

Jicamarca  
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What is special about Jicamarca?

- 1- Large Power-Aperture  $360,000 \text{ MW} \cdot \text{m}^2$
- 2- Location - Magnetic Equator (Eq. phenomena)
  - Point  $\perp$  to  $\vec{B}$  - IS  
 $\backslash$  FAI.
- 3- Wavelength - 3 meter
  - $\gg r_s$
  - $\gg l_0$

Common to other radars.

- a - IS facility
- b - Coherent radar (ionospheric)
- c - MST radar

Outline

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	1	2	3	4	
a	✓	✓	✓		
b		✓	✓		
c	✓		✓		

## Incoherent Scatter radar

(3)(1) - High Altitude  $N_e$  profiles

(2) - Ion-gyro effects

- Plasma theory
- Mass spectrometer
- Drifts. ( $0.1 \text{ mV/m}$ )
- $\vec{B}$

(2)(3) - Faraday rotation exp.

(2) - Equatorial Anomaly studies.

## Cohesent radar - Ionospheric

- Spread-F
- Electrojet
- 150 km "junk"

## MST radar

- Only real MST
- Examples : Stratospheric  
Mesospheric -

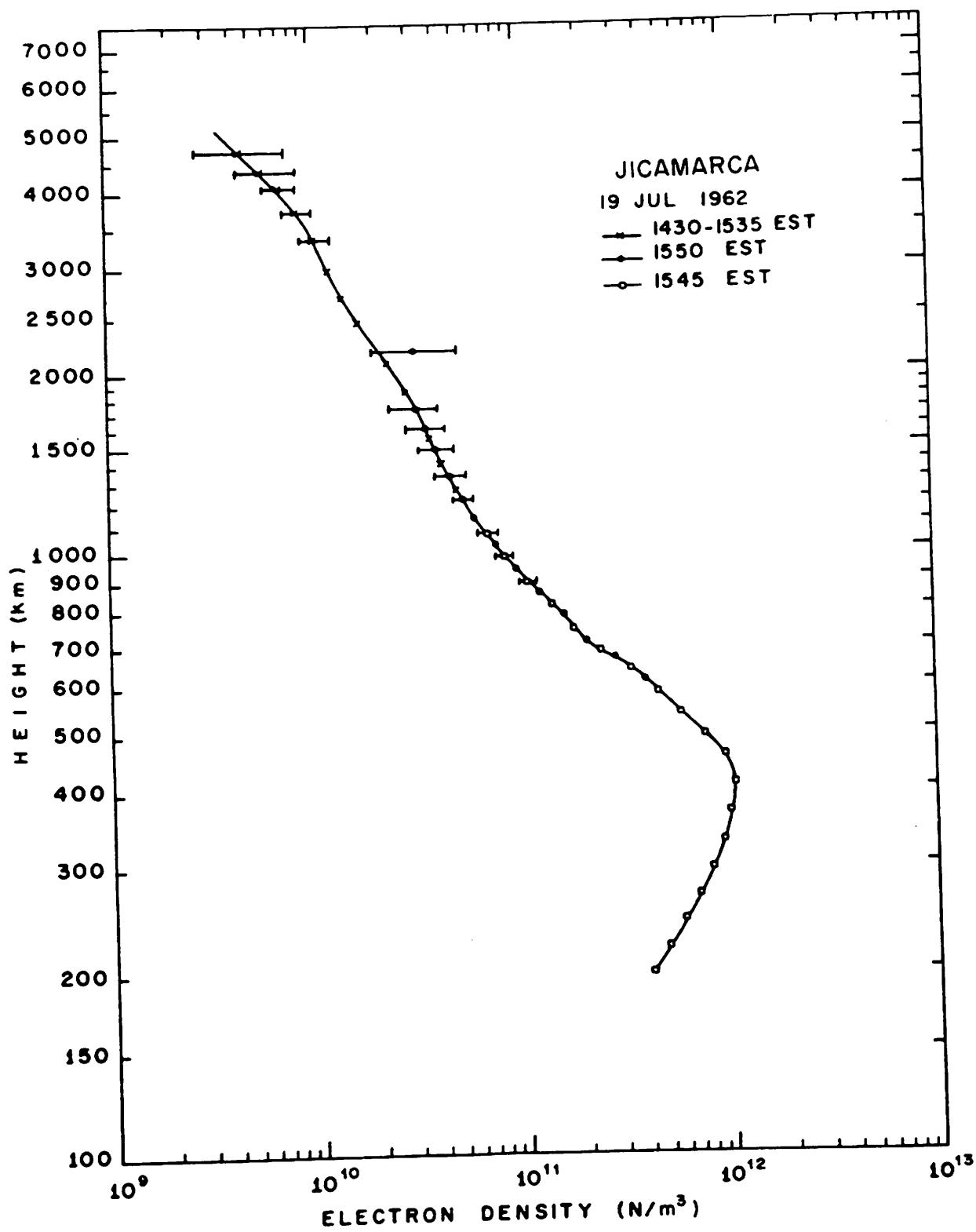


Figure 12

C31-542A

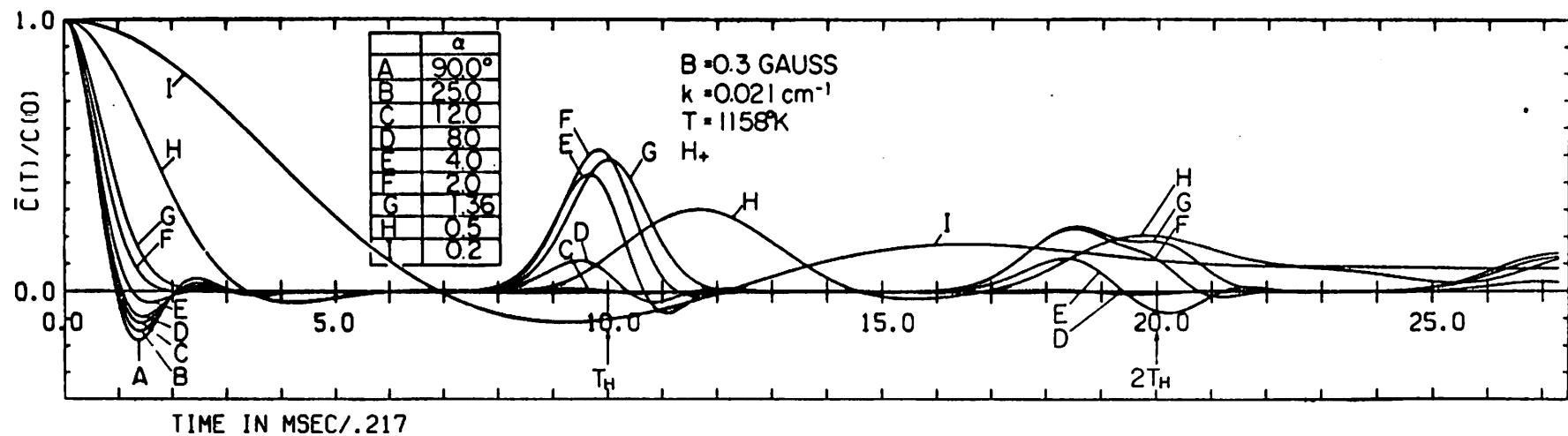


FIG. 7a. Autocorrelation function envelope  $\bar{C}(\tau)^+$  for a  $[H]^+$  plasma showing the effect of variations in the direction of the wave vector  $\vec{k}$  as it approaches the perpendicular to magnetic field ( $\alpha = 0$ ). Collisionless model, other parameters as shown in the figure.

