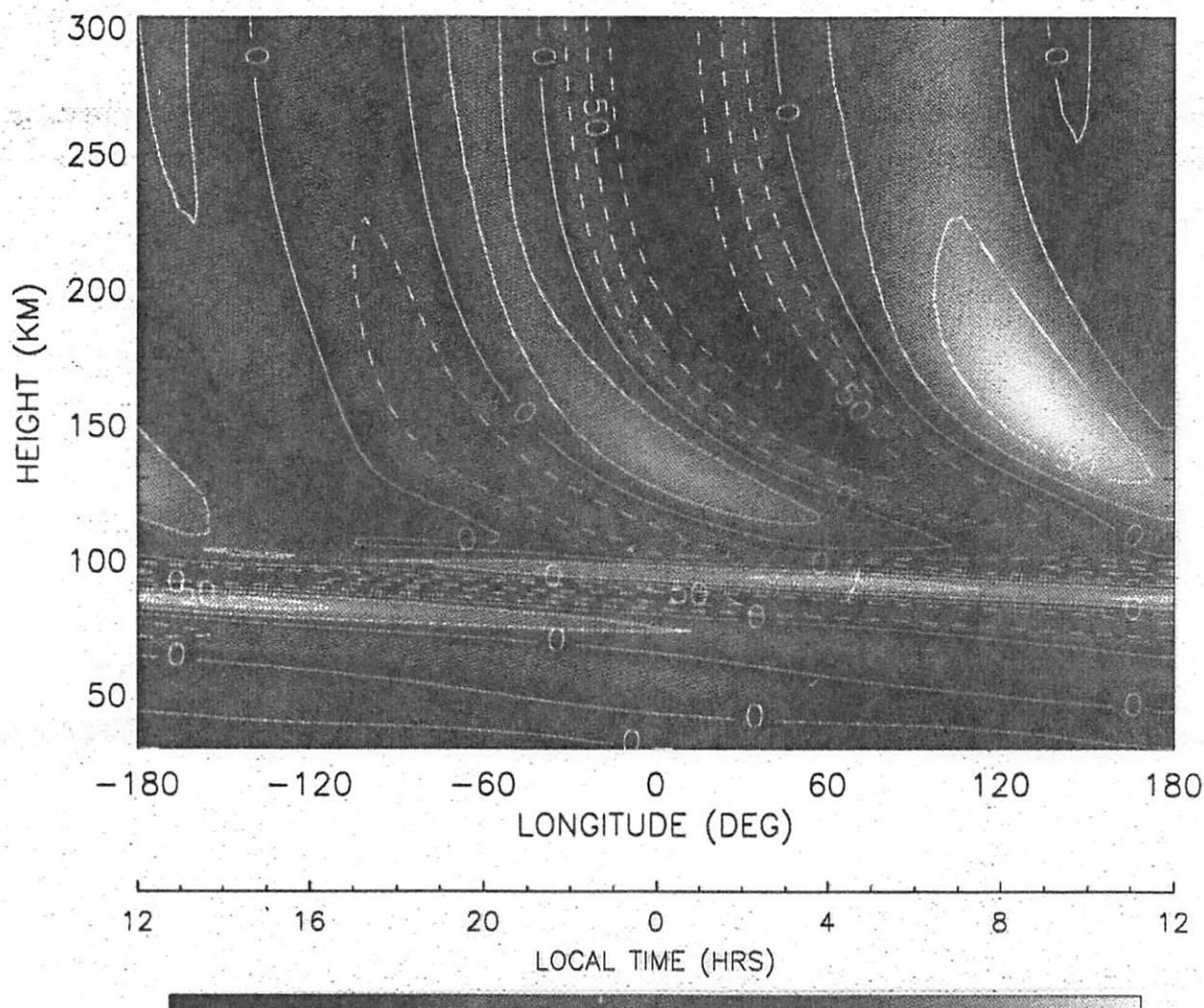
VU ZONAL MEANS UT = 0.00 (TISOG9)



V_n (meridional
wind m/s)

TIMESGCM VN

LAT = 17.5 UT = 0.0



CONTOUR -100. TO 100. BY 25.0 (MIN,MAX=-99.2,69.2)

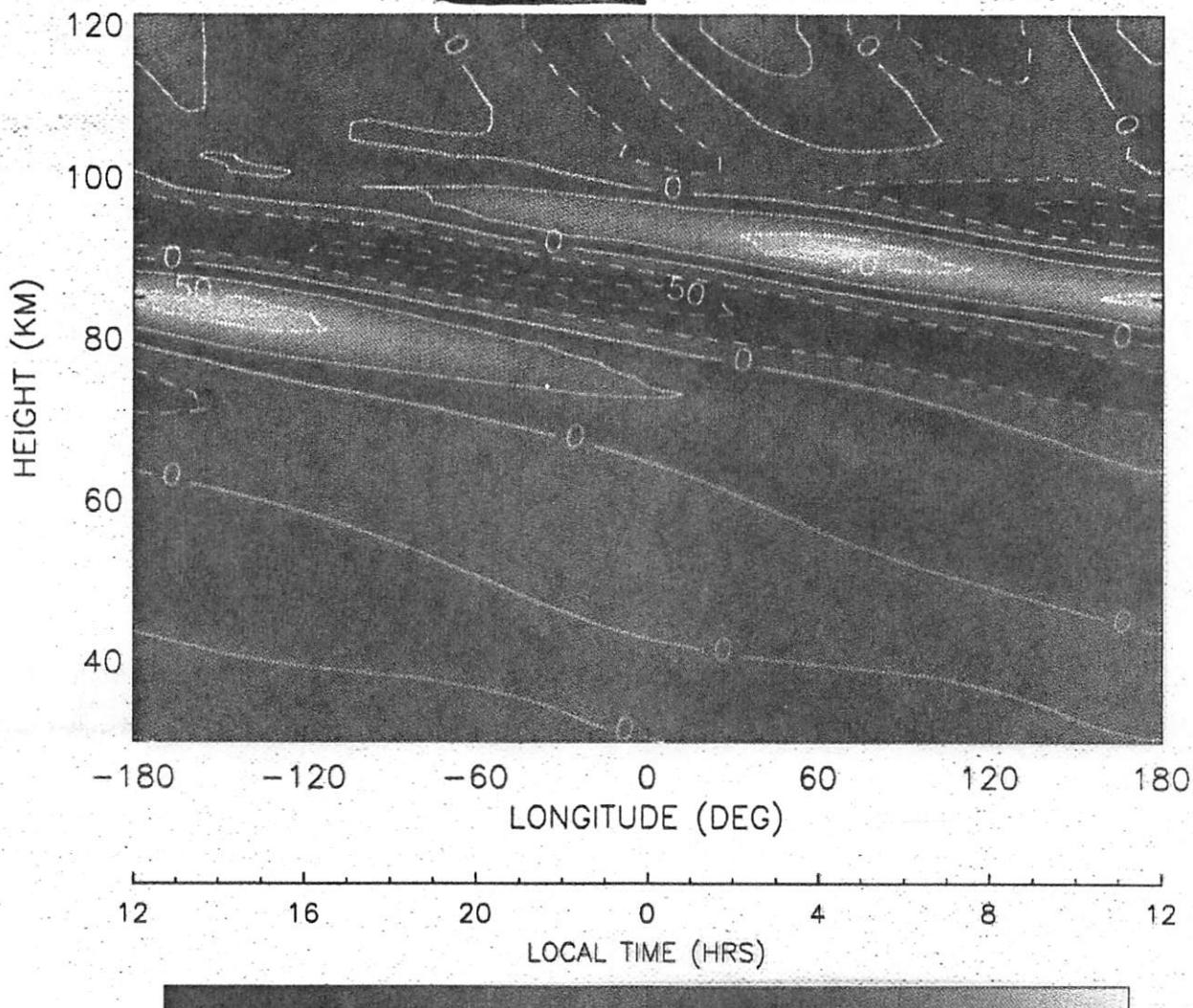
HISTORY /ROBLE/RGR93/TSEQBZ4 (6: 0: 0)

EQUINOX SMIN

V_u (meridional
wind m/s)

TIMESGCM VN

LAT = 17.5 UT = 0.0



-59.36

0.7361

60.84

CONTOUR -100. TO 100. BY 25.0 (MIN,MAX=-59.4,60.8)

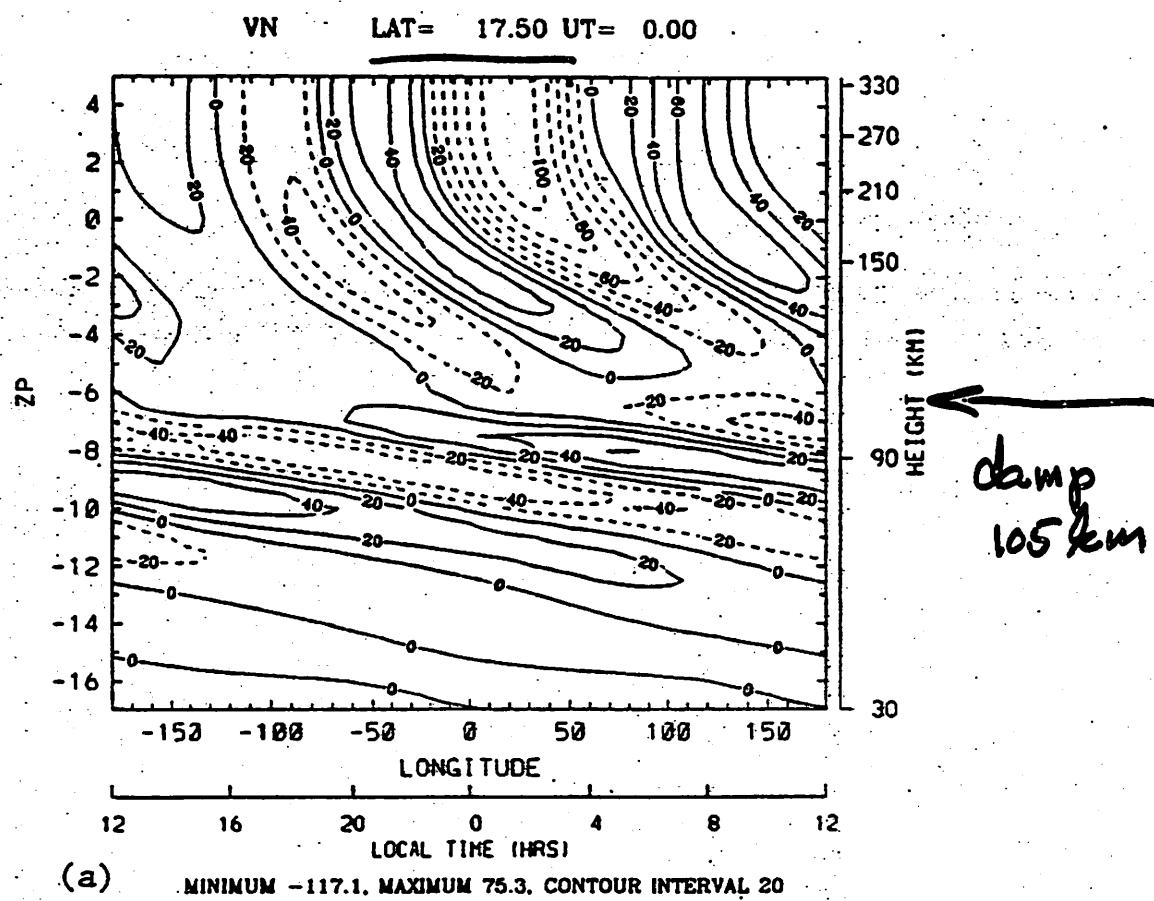
HISTORY /ROBLE/RGR93/TSEQBZ4 (6: 0: 0)

EQUINOX SOLAR MIN.

WEAK G.W.

SOURCE

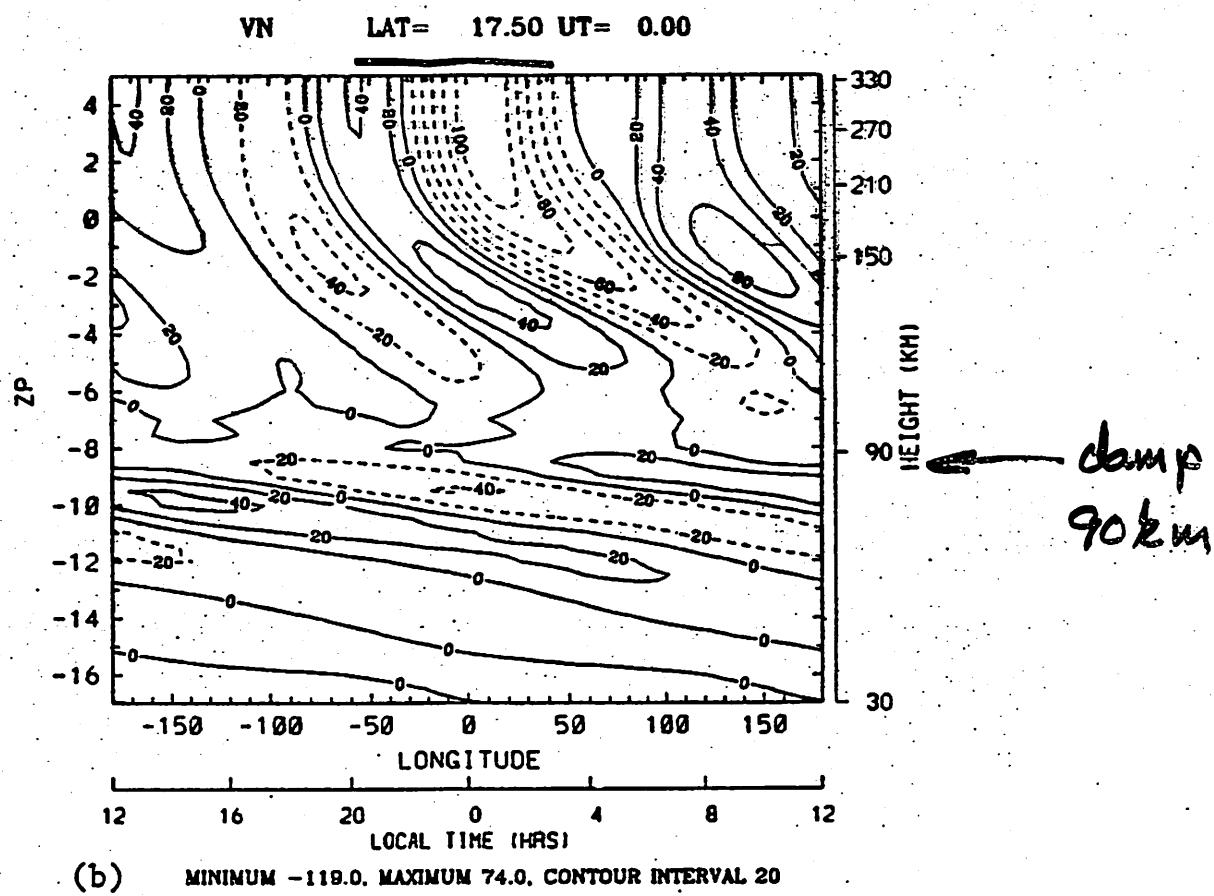
2 J/kg.



