

CEDAR MEETING, 21–26 June 92, Boulder CO

EFFECTS OF SOLAR VARIABILITY
ON GLOBAL CHANGE

Judith Lean
Naval Research Laboratory

- * *Global Change Research*
- * *Solar Forcing of Climate*
- * *Solar Variability and Ozone*
- * *Solar Effects on the Upper Atmosphere*

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List of Viewgraphs

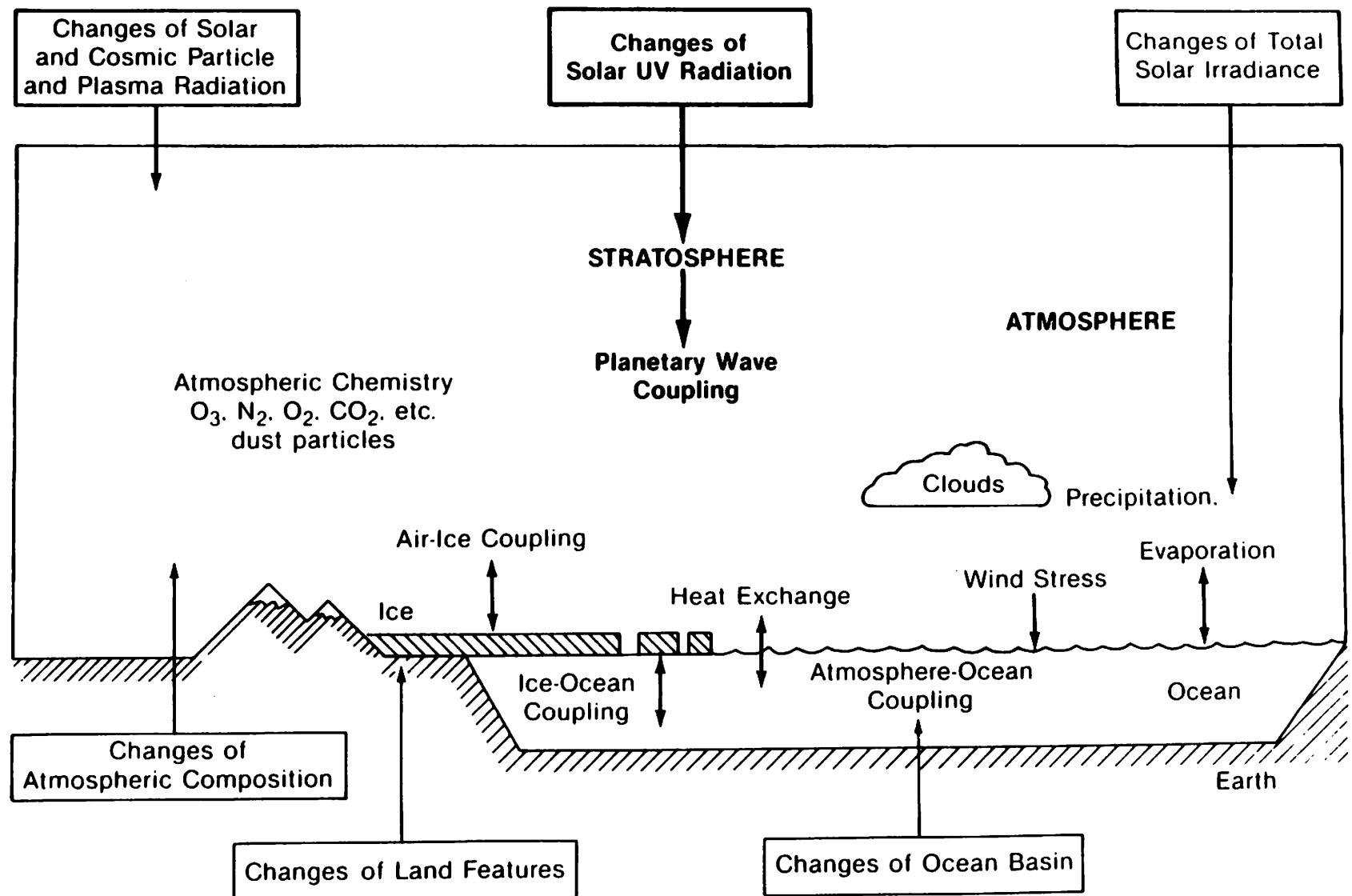
1. Overview of talk
2. Statement of "global change" from USGCRP at-a-glance
3. Schematic of the climate system
4. Summary of global change forcing agents
5. NAS 1982 statement about solar influences on weather and climate
6. Historical ^{14}C solar variability
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8. Rz since 1600
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10. Sun, weather, QBO (troposphere)
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19. Stellar distribution, including S estimates
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21. GISS Maunder Minimum temperature calculations
22. Wigley and Kelly Little Ice Age (LIA) vs greenhouse gases forcing and temperature response
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24. S and UV irradiance variations
25. Energy deposition in Earth's atmosphere (plus temperature)
26. Solar radiation, chemistry and dynamics
27. Observed total ozone change with solar activity

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29. Scenario for the future ..UV forcing of ozone
30. Callis et al. results..alternative interpretation of O₃ record
31. Connections between solar influences and other elements of USGCRP
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33. EUV irradiance variations ..measurements vs empirical model
34. Particle inputs to upper atmosphere
35. Increase in upper atmosphere temperature from solar min to max
36. Satellite drag
37. Effects of global change on upper atmosphere (from Roble & Dickinson)
38. Present and future irradiance measurement programs timeline

THE U.S. GLOBAL CHANGE RESEARCH PROGRAM (USGCRP) AT-A-GLANCE

(from Our Changing Planet: The FY 1990 Research Plan, The U.S. Global Change Research Program, A Report by the Committee on Earth Sciences, July 1989)

- * ***Many global changes can have tremendous impact on the welfare of humans. These events may stem from natural processes that began millions of years ago or from human influence.***
- * ***The goal of the USGCRP is to provide a sound scientific basis for national and international decision making on global change issues.***
- * ***The scientific objectives of the USGCRP are to monitor, understand and ultimately predict global change.***
- * ***The USGCRP is broad in scope, encompassing the full range of earth system changes, including physical, chemical, geological, social and biological changes. It addresses both natural phenomena, as well as the effects of human activity.***
- * ***The particular research activities which comprise the USGCRP are grouped into seven interdisciplinary scientific elements:***
 - 1. Climate and Hydrologic Systems***
 - 2. Biogeochemical Dynamics***
 - 3. Ecological Systems and Dynamics***
 - 4. Earth System History***
 - 5. Human Interactions***
 - 6. Solid Earth Processes***
 - 7. Solar Influences***



GLOBAL CHANGE FORCING AGENTS

Hansen & Lacis, Nature, 1990

Anthropogenic

- CO₂ carbon dioxide
- CH₄ methane
- N₂O nitrous oxide
- CFCs chlorofluorocarbons
- aerosols sulphates from SO₂
- albedo change desertification, deforestation

Natural

- solar variability
- volcanos sulphates from SO₂

National Academy of Sciences, 1982

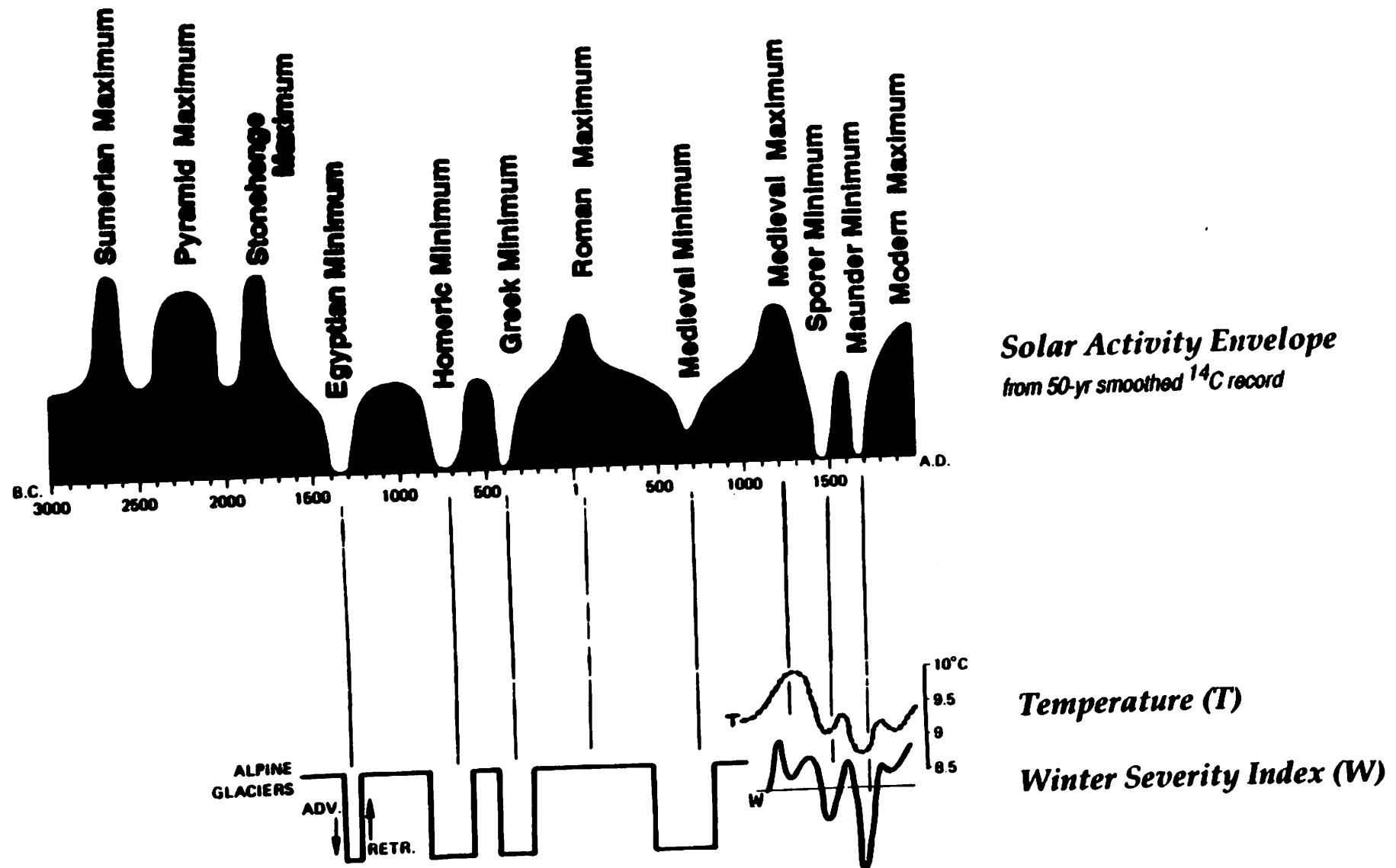
Panel on Solar Variability,
Weather and Climate

J. Eddy (Chair)

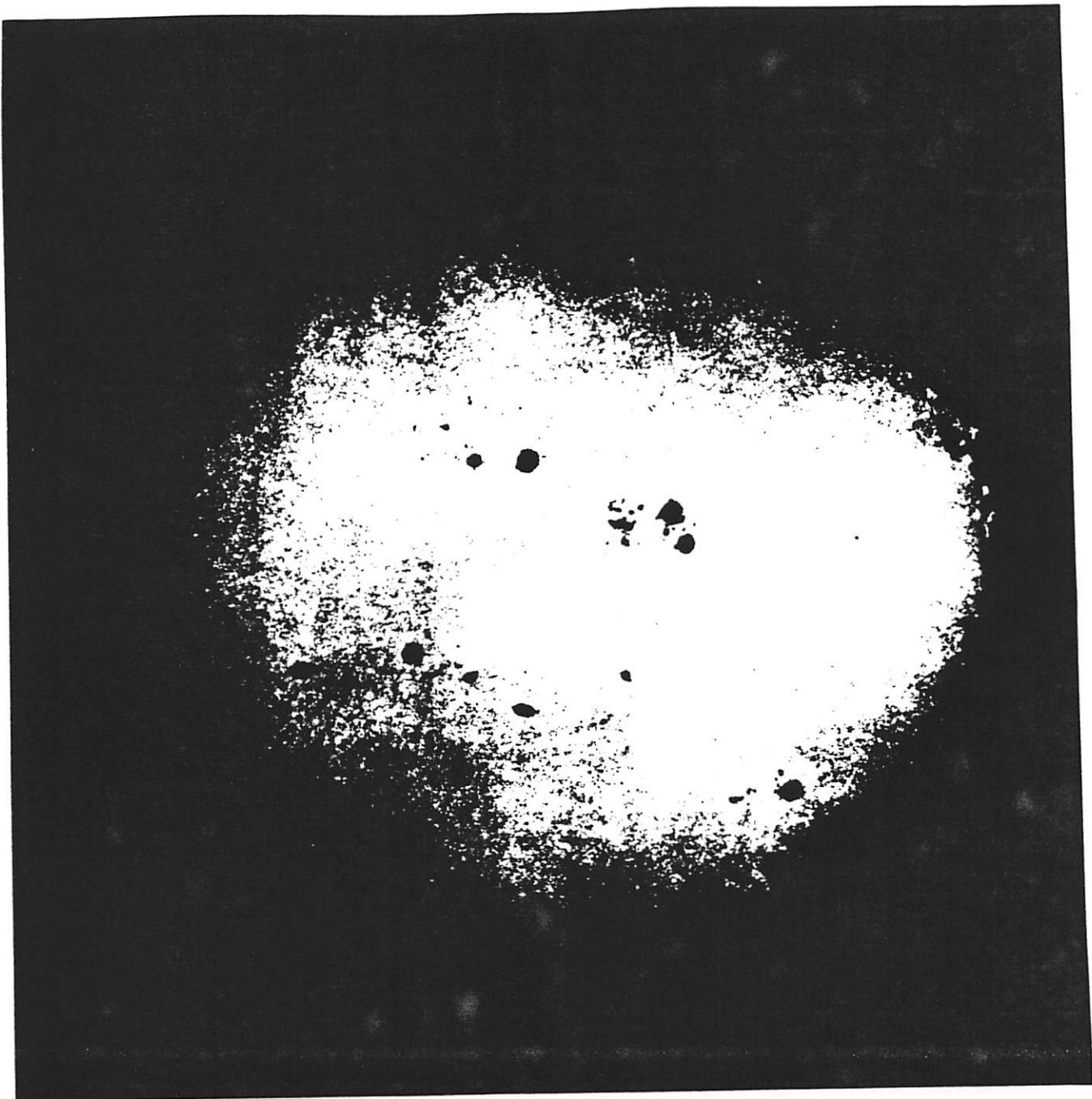
*...it is conceivable that solar variability plays
a role in altering weather and climate at some
yet unspecified level of significance ...*

