

WACCM studies of the upper atmosphere

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Outline

- WACCM-3 description
- A summary of current research:
 - ▶ Tides in WACCM
 - ▶ WACCM-X
 - ▶ PMC studies
 - ▶ Sodium Chemistry
- How to obtain WACCM
- Summary

UCAR Quarterly

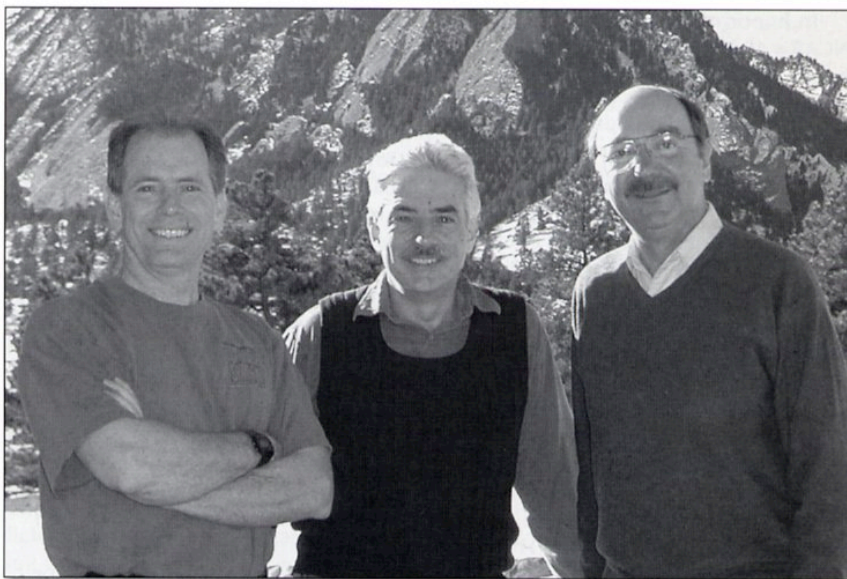
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It's about TIME: A new upper-atmosphere model

by Carol Rasmussen

Ray Roble, a senior scientist in the High Altitude Observatory, has spent the last 20 years or so creating and refining a general circulation model of the upper atmosphere—the transition zone between the bulk of the atmosphere below 30 kilometers and the void above 500 km. Now, the knowledge Roble has amassed in developing his thermosphere-ionosphere-mesosphere electrodynamics general circulation model (TIME-GCM) is being incorporated into the middle atmosphere community climate model, an offspring of NCAR's CCM family that was created by Byron Boville (Climate and Global Dynamics Division). A third co-principal investigator, Rolando Garcia, and colleagues in NCAR's Atmospheric Chemistry Division (ACD) will be adding chemistry components to the model.

"A year from now, we expect to have a model that extends from the surface to 120 kilometers and includes ozone chemistry," says Boville. The yet-unnamed composite model will be one of only two or three general circulation models to reach that altitude. Some GCMs only extend to 30 km, including the



Byron Boville, Rolando Garcia, and Ray Roble. (Photo by Carlye Calvin.)

troposphere and part of the lower stratosphere; a few have been extended into the mesosphere, with upper boundaries in the range of 60–80 km. The new model will include the entire mesosphere as well as the lower thermosphere. But the plan doesn't stop there. Further into the future, the scientists will add thermos-

electrically charged atoms and molecules of the ionosphere are ionized by solar radiation and aurora particle precipitation. This happens throughout the atmosphere, but in the denser air lower down, charged particles quickly bump into oppositely charged particles and recombine. In the sparsely populated ionosphere, ionized parti-

WACCM Project Goals

Develop a state of the art, “high-top”, numerical model with coupled chemistry that will serve the needs of the scientific community

Conduct research to understand the coupling between atmospheric layers, the role of chemical and physical processes in defining these couplings, and the interaction between the Earth’s atmosphere and the Sun

Acknowledgement

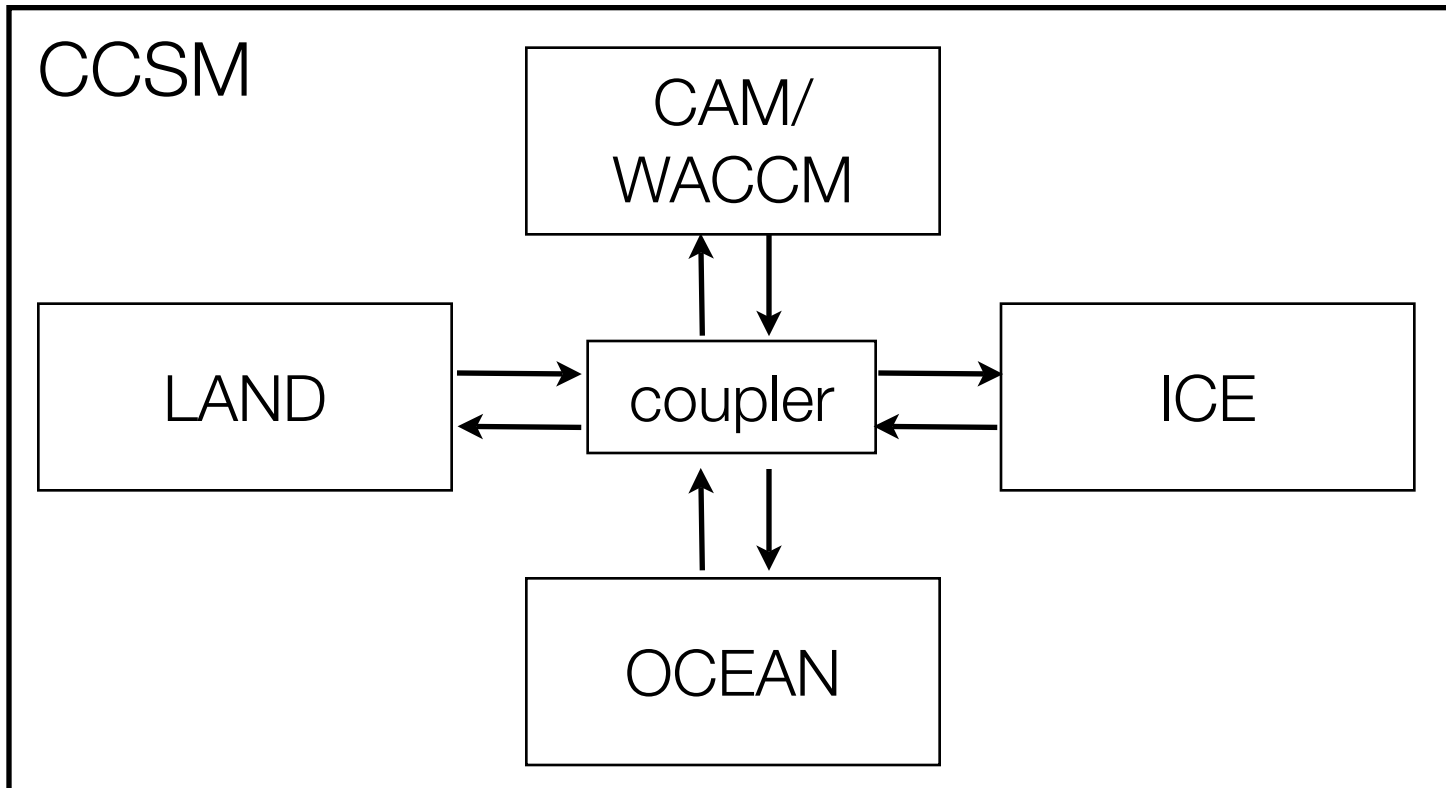
WACCM Development Contributors:

Charles Bardeen, Byron Boville, Chris Fischer, Benjamin T. Foster, Rolando R. Garcia, Andrew Gettelman, Maura E. Hagan, Charles Jackman, Douglas E. Kinnison, Han-Li Liu, Joseph M. McInerney, Daniel R. Marsh, Astrid Maute, Aimee Merkel, Mike Mills, Liying Qian, Cora Randall, Arthur D. Richmond, Jadwiga H. Richter, Raymond G. Roble, Fabrizio Sassi, Anne K. Smith, Stanley C. Solomon, Francis Vitt

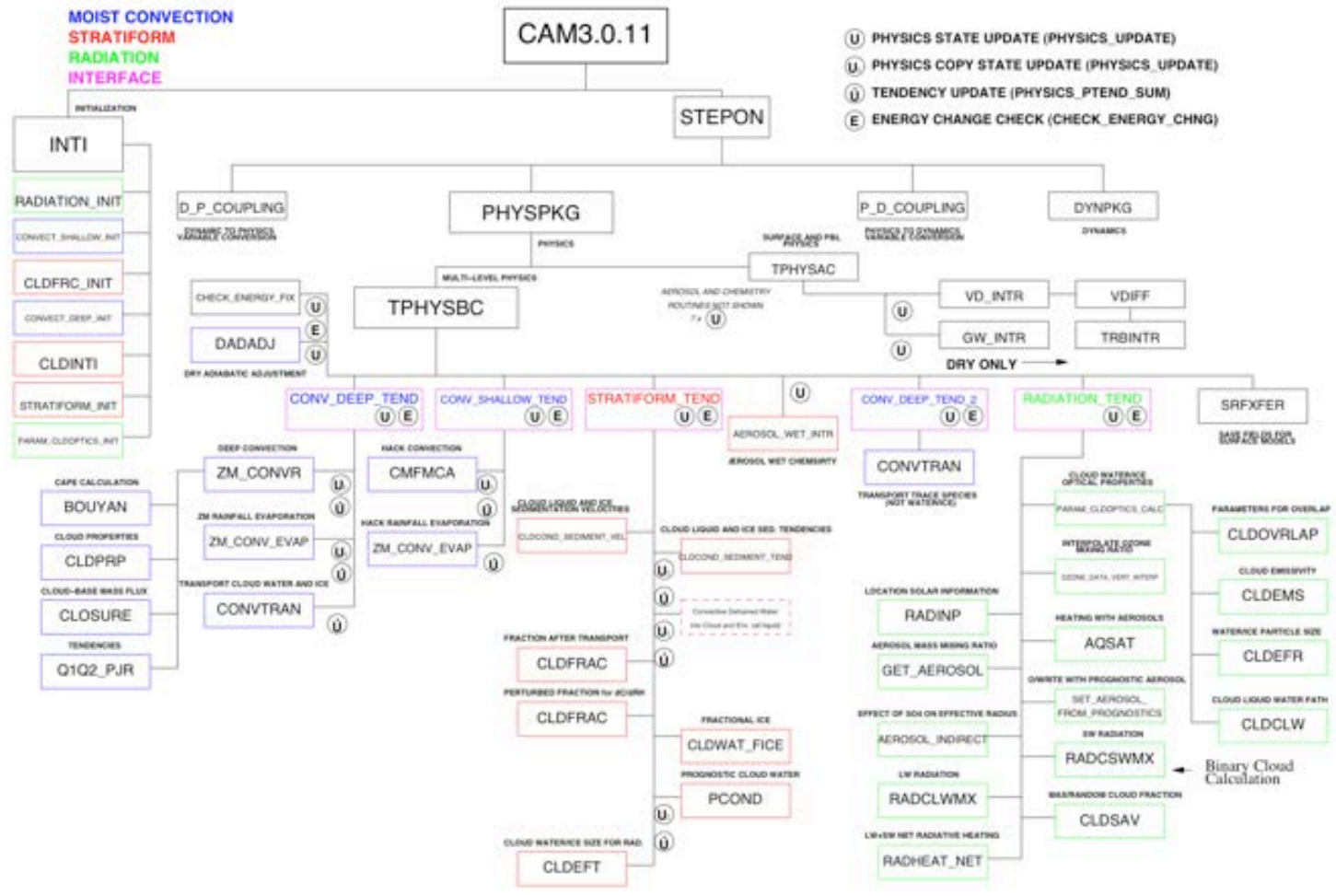


WACCM description

- WACCM is an extension of the Community Atmosphere Model (CAM) - the atmospheric component of the Community Climate System Model



Community Atmosphere Model Flowchart



WACCM-3.5

Model Framework	Dynamics	Tracer Advection	Resolution	Chemistry	Other Processes
Extension of the NCAR Community Atmosphere Model version 3 (CAM3) Based upon CAM3.5.48	Finite Volume Dynamical Core (Lin, 2004) Fully-interactive with chemistry, i.e., consistent with model-derived, radiatively active gases: O3, CO2, CH4, N2O, H2O, CFC11, CFC12, O2, NO	Flux-form Finite Volume (Lin, 2004)	Horizontal: 1.9° x 2.5° (lat x lon) Vertical: 66 levels 0-140km < 1.0km in UTLS 1-2 km in stratosphere ~3 km in MLT	Middle Atmosphere Mechanism: 57 Species including Ox, HOx, NOx, BrOx, and ClOx Heterogeneous chemistry E-region Ion Chemistry	GW Param.: convection-, frontal-, and orographically-generated Molecular Diffusion: Banks and Kockarts, 1973 Auroral processes, including ion drag, and Joule heating LW/SW and chemical potential heating