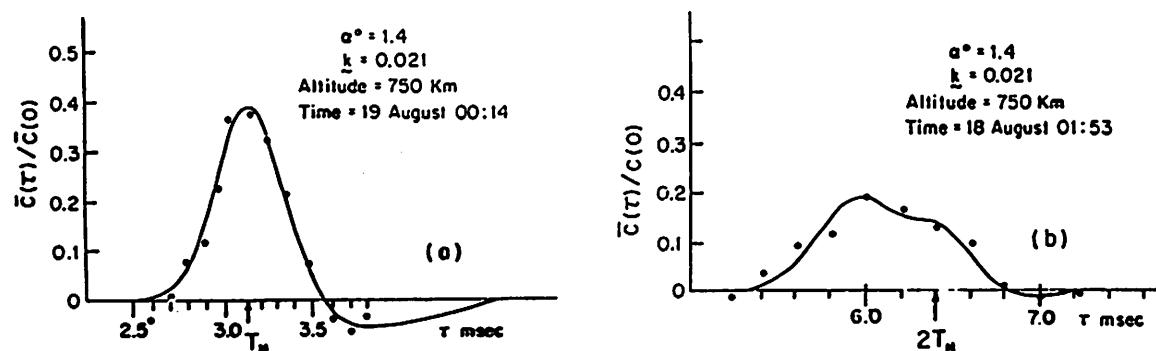
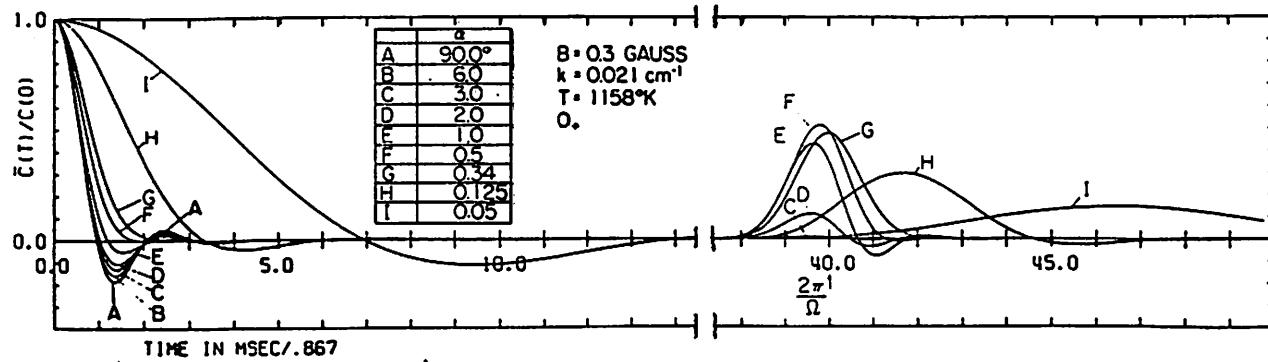


Figure 2. Comparison between measured and experimentally determined values of the correlation function close to the first and second hydrogen gyro-peak. Temperature and composition was adjusted for best fit. (After Woodman, 1967, with experimental points from Farley, 1967).

## GEOMAGNETIC FIELD BY INCOHERENT SCATTER

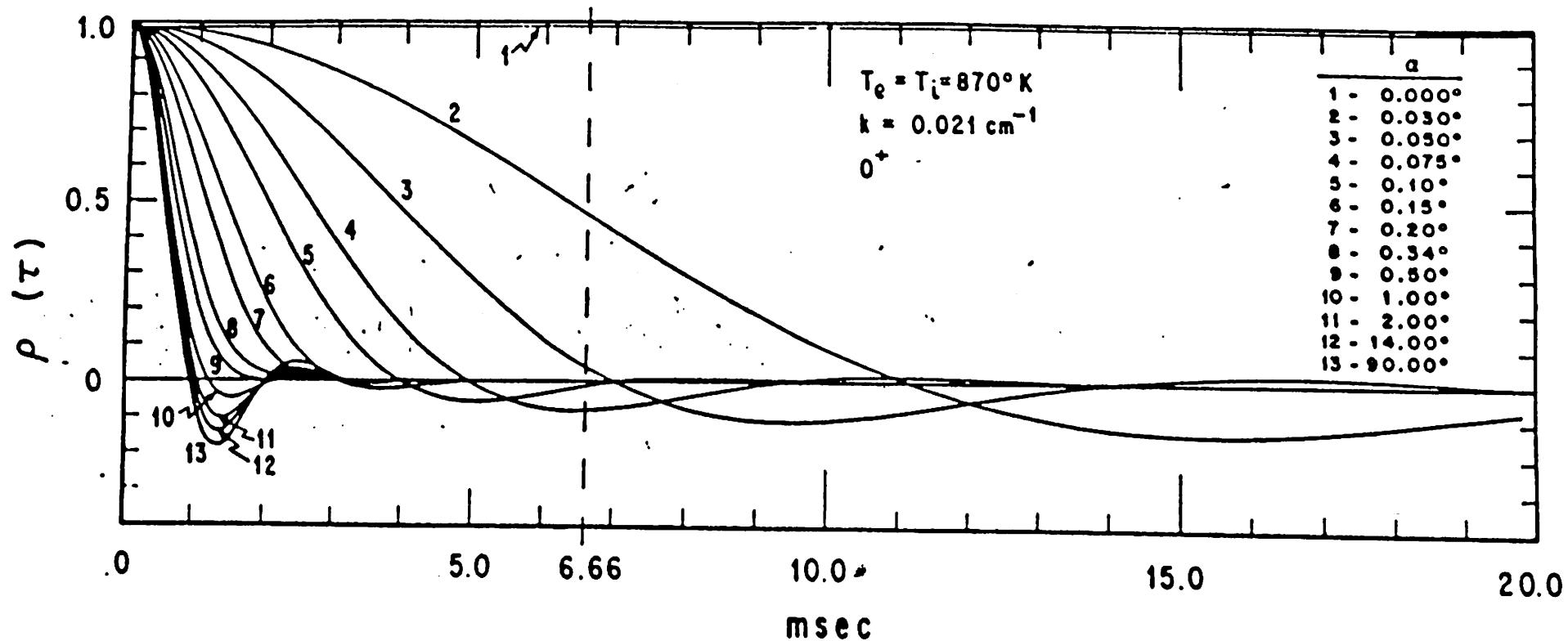
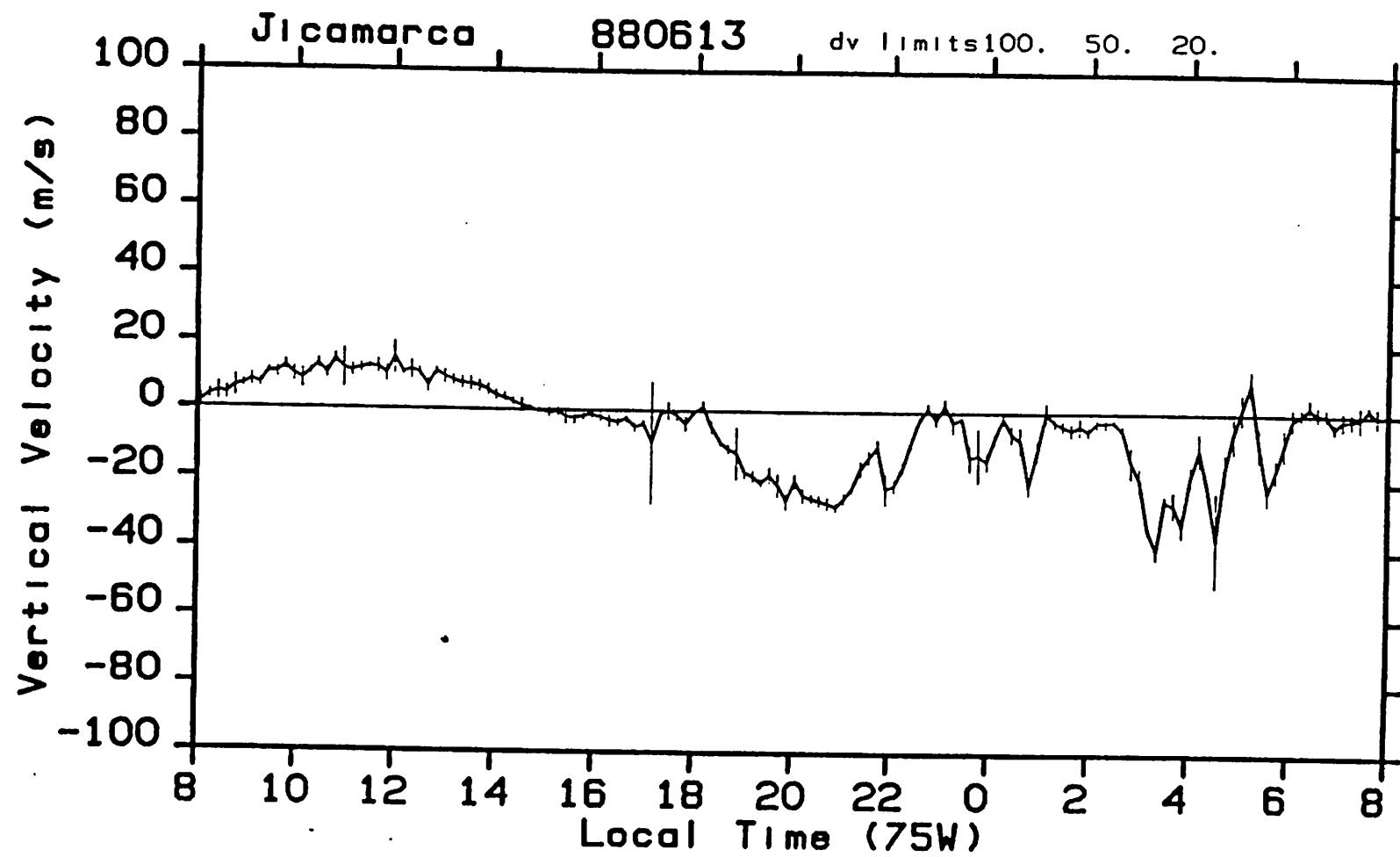


Fig. 1. Theoretical autocorrelation functions  $\rho(\alpha, \tau)$  for different angles  $\alpha$ , where  $\alpha$  is the complement of the angle between  $\mathbf{k}$  and  $\mathbf{B}$ .