Strong G.W.

SOURCE

4 J/kg.



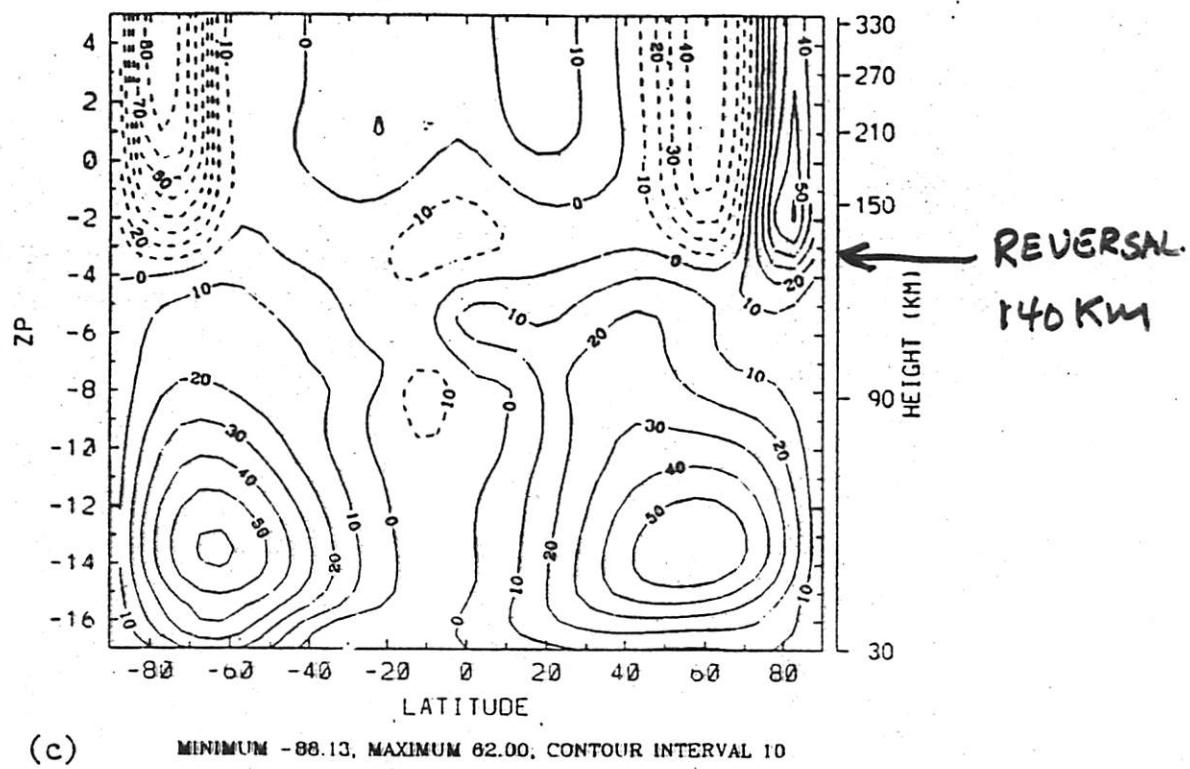
U_n ZONAL MEAN EQUINOX SOLAR MIN,
(m/s)

UN ZONAL MEANS UT= 0.00

WEAK G.W.

SOURCE

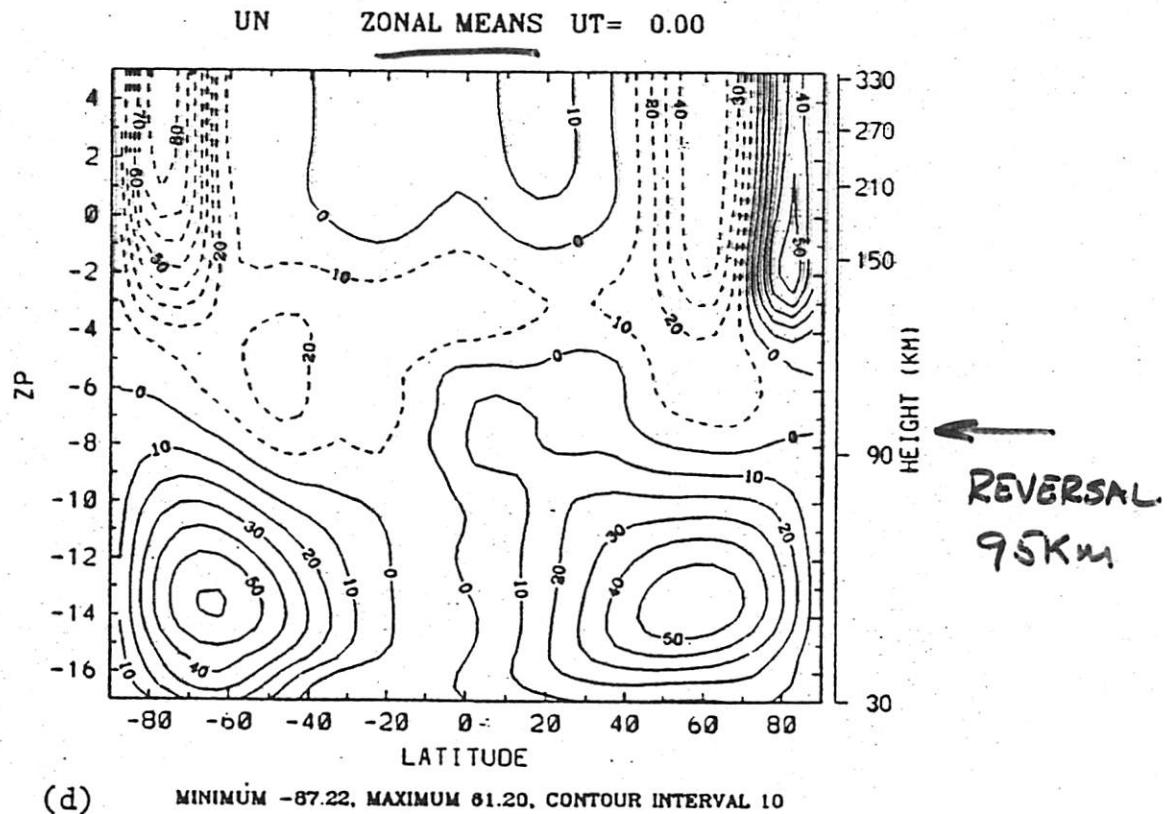
2 J/kg.



Strong G.W.

SOURCE

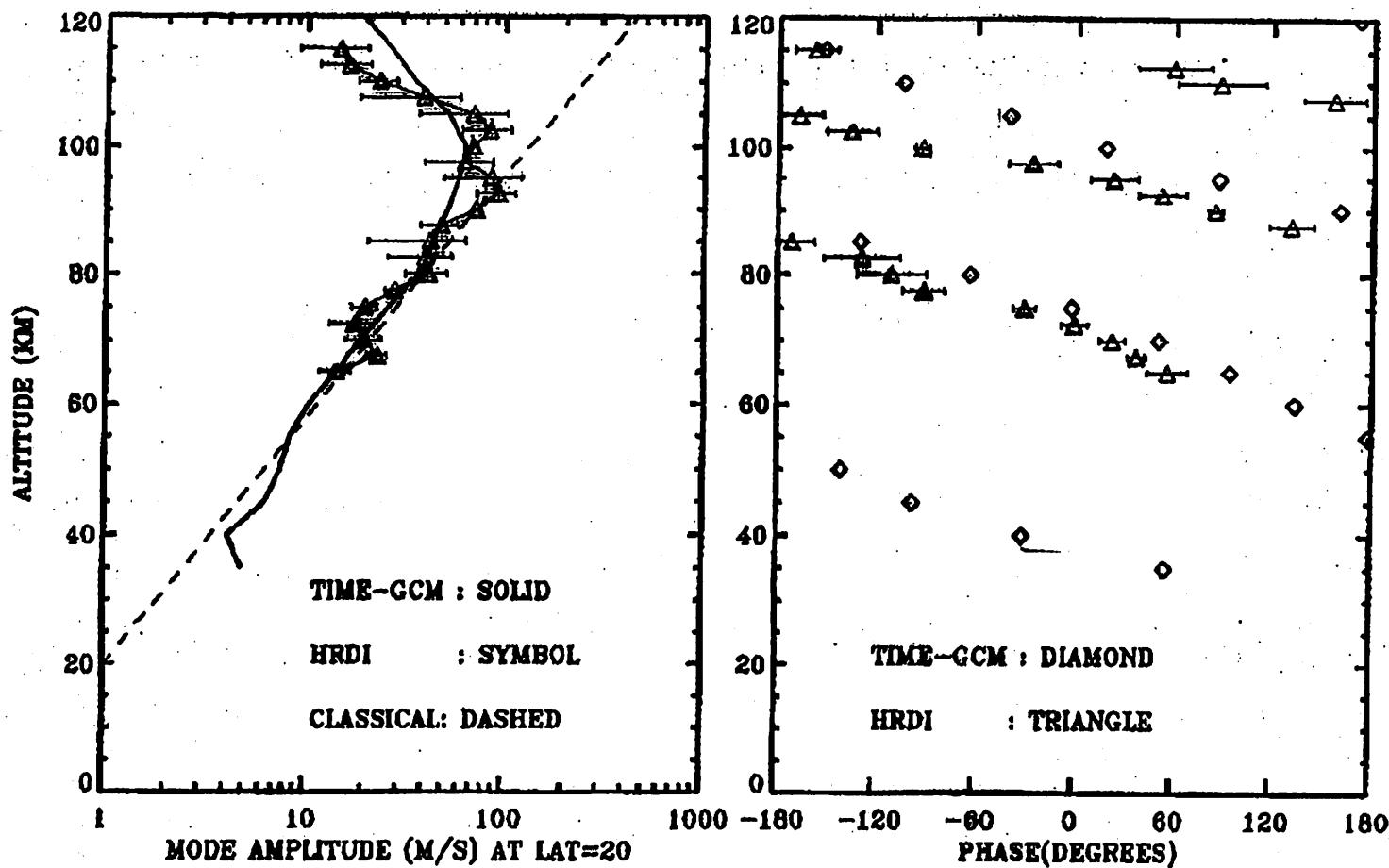
4 J/kg.



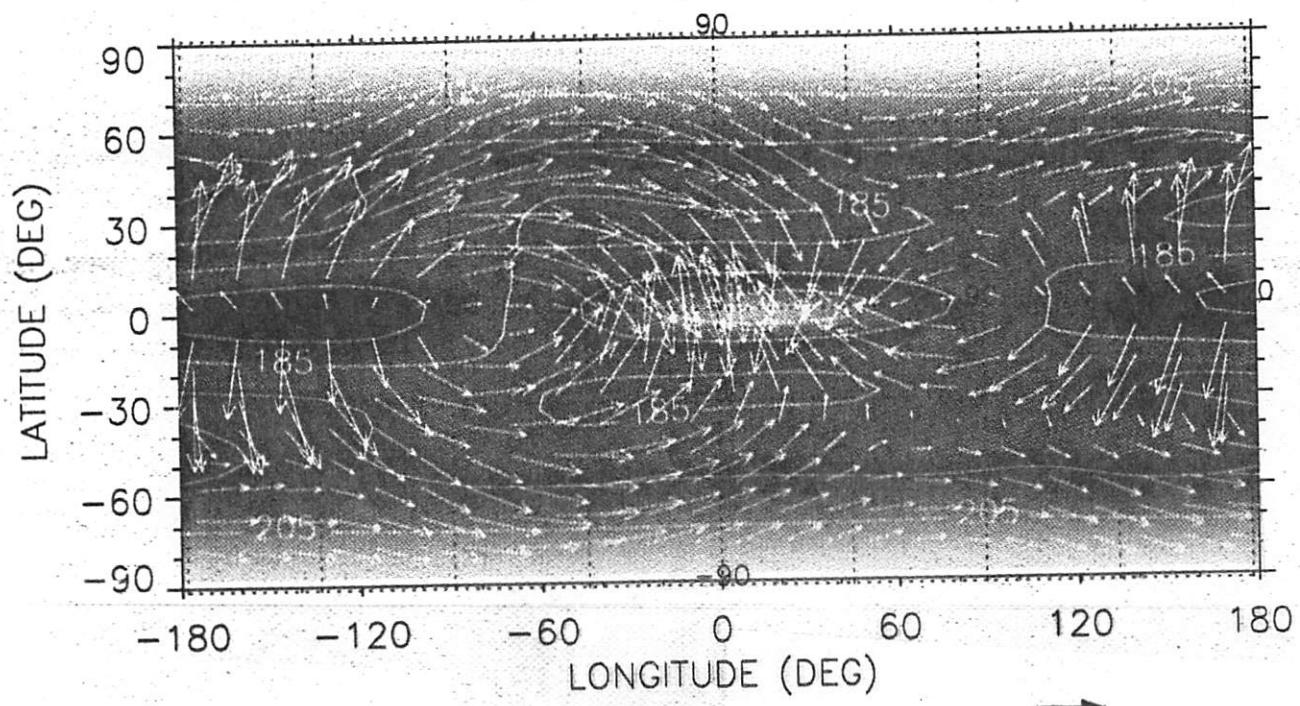
D Wu - U. of Michigan

(2)

DIURNAL (1,1) TIDE AT SPRING EQUINOX
(MERIDIONAL COMPONENT)