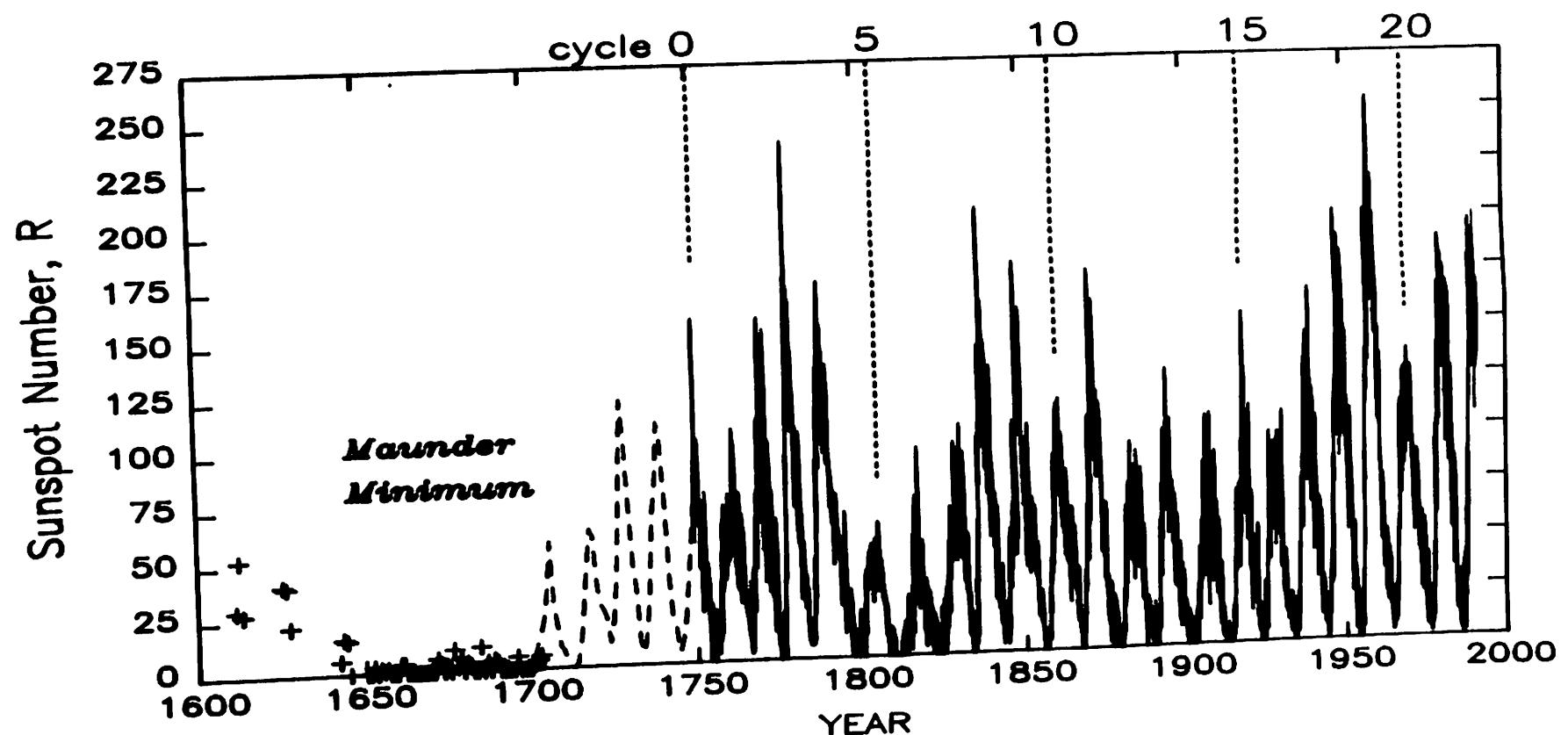
Historical Solar Activity

Eddy (1978)



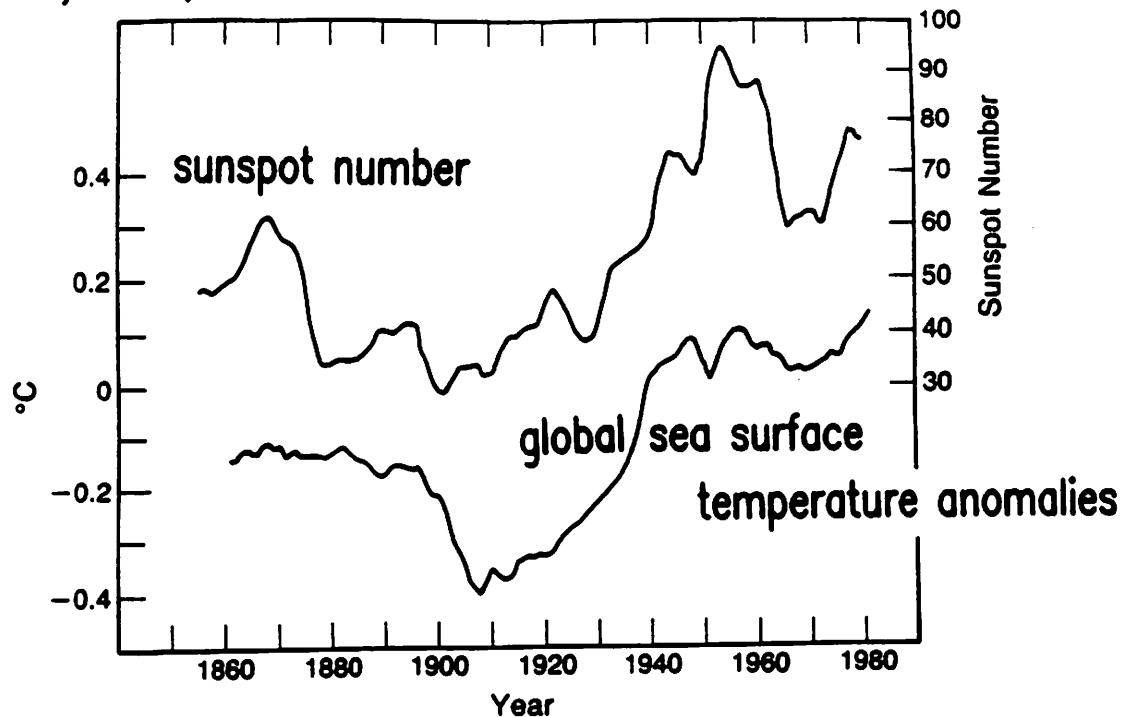
M. WILSON
Palomar Obs.
near sunspot max.
July 13, 1937



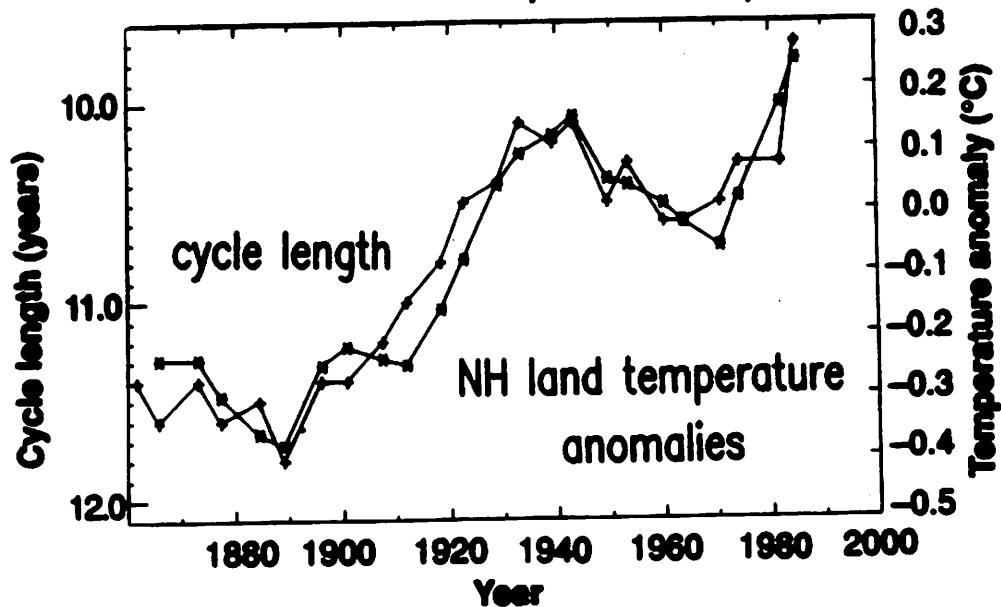


SOLAR VARIABILITY AND TEMPERATURE

Reid, JGR, 1991

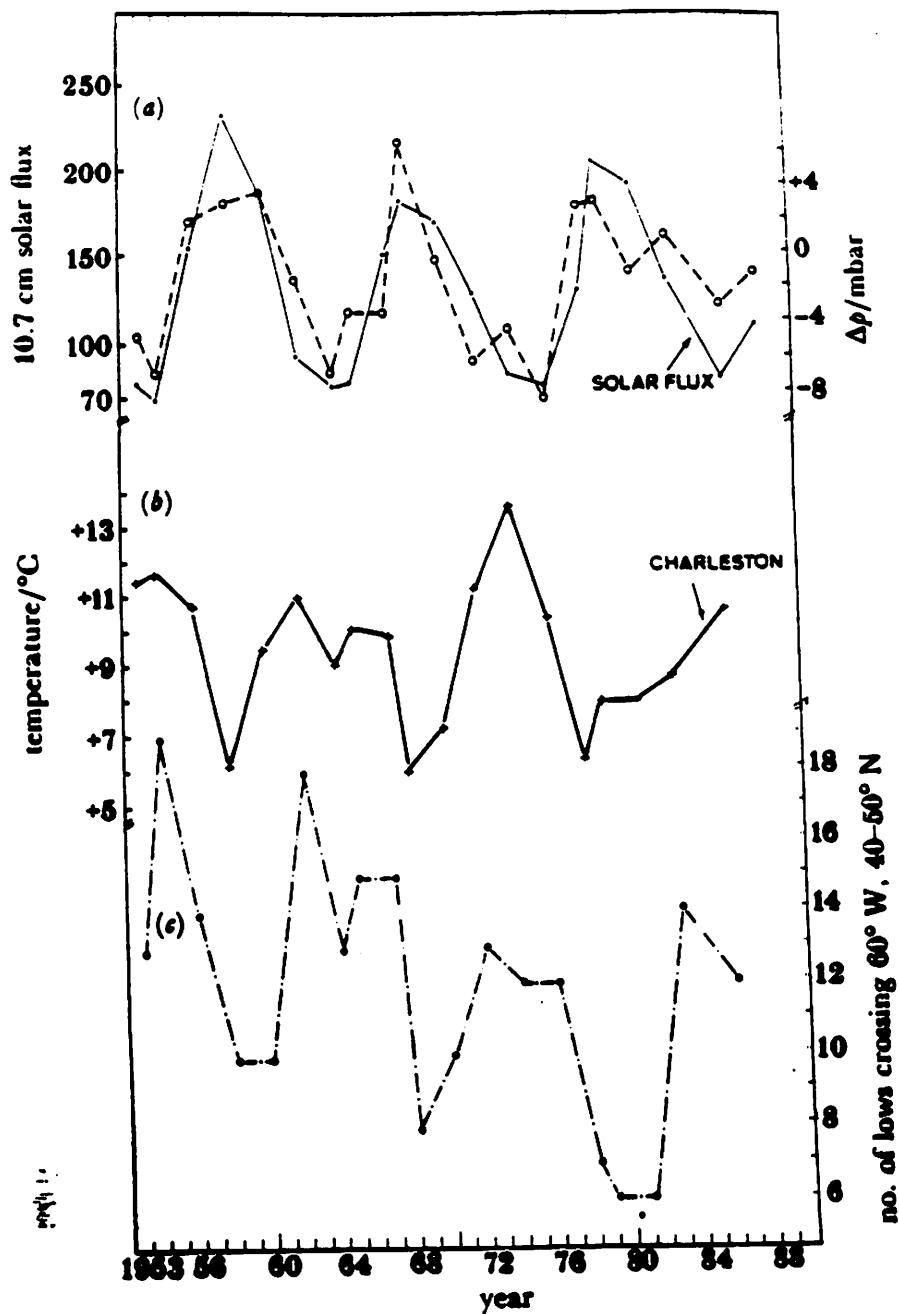


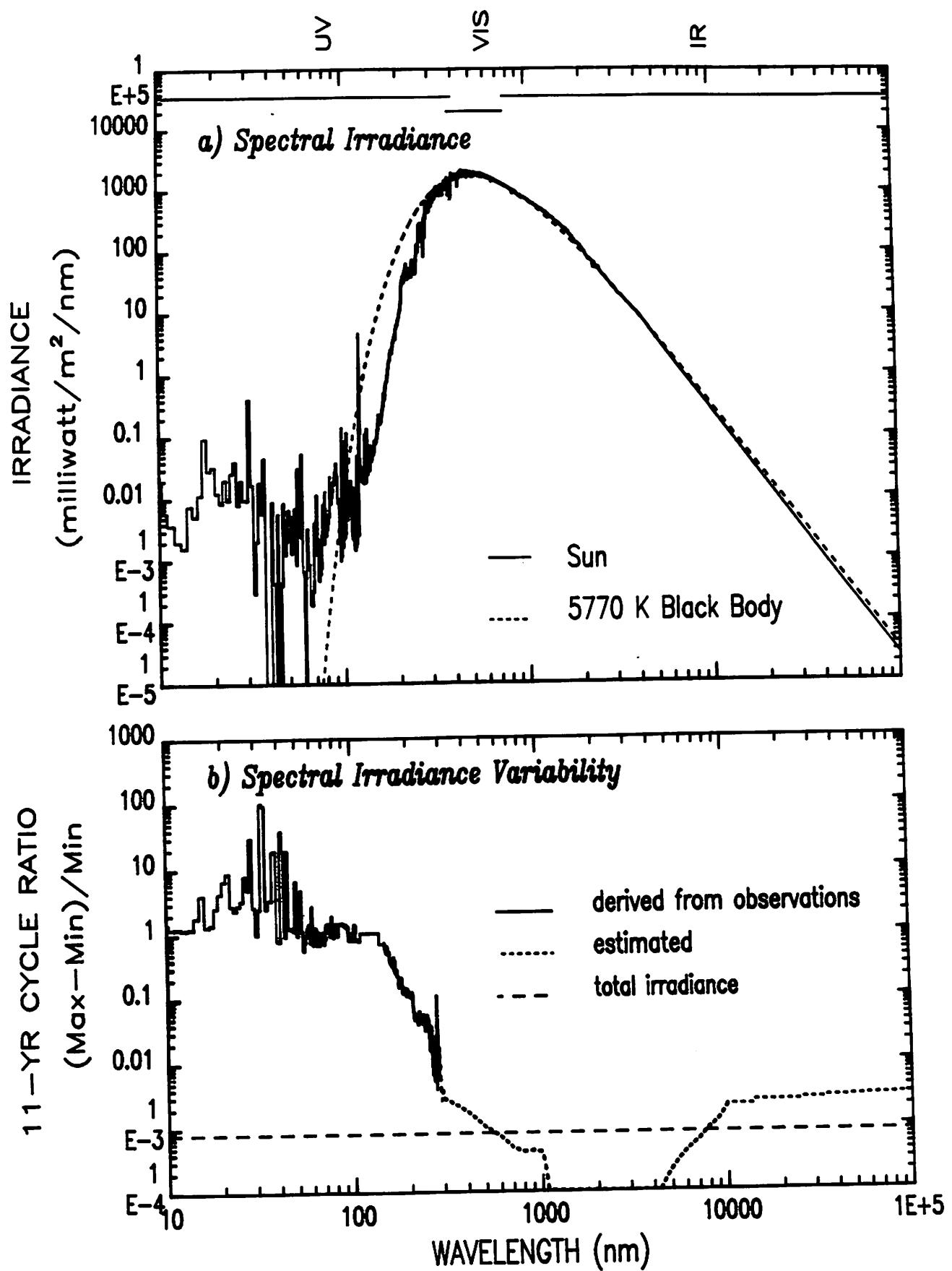
Friis-Christensen & Lassen, Science, 1991