WACCM-3 Chemical Constituents in MA Mechanism

Long-lived Species: (17-species)

Misc: CO_2 , CO , CH_4 , H_2O , N_2O , H_2O_2

CFCs: CCl_4 , CFC-11, CFC-12, CFC-113

HCFCs: HCFC-22

Chlorocarbons: CH_3Cl , CH_3CCl_3 ,

Bromocarbons: CH_3Br

Halons: H-1211, H-1301

Constant Species: N_2

Short-lived Species: (34 neutrals + 5 ions + electrons)

O_x : O_3 , O , $\text{O}(^1\text{D})$

NO_x : N , NO , NO_2 , NO_3 , N_2O_5 , HNO_3 , HO_2NO_2

ClO_x : Cl , ClO , Cl_2O_2 , OCIO , HOCl , HCl , ClONO_2 , Cl_2

BrO_x : Br , BrO , HOBr , HBr , BrCl , BrONO_2

HO_x : H , OH , HO_2 , H_2O

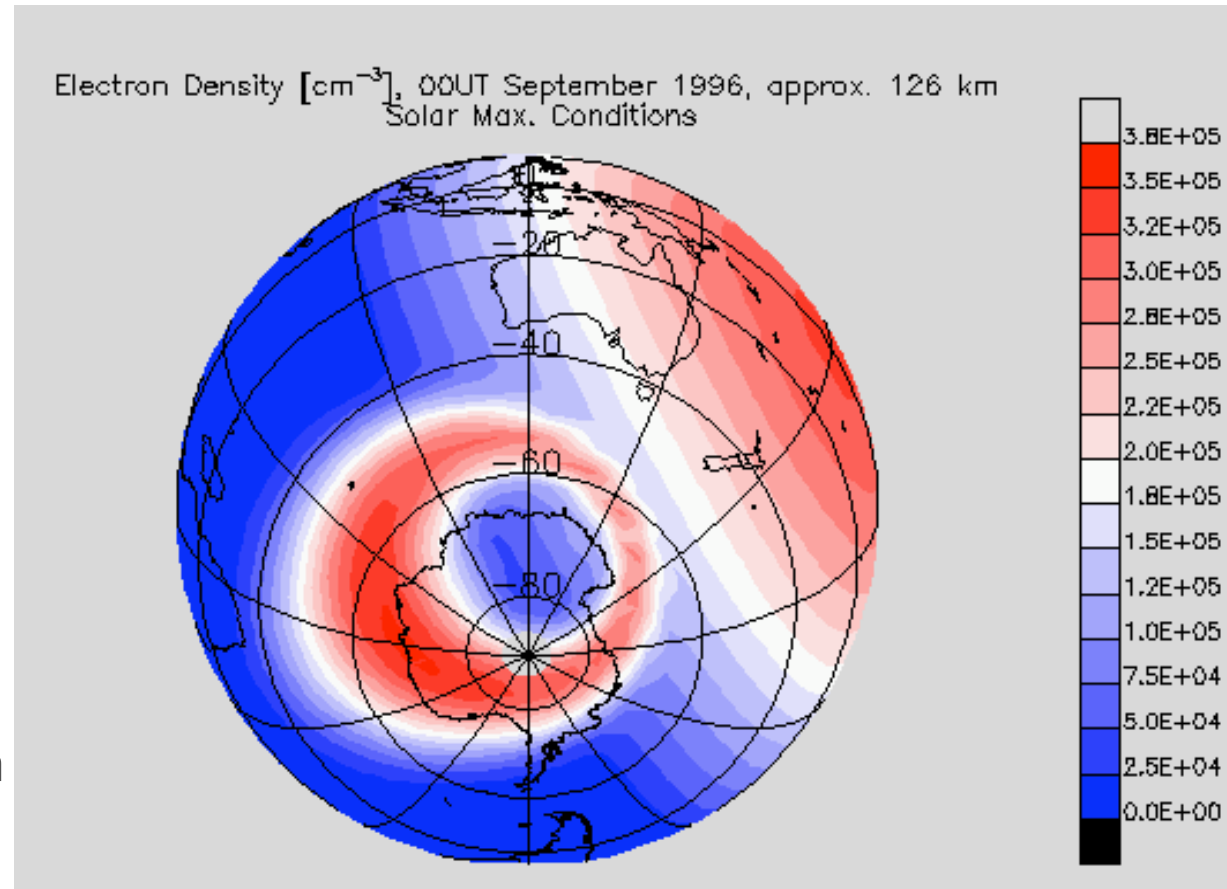
HC Species: CH_2O , CH_3O_2 , CH_3OOH

Ions: N^+ , N_2^+ , NO^+ , O^+ , O_2^+

Other: $\text{O}_2(^1\Sigma)$, $\text{O}_2(^1\Delta)$, $\text{N}(^2\text{D})$, e

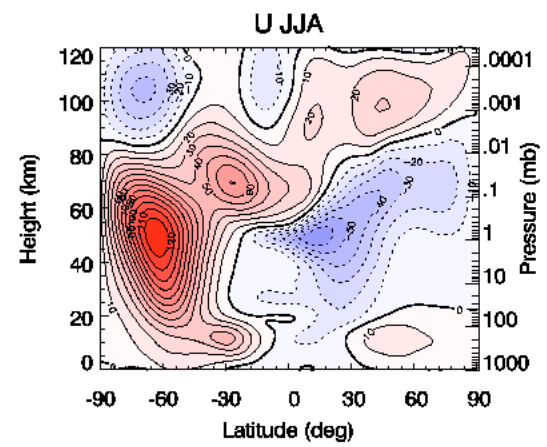
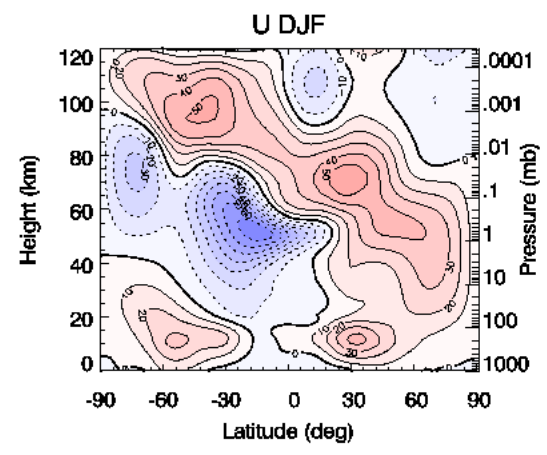
WACCM-3

- WACCM modifications for the upper atmosphere:
 - ▶ Variability of solar spectrum
 - ▶ Chemical heating
 - ▶ Airglow
 - ▶ E-region ion chemistry (5 ions & electrons)
 - ▶ EUV and x-ray ionization
 - ▶ Auroral processes
 - Particle precipitation
 - Ion drag
 - Joule heating

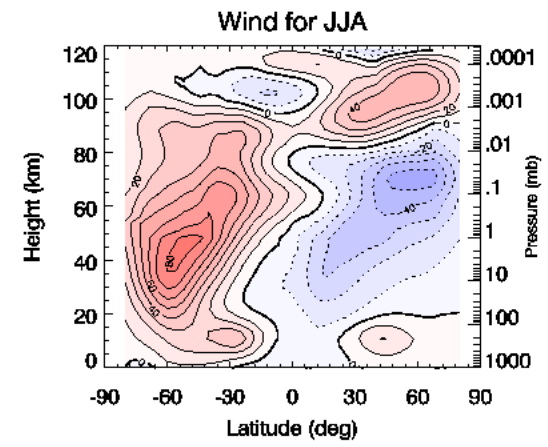
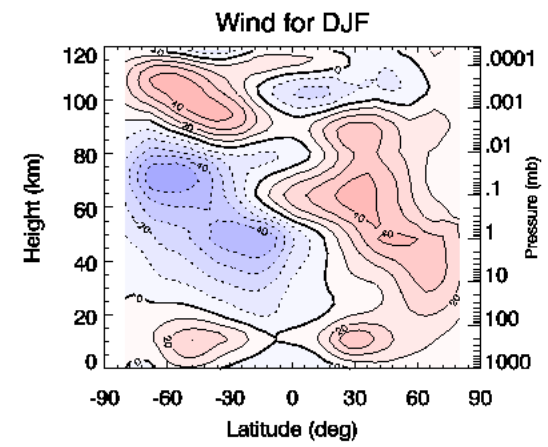


Zonal winds 0-120 km

WACCM3.6



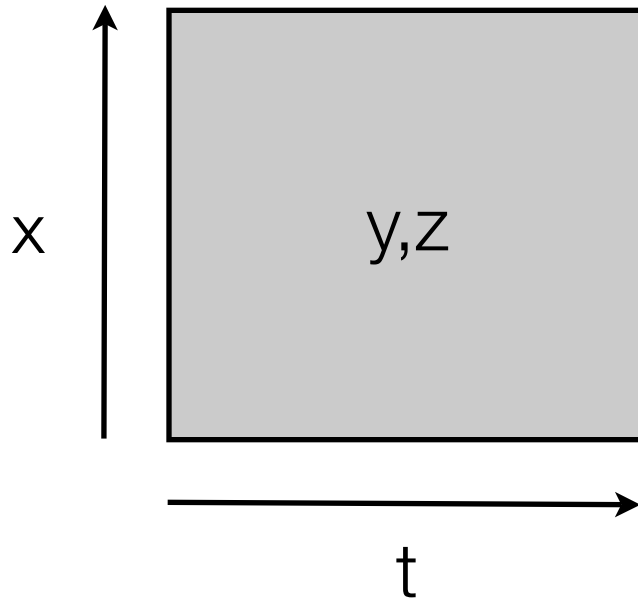
URAP



Atmospheric Tides in WACCM

with R.R. Garcia and A.K. Smith

Monthly mean tidal diagnostics



Applied to u , v , T , and PS . Calculated and output once per month

Amplitude and phase calculated at each height (z) and latitude (y)

Post-processing:
 migrating and non-migrating modes can be separated by FFT expansion in longitude of the tidal coefficients

Example for the diurnal tide:

The diurnal tide for each month of output can be represented with two coefficients, C_{24} and S_{24} , at each gridpoint (x , y , z):

$$T_d = C_{24} \cdot \cos\left(\frac{2\pi t}{24}\right) + S_{24} \cdot \sin\left(\frac{2\pi t}{24}\right)$$

t = universal time (hr)

During the model run, calculate C_{24} and S_{24} by accumulating the following sums at each time step:

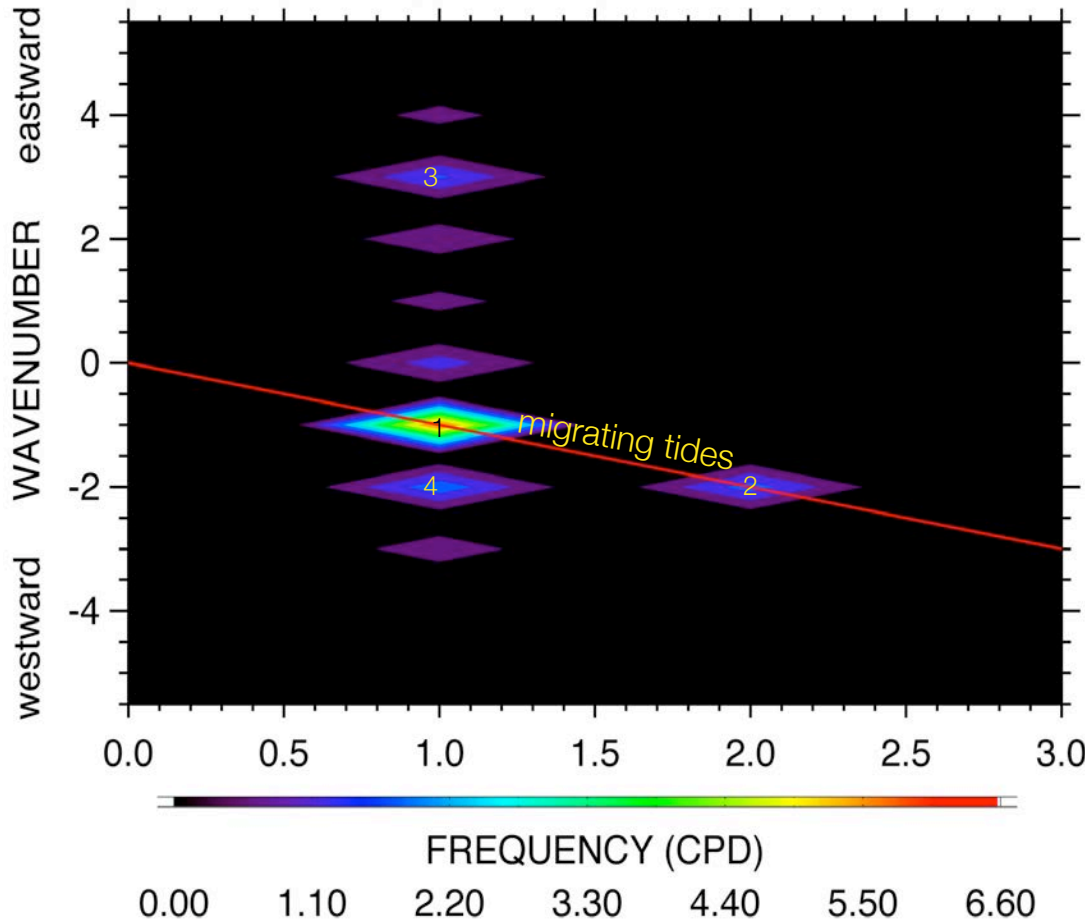
$$C_{24} = \frac{2}{N} \sum_1^N T(t_n) \cos\left(\frac{2\pi t_n}{24}\right)$$

$$S_{24} = \frac{2}{N} \sum_1^N T(t_n) \sin\left(\frac{2\pi t_n}{24}\right)$$



Equatorial spectrum at 12 SH (~85 km)