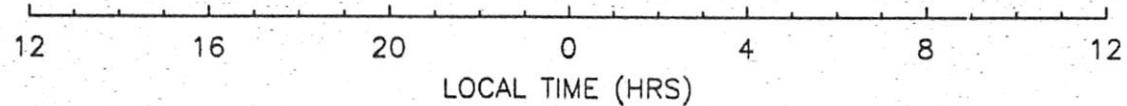
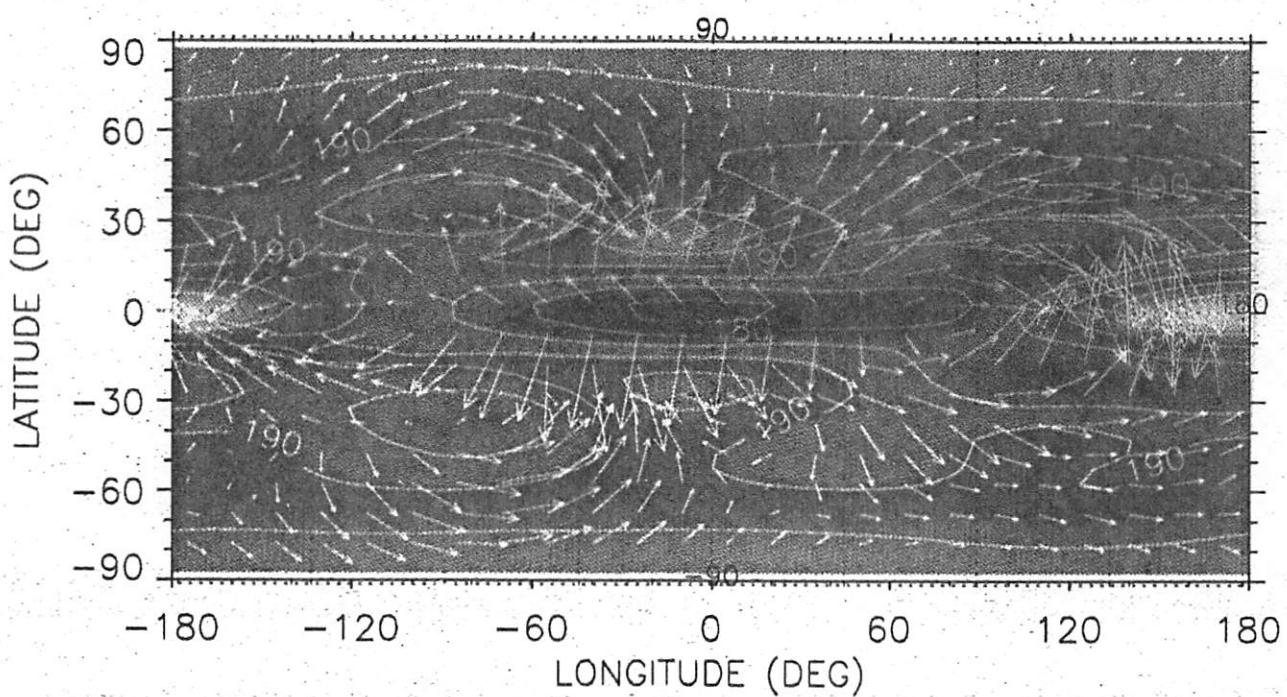


TIMESGCM TN (UN+VN) ZP = -9.0 UT = 0.0



95 Km \vec{T}_n, \vec{V}_n

TIMESGCM TN (UN+VN) ZP = -7.0 UT = 0.0



173.4

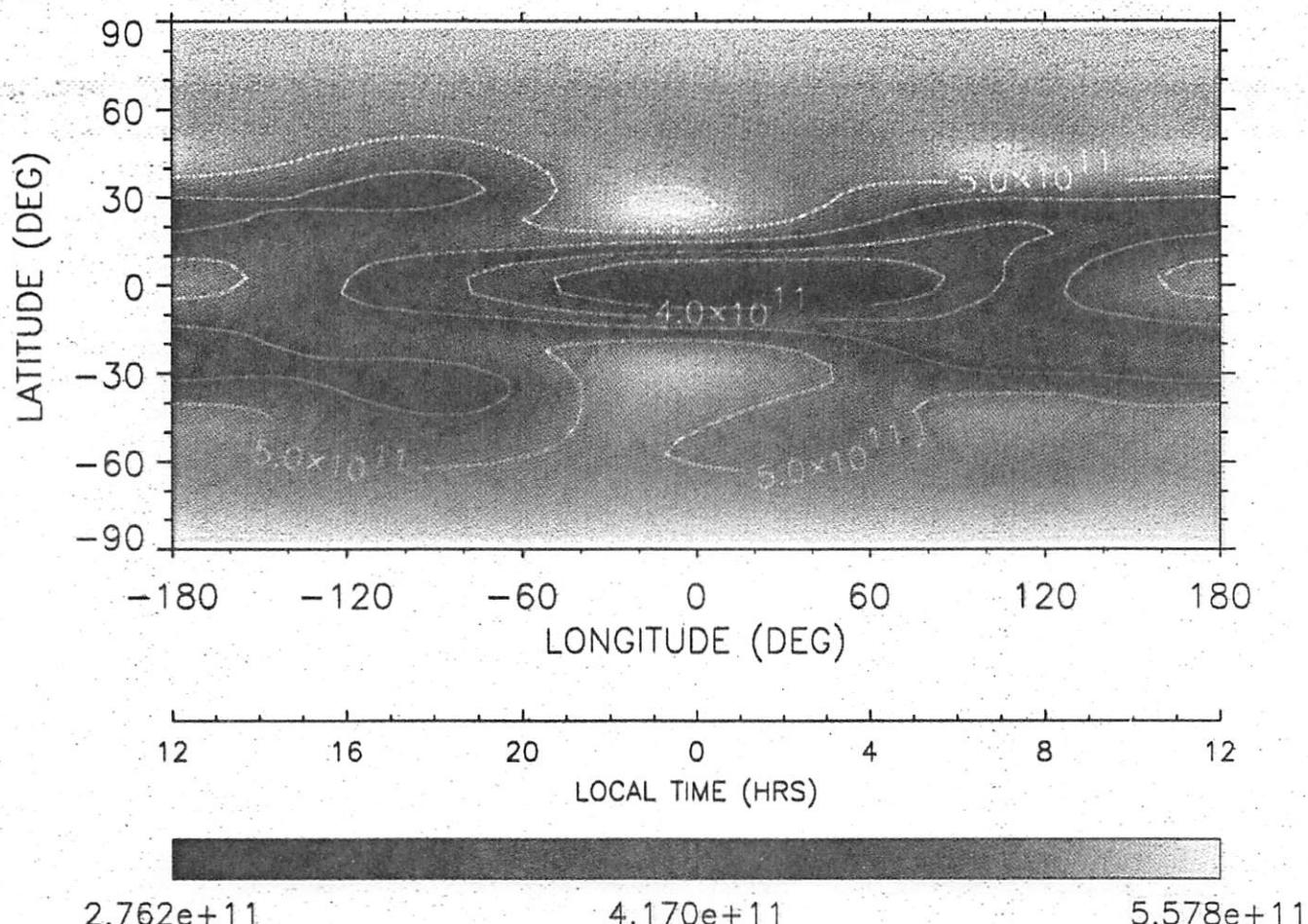
188.7

203.9

97 Km

$n(0, \text{cm}^{-3})$

TIMESGCM 01 HT = 97.0 UT = 0.0



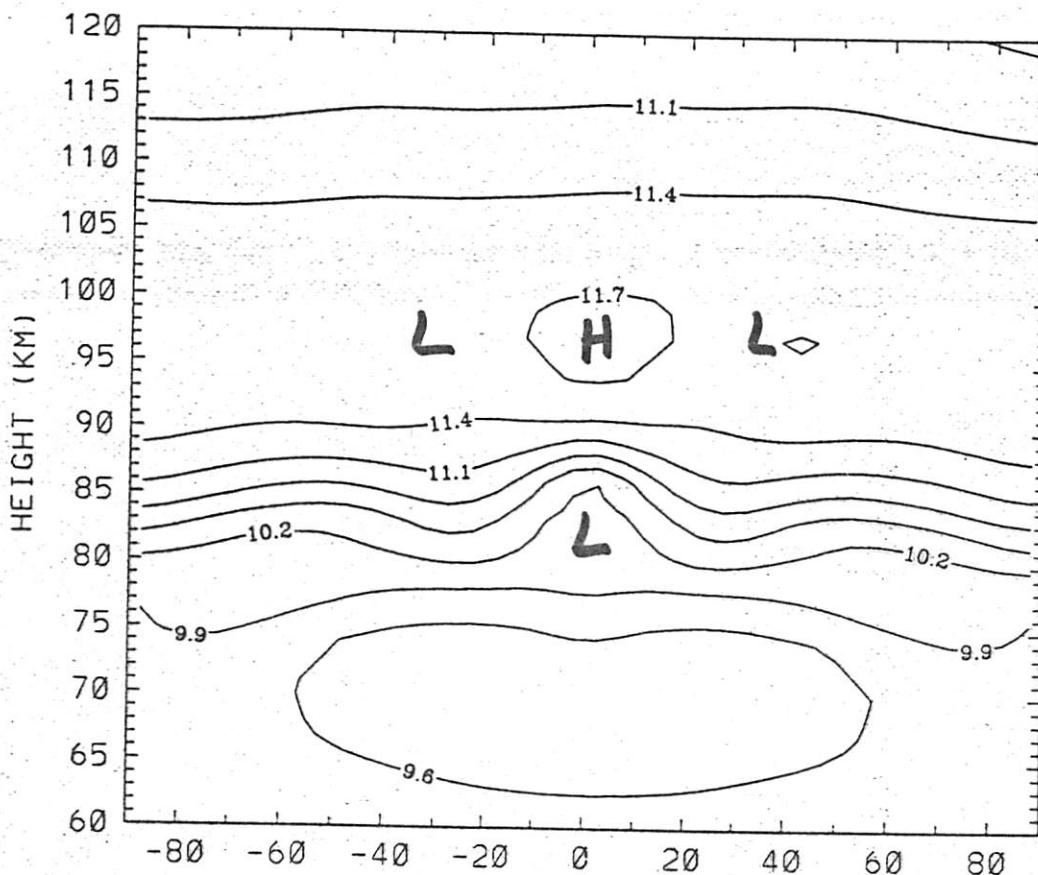
CONTOUR 3.00E+11 TO 6.00E+11 BY 5.00E+10 (MIN,MAX=2.76E+11,5.58E+11)

HISTORY /ROBLE/RGR93/TSEQBZ4 (6: 0: 0)