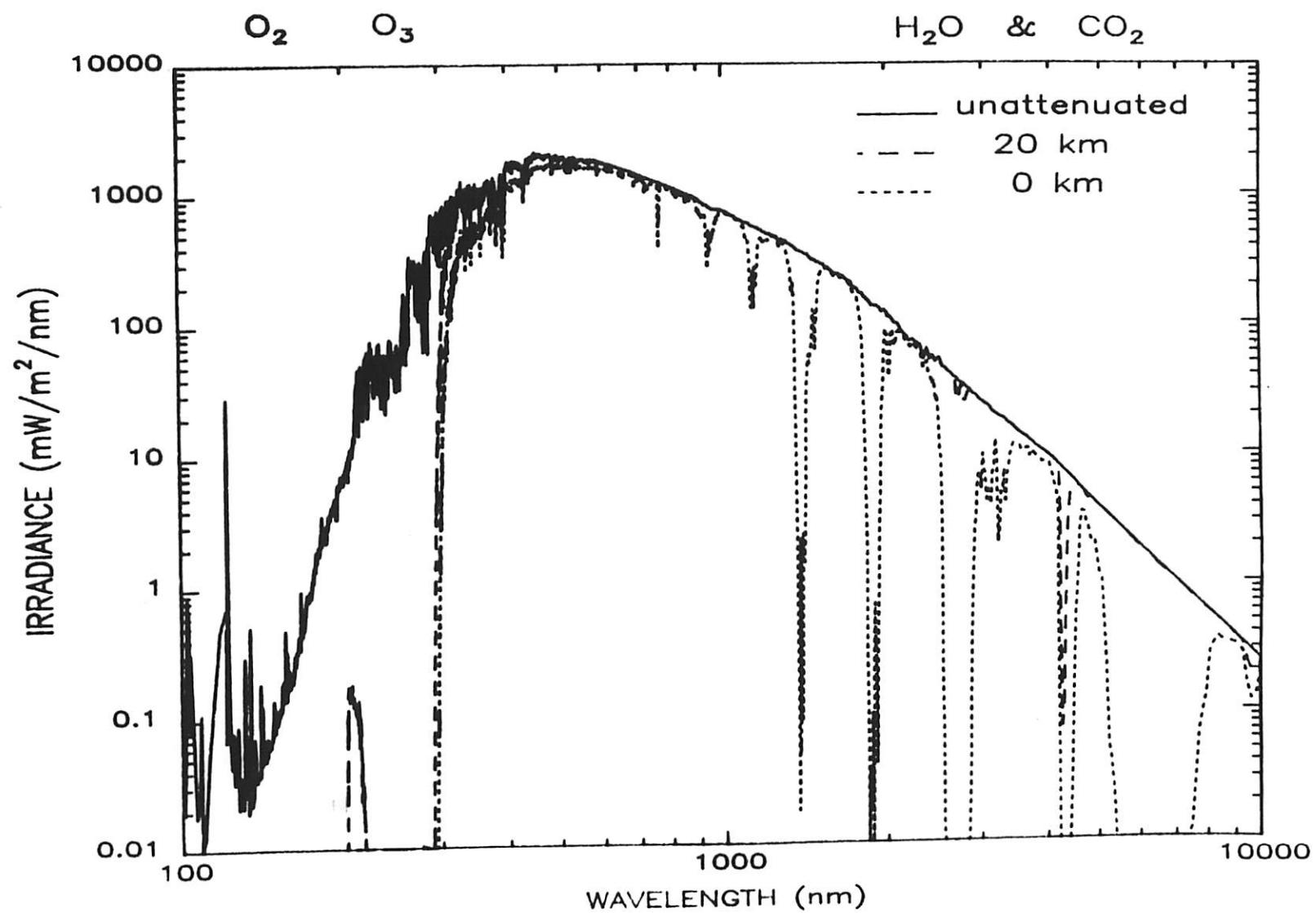


Solar Variability, the Quasi Biennial Oscillation and the Weather

Labitzke and van Loon, 1990

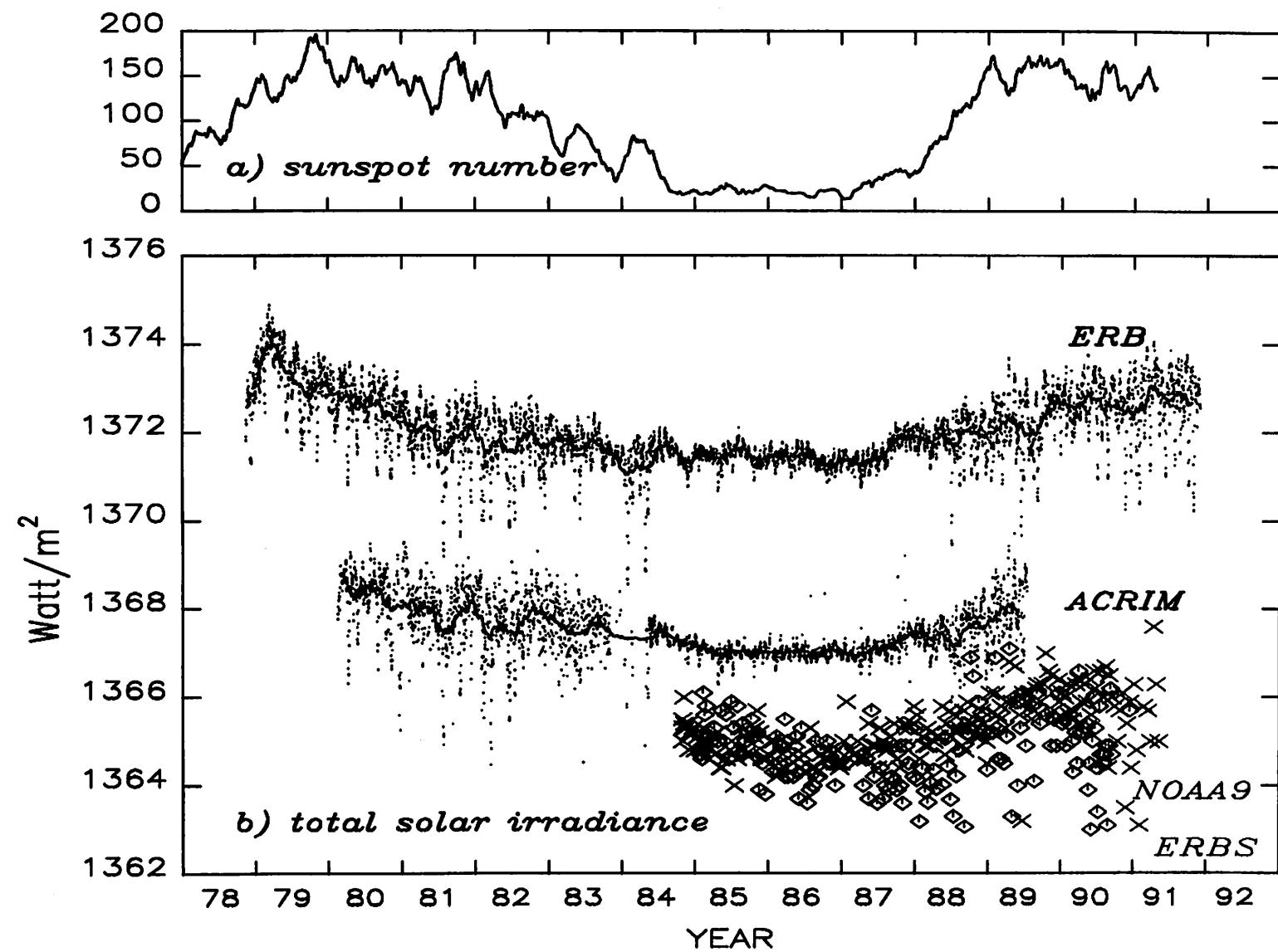






SOLAR-TERRESTRIAL ENERGY SOURCES

Source	Energy (Watt/m ²)	Solar Cycle Change (Watt/m ²)	Deposition Altitude
<i>Solar Radiation</i>			
total solar irradiance	1370	1.2	surface
UV 200–300 nm	16	0.14	0–50 km
VUV 0–200 nm	0.15	0.15	50–500 km
<i>Particles</i>			
Solar protons	0.002		30–90 km
Proton aurora, disturbed	0.036		30–90 km
Galactic cosmic rays	0.0000007		0–90 km
<i>Joule Heating of Thermosphere</i>			
E=100 mV/m	0.14		100–500 km
<i>Solar Wind</i>			
	0.0003		above 500 km



SOLAR FORCING OF GLOBAL SURFACE TEMPERATURE

Change in equilibrium global temperature ΔT for change in input solar energy ΔS is

$$\Delta T = \lambda^{-1} 0.7\Delta S/4$$

where

$$\lambda^{-1} \sim 0.3 \text{ to } 1.0 \text{ }^{\circ}\text{C}/\text{Wm}^{-2}$$

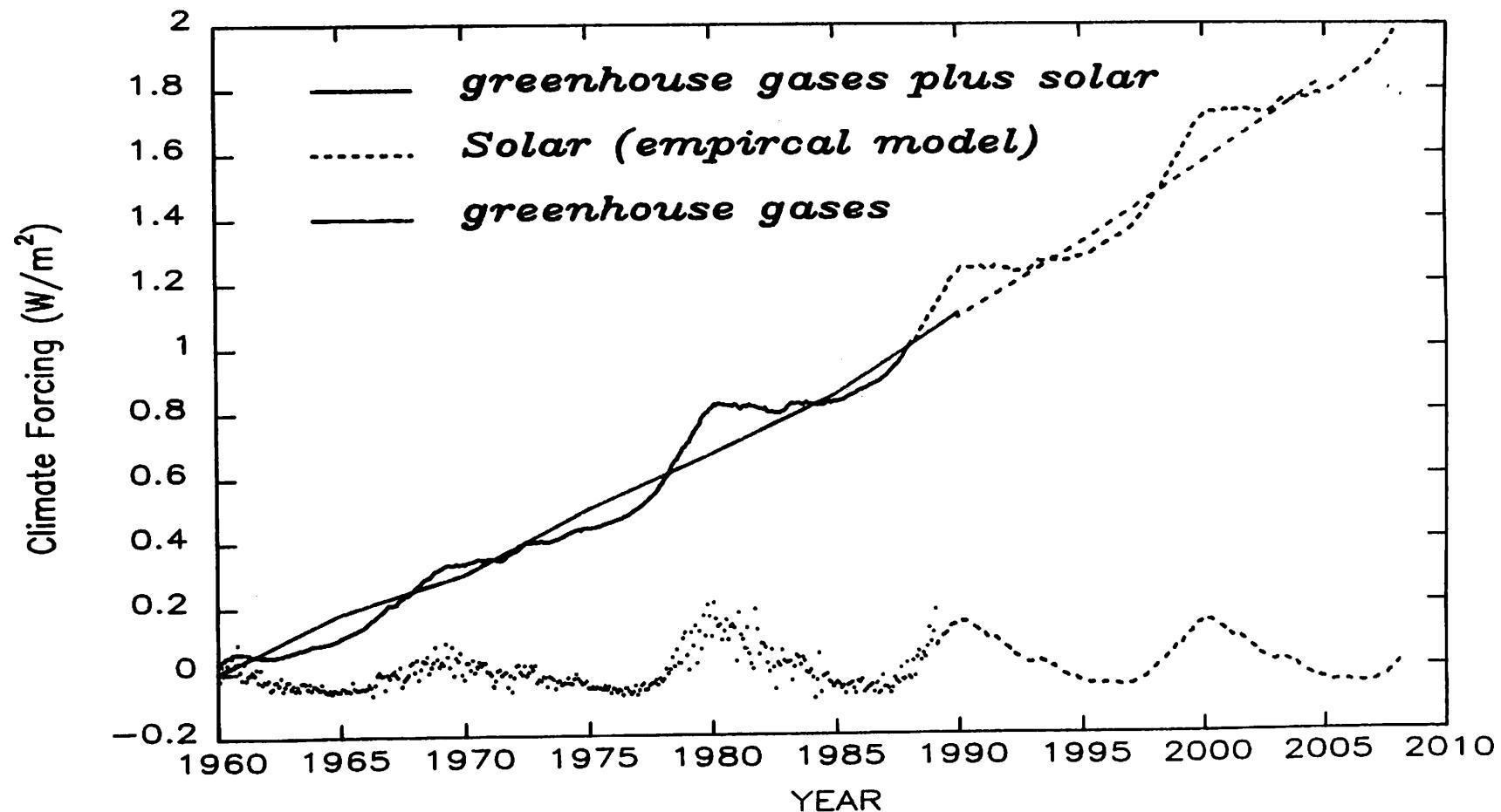
*is the climate sensitivity
(Wigley & Raper, GRL, 1990).*

Cycle 21 solar max to solar min

$$\Delta S = 1.2 \text{ Wm}^{-2} \quad \Delta T = 0.06 - 0.21 \text{ }^{\circ}\text{C}$$

SOLAR vs. GREENHOUSE CLIMATE FORCING

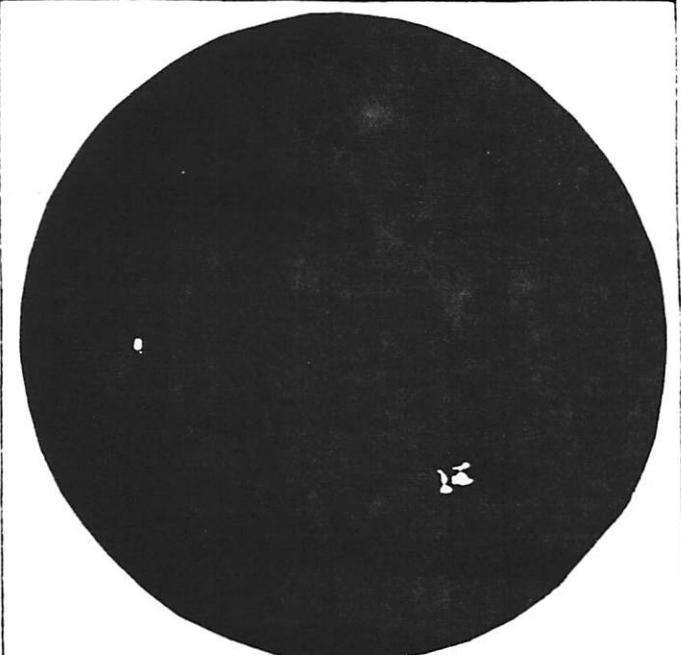
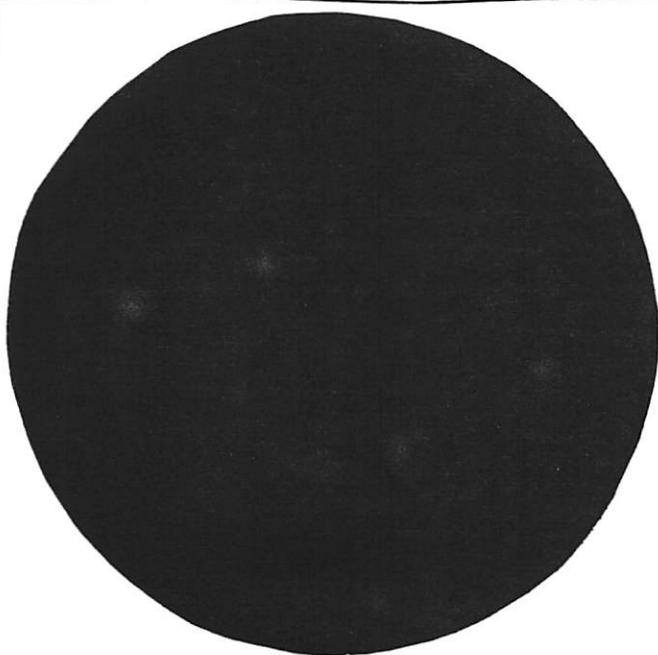
adapted from Hansen & Lacis (1990) and Foukal & Lean (1990)