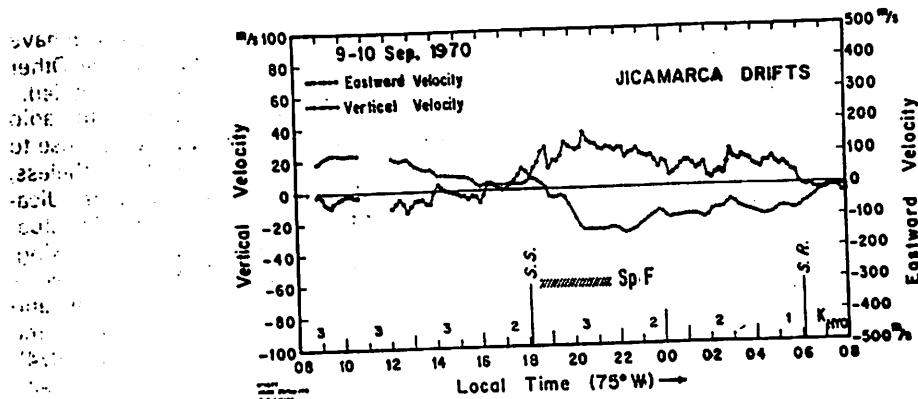


Figure 3a

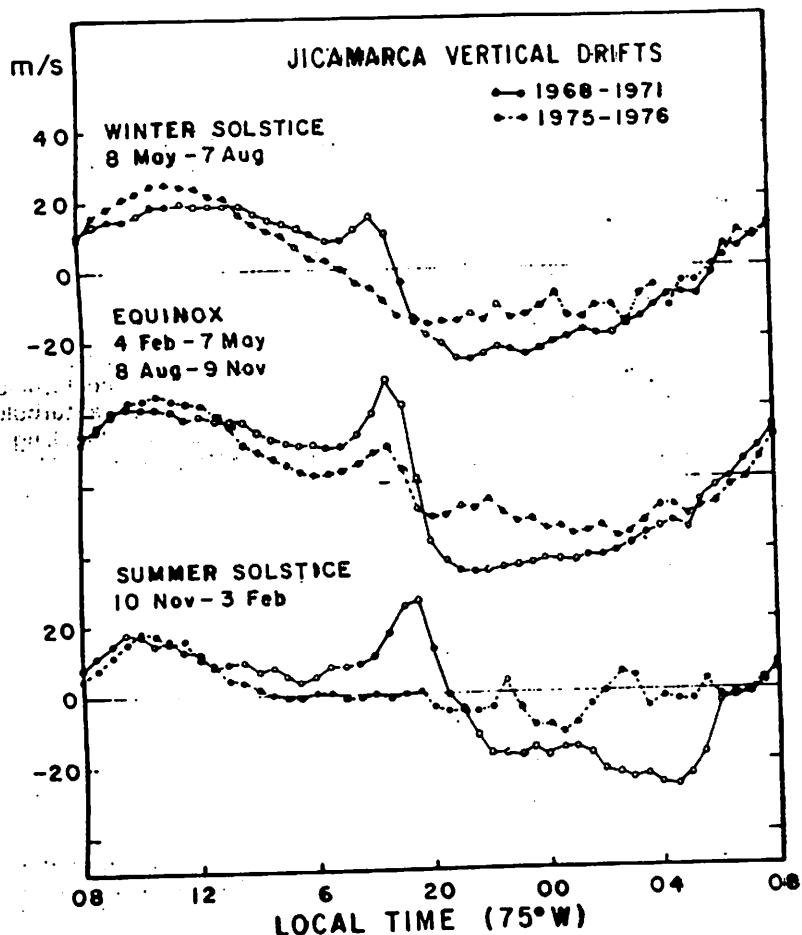


Figure 3b

Figure 3. Example of the diurnal variations of the vertical and E-W drift velocities. 3b) Vertical drift velocities for different seasons and solar activity.