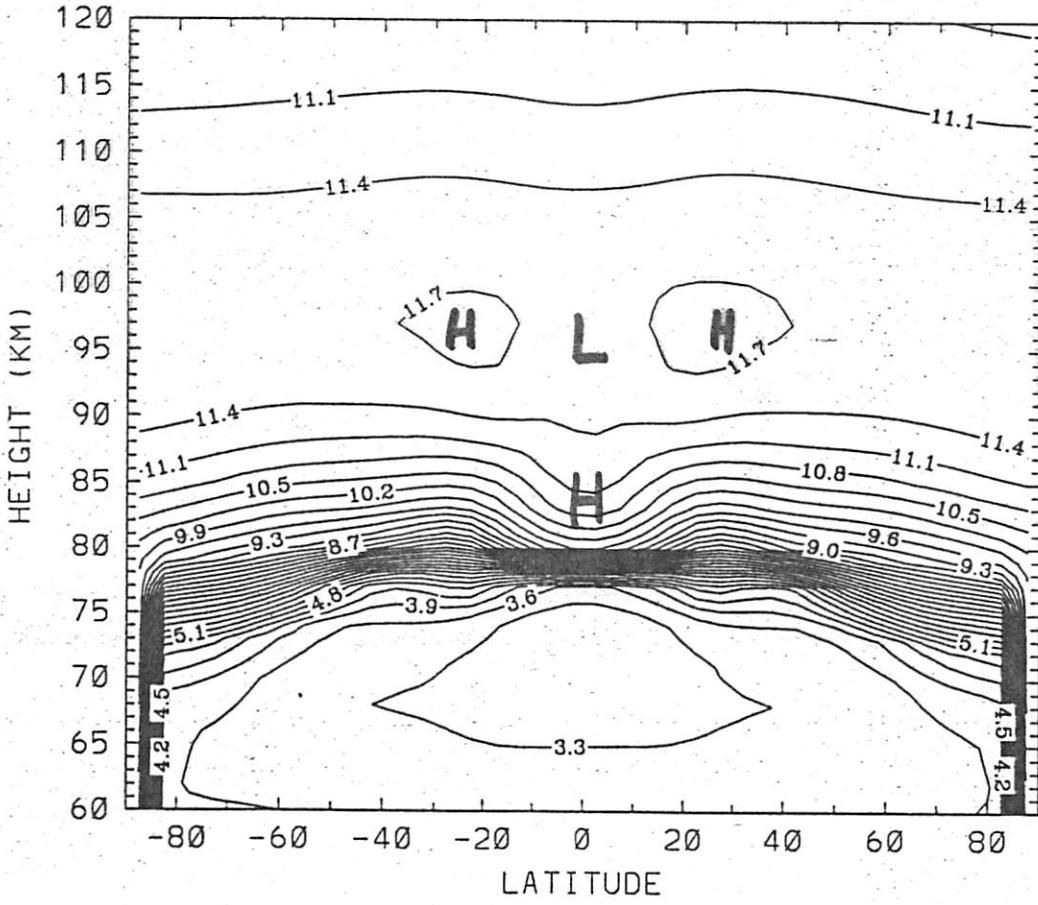
ATOMIC OXYGEN

$\log_{10} [n(0); \text{cm}^{-3}]$

NOON



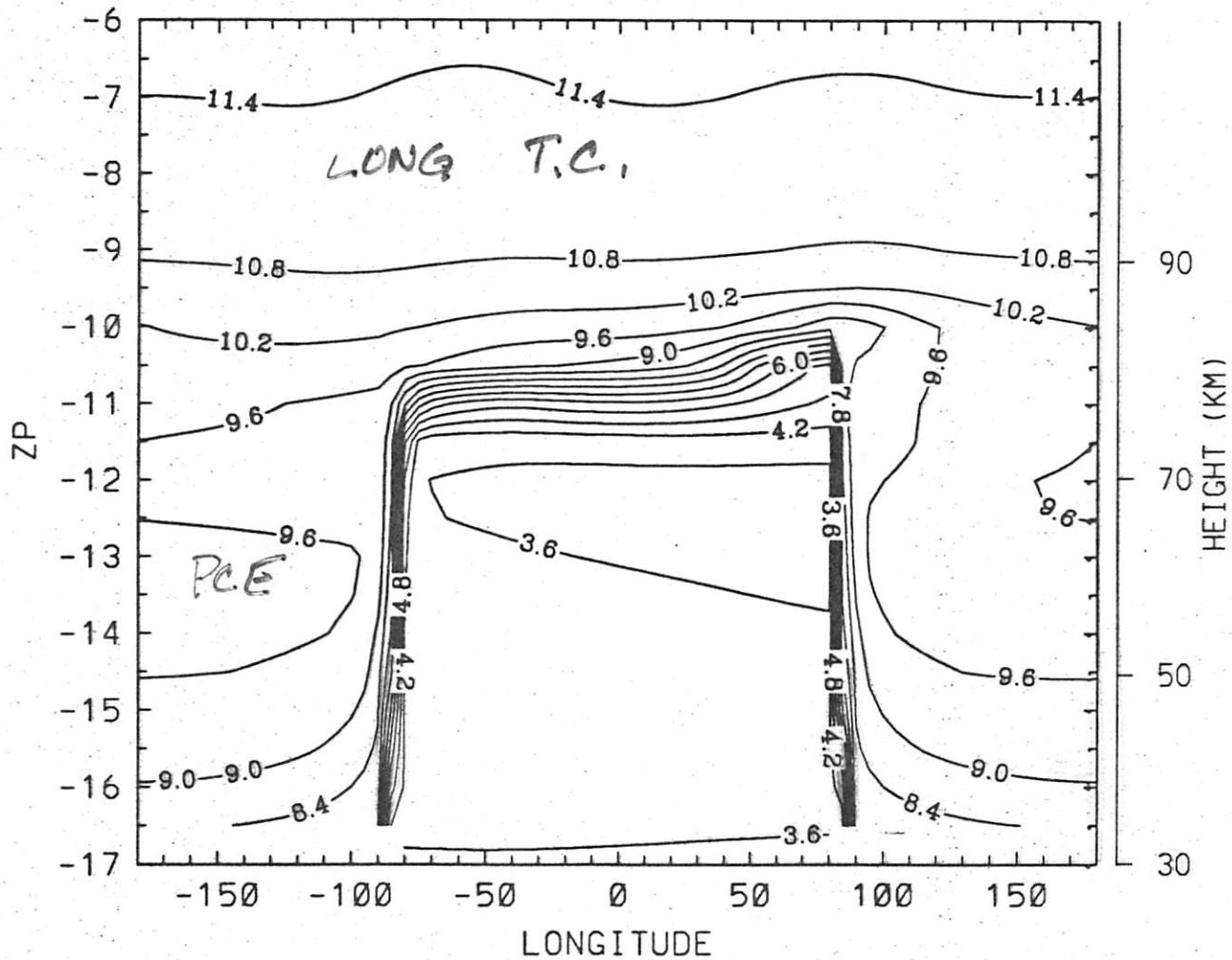
MIDNIGHT



$\log_{10} [n(0), \text{cm}^{-3}]$

LOG10 01

LAT= 2.50 UT= 0.00

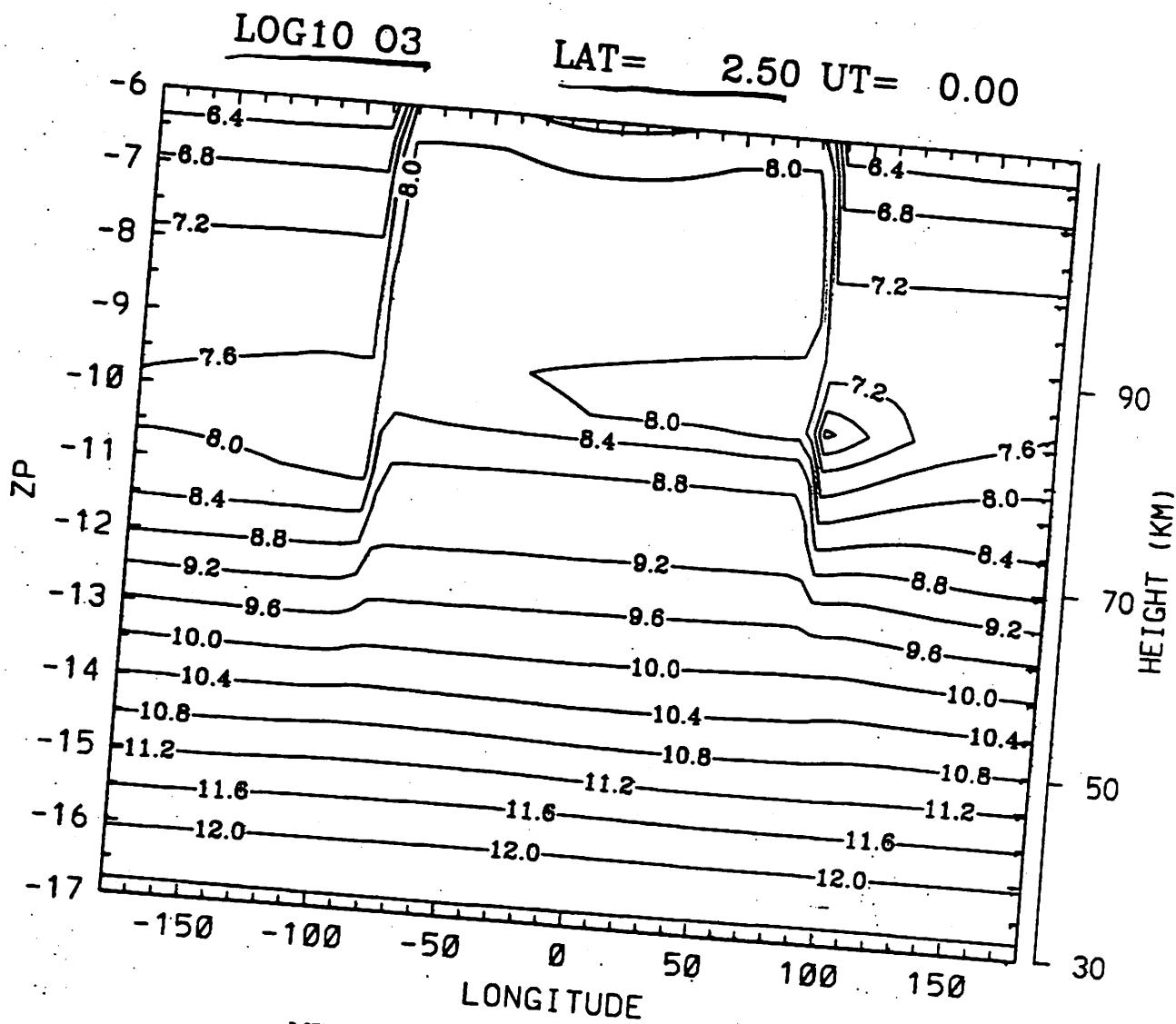


HISTORY=/ROBLE/RGR93/SOLGU8

SUNSET

SUNRISE

$\log_{10}[n(O_3), \text{cm}^{-3}]$



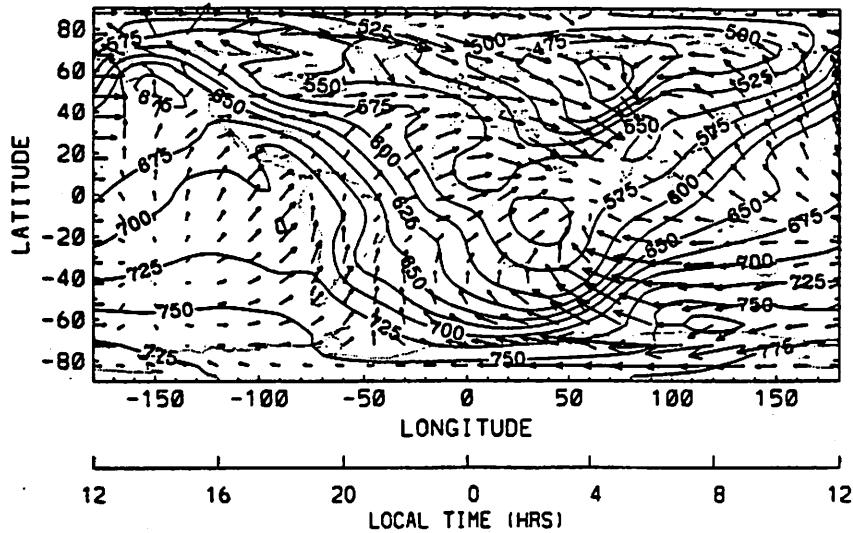
SUNSET

RISE
SUNSET

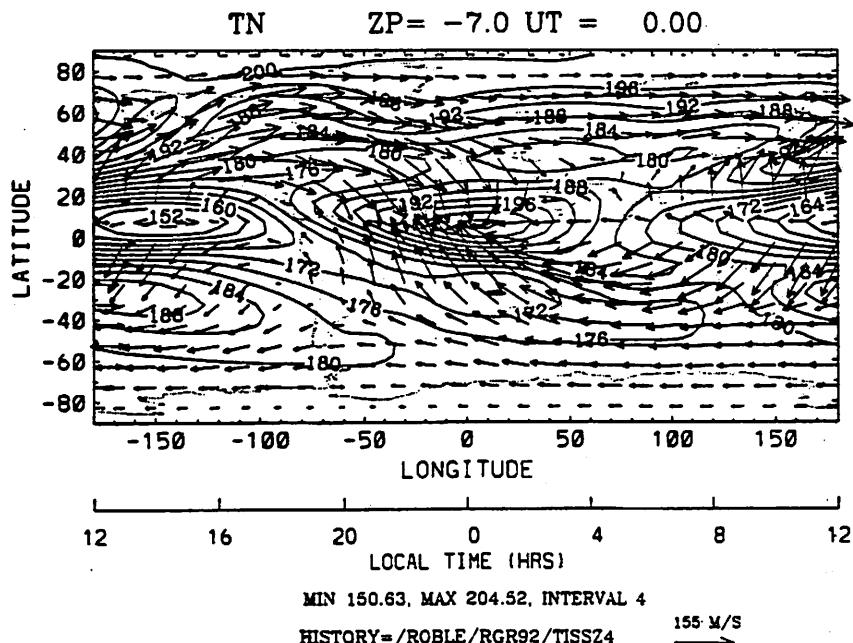
DEC SOLSTICE

TN ZP = 2.0 UT = 0.00

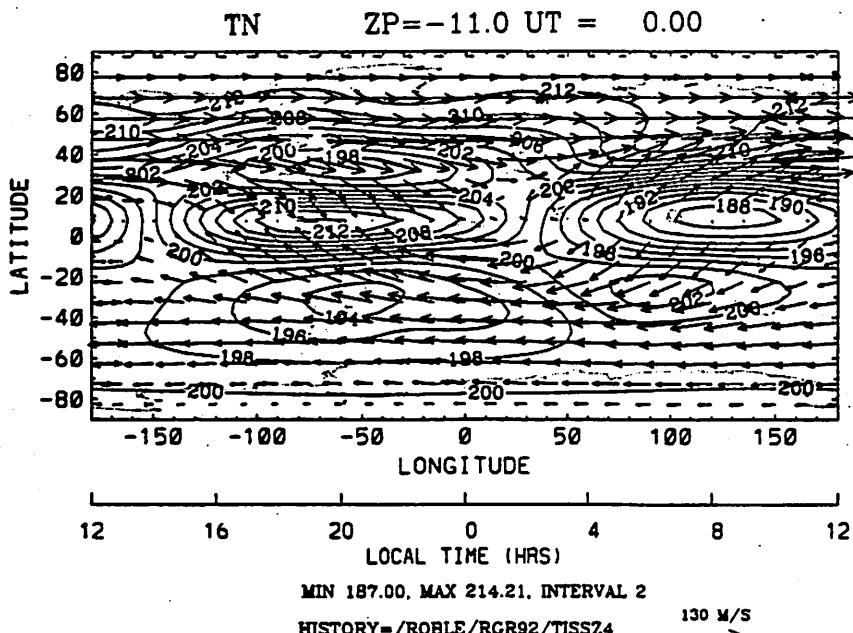
~300 KM



~90 KM



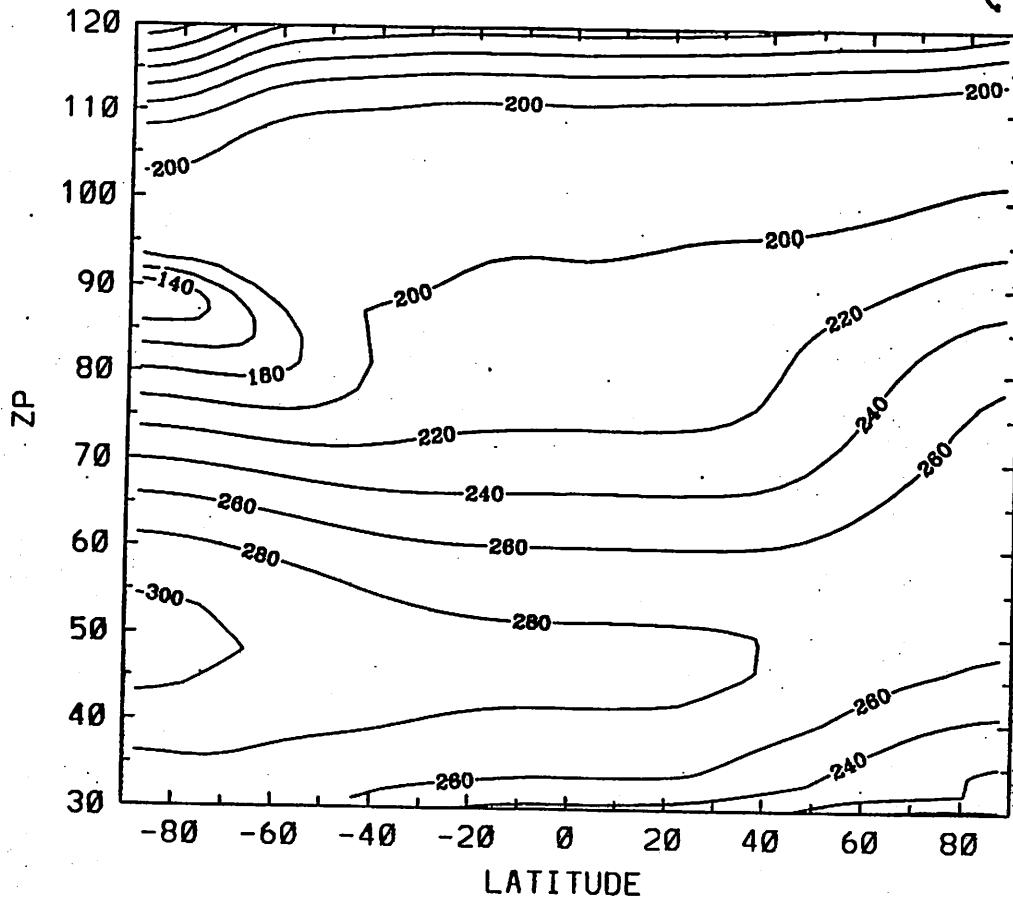
~70 KM



TIME-GCM

TN ZONAL MEANS UT= 0.00

(K)