10 MAY 75

photosphere

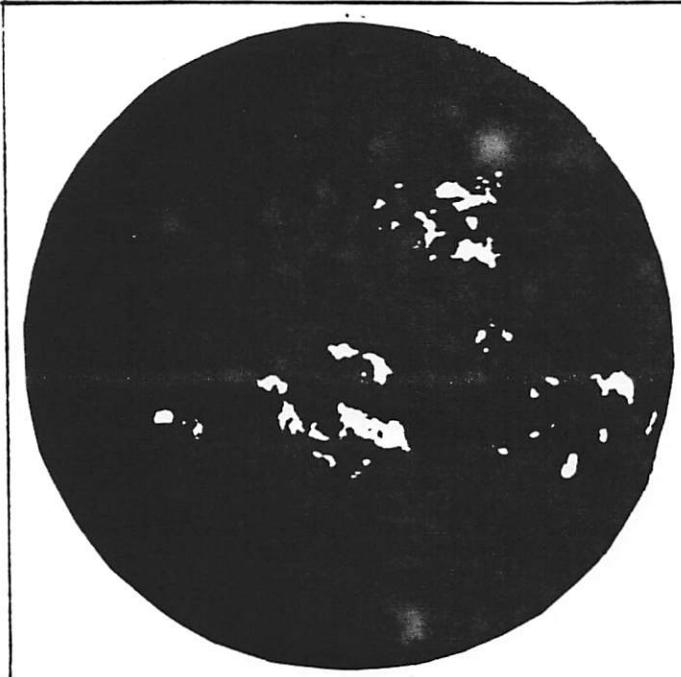
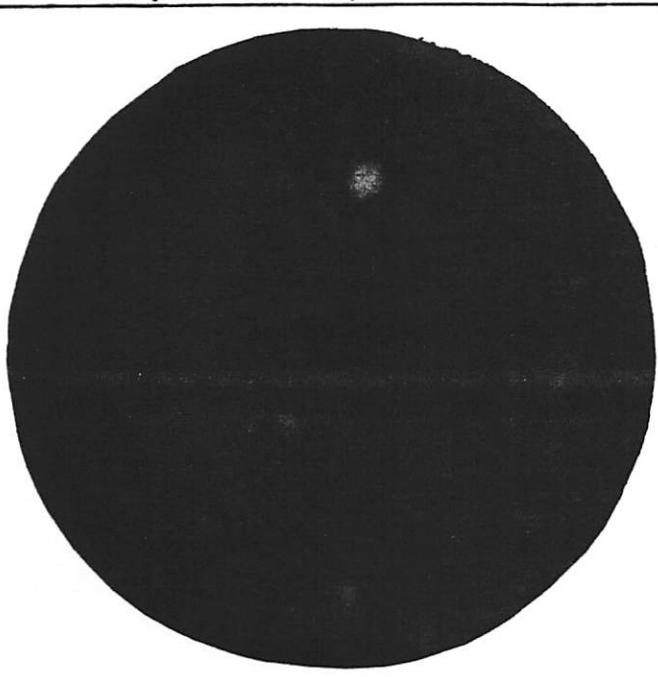
chromosphere

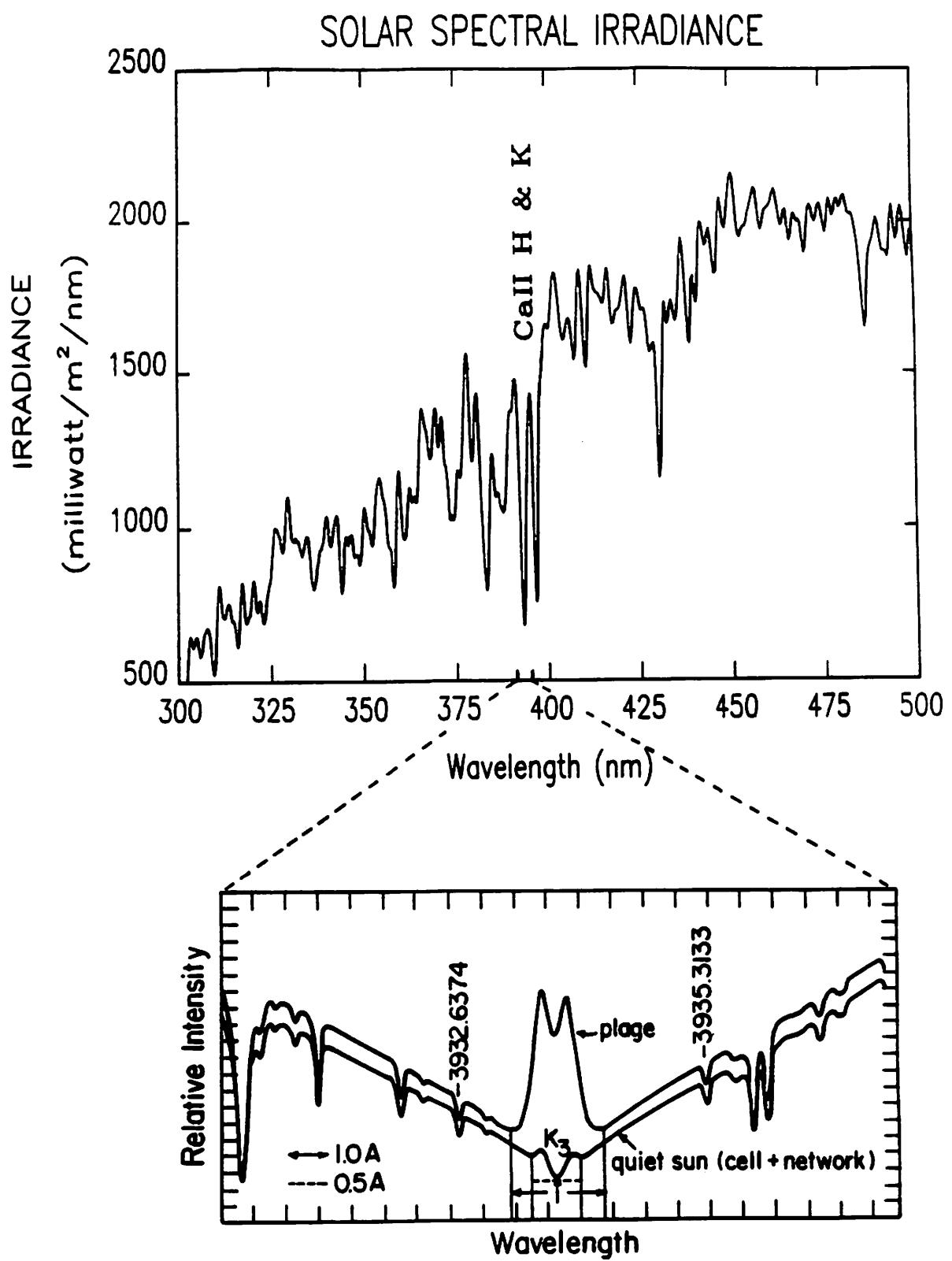


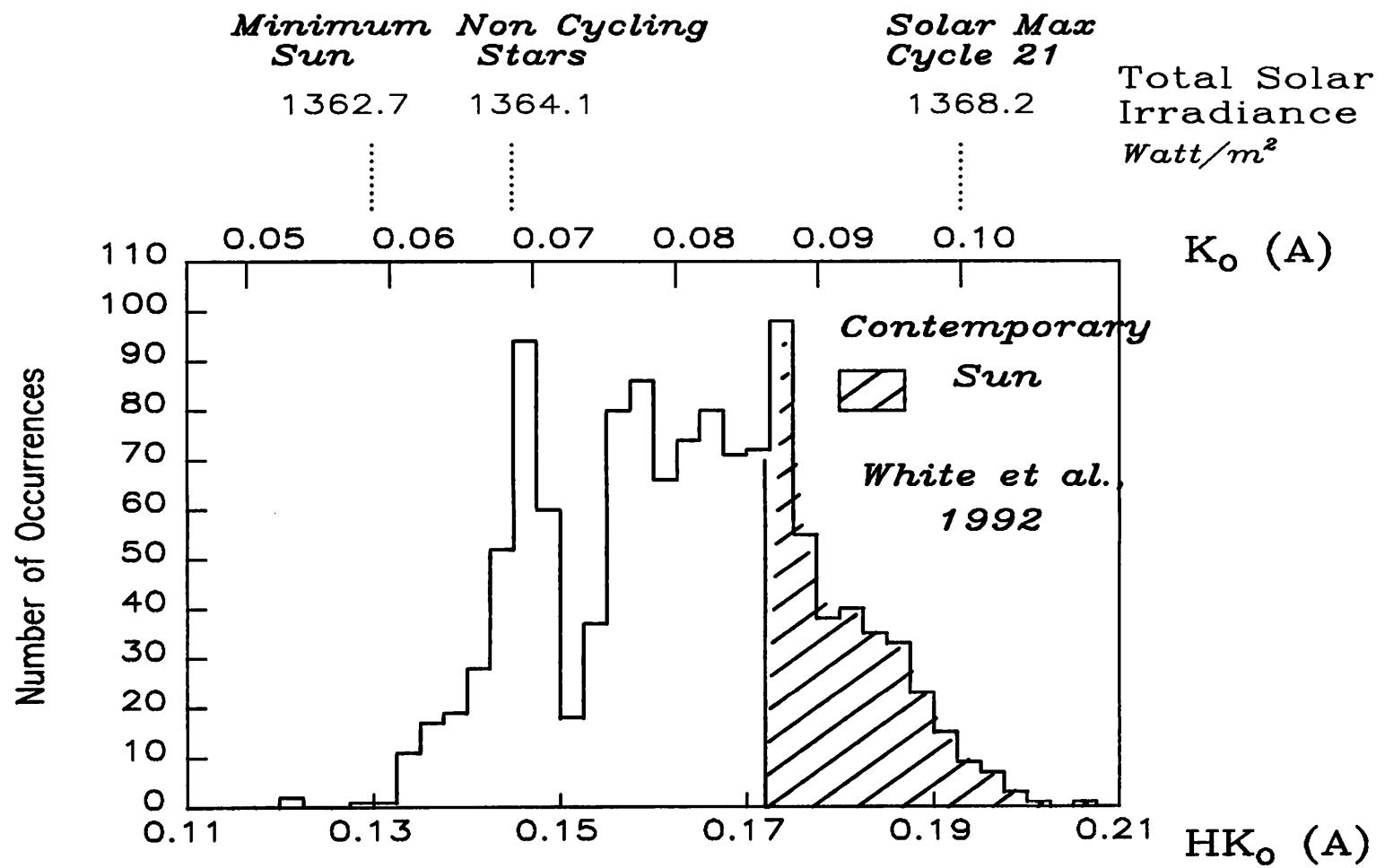
7 MARCH 79

photosphere

chromosphere







SOLAR FORCING OF GLOBAL SURFACE TEMPERATURE

Change in equilibrium global temperature ΔT for change in input solar energy ΔS is

$$\Delta T = \lambda^{-1} 0.7\Delta S/4$$

where

$$\lambda^{-1} \sim 0.3 \text{ to } 1.0^{\circ}\text{C}/\text{Wattm}^{-2}$$

is the climate sensitivity

(Wigley & Raper, GRL, 1990)

Cycle 21 solar max to solar min

$$\Delta S = 1.2 \text{ Watt/m}^2 \quad \Delta T = 0.06 - 0.21^{\circ}\text{C}$$

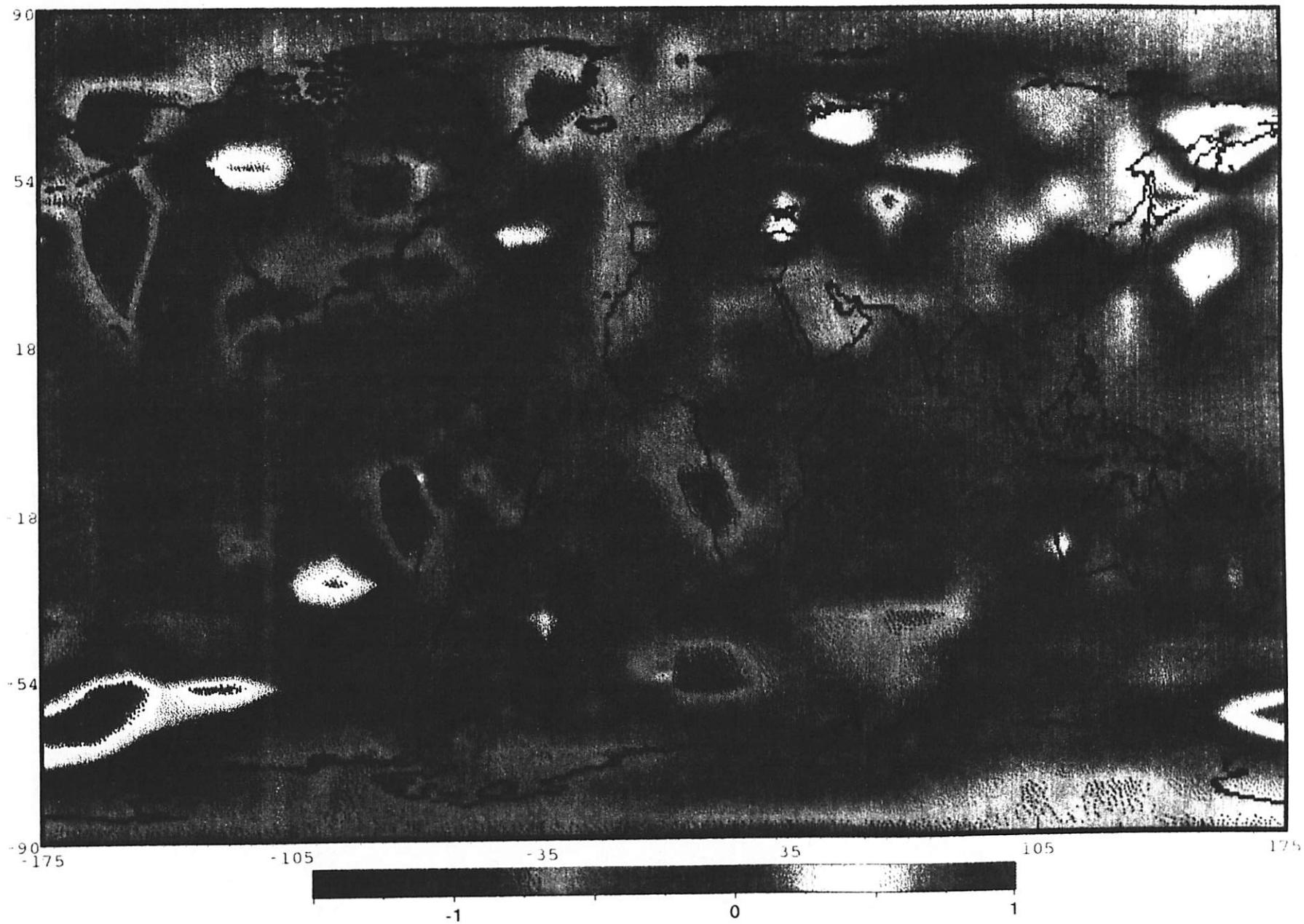
Cycle 21 solar max to Maunder Minimum

$$\Delta S = 3.9 \text{ Watt/m}^2 \quad \Delta T = 0.2 - 0.68^{\circ}\text{C}$$

Δ SURF AIR TEMP ($^{\circ}$ C)