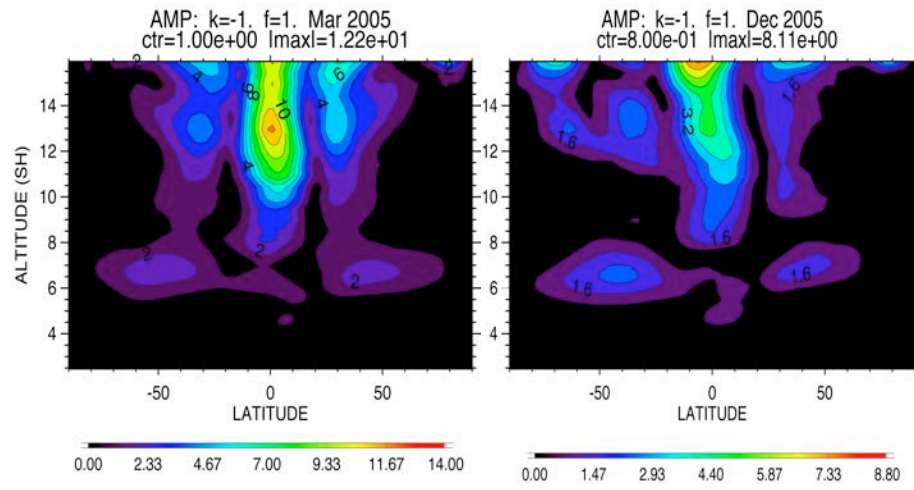
annual composite lat = 1. lev = 12.
ctr = 0.60 lmaxl = 6.6



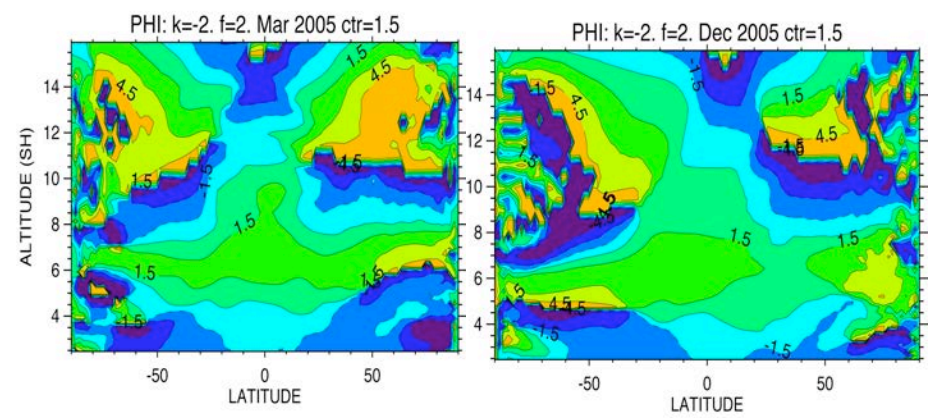
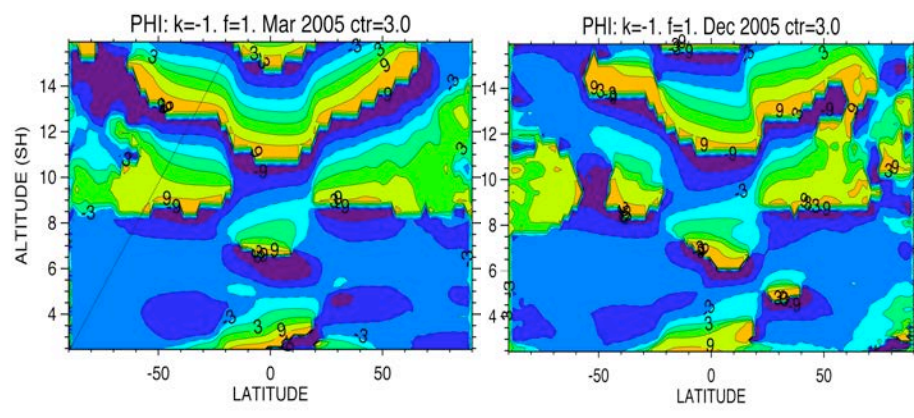
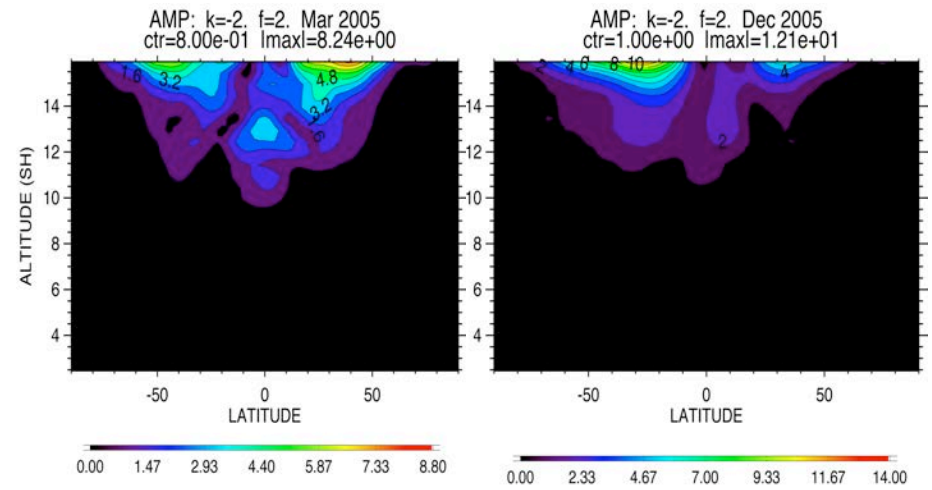
- 1: m=1 migrating diurnal
- 2. m=2 migrating SD
- 3. m=3 diurnal eastward
- 4. m=1 SD westward

Structure of modeled migrating temperature tides

T diurnal westward 1
March December



T semidiurnal westward 2
March December



max at lat/height 0.95, 12.43 = 11.29

max at lat/height 0.95, 12.43 = 4.58

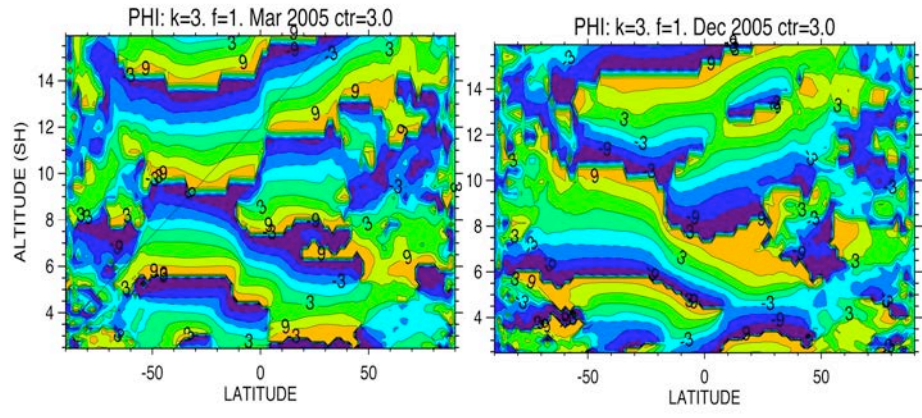
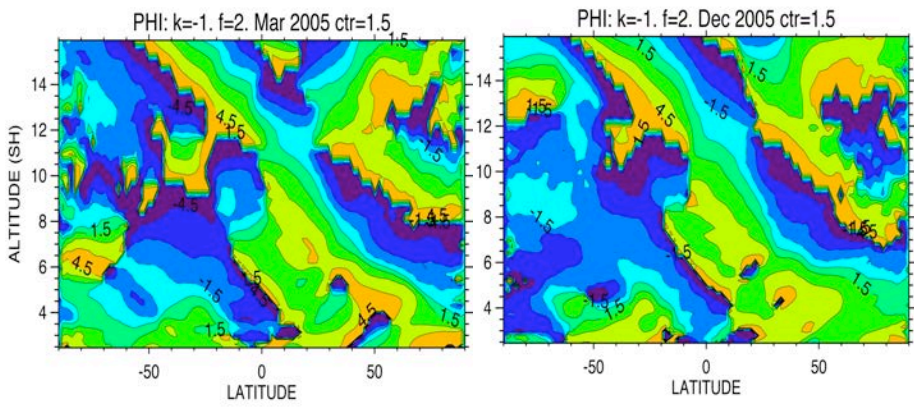
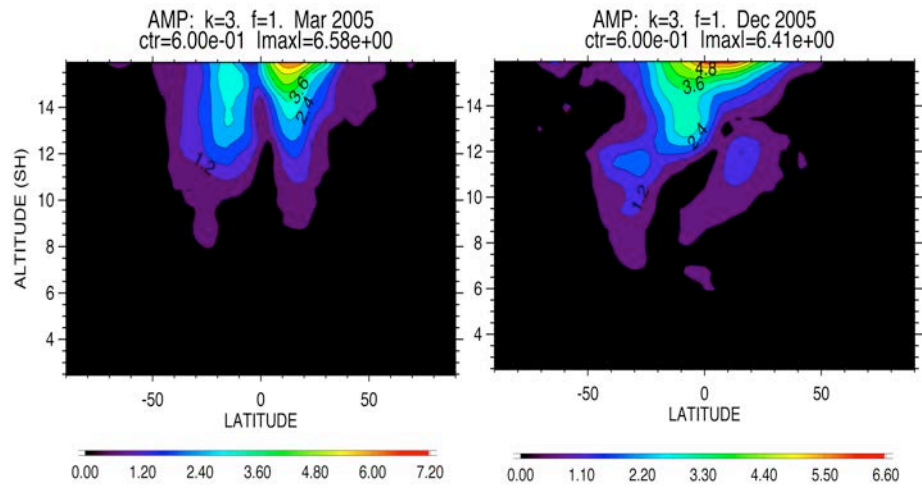
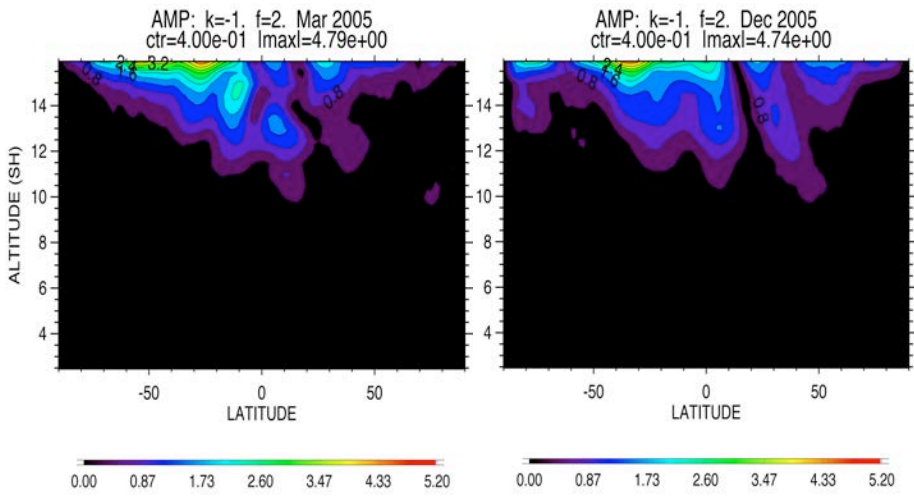
max at lat/height 0.95, 12.43 = 3.13