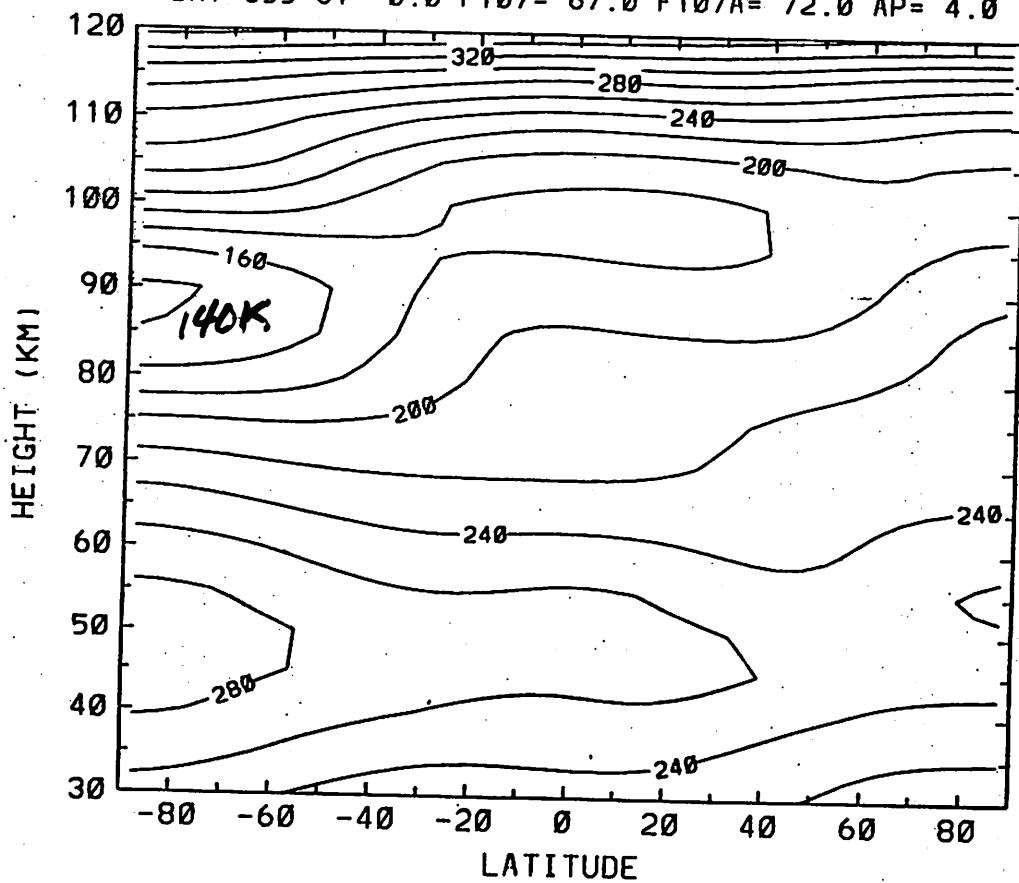


MINIMUM 120.3, MAXIMUM 332.6, CONTOUR INTERVAL 20

MSIS-90

MSIS90 TN (DEG K) ZONAL MEANS
DAY=355 UT= 0.0 F107= 67.0 F107A= 72.0 AP= 4.0

(K)



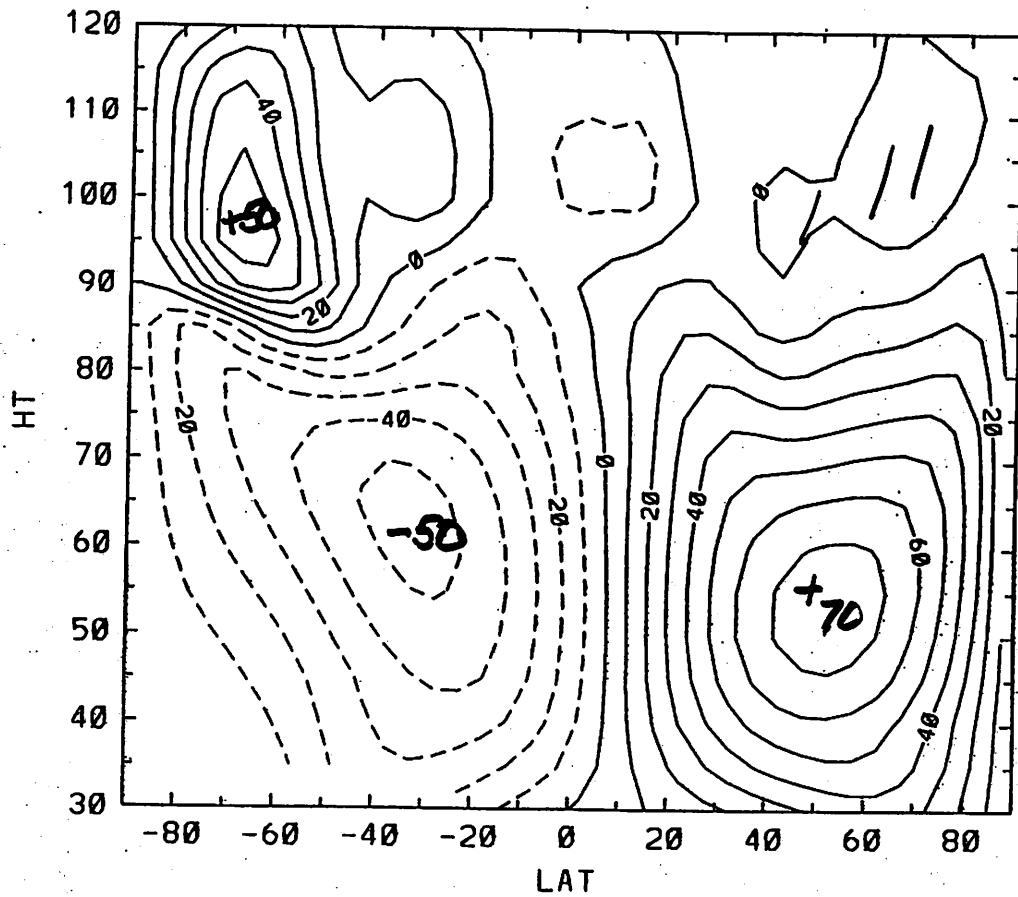
MIN 137.4, MAX 367.3, INT 20

TIME-GCM

UN

ZONAL MEANS UT= 0.00

(m/s)

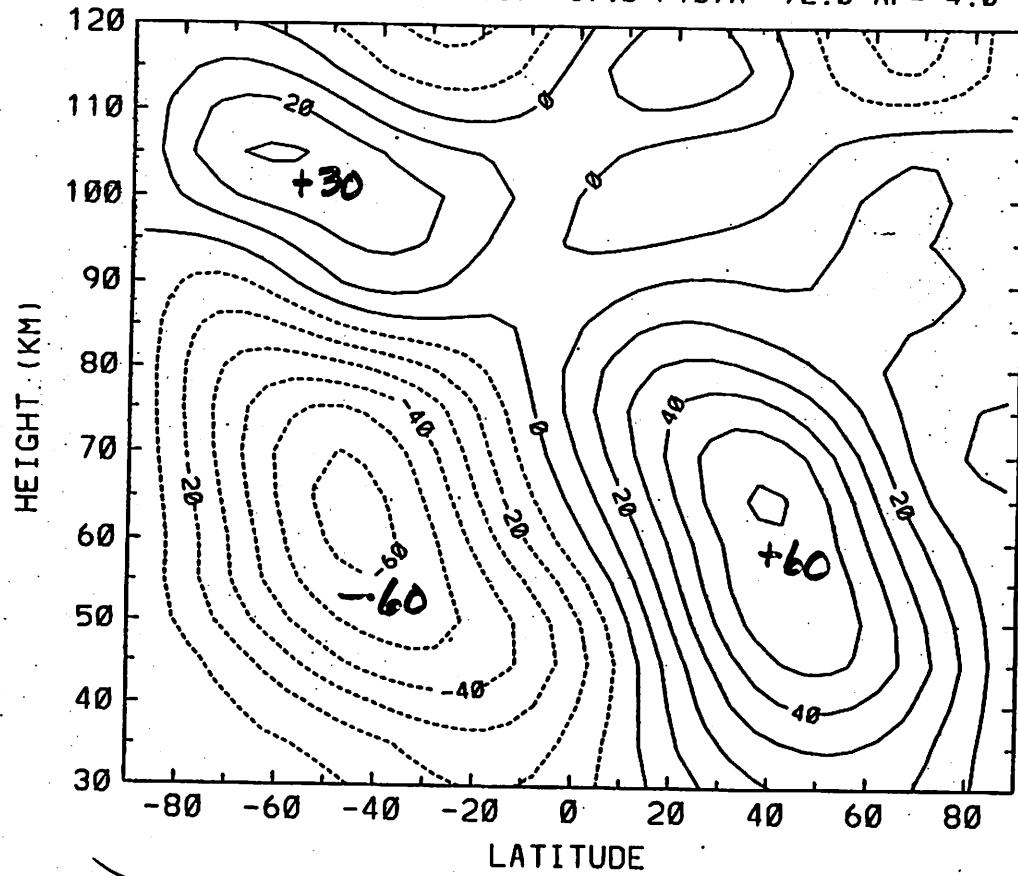


MIN -53.56, MAX 75.90, INT 10

M.S15-90

HWM93 UN (m/s) ZONAL MEANS
DAY=355 UT= 0.0 F107= 67.0 F107A= 72.0 AP= 4.0

(m/s)



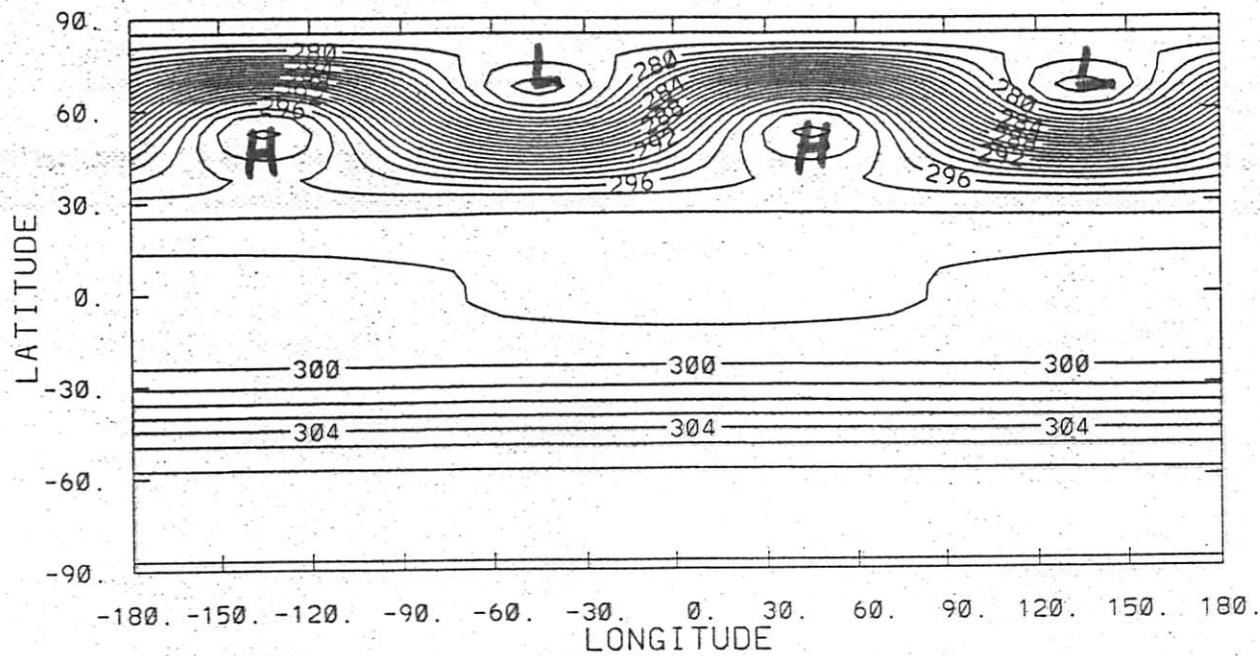
MIN -63.42, MAX 61.38., INT 10

GEO POTENTIAL (10mb, 30 Km)

UTIME=12: 0

FIELD=H

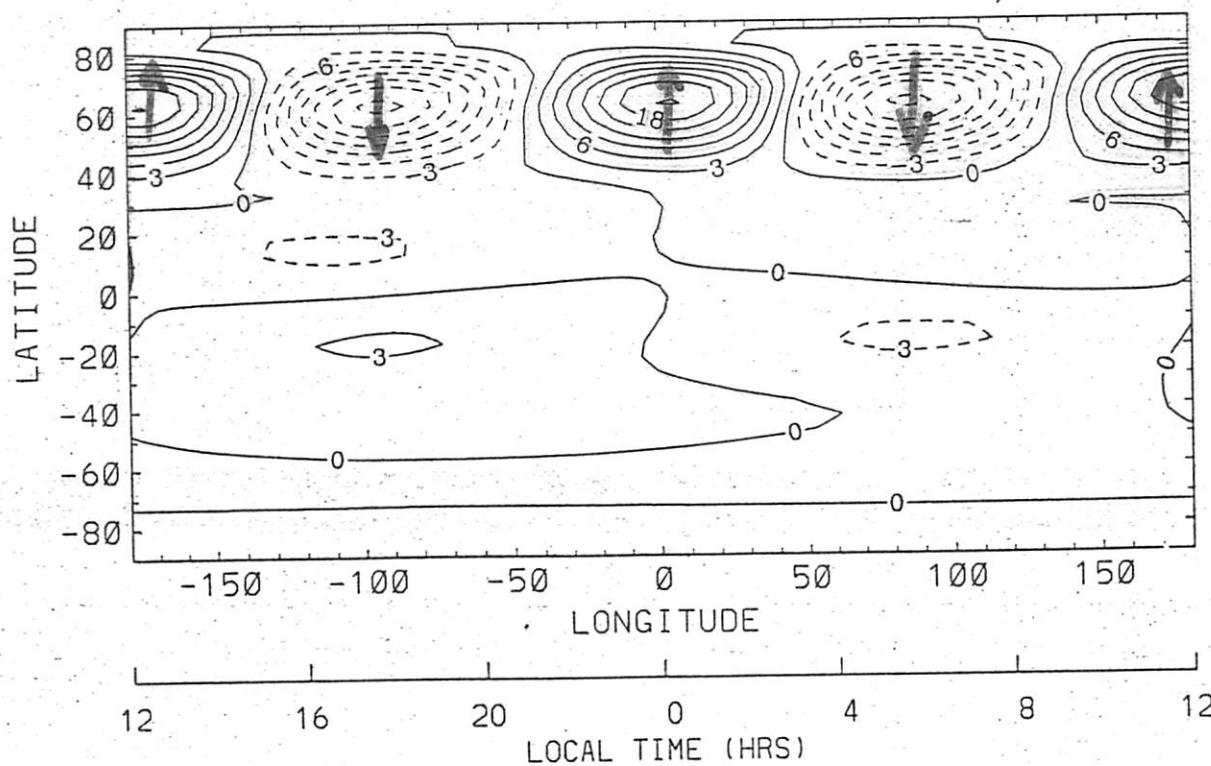
Z=-17.0



MERIDIONAL WIND (m/s) [10mb, 30km]