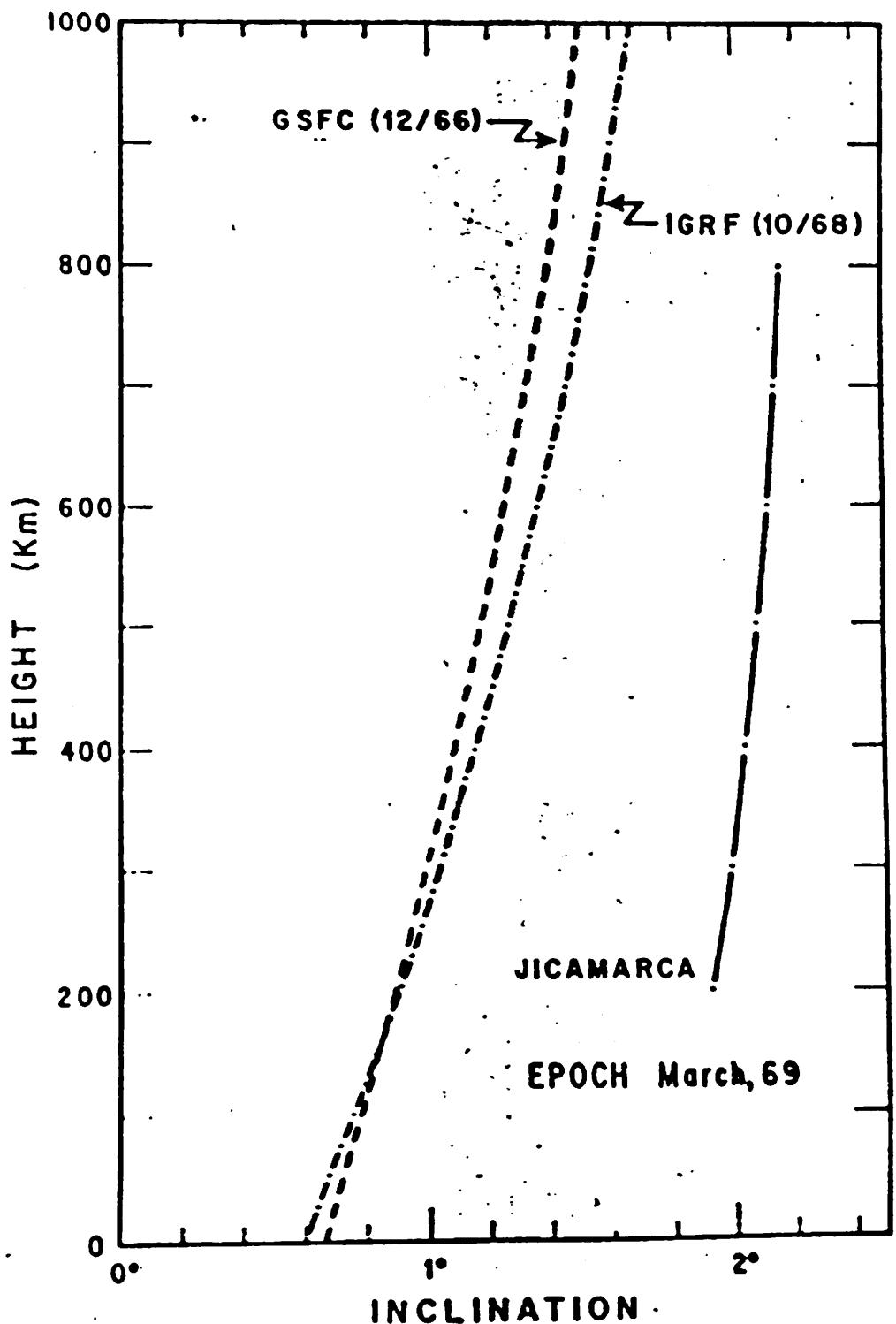
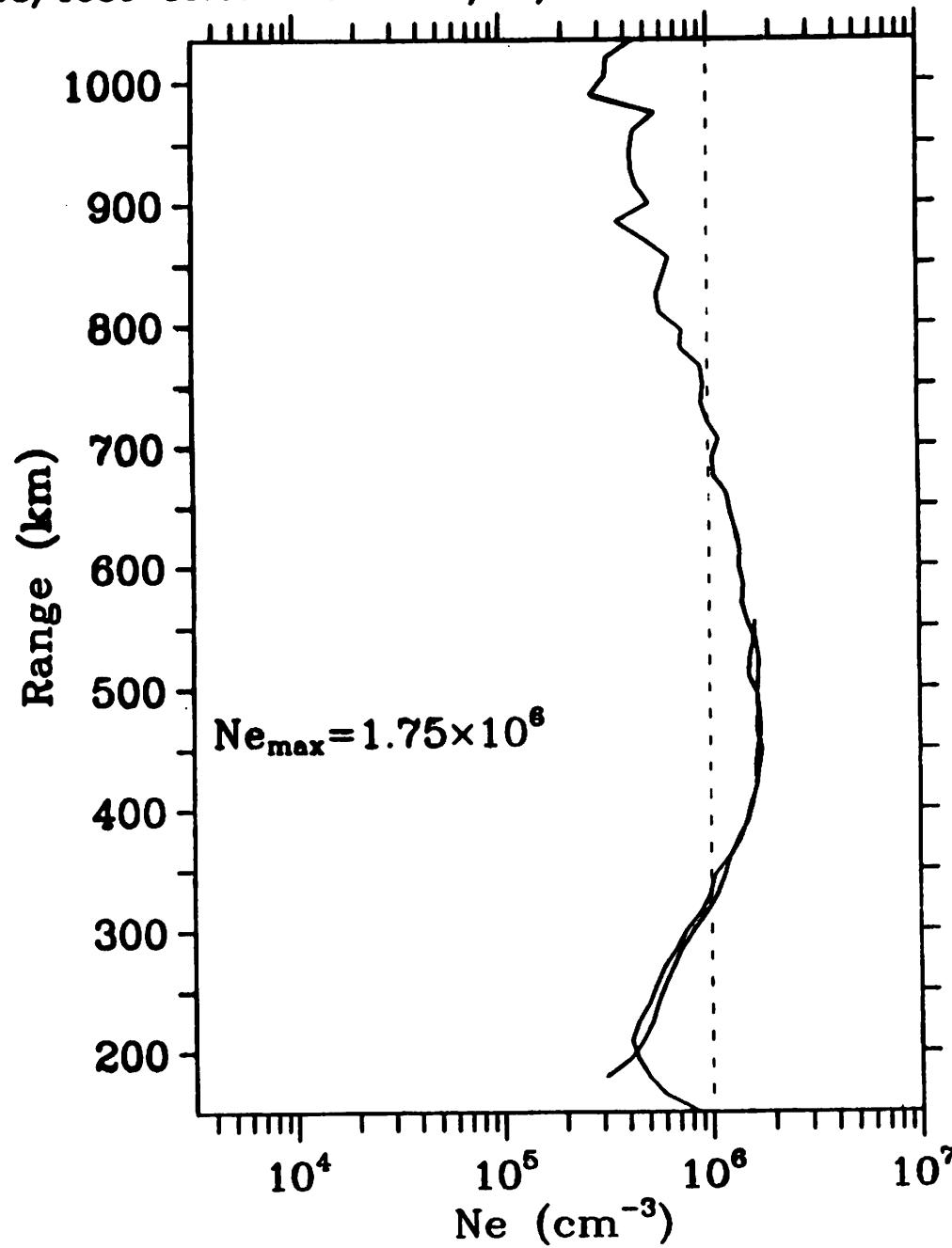
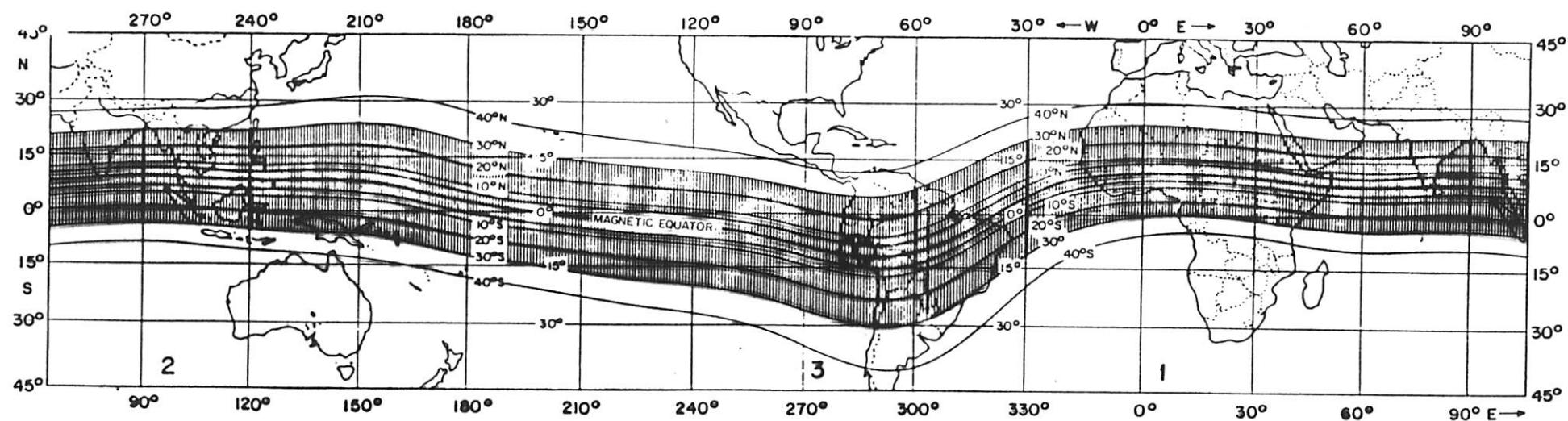
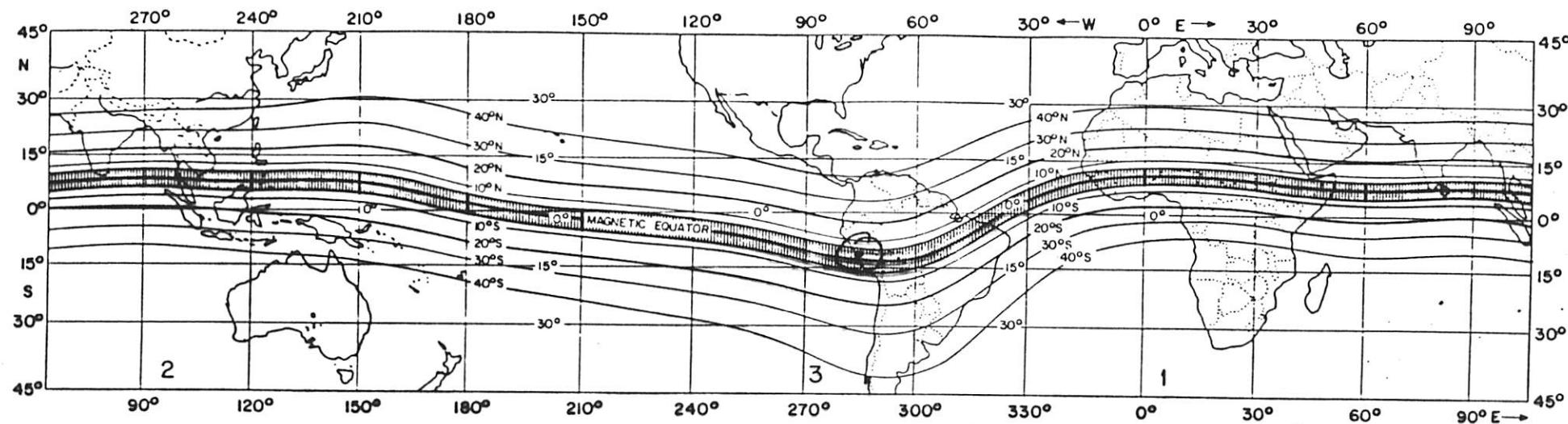


Fig. 6. Inclination of the magnetic field along Jicamarca vertical ( $-11.95^{\circ}$  latitude  $76^{\circ}52'20''$  longitude). The interrupted solid line corresponds to values determined experimentally at Jicamarca; the other two correspond to two of the latest earth magnetic field models, GSFC 12/66 and IGRF 10/68.