max at lat/height 0.95, 12.43 = 1.93

Structure of two modeled "non-migrating" tides

T semidiurnal westward 1
March December

T diurnal Eastward 3
March December



max at lat/height 0.95, 12.43 = 1.14

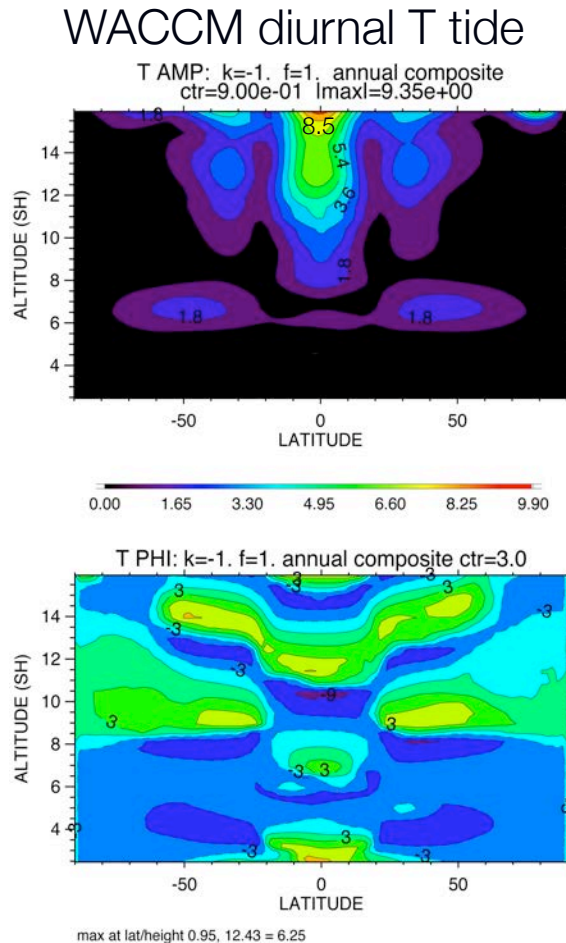
max at lat/height 0.95, 12.43 = 1.13

max at lat/height 0.95, 12.43 = 0.77

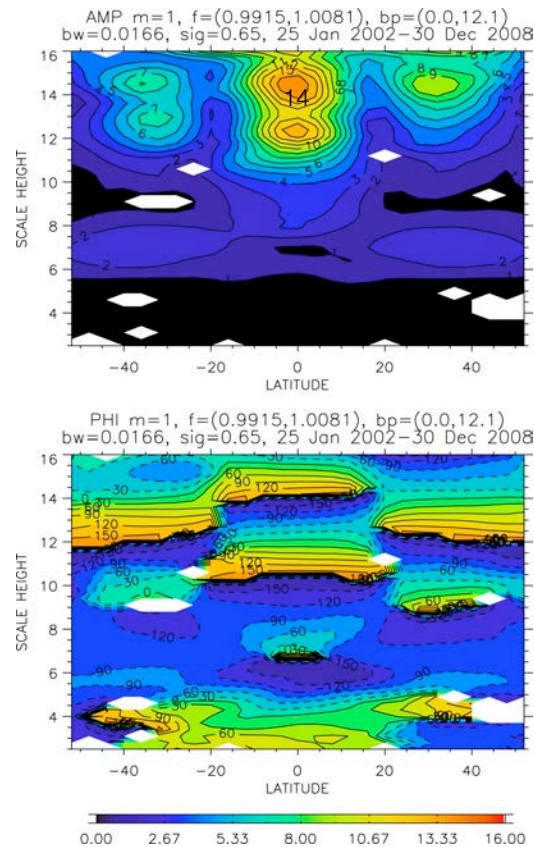
max at lat/height 0.95, 12.43 = 1.56

Diurnal migrating T tide: WACCM vs. SABER

WACCM diurnal tide obtained from monthly amplitude/phase over period of simulation (50 yr) Structure shown here is the time-mean over the entire simulation period



SABER diurnal T tide



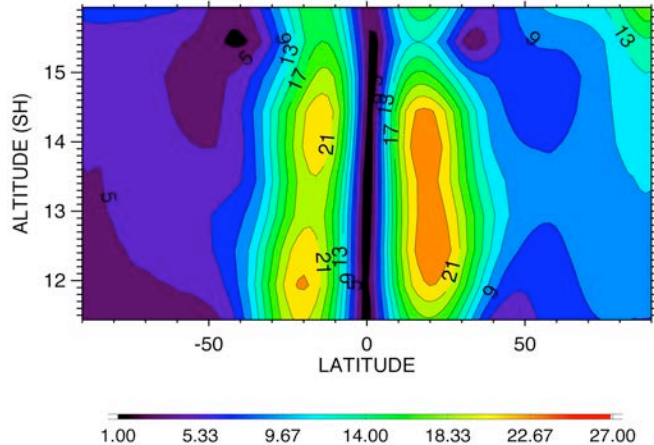
SABER diurnal tide from Salby spectral analysis of SABER data Tide structure determined from coherence of spectrum over a band centered at 1 cpd westward Results may be viewed as long-term mean over SABER period, 2002-2007

The morphology of the tide is generally consistent between WACCM and SABER WACCM amplitudes are considerably smaller, especially in the mesosphere and lower thermosphere

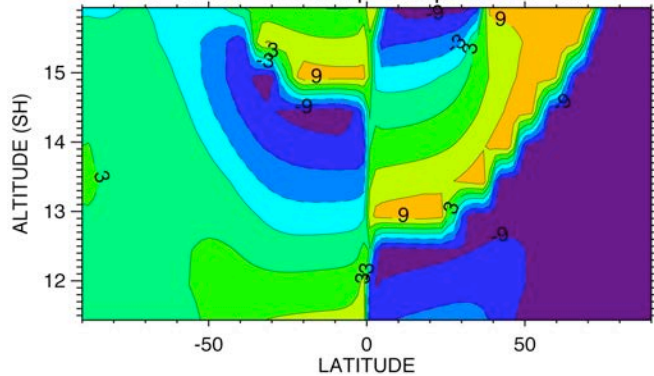
Diurnal migrating V tide structure : TIDI vs. WACCM

WACCM may composite

V AMP: $k=-1$, $f=1$, Apr composite
ctr=2.00e+00 lmaxl=2.48e+01



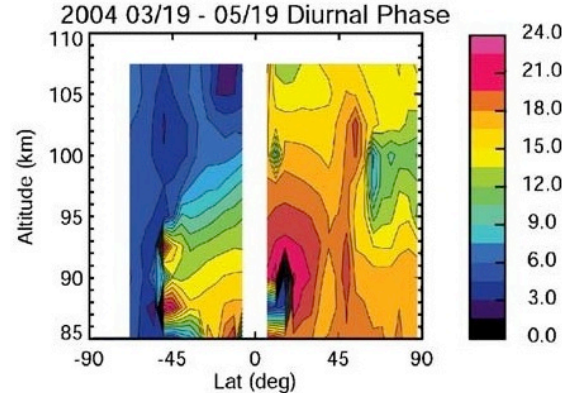
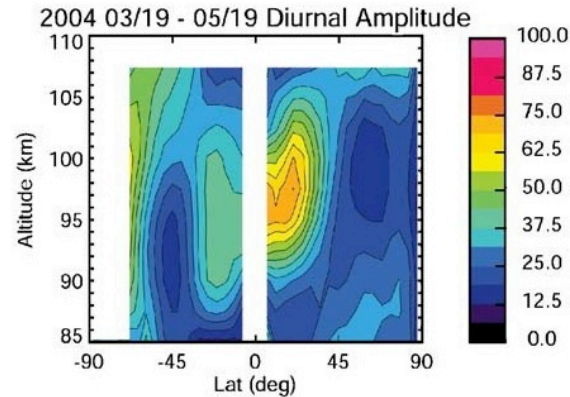
V PHI: $k=-1$, $f=1$, Apr composite ctr=3.0



max at lat/height 0.95, 12.43 = 3.06

TIDI march-may 2004

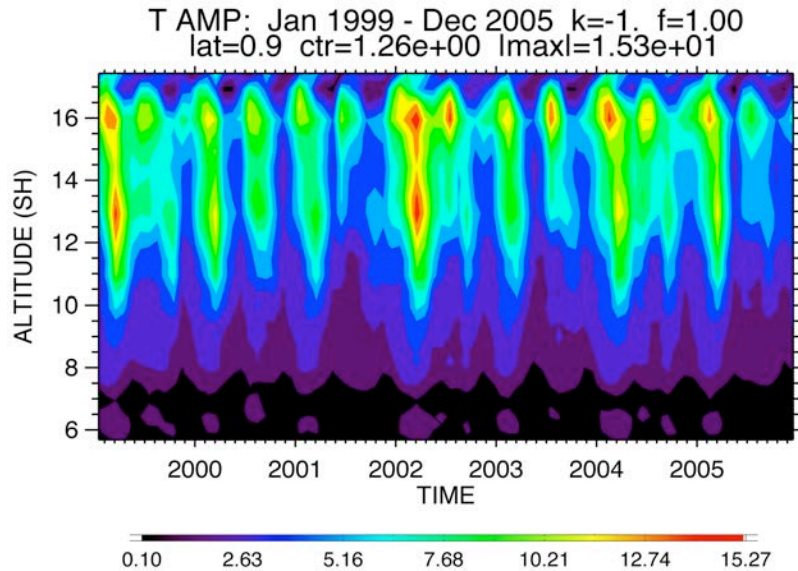
TIDI Diurnal Migrating V Tide



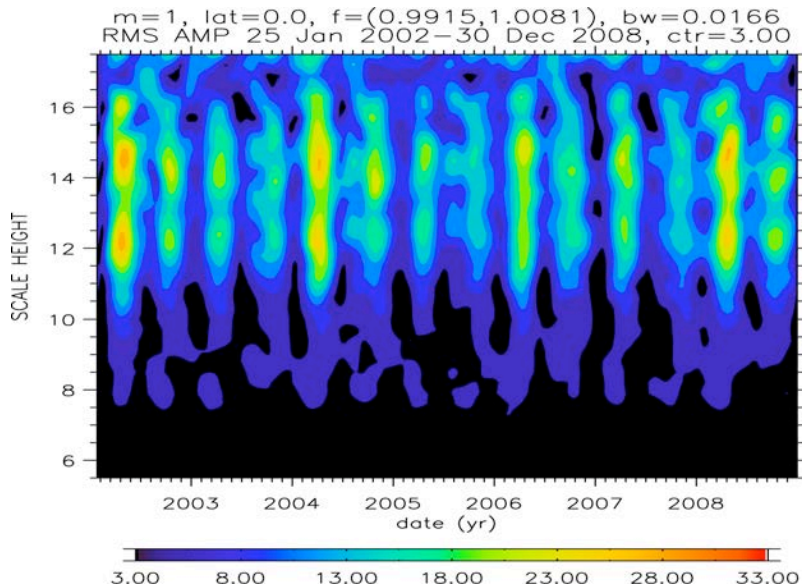
Liu et al. (JASTP, 2006)

- Similar structure in WACCM and TIDI observations, but smaller amplitudes in the model

Altitude-time variability of the diurnal tide at the Equator

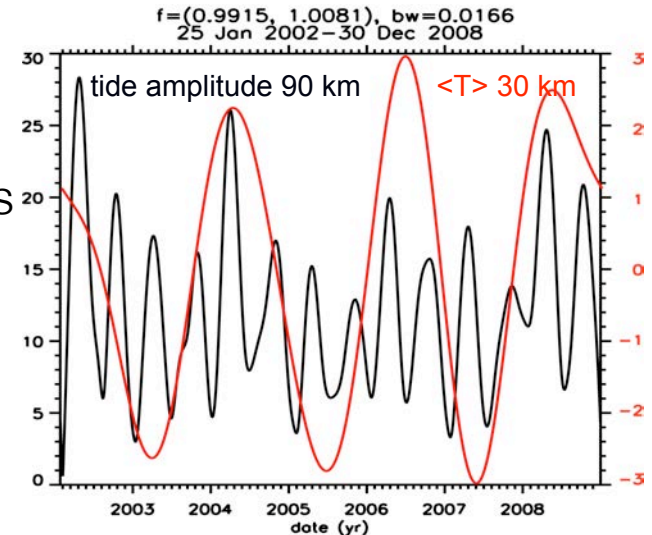


model and observations display a clear semi-annual variation, with maxima at the equinoxes
 amplitudes observed by SABER are about 2X larger than calculated with WACCM there is also considerable interannual variability (quasi-biennial)