VN

ZP=-17.0 UT = 0.00



MIN -21.90, MAX 21.21, INTERVAL 3

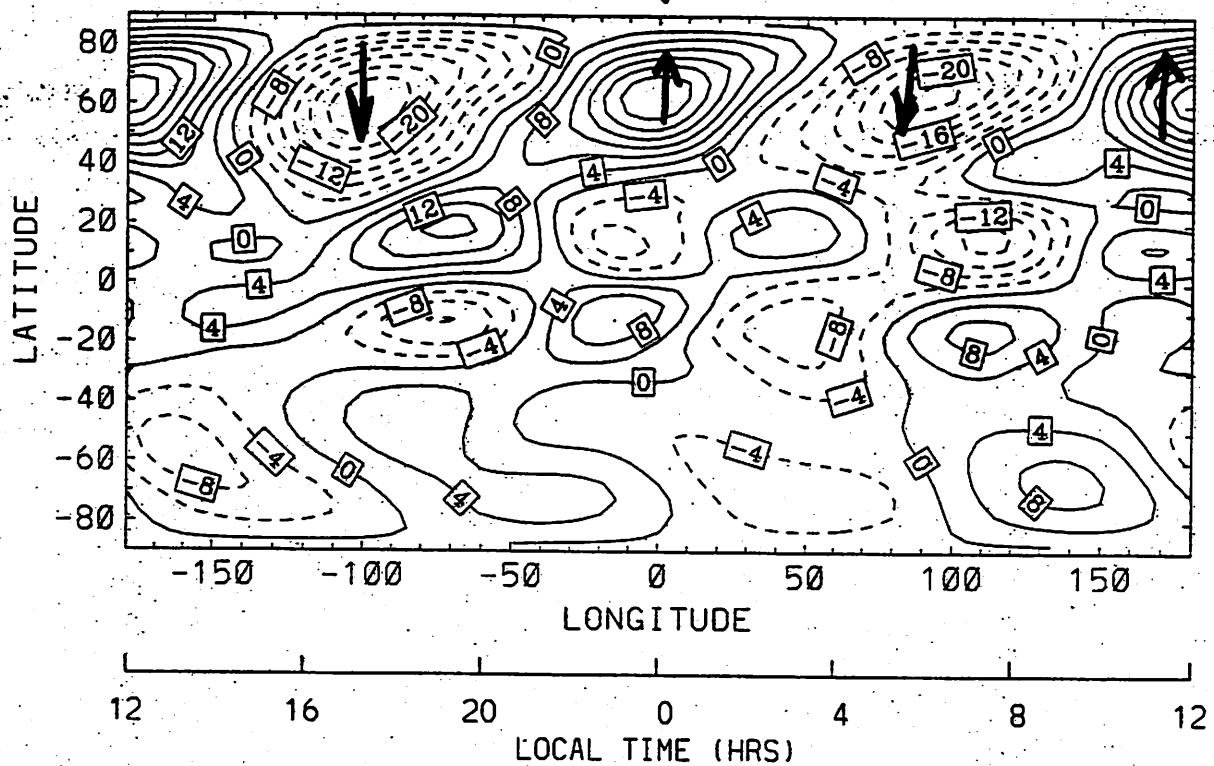
MERIDIONAL WIND (m/s)