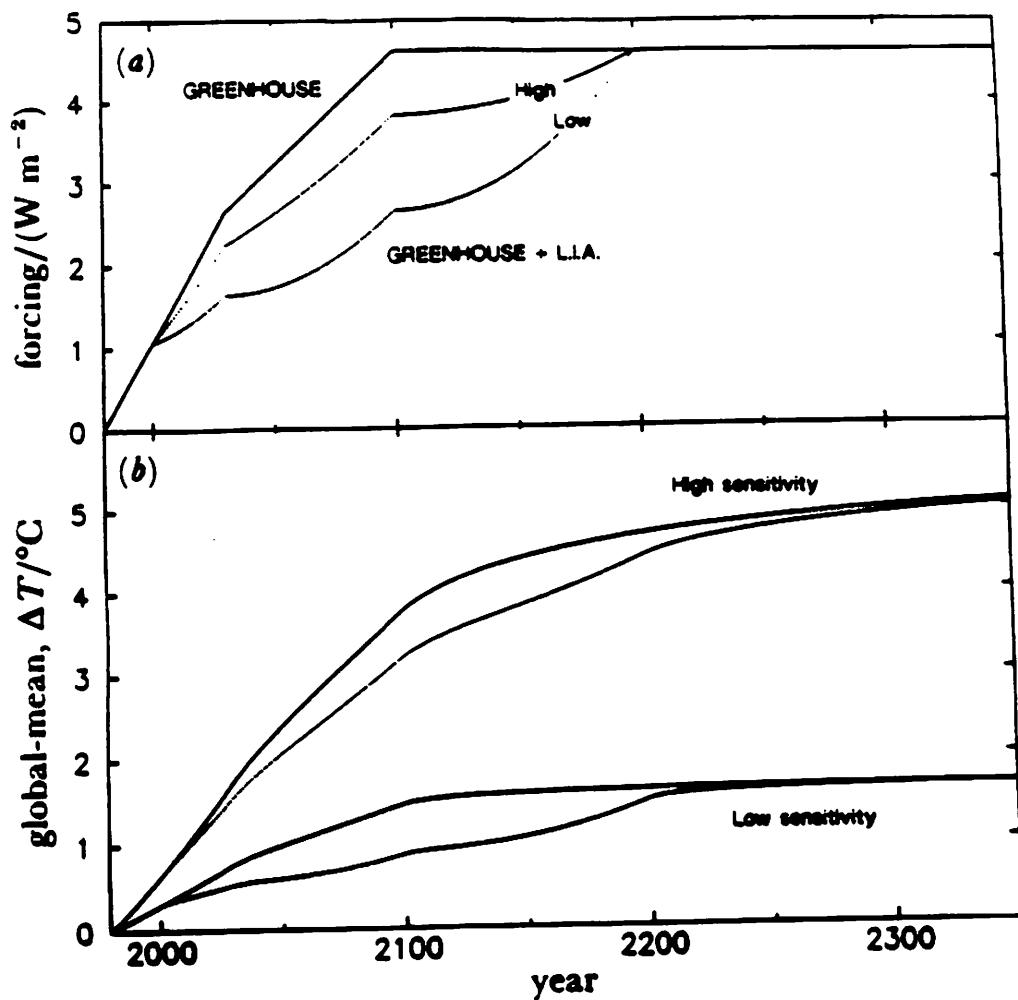
ANNUAL

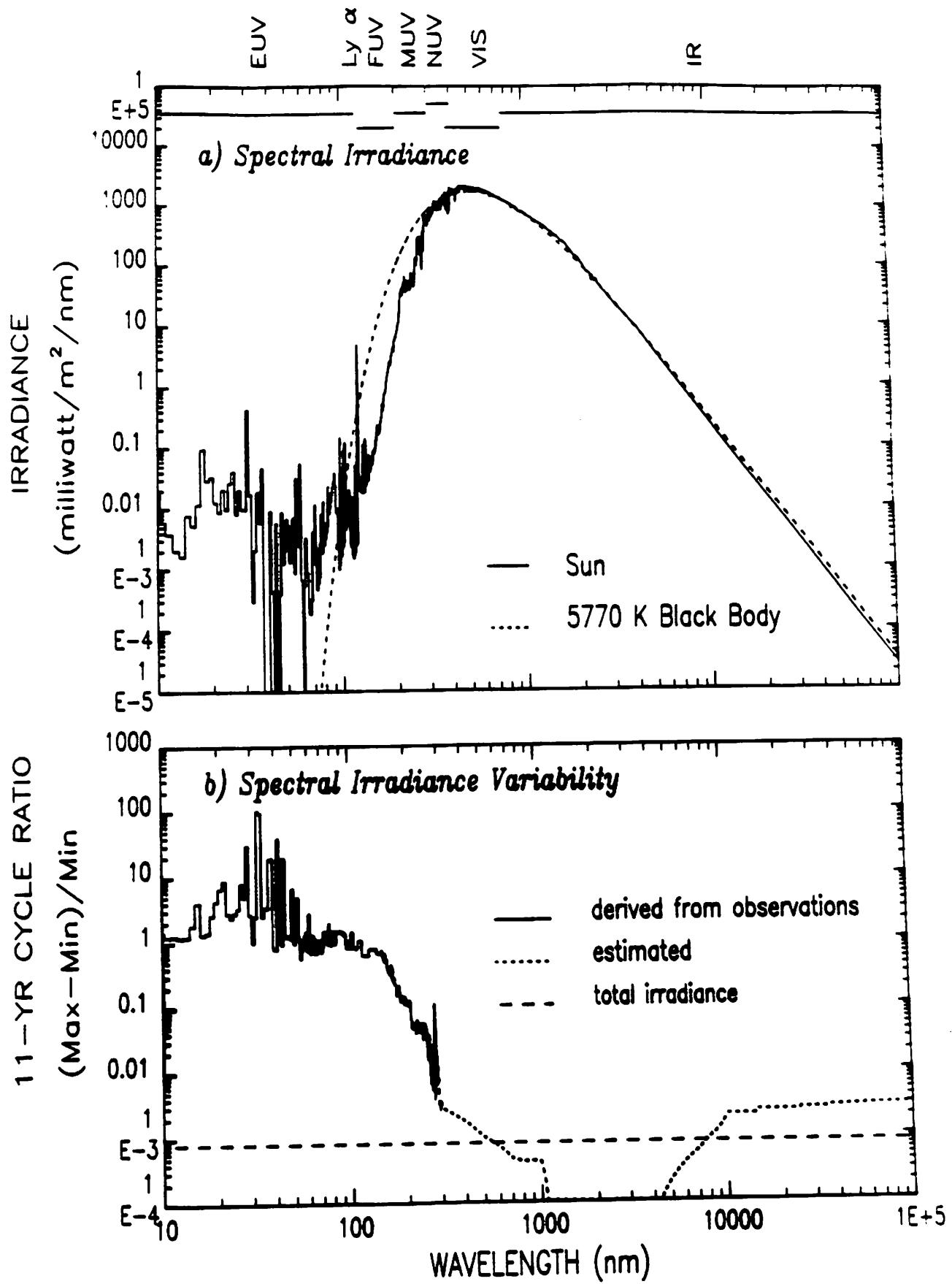
-25% S.C.-CONTROL



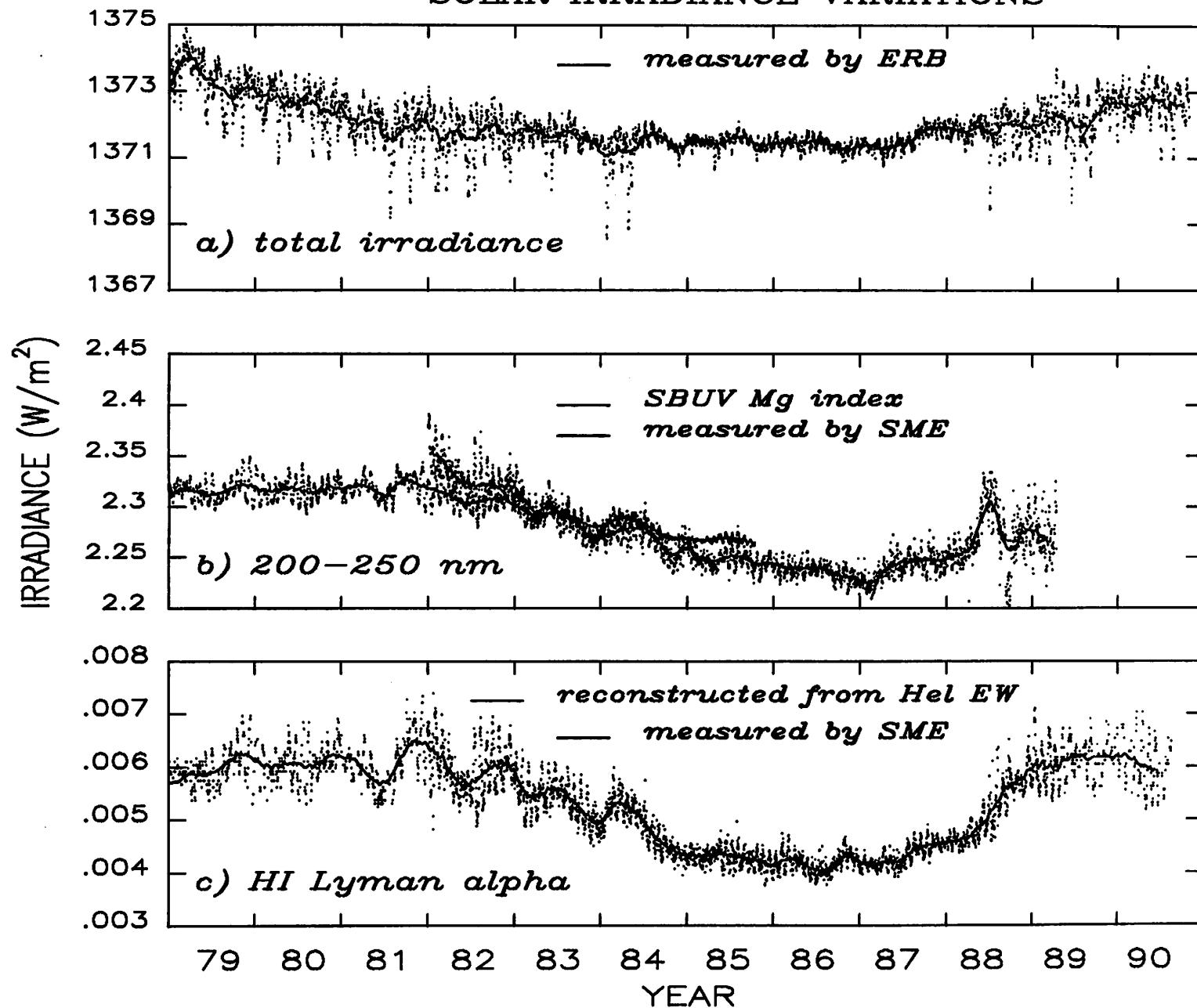
Scenarios for Future Climate Change: Greenhouse Gases vs. Solar "Little Ice Age"

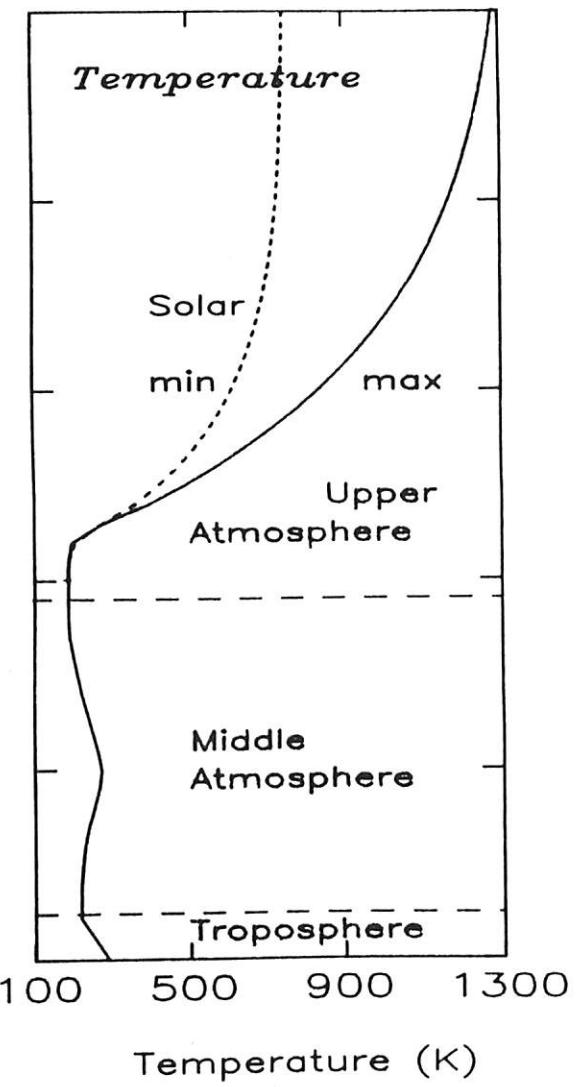
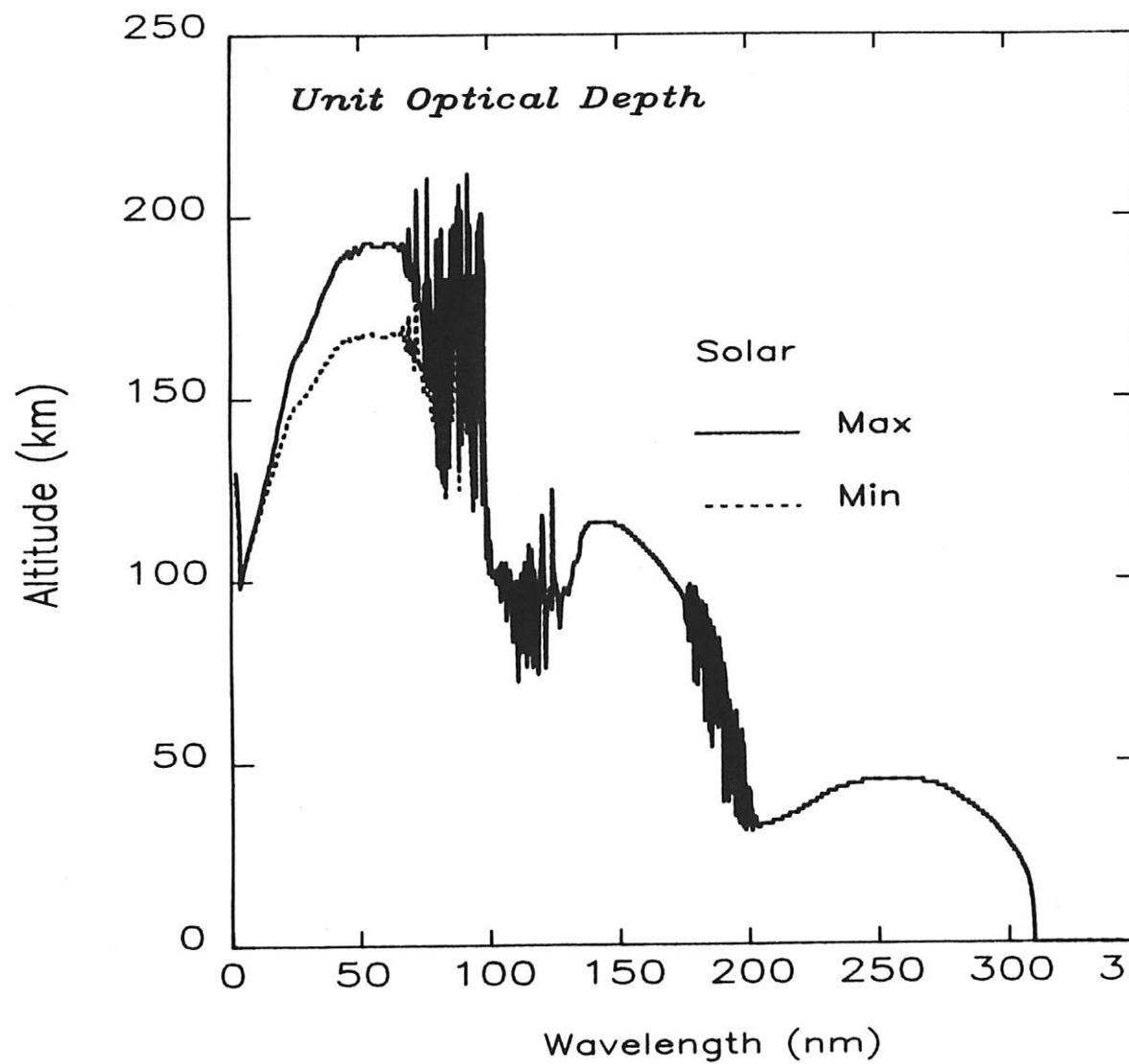
Wigley and Kelly, 1990