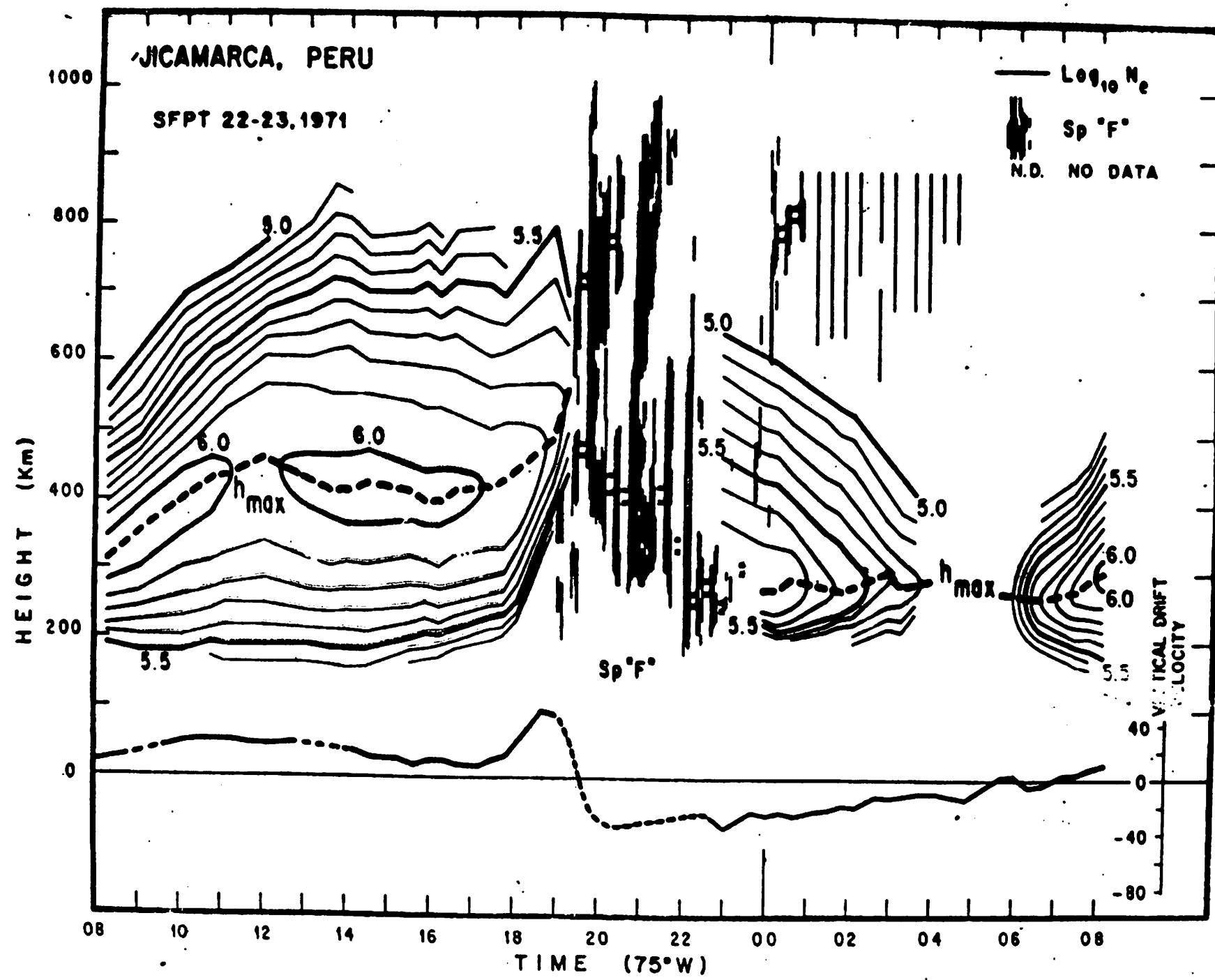
2/03/1989 15:25:08.4 To 2/03/1989 15:29:36.5 Nint= 1280 Tx= 15