Similar behavior is seen in SABER data

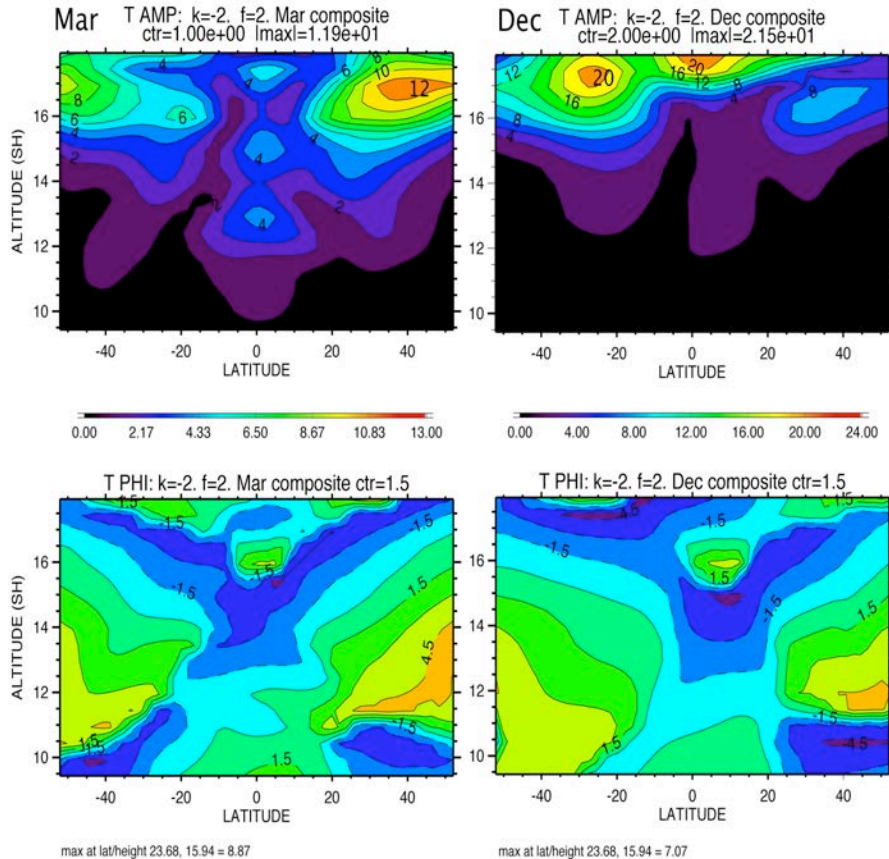
SABER Diurnal T tide @ Equator, 90 km



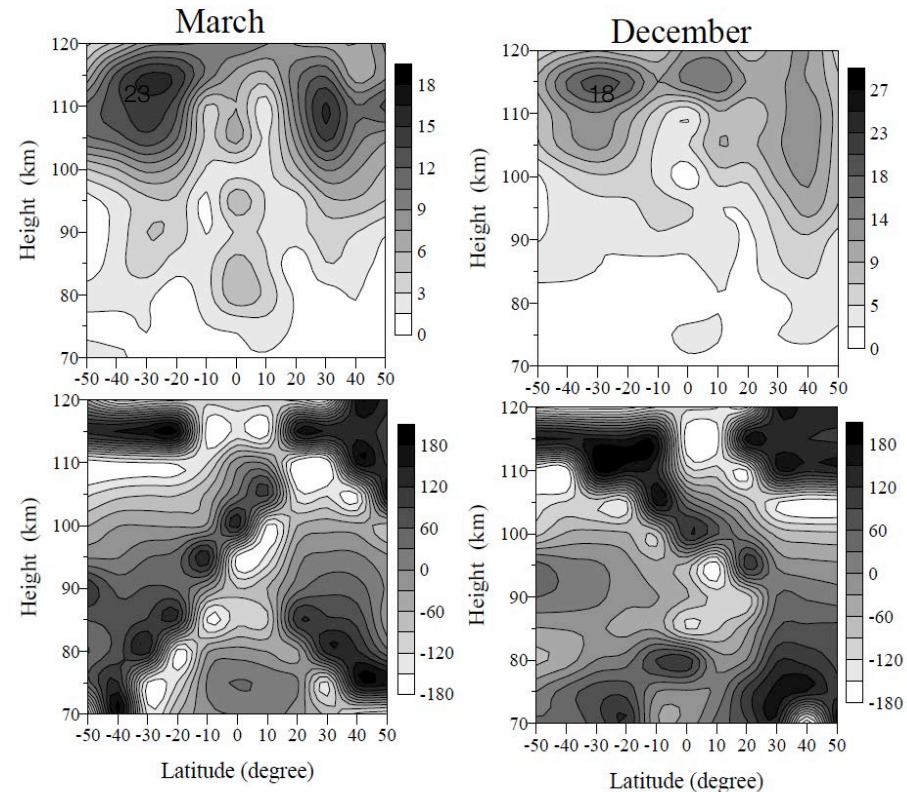
QBO at ~35 mb in red

Semidiurnal migrating T tide: WACCM vs. SABER

WACCM Semidiurnal migrating T tide



SABER Semidiurnal migrating T tide



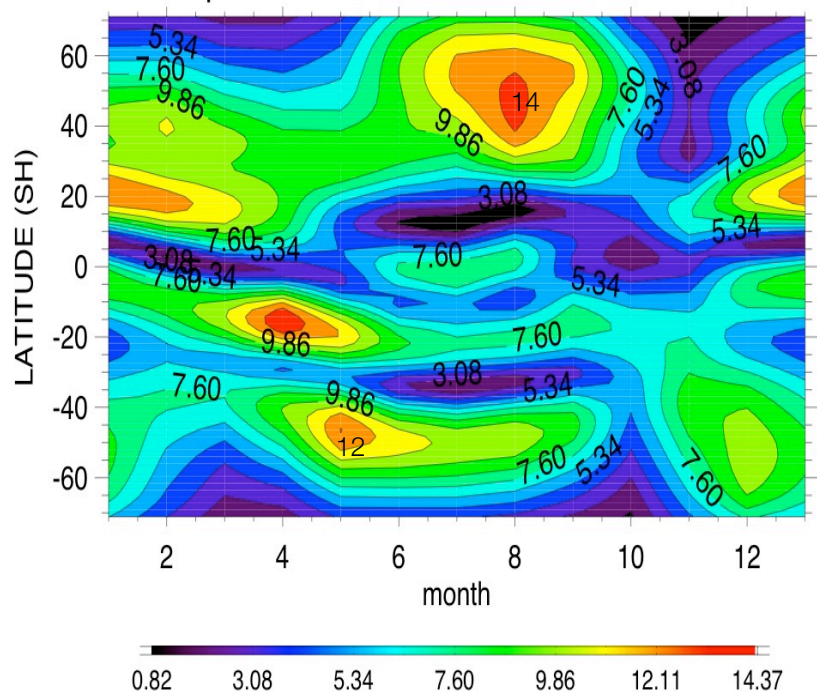
Pancheva et al., 2009

As with diurnal tide, structures are generally consistent; but WACCM amplitudes are smaller

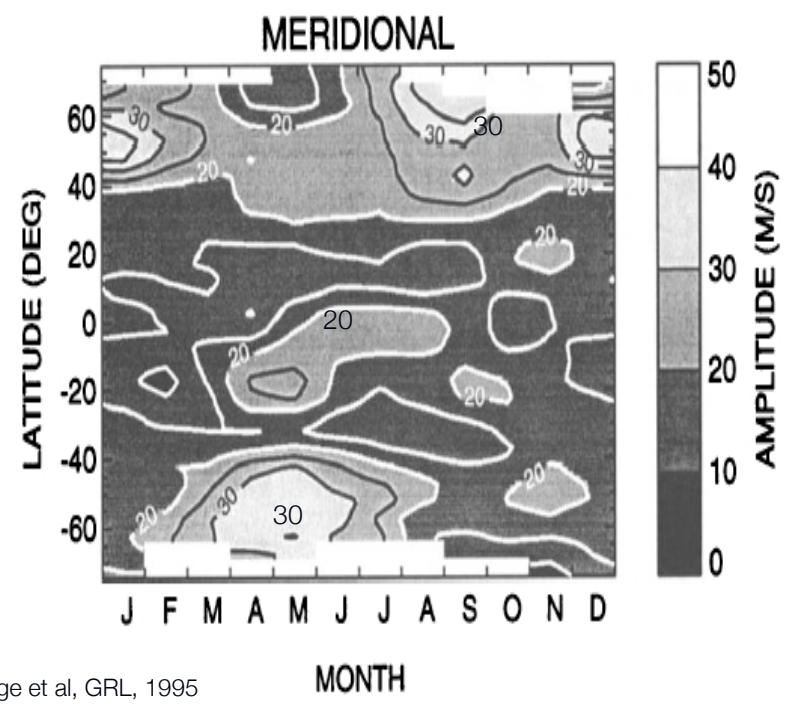
Seasonal variability of the semidiurnal V tide, 95 km

WACCM composite V seasonal cycle, ~95 km

AMP: Seasonal Composite $k=-2$. $f=2.00$
 $z_p=13.9$ $ctr=1.13e+00$ $lmax=1.44e+01$



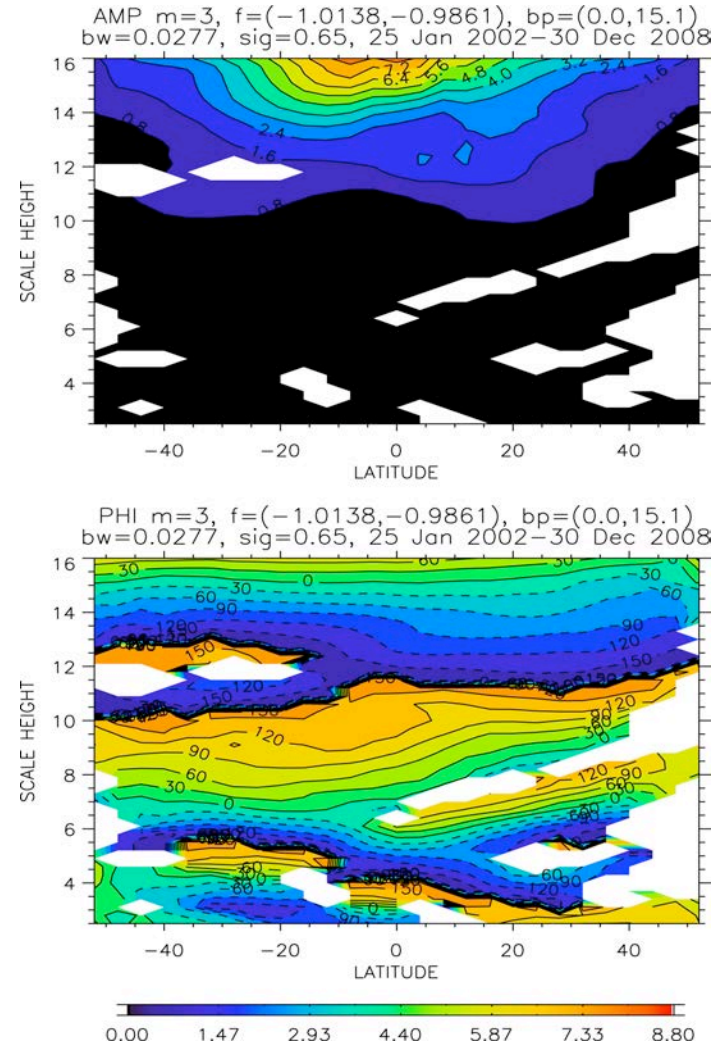
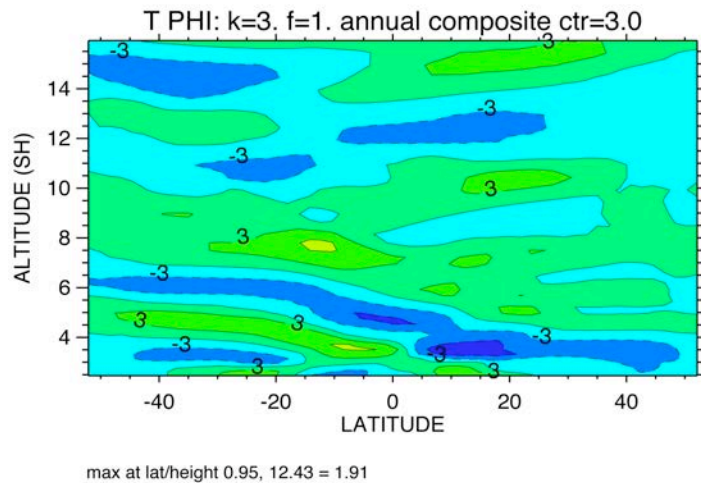
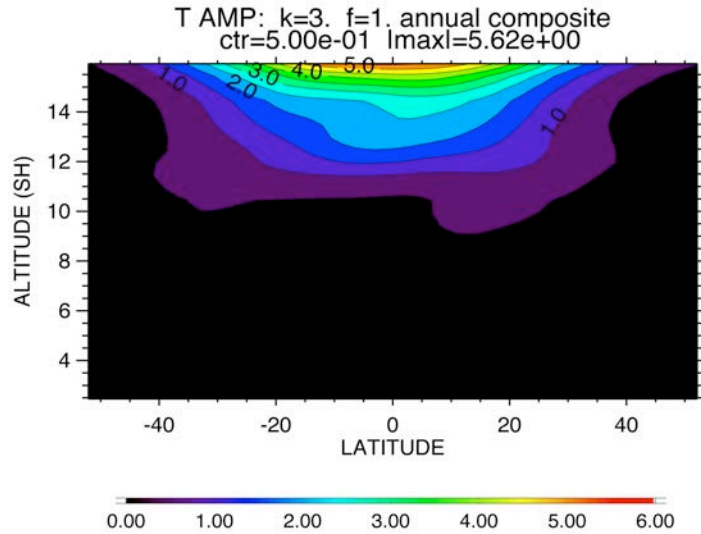
HRDI composite V seasonal cycle @ 95 km



Burrage et al, GRL, 1995

WACCM: 50-year mean
 UARS/HRDI: Nov 1991 – July 1994

Diurnal eastward-3 T tide: WACCM vs. SABER



WACCM: 50 year mean

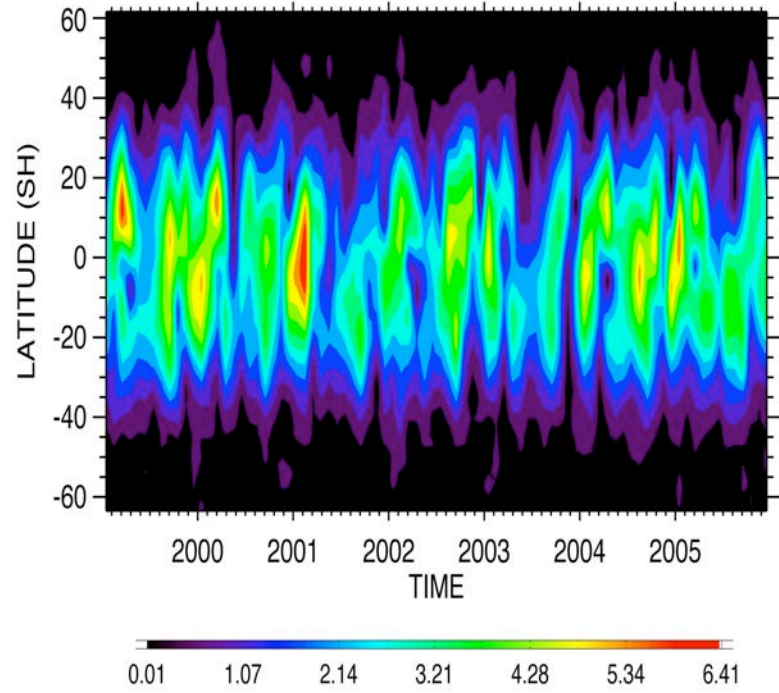
SABER: 2002-2007

large-amplitude above ~ 10 sh (70 km) has the structure of a Kelvin wave
 phase behavior suggests also RG structure at lower altitudes

Seasonal variation of eastward-3 T tide at equator

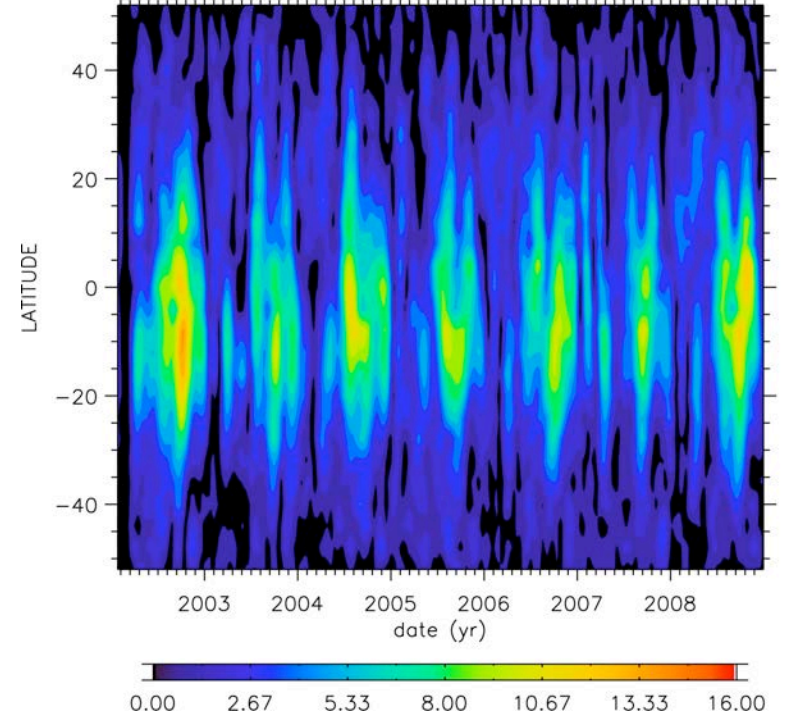
WACCM m=3 diurnal at 105 km

T AMP: Jan 1999 - Dec 2005 k=3. f=1.00
 zp=14.9 ctr=5.34e-01 lmaxl=6.41e+00



SABER m=3 diurnal at 105 km

m=3, z=15.1 sh, f=(-1.0138,-0.9861), bw=0.0277
 RMS AMP 25 Jan 2002-30 Dec 2008, ctr=1.00



SABER: regular seasonal, maxima always in late NH summer (~August)
 WACCM: semiannual variation, with maxima in August and January

WACCM-X: Thermosphere and Ionosphere Extension of WACCM

led by Han-Li Liu (HAO)

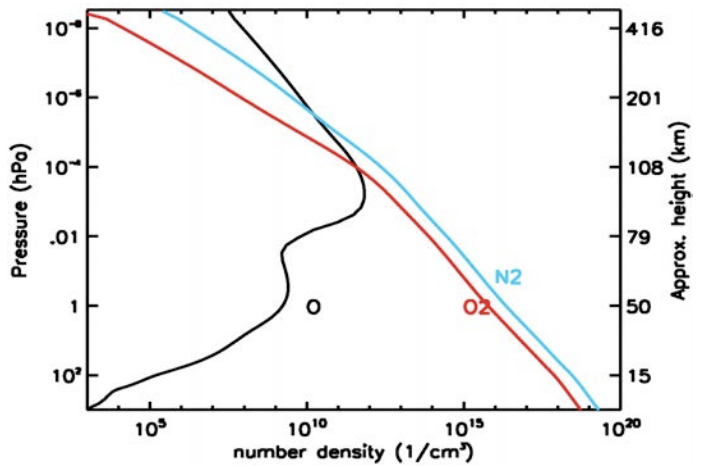


WACCM-X modifications

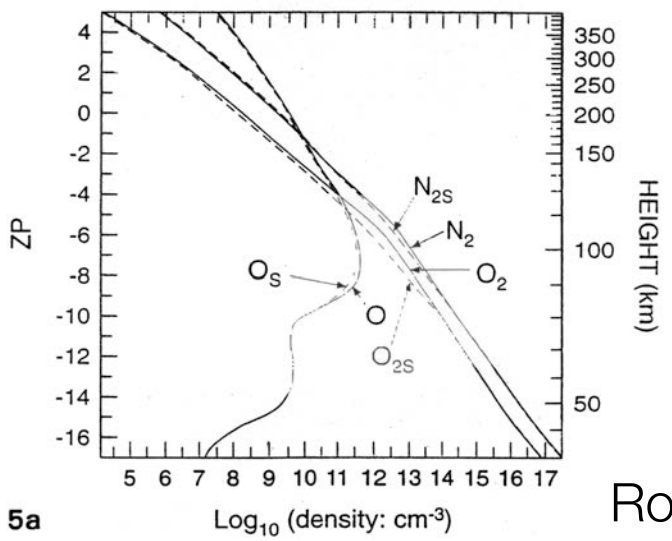
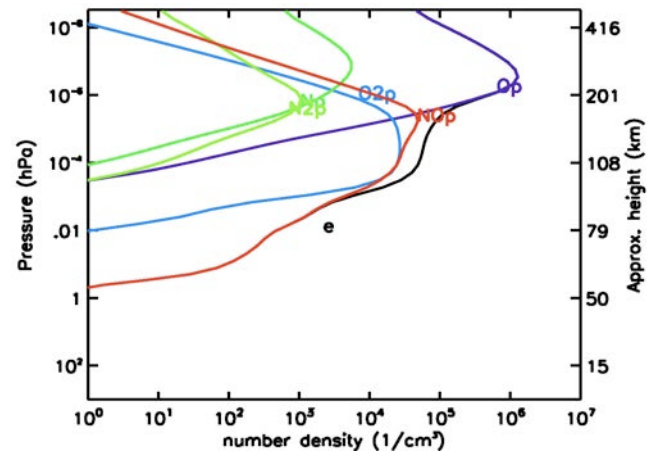
- 81 or 125 vertical levels (0-~500 km) ✓
- Major/minor species diffusion ✓
- Species dependent C_p , R , m ✓
- Ambipolar diffusion
- Ion/electron transport due to Lorentz force
- Ion/electron energy equations
- Coupling with plasmasphere (GIP) / magnetosphere (GAIM/CISM)