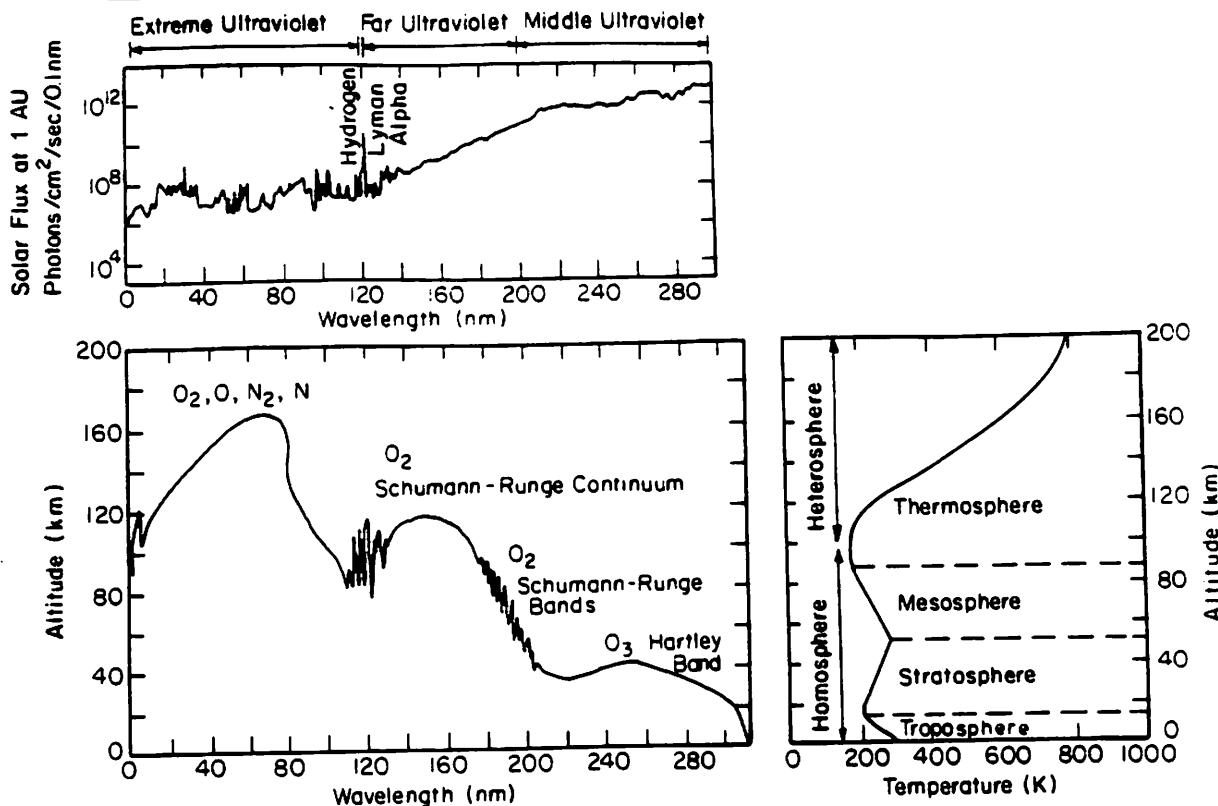


SOLAR IRRADIANCE VARIATIONS



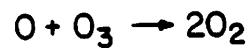
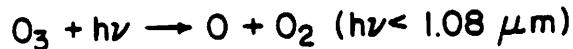
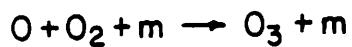
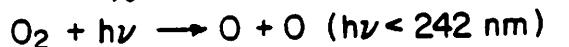