JICAMARCA, 6 OCTOBER 1984



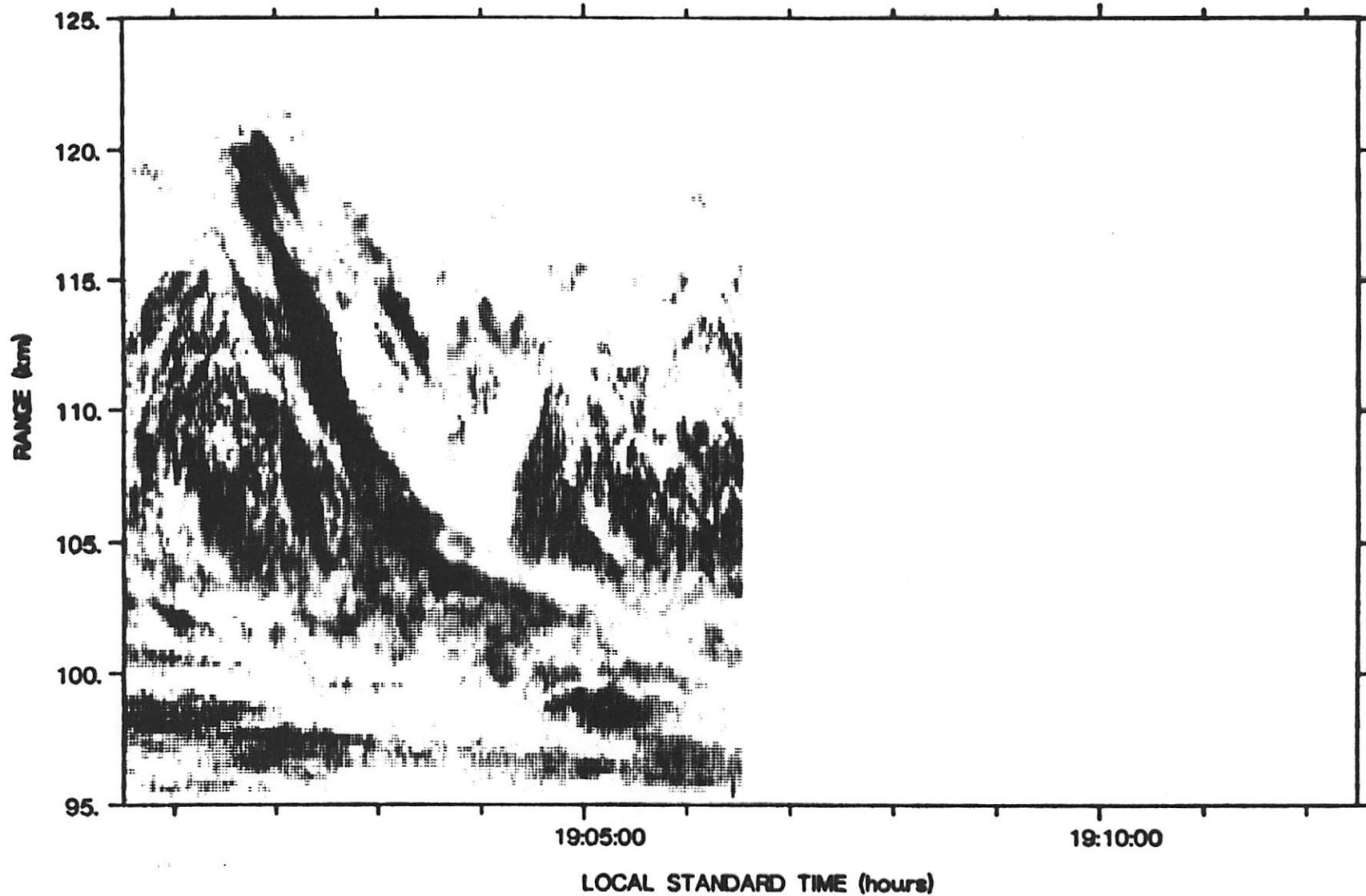


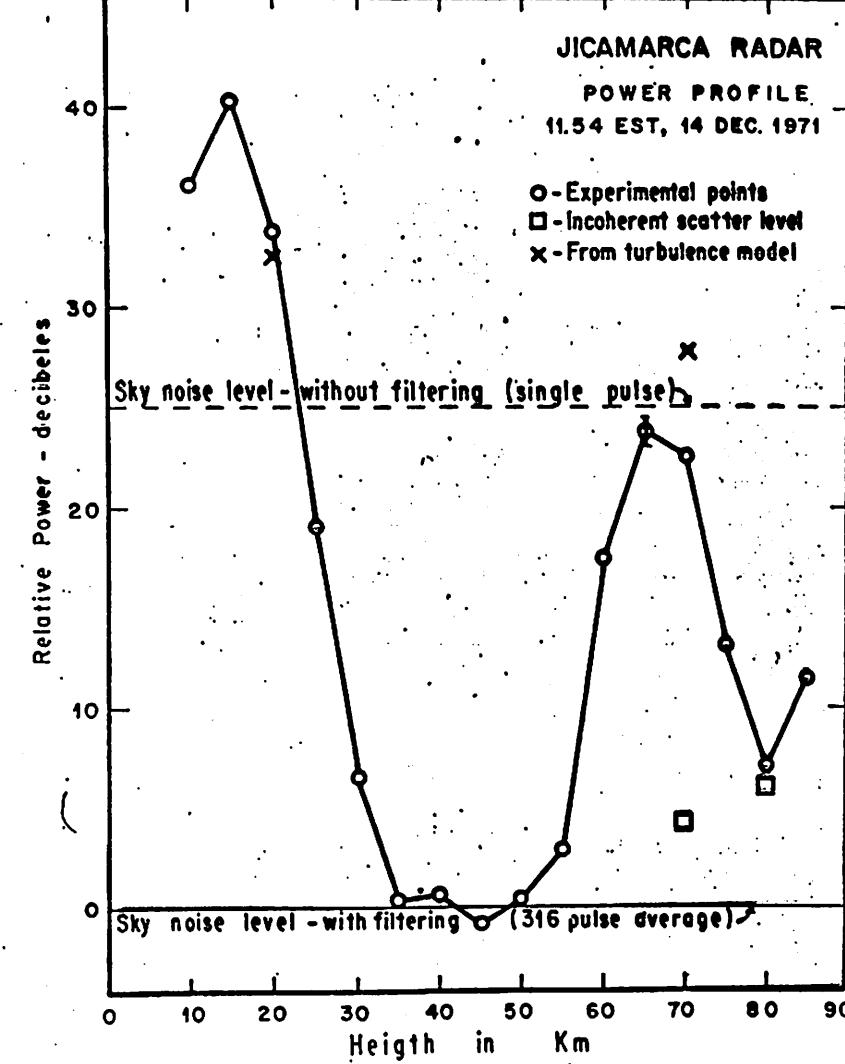
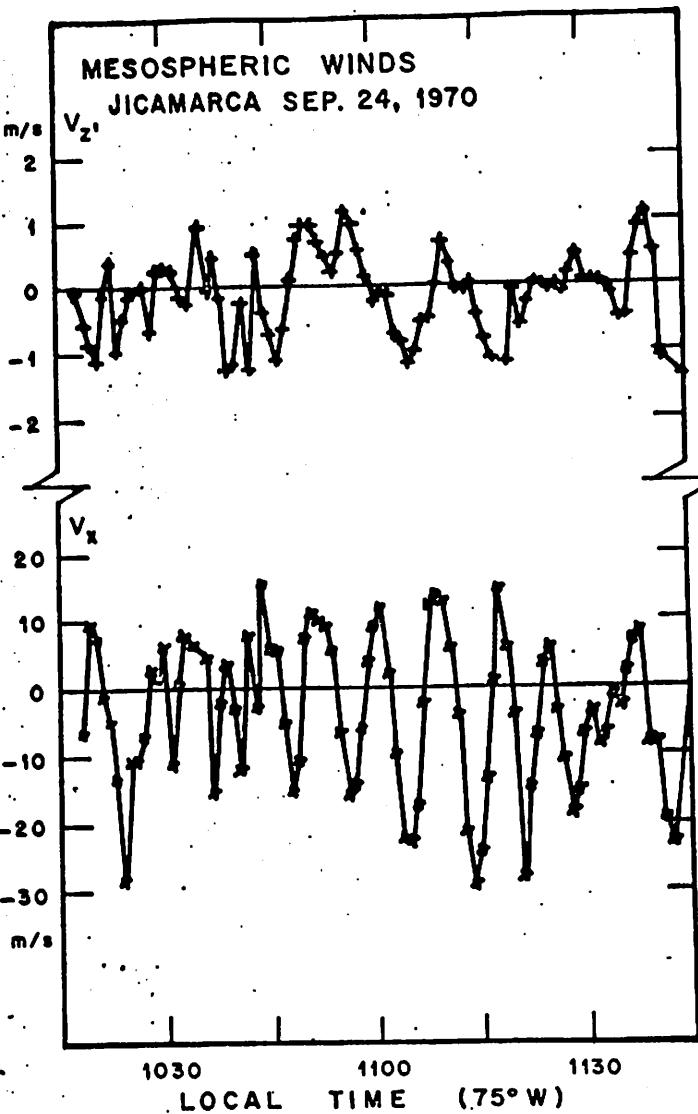
Convection with (?)  
with a "linearity"  
break?  
(convection maxima below now)

68

JICAMARCA, PERU

90/01/15 S/N





HEIGHT

40

35

30

25

20

15

11  
12  
13  
14  
15  
16  
17  
DATE: 7 AUG 89

12

13

14

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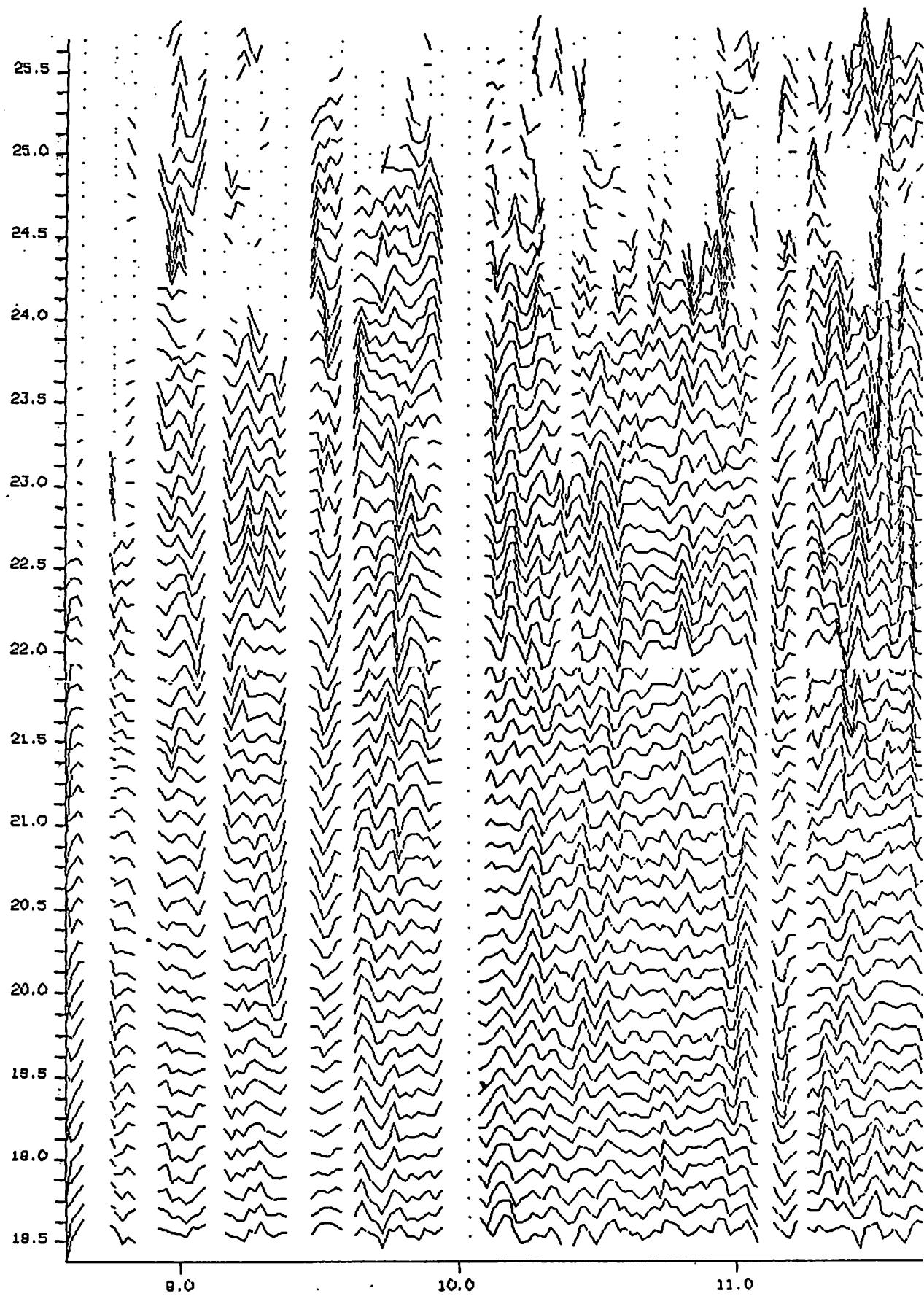
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17