VN

RAW DIFFERENCE FIELD

ZP = -7.0 UT = 0.00

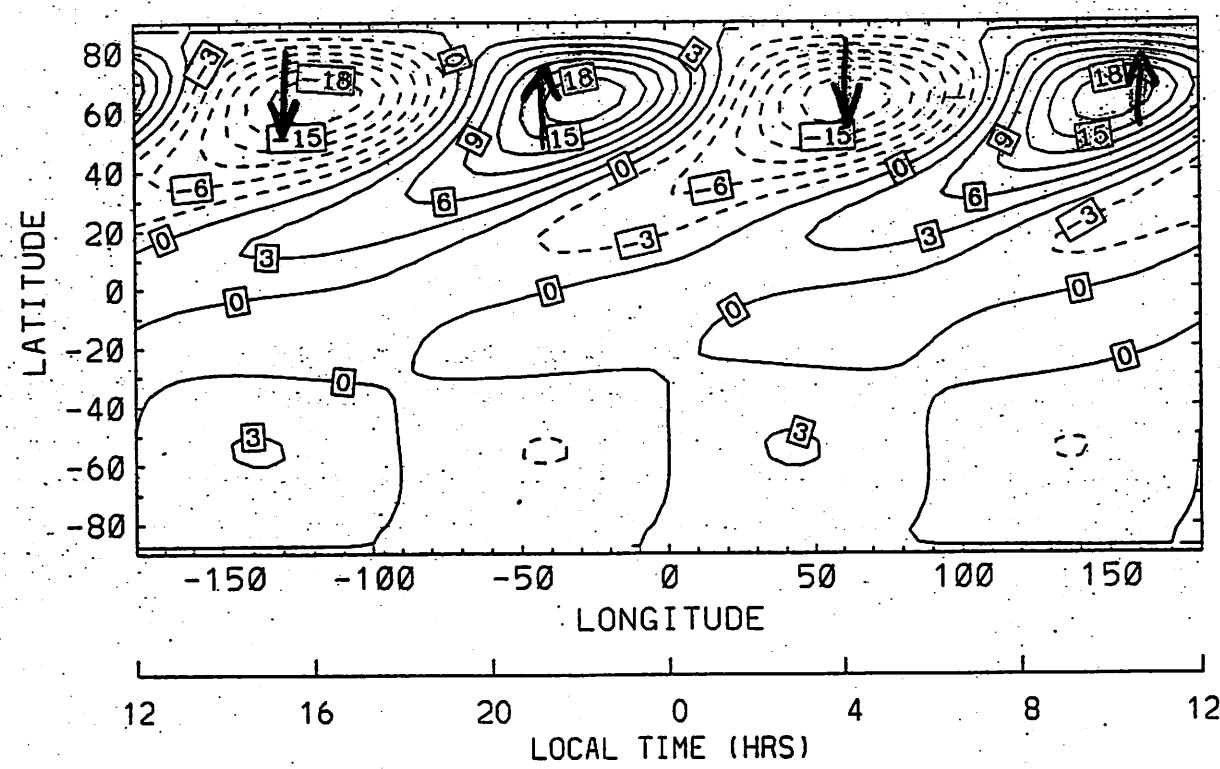
$\sim 97 \text{ KM}$



MIN -27.93, MAX 30.70, INTERVAL 4

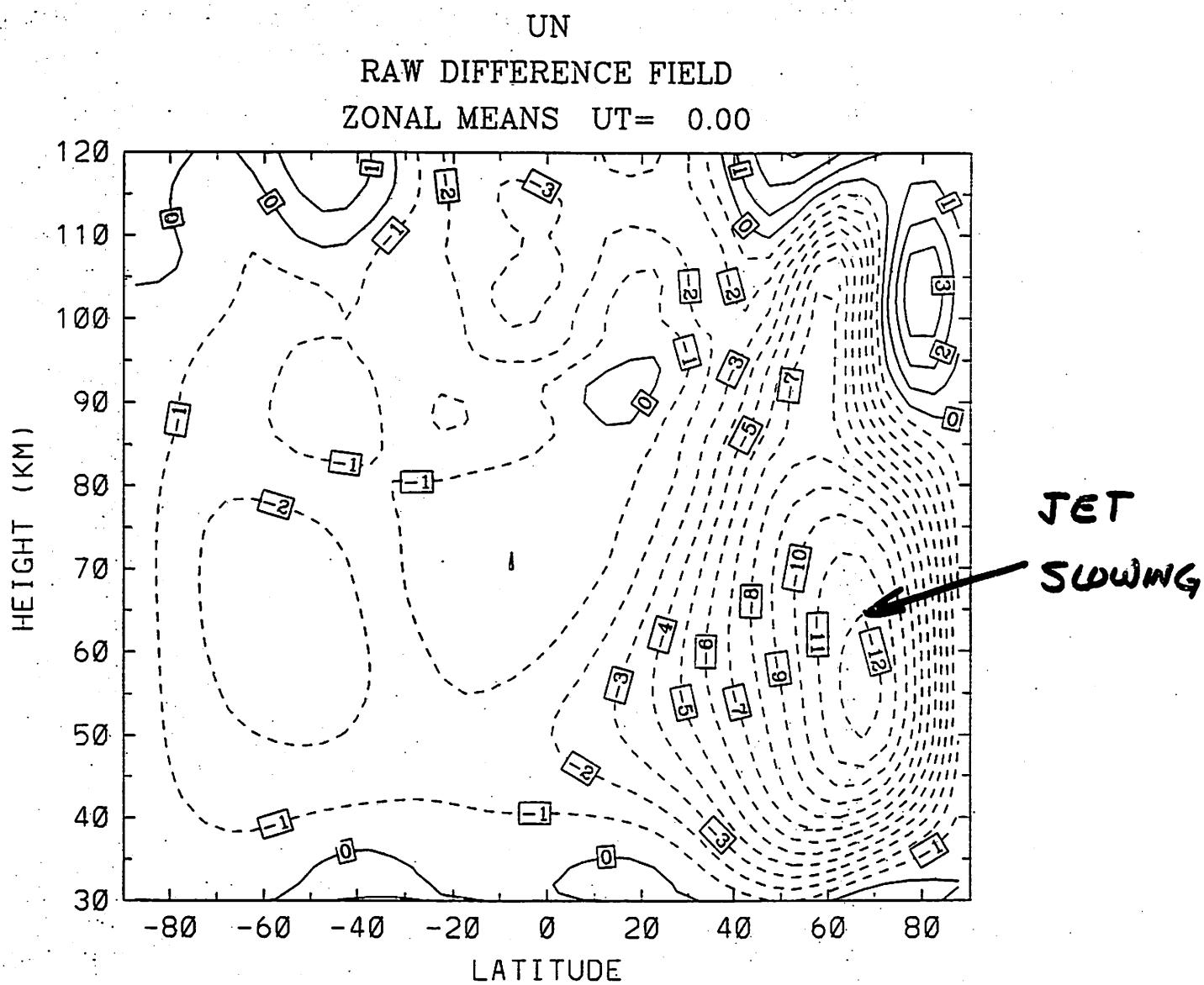
ZP = -15.0 UT = 0.00

$\sim 50 \text{ KM}$



ZONAL MEAN ZONAL WIND DIFF

$$\Delta \bar{U} = \bar{U}_{pw} - \bar{U} \text{ (m/s)}$$



MIN -12.53, MAX 3.78, INTERVAL 1

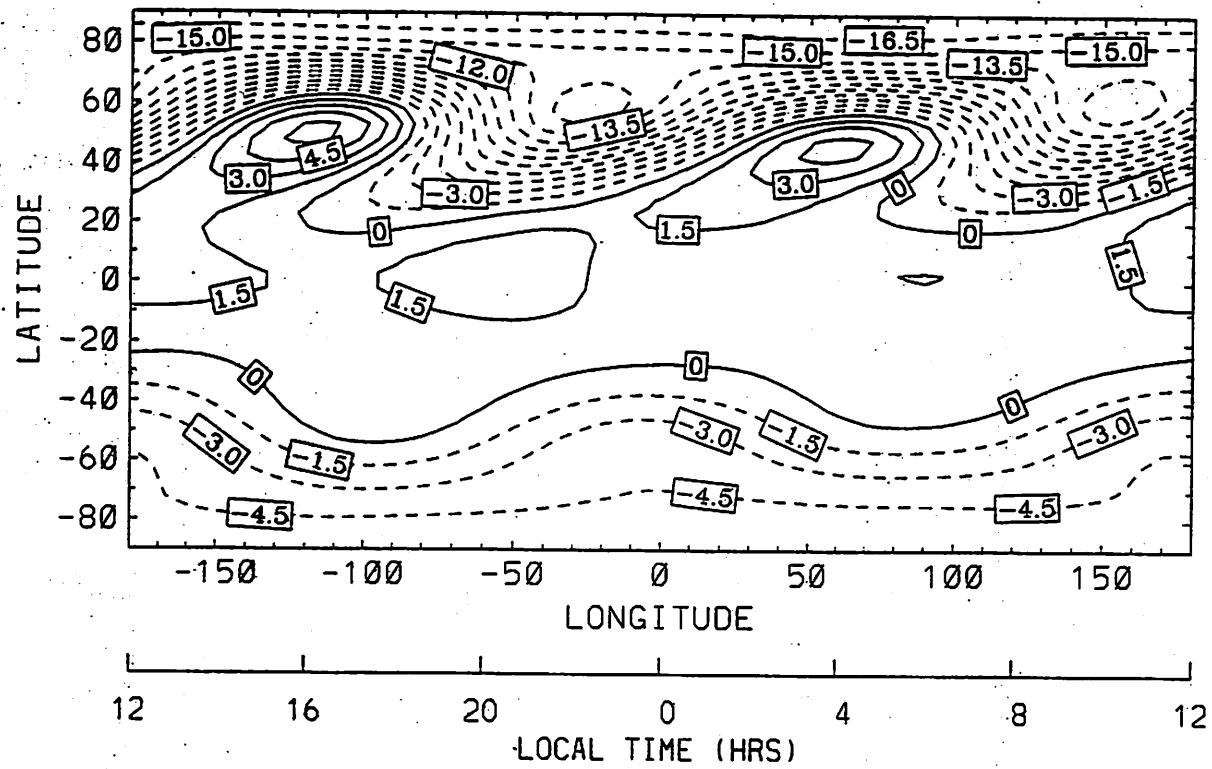
DIFFS FROM T17PW9 MINUS T17LB2

03

PERCENT DIFFERENCE FIELD

ZP=-15.0 UT= 0.00

50 Km

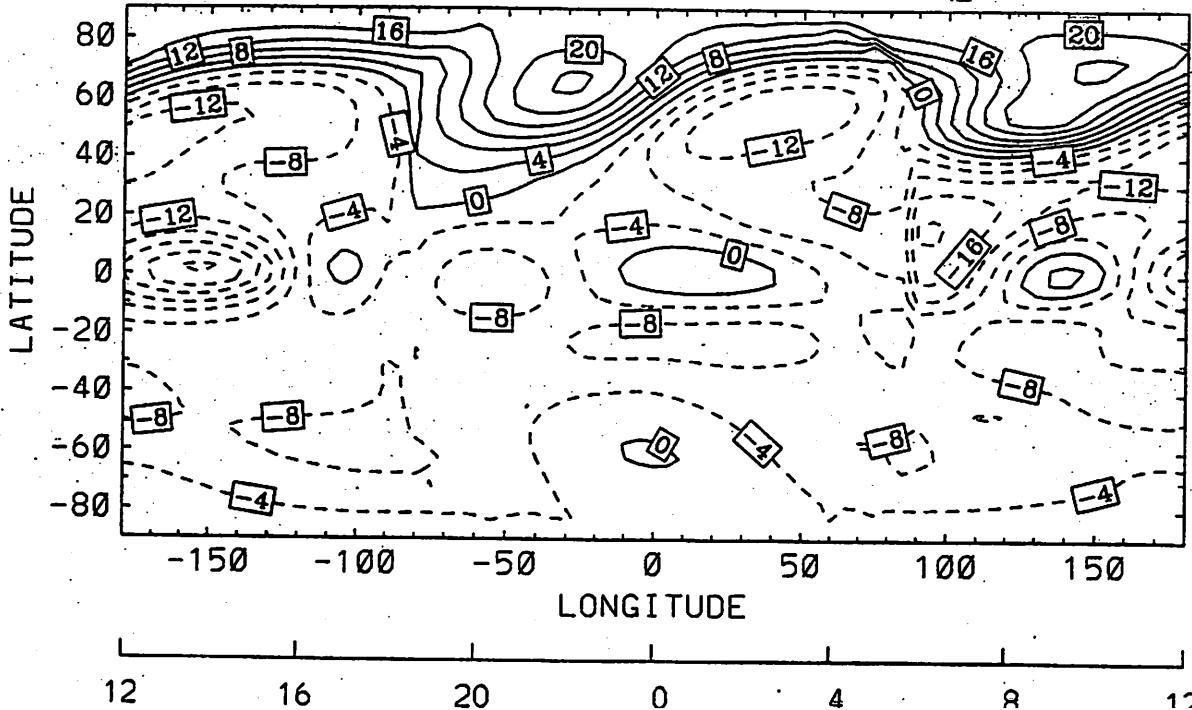


OH

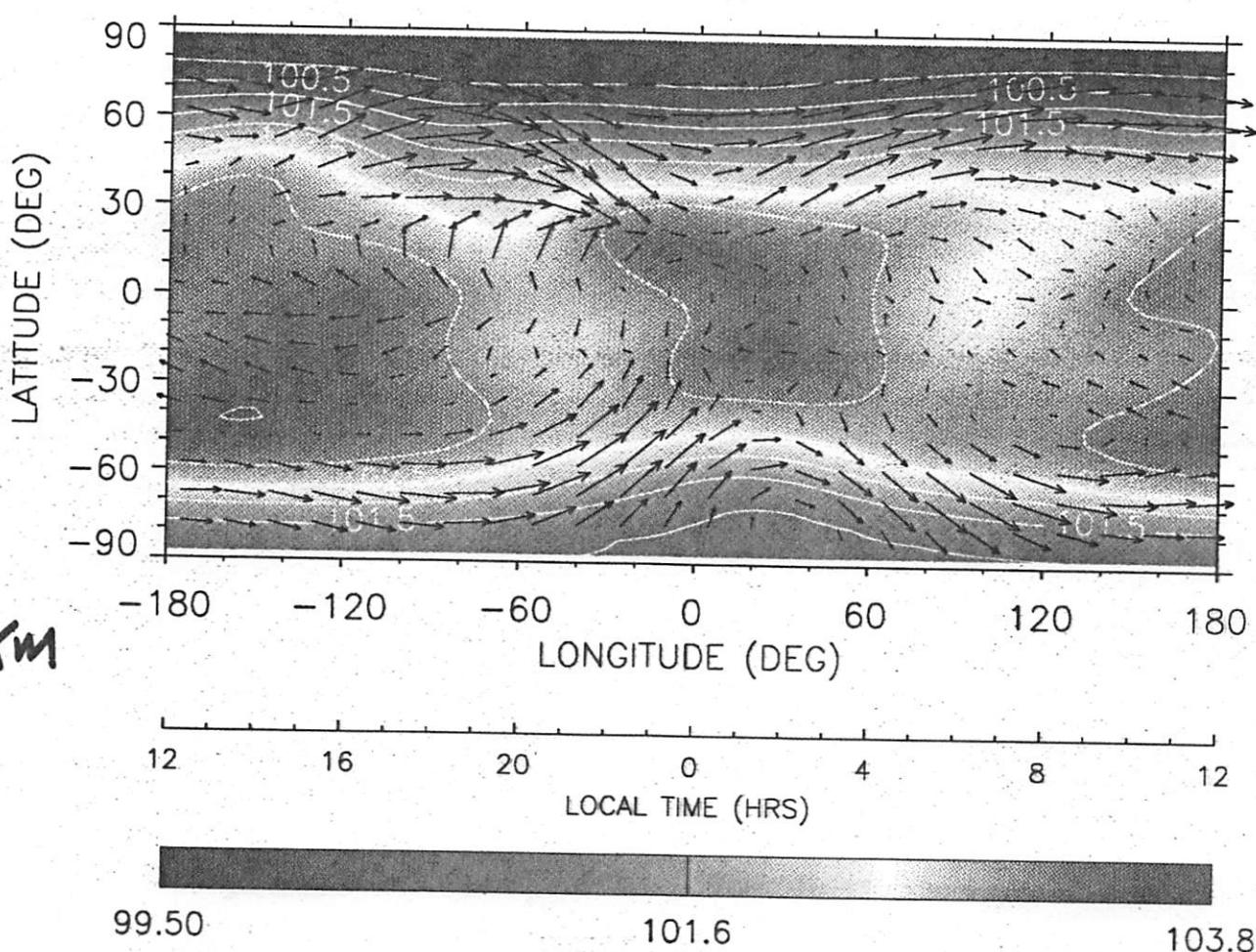
PERCENT DIFFERENCE FIELD

ZP= -9.0 UT= 0.00

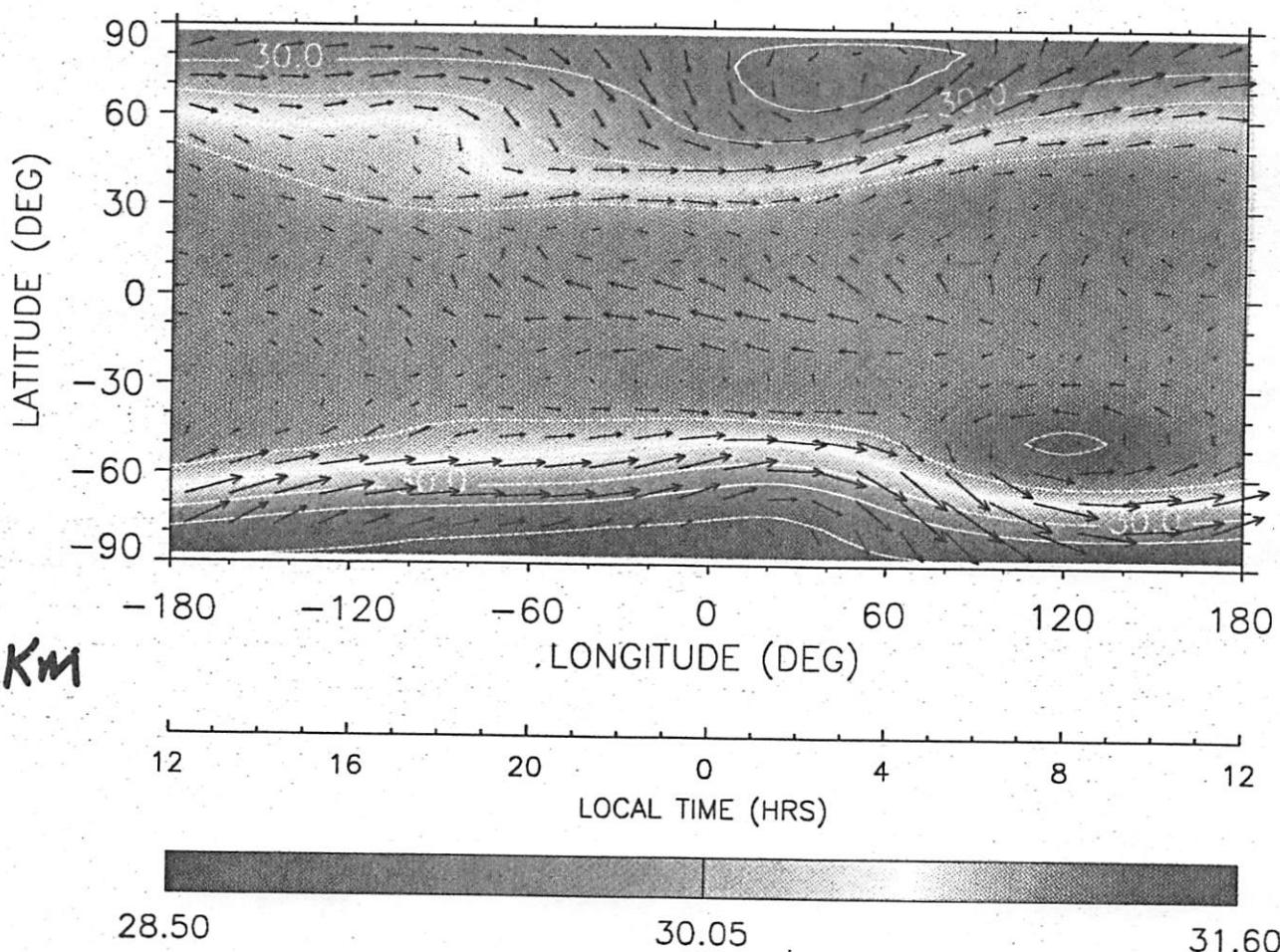
~ 85 Km

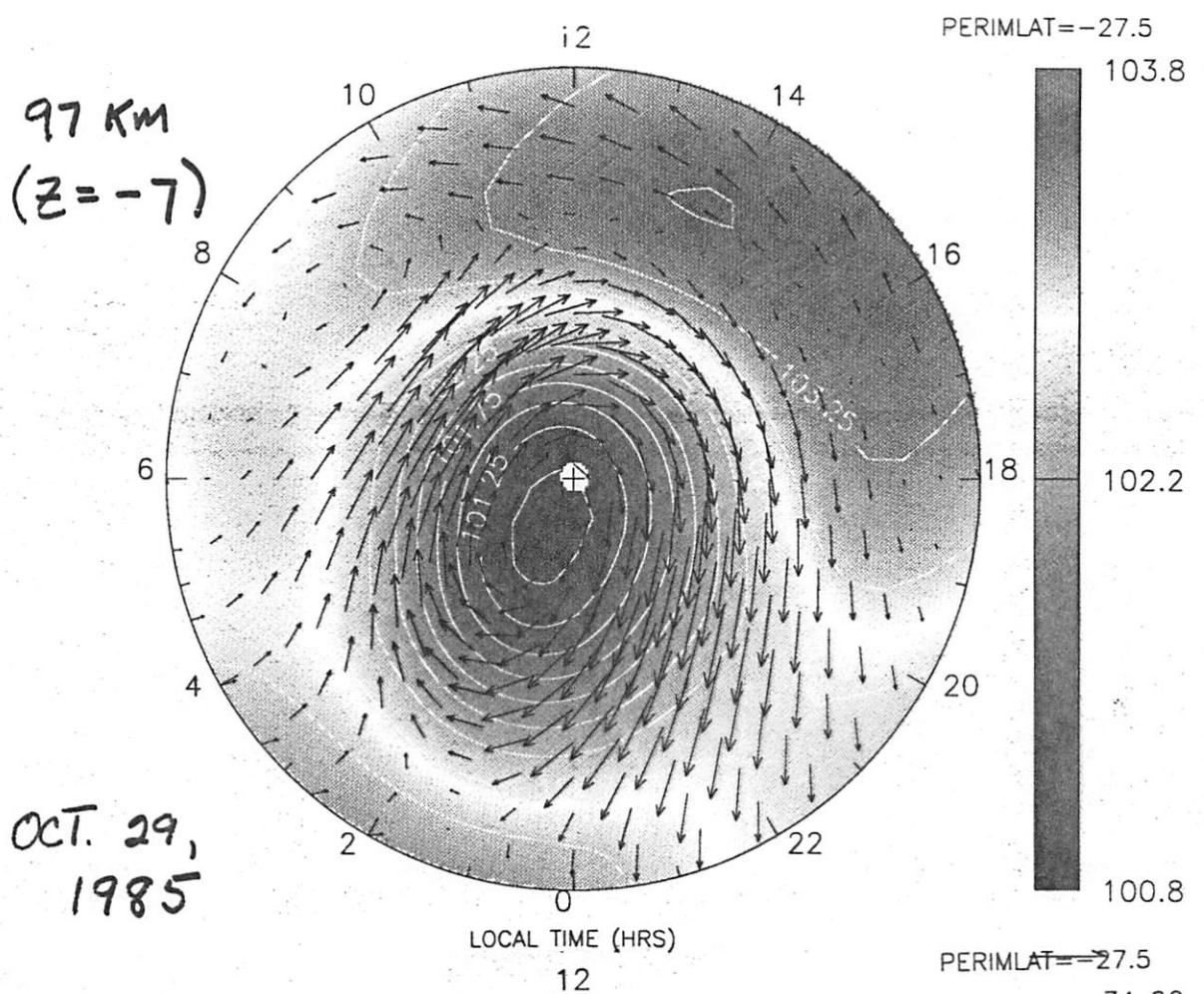


TIMESGCM Z (UN+VN) ZP = -7.0 UT = 0.0