WACCM-X: Compositional Structures

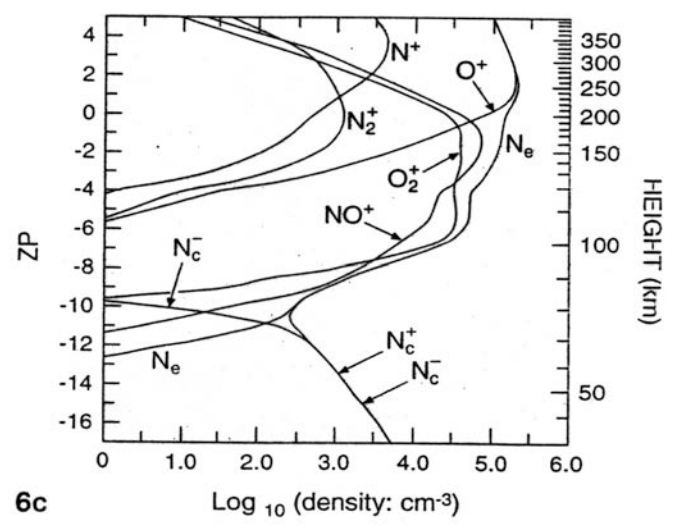
Major Species



Ions and Electrons



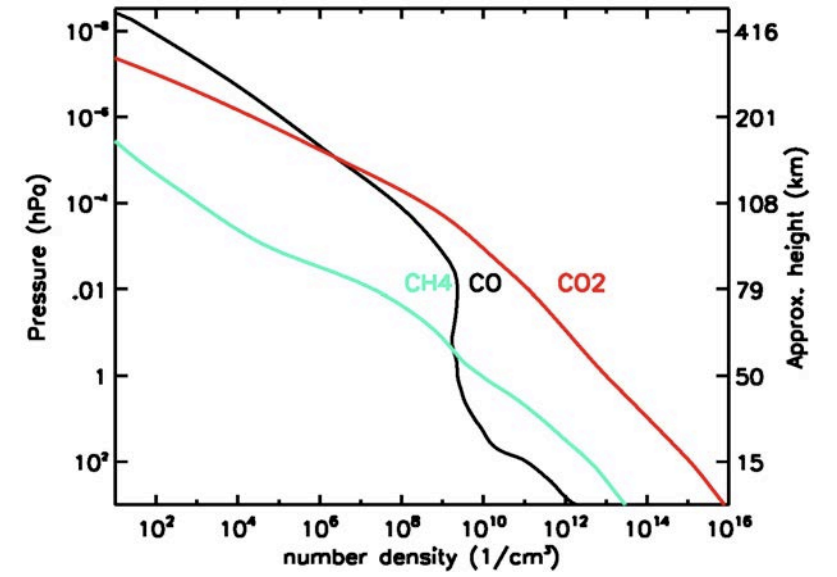
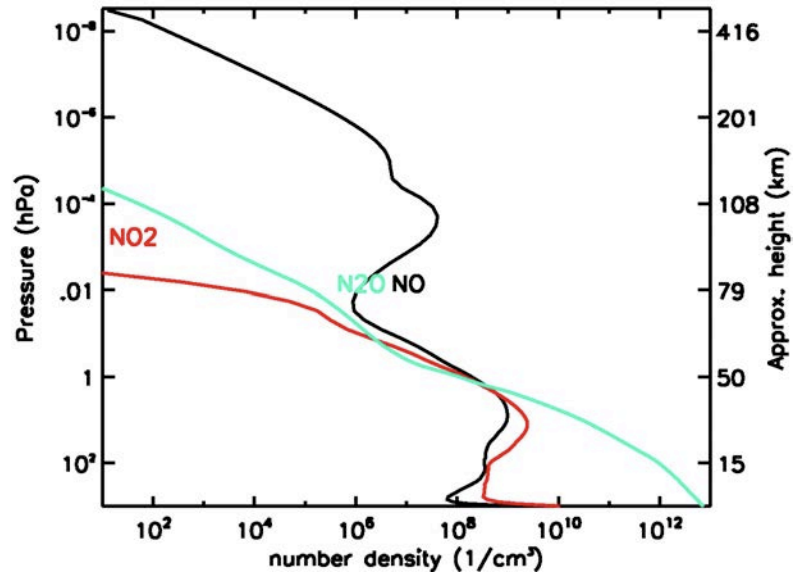
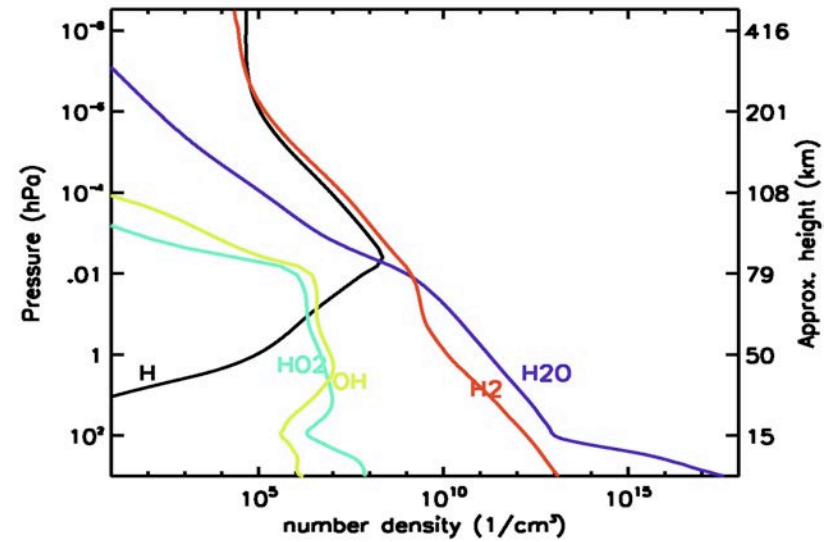
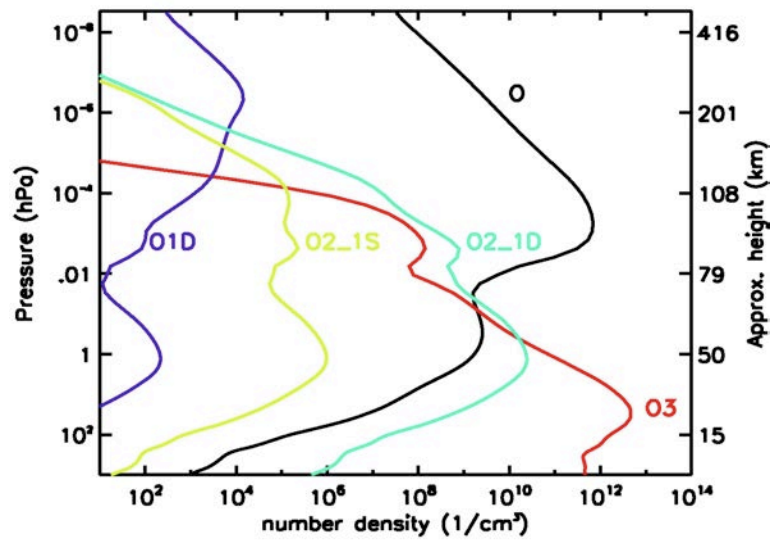
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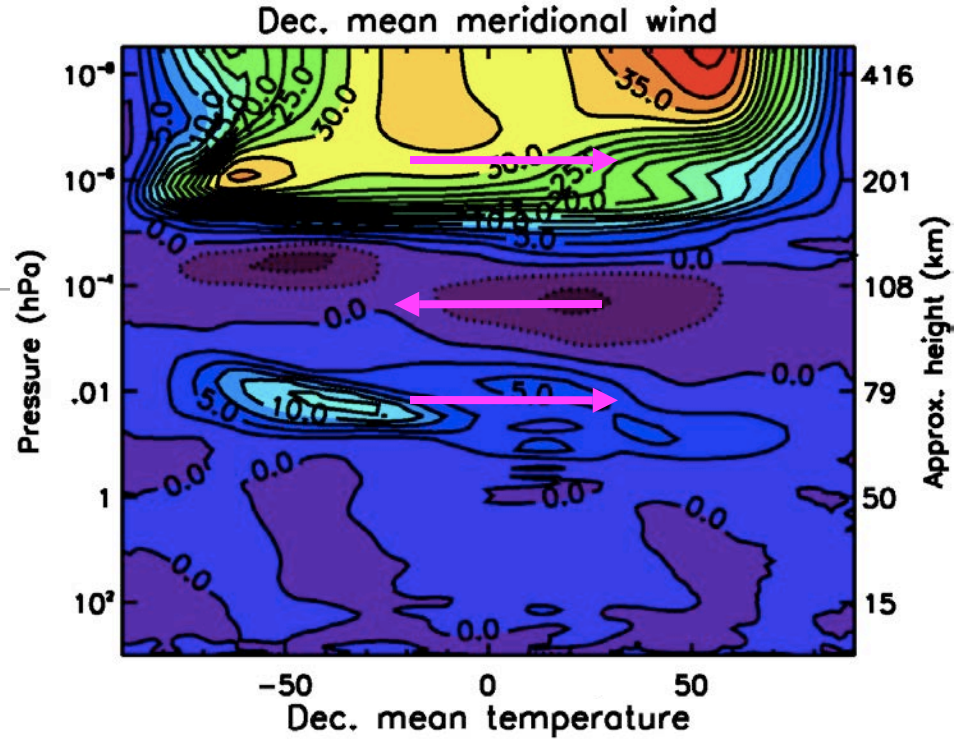
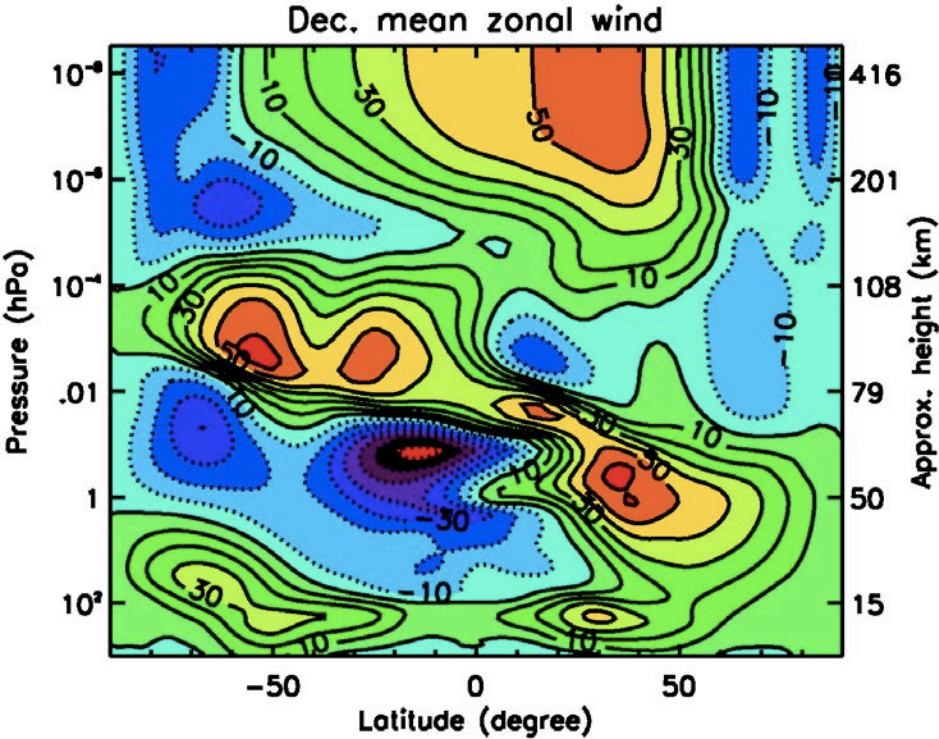


6c

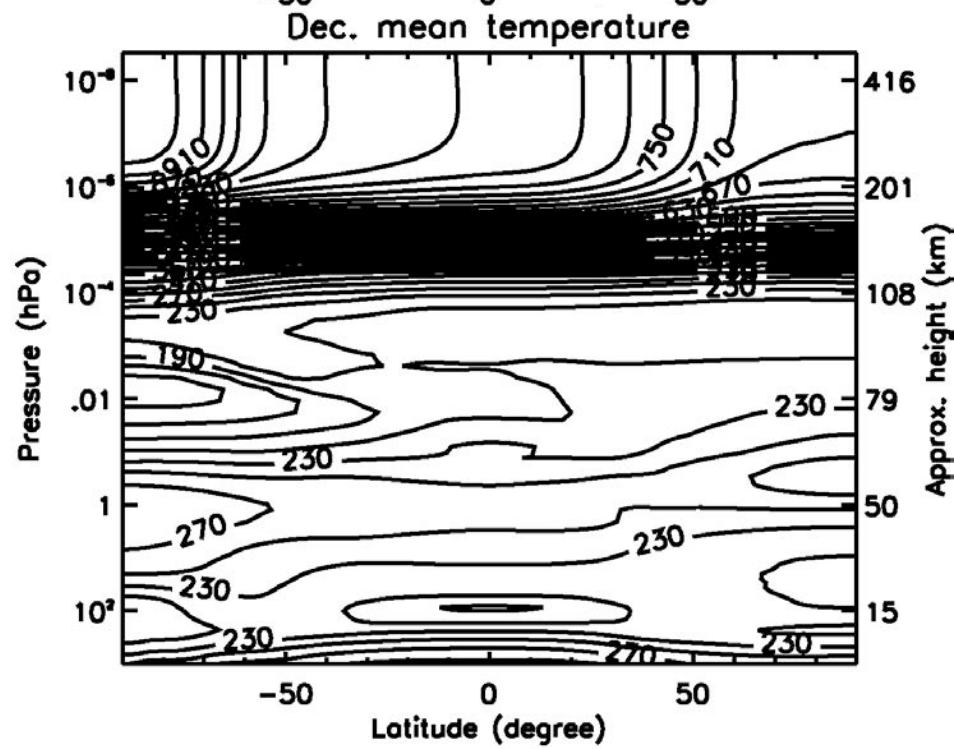
Roble, 1995

Minor species

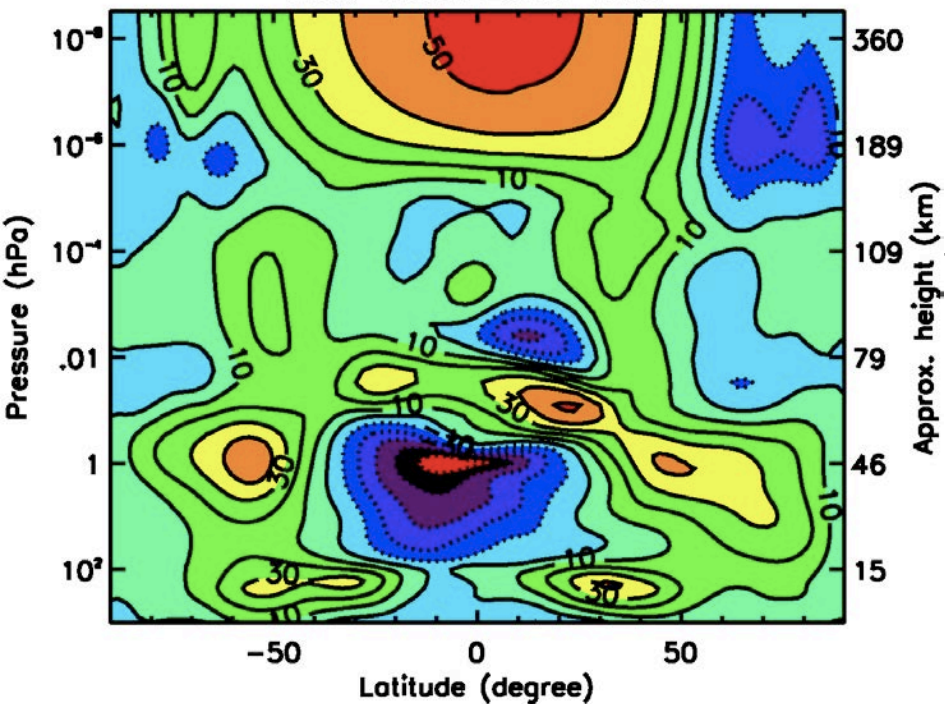




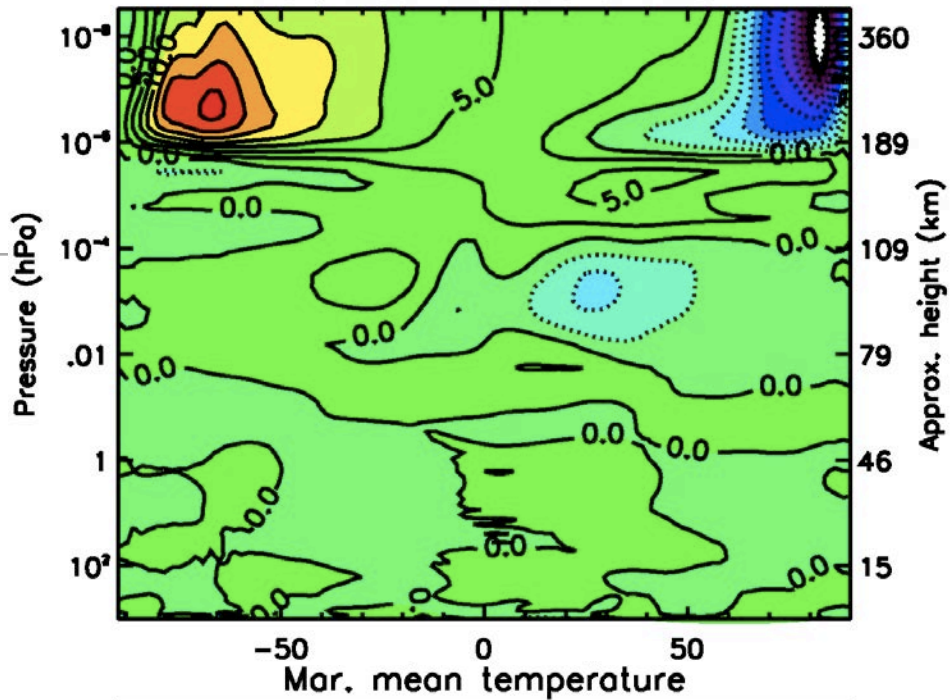
WACCM-X: Winds and
Temperature (December)