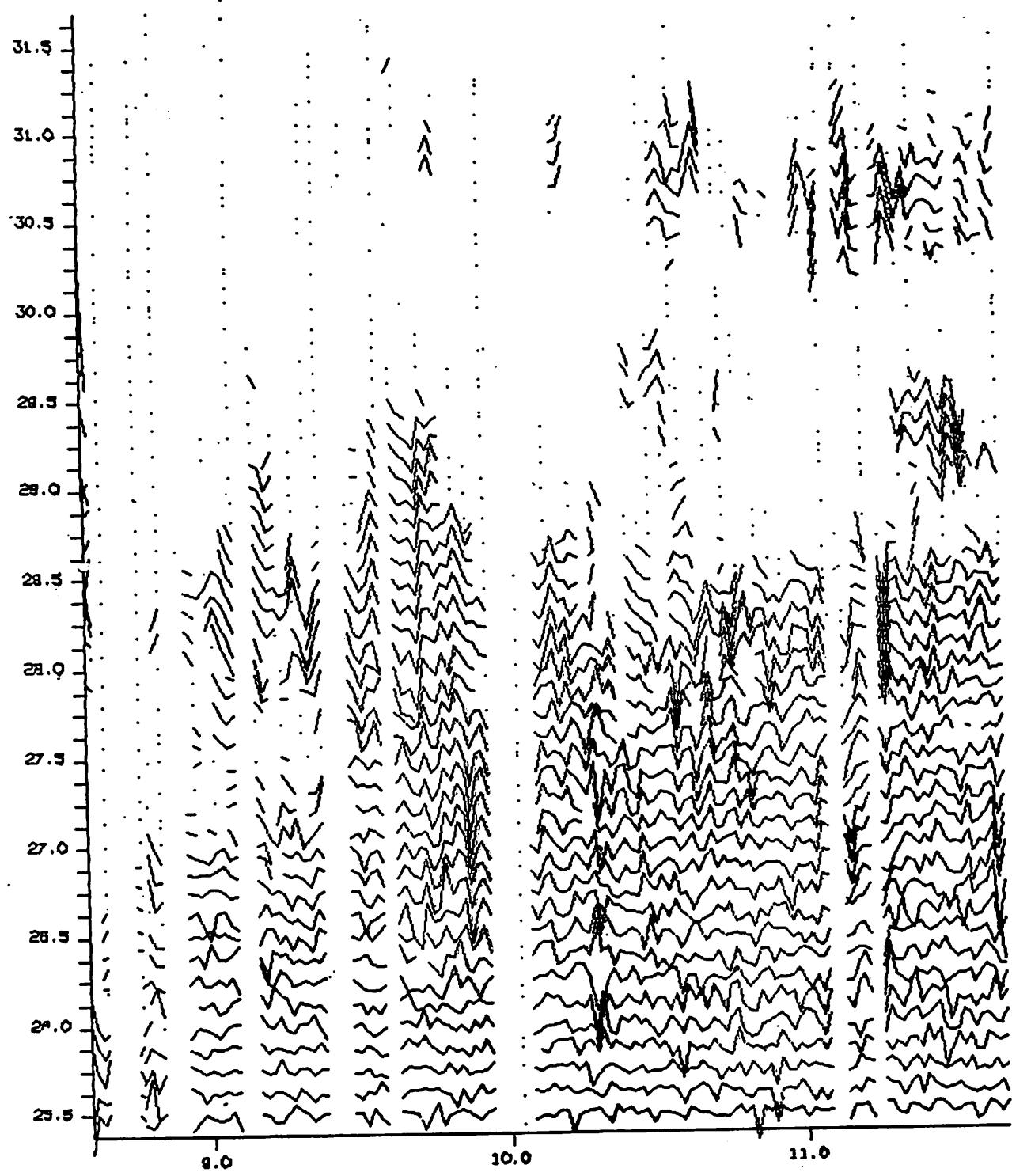
LOCAL TIME (HORAS)

JICAMARCA - POWER MAP  
NORMALIZACION: C

ANTENA: -11.95 DEG., 0m 0s



29 SEP 88 - Rx "B"



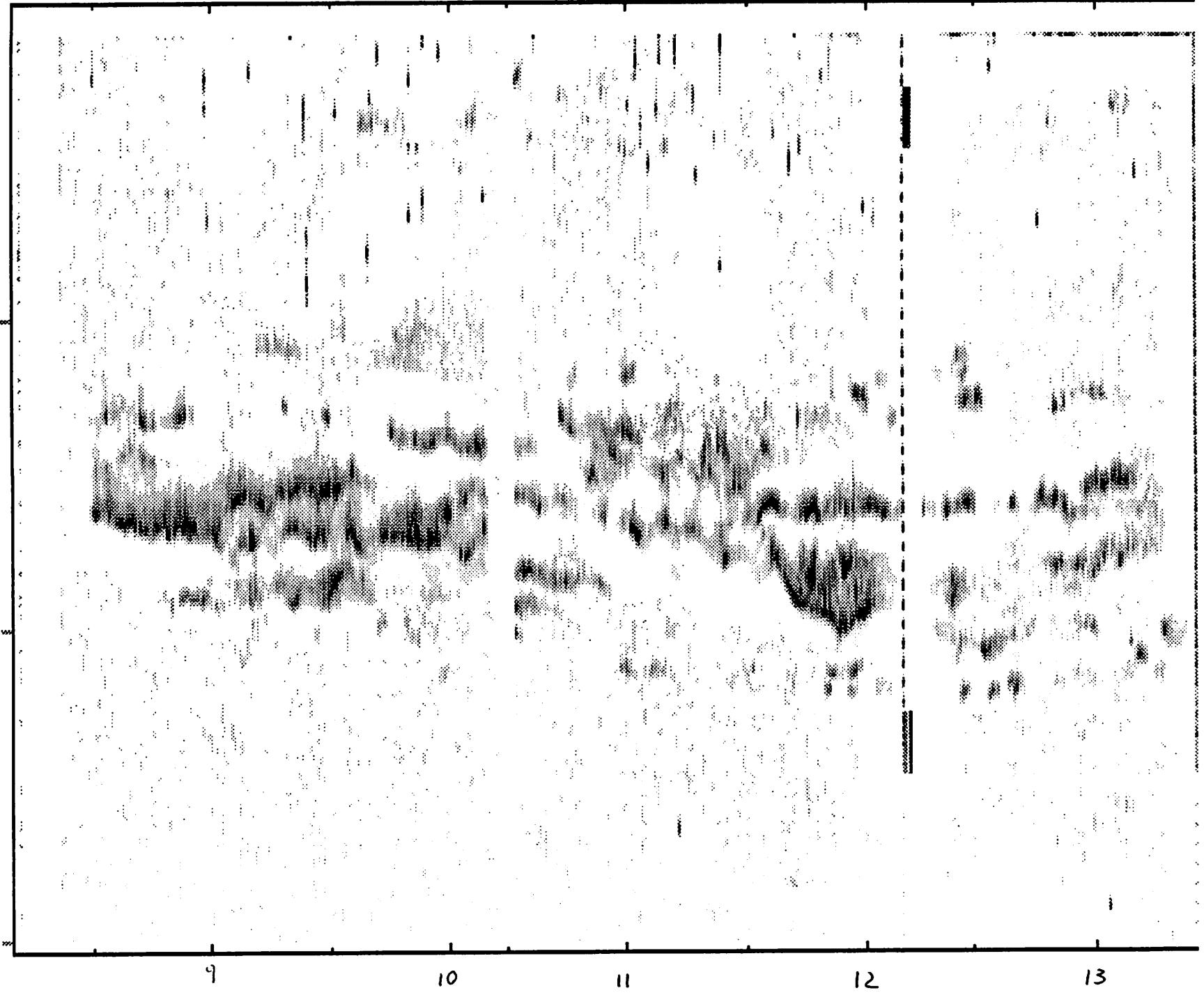
29 SEP 88 - Rx "A"