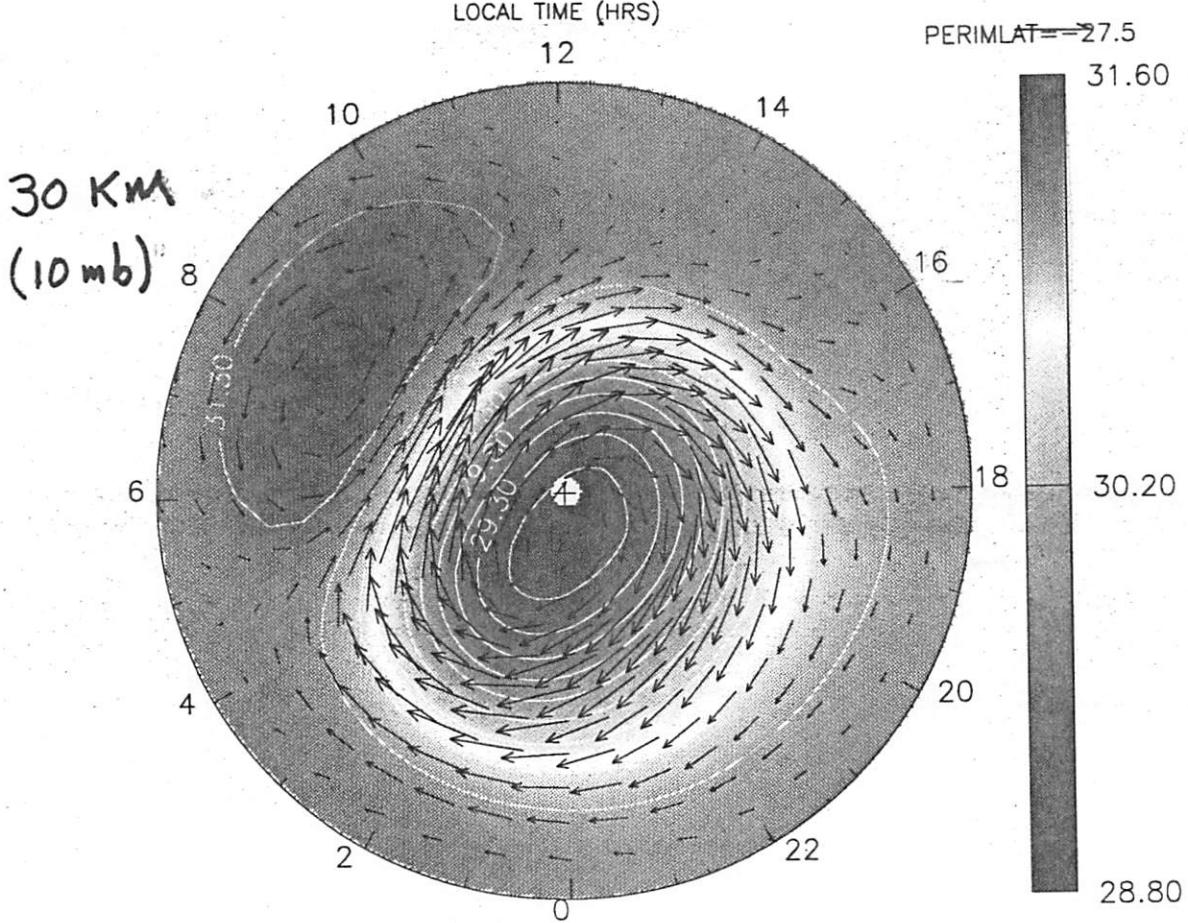


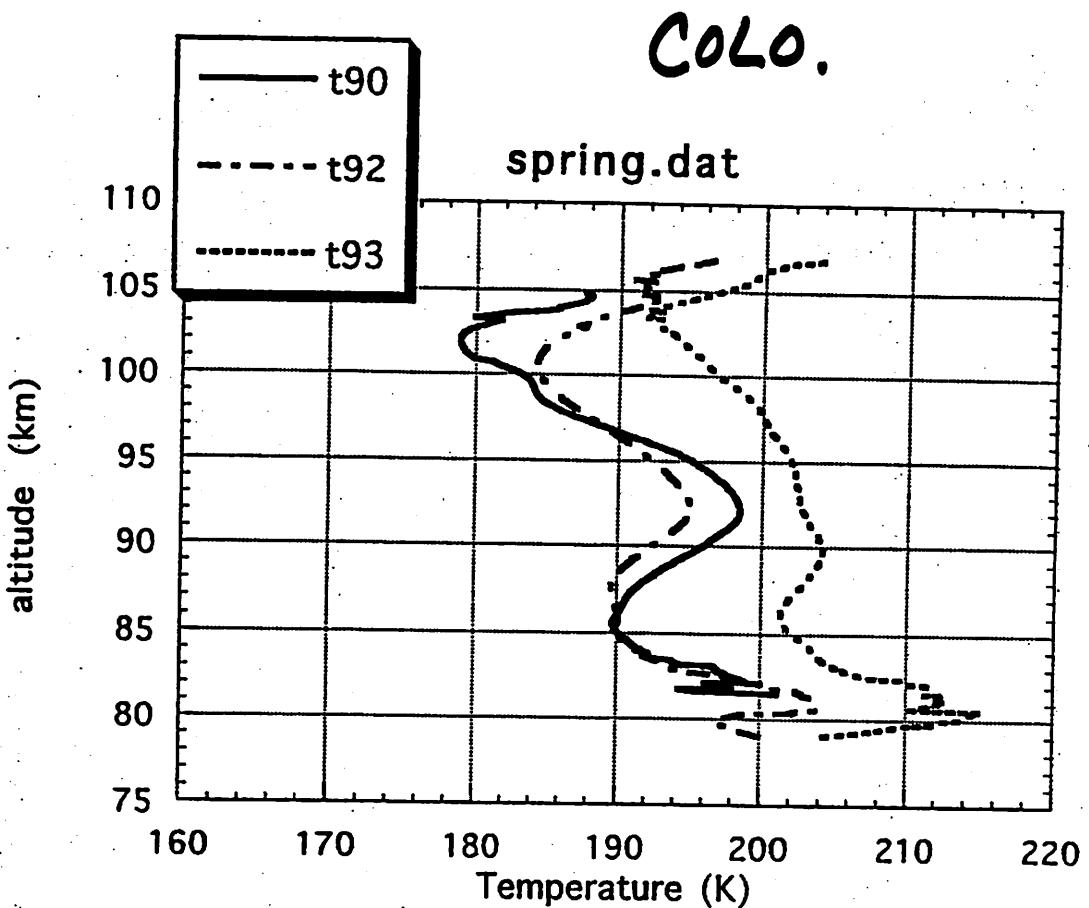
TIMESGCM Z (UN+VN) ZP = -17.0 UT = 0.0





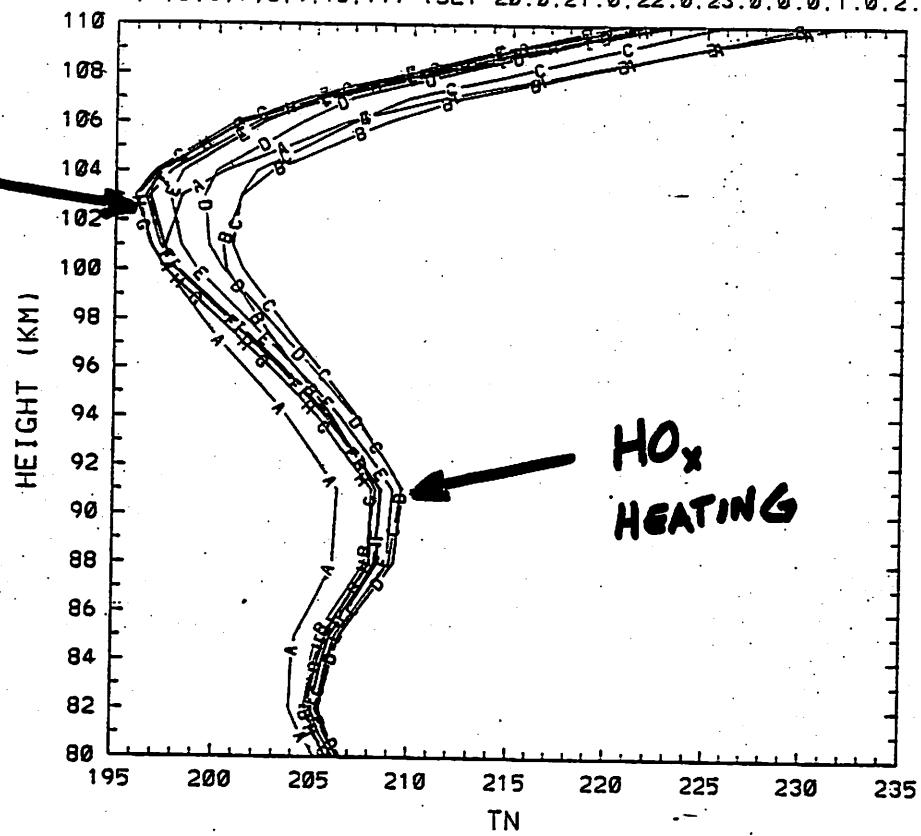
GEO POTENTIAL HEIGHT (KM)





LAT. LON = 42.50, -105.00 (FT. COLLINS)
 $A \rightarrow I = UT\ 3, 4, 5, 6, 7, 8, 9, 10, 11\}$ (SLT 20.0, 21.0, 22.0, 23.0, 0.0, 0.1, 0.2, 0.3, 0.4, 0.1)

O- CO_2
COOLING

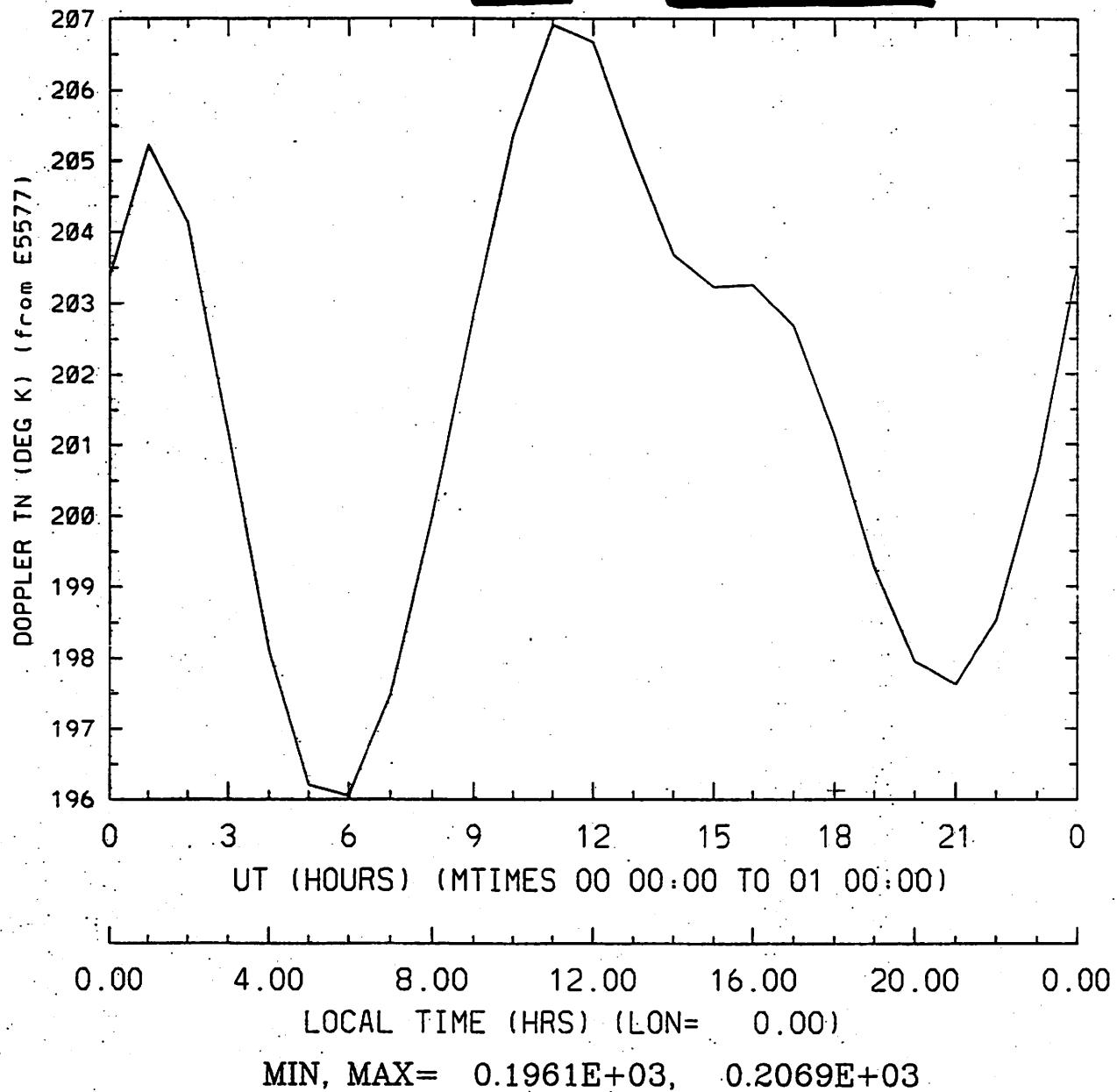


~ 100 KM

$\lambda 5577 \text{ \AA}^{\circ}$

DOPPLER TN (DEG K) (from E5577)

LAT,LON = 2.50 0.00 (EQUATOR)



SUMMARY

- TIME-GCM (30-500 Km), UP AND
RUNNING → CEDAR STUDIES
- + CONTROLLED EXPERIMENTS
- + NMC-10MB, SOLAR, AURORAL FORCINGS
- + PROCESSORS FOR ANALYSIS OF CEDAR,
UARS, ETC. DATA
- + COMMUNITY MODEL