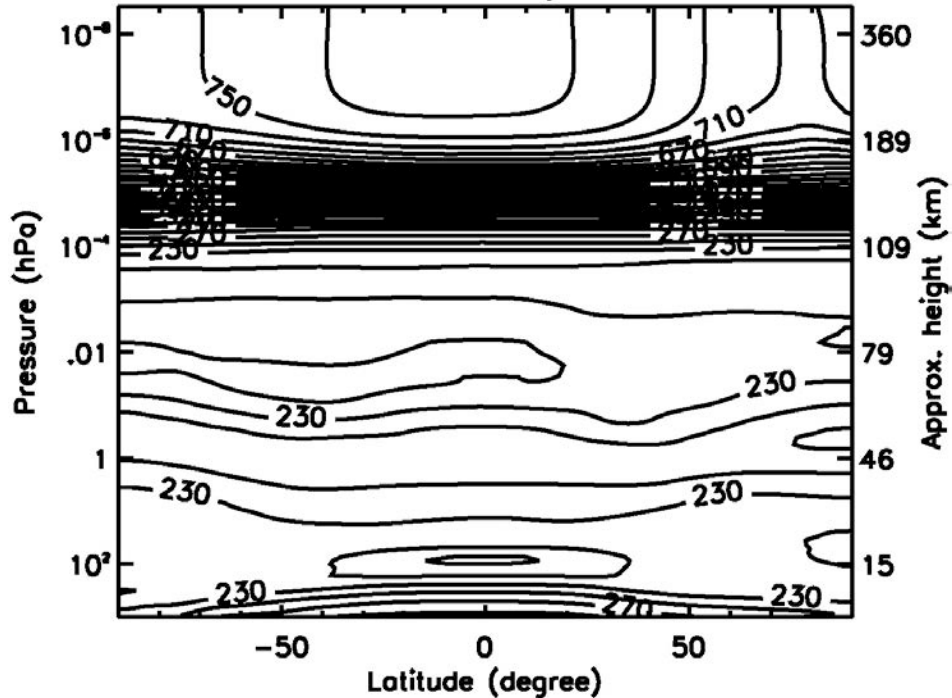
Mar. mean zonal wind



Mar. mean meridional wind



Mar. mean temperature

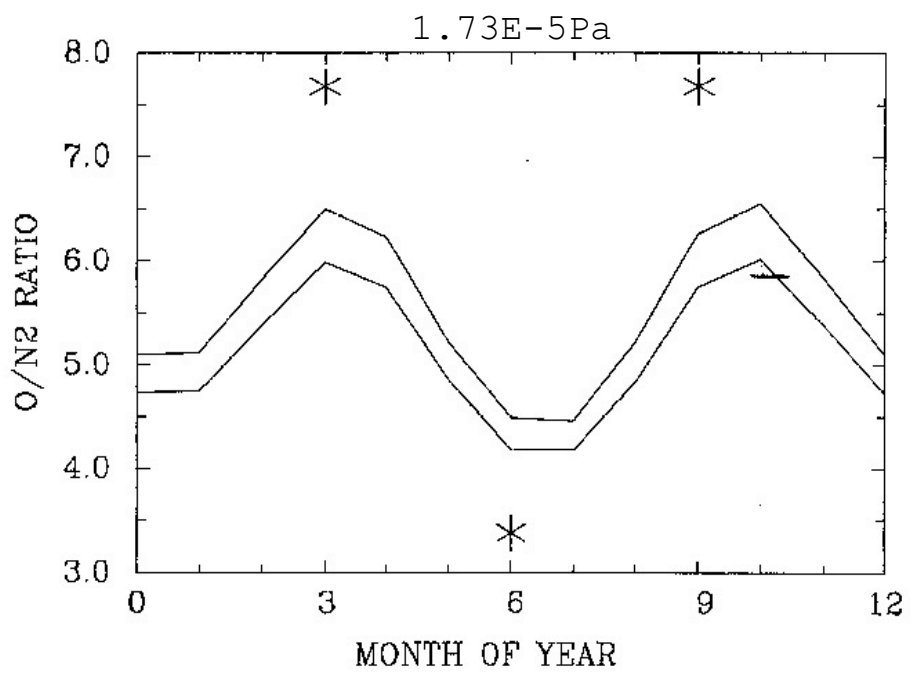
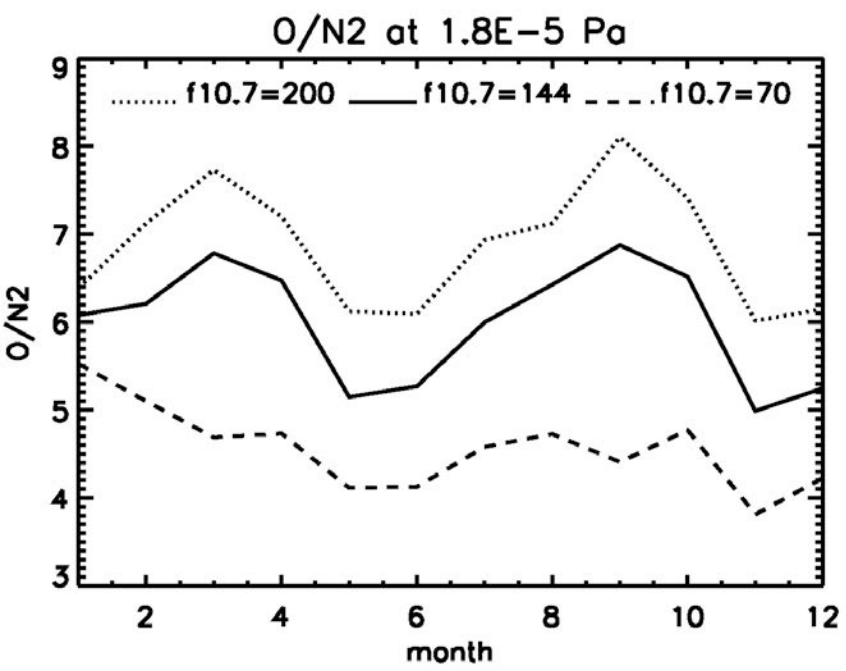


WACCM-X: Winds and Temperature (March)

Thermosphere O/N₂ Semi-annual Variation (~250 km)

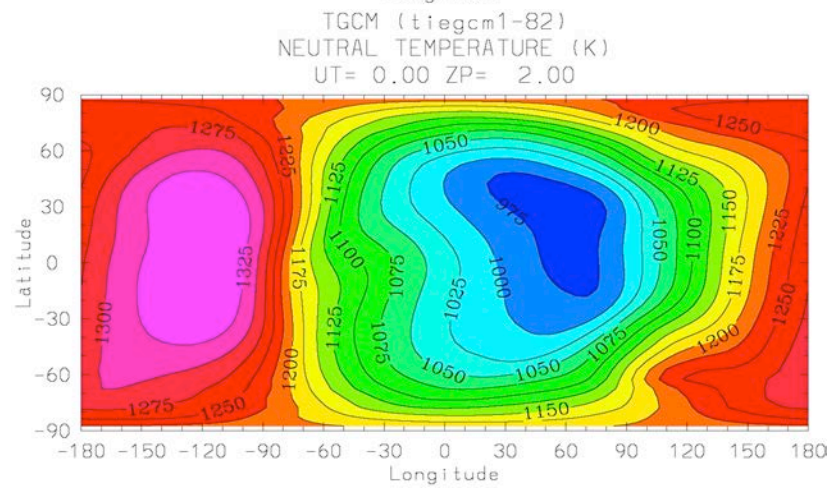
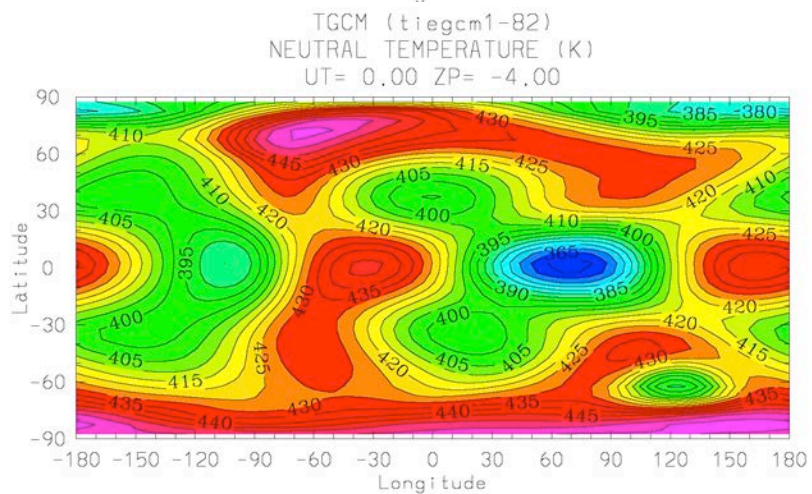
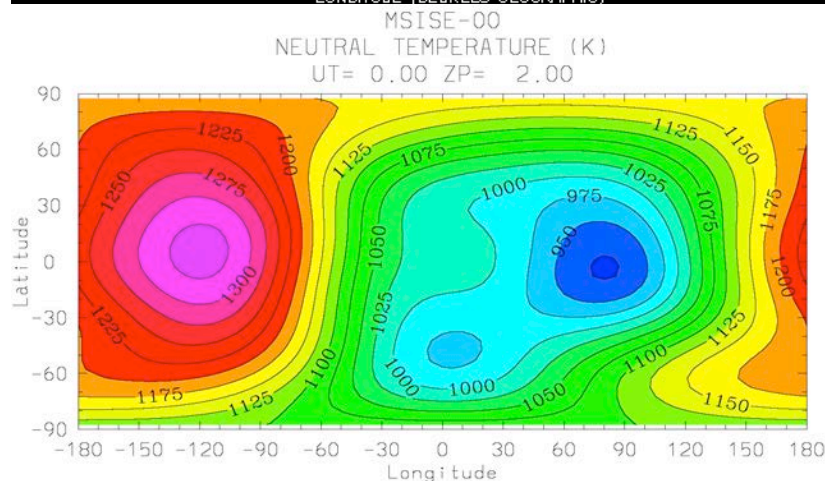
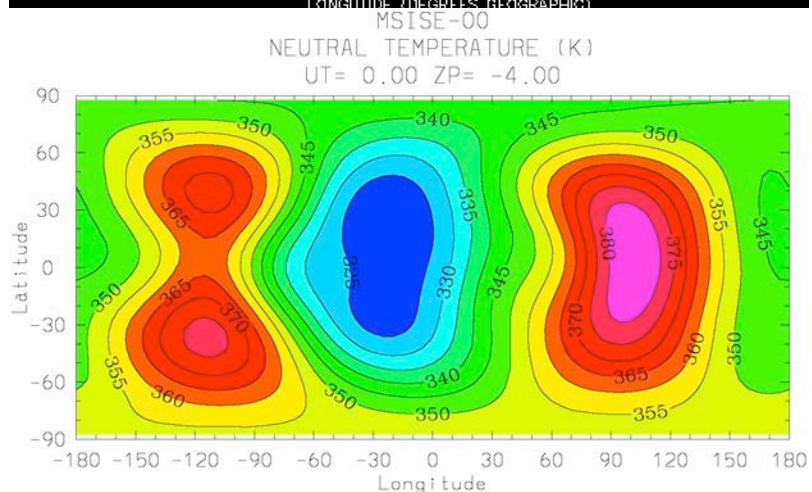