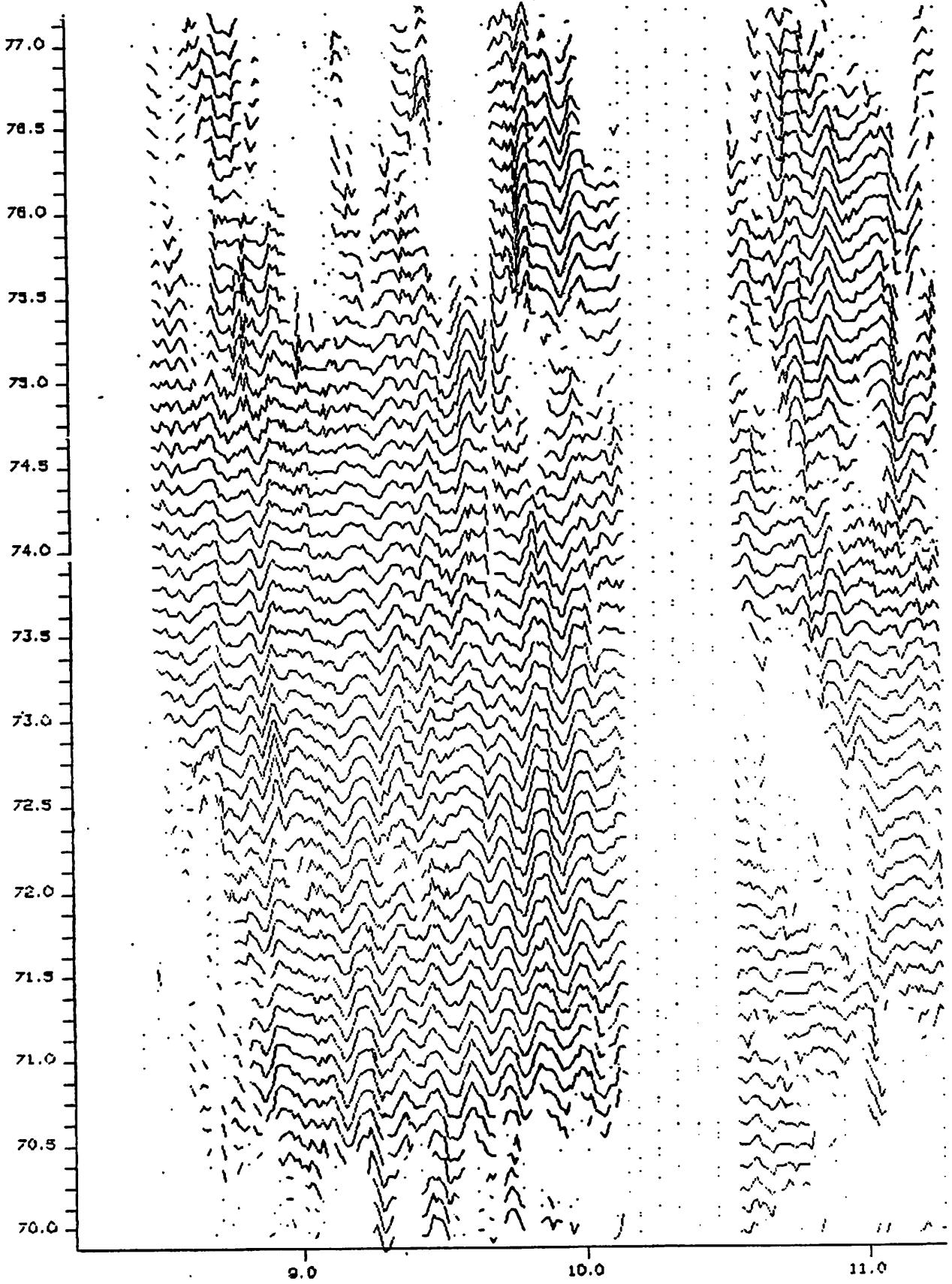
11 Sept 89 AUTO - 12.56 WEST side.

80

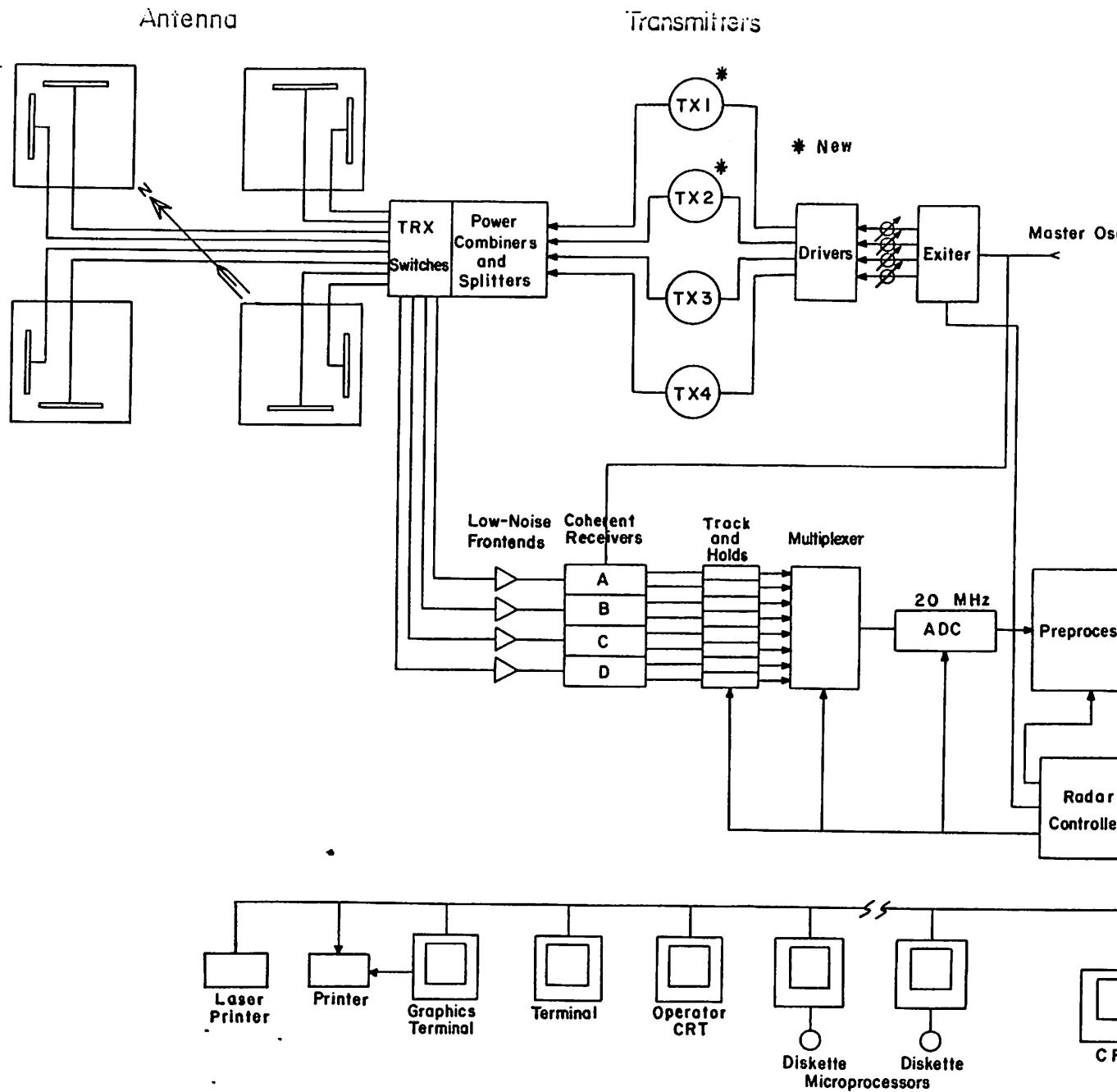
70

60





11 SEP 89 - Rx "A"



Schematic Diagram of the Jicamarca Radar