1. Radiation

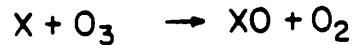


2. Chemistry

O_x Photochemistry

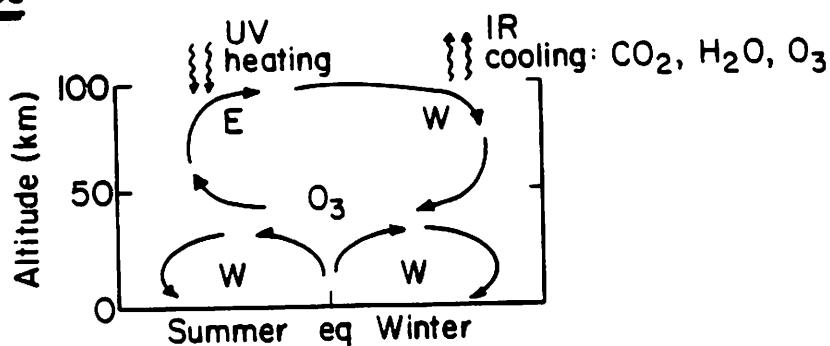


Catalytic Ozone Destruction



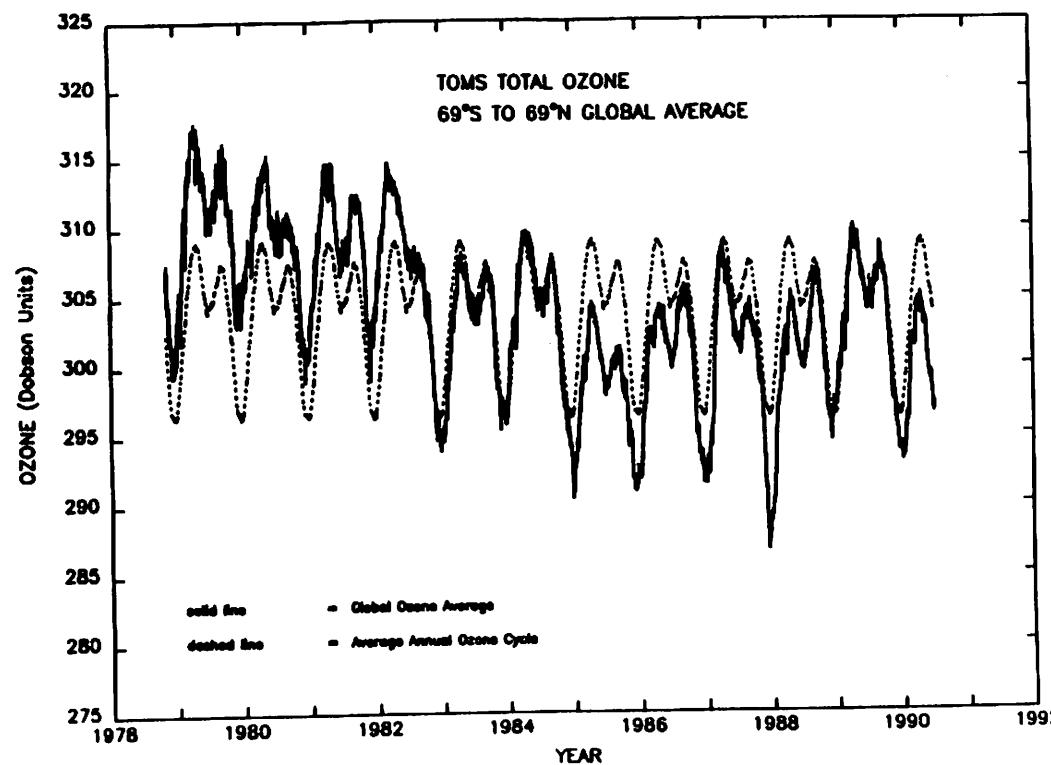
X: HO_x, NO_x, Cl_x

3. Dynamics

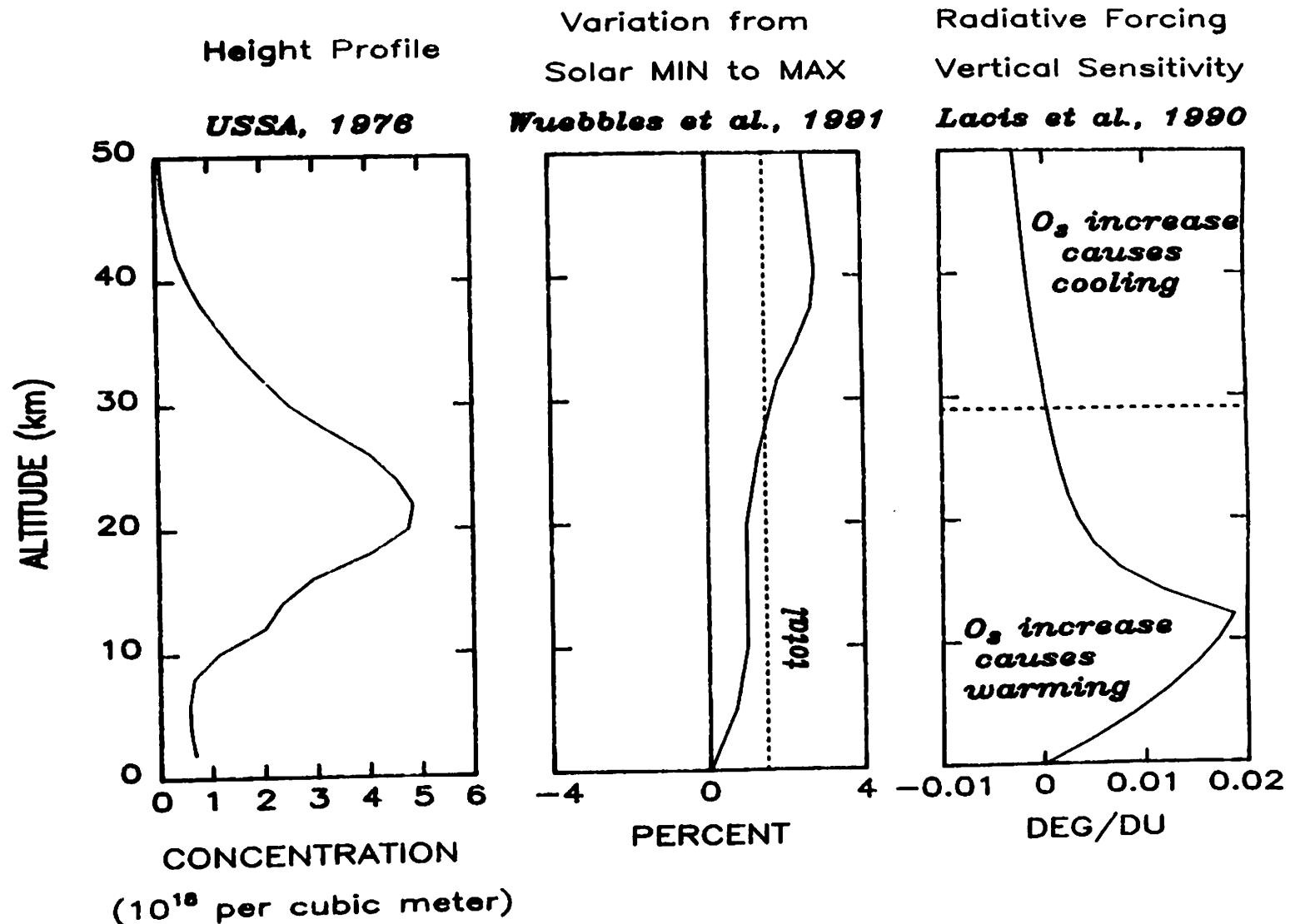


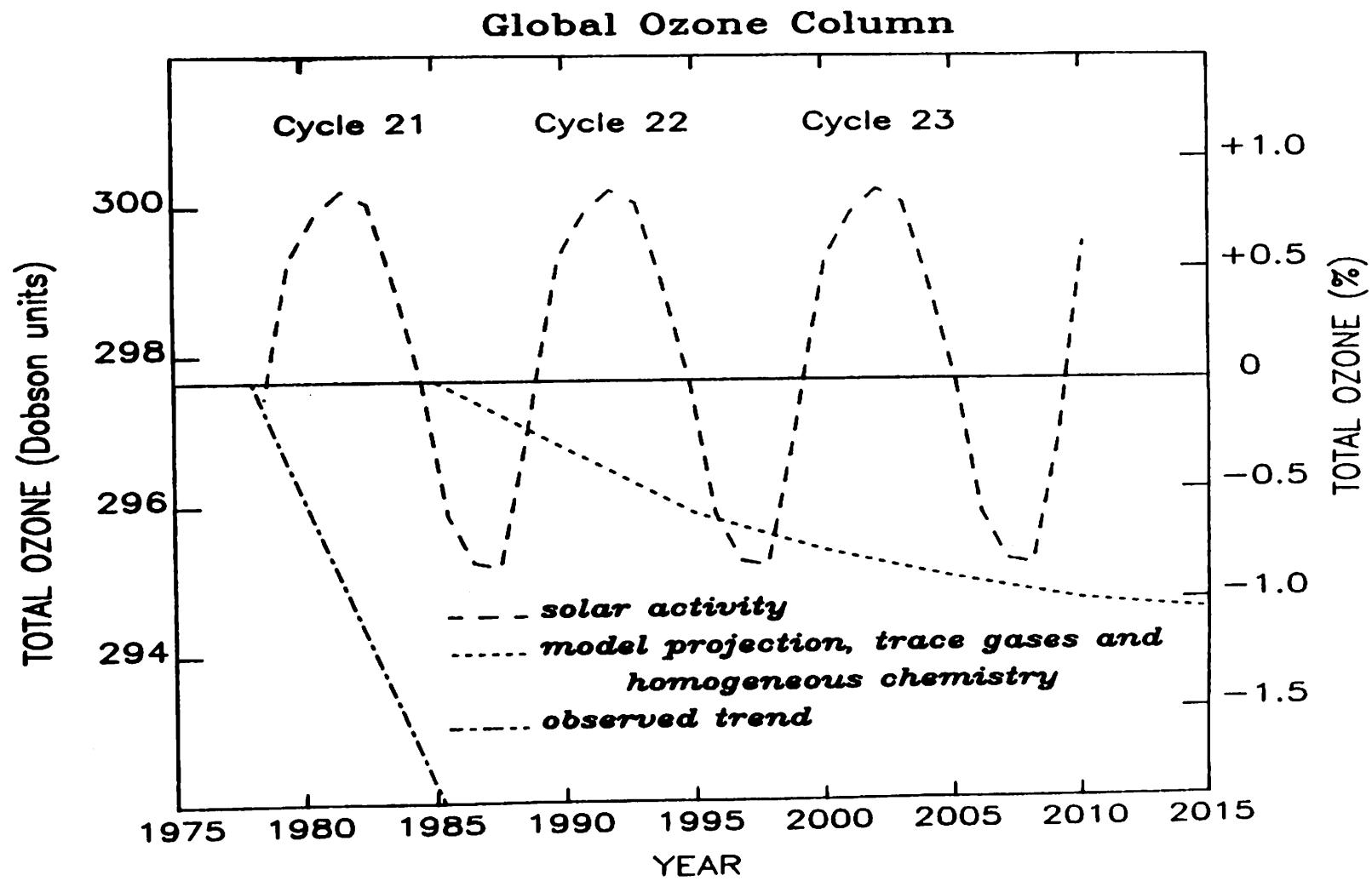
LONG TERM VARIATIONS IN TOTAL OZONE

Herman et al., J. Geophys. Res., 1991

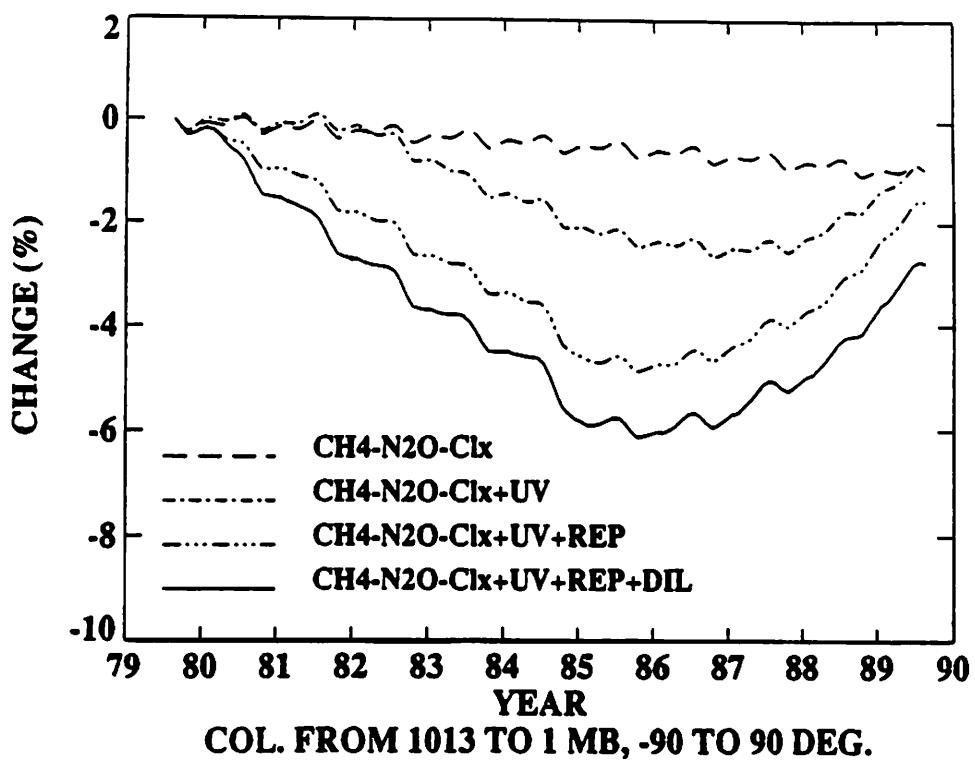


ATMOSPHERIC OZONE





Callis et al., J. Geophys. Res., 1991

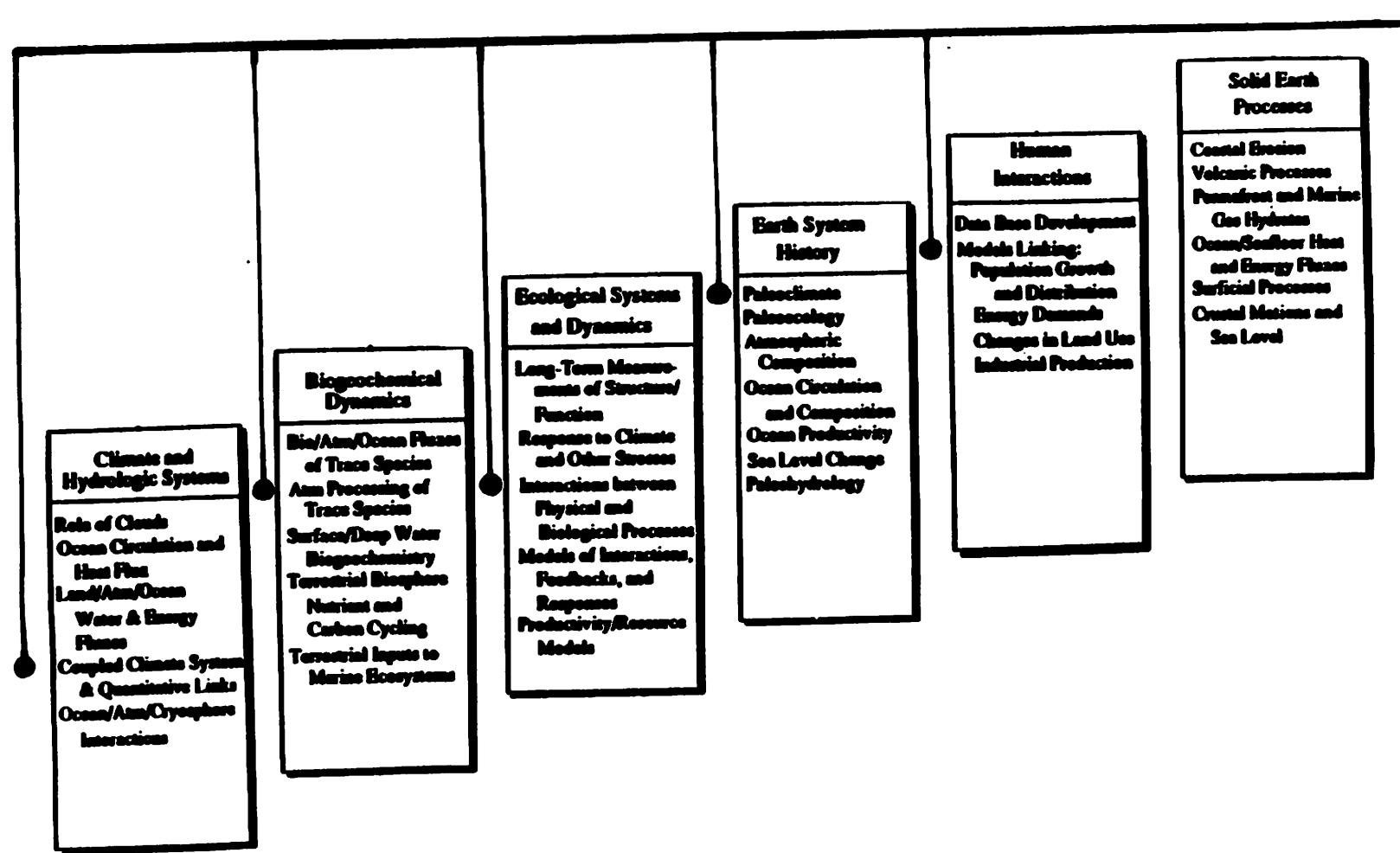


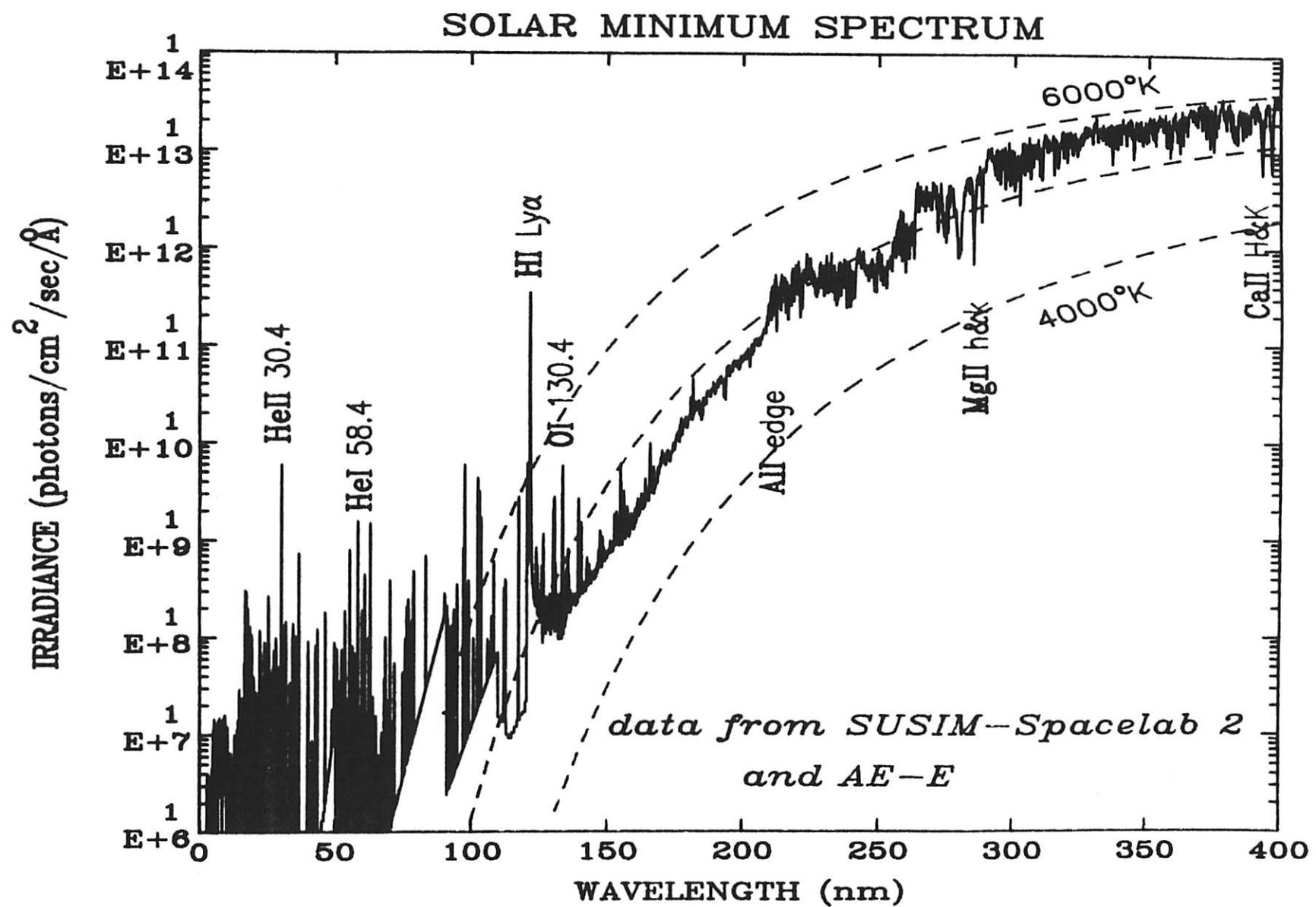
Percentage changes in global (90 S to 90 N) total column ozone as calculated for four scenarios.

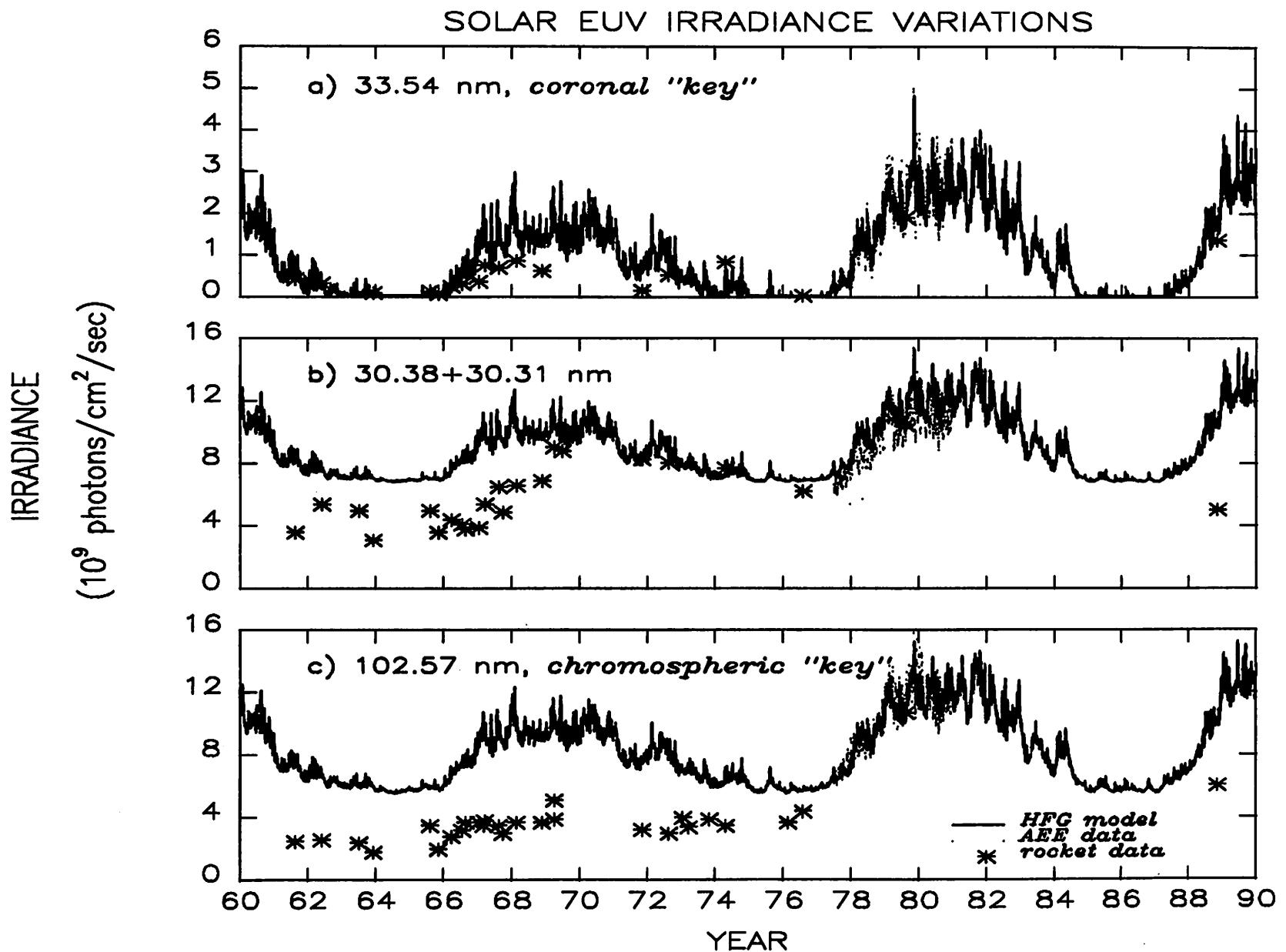
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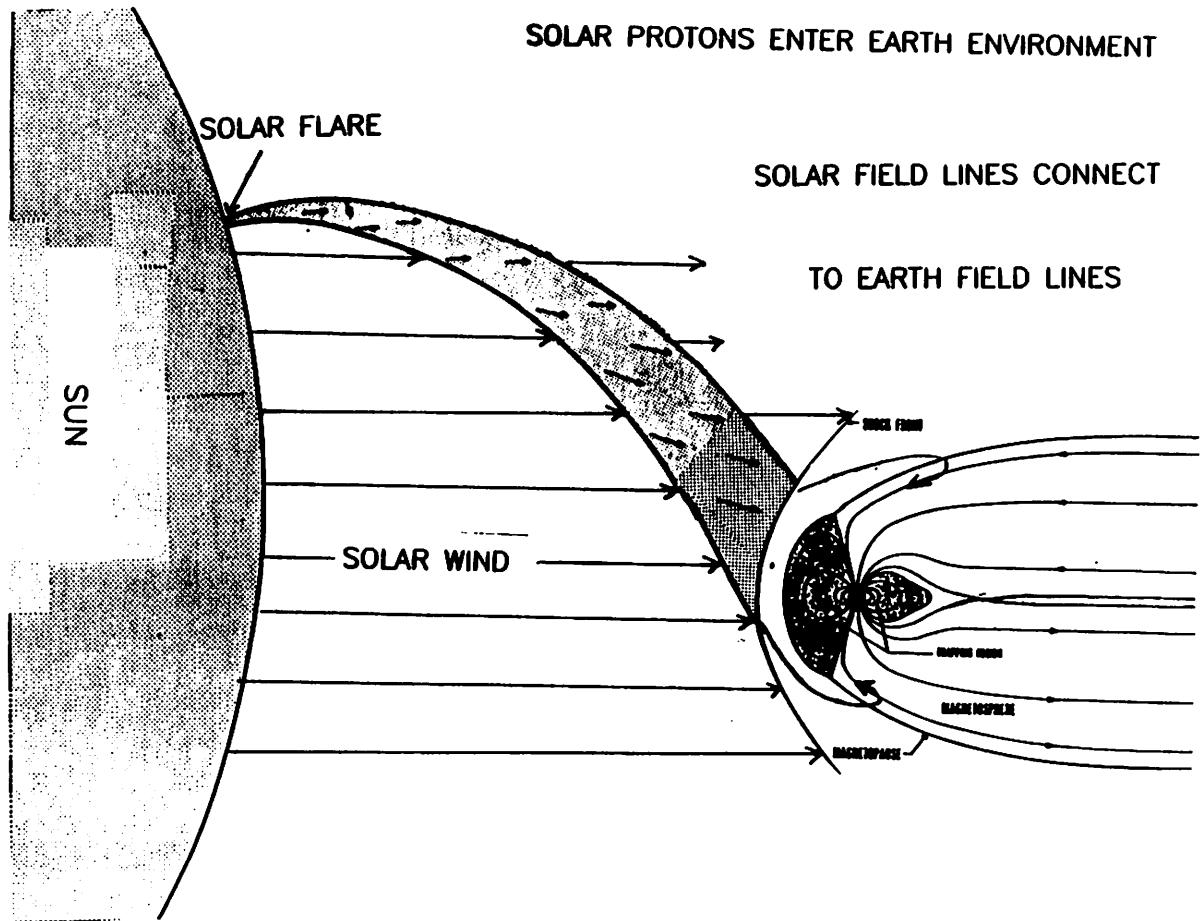
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Connections between Solar Influences on Global Change and other elements of the US Global Change Research Program

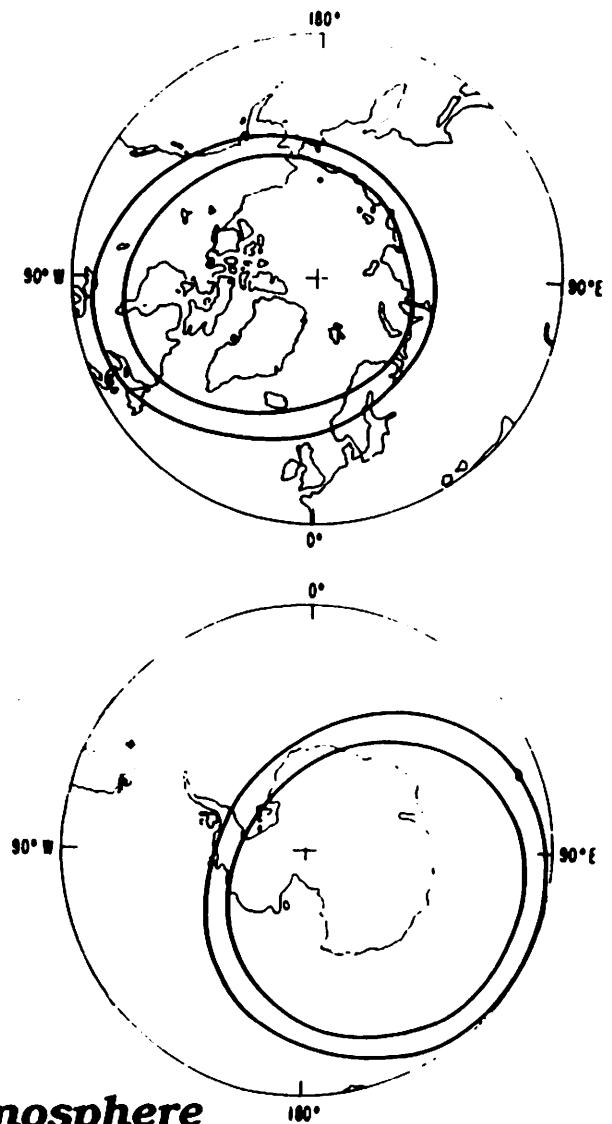






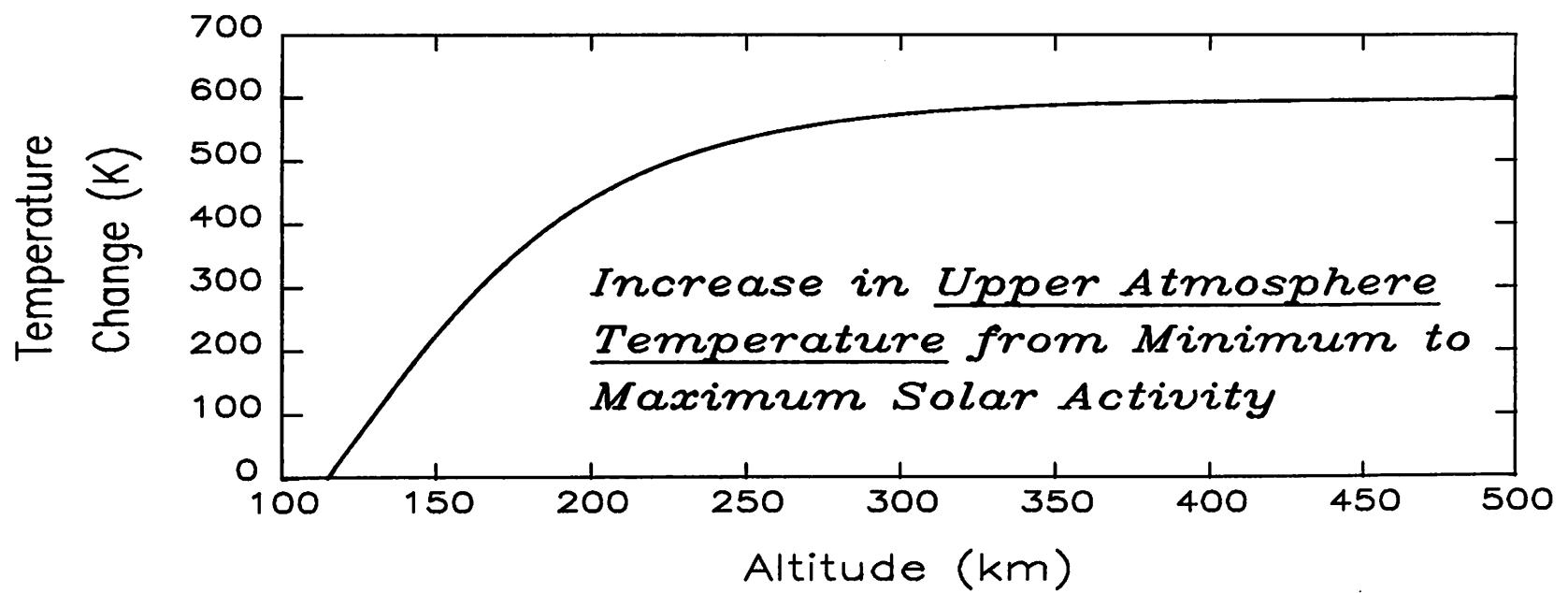
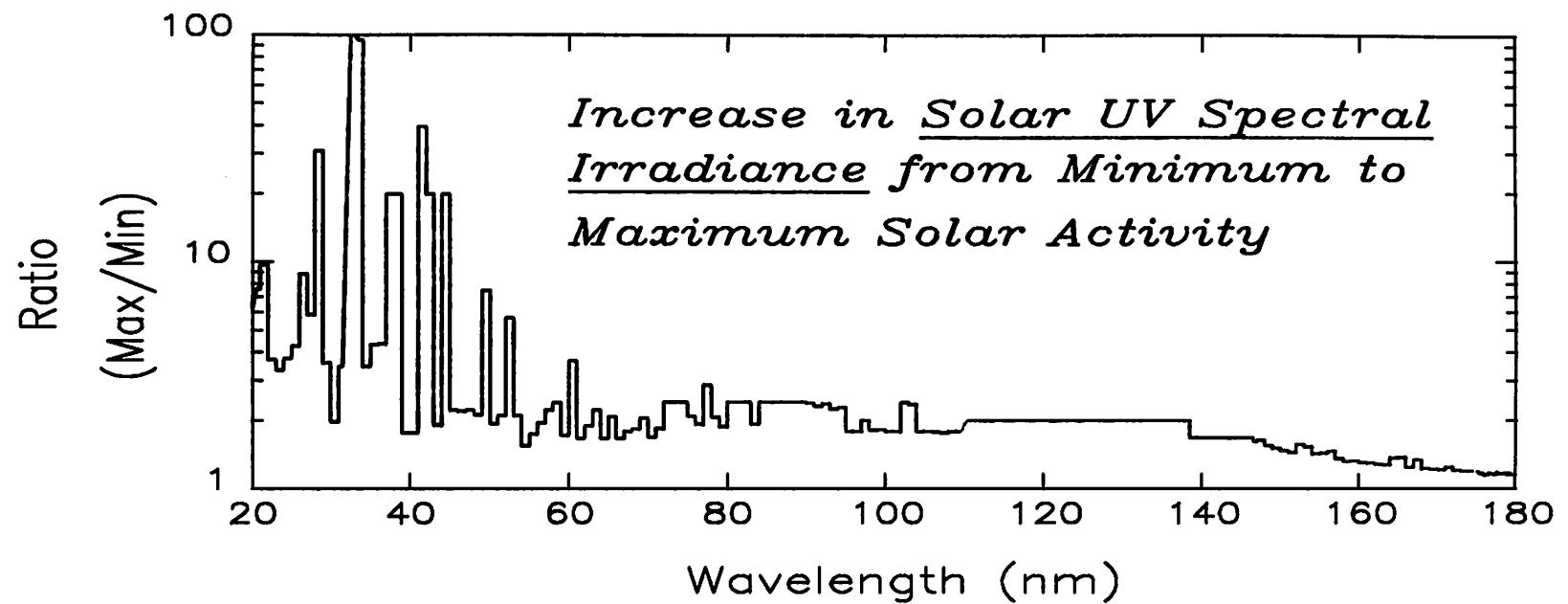


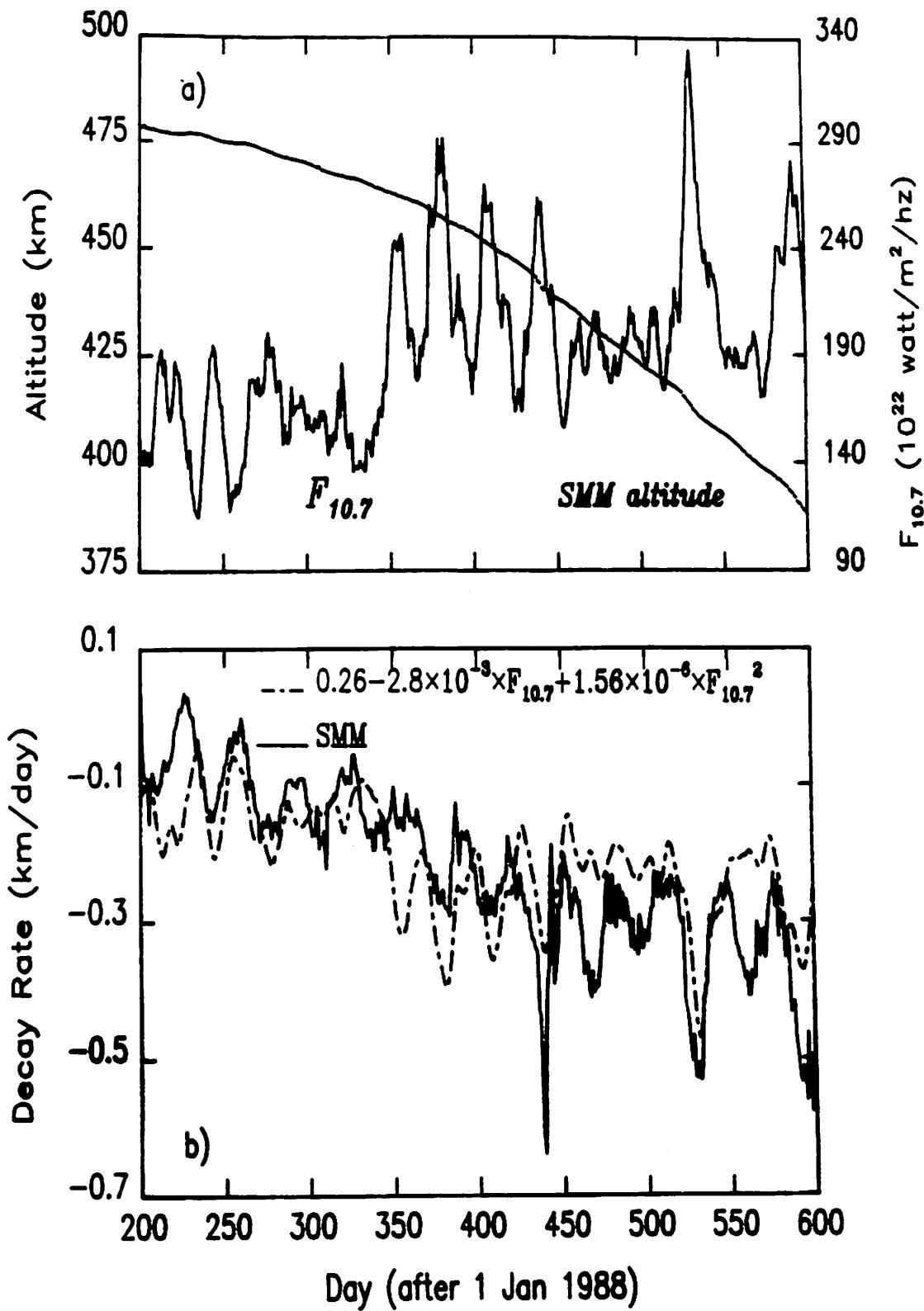
PROTONS ENTER OVER POLAR CAPS



Penetration of energetic particles into the Earth's atmosphere

Adapted from C. Jackman and G. Reid





UPPER ATMOSPHERE SENSITIVITY TO CO₂ & CH₄

Roble & Dickinson, Geophys. Res. Lett., 1989

solid=double

dashed=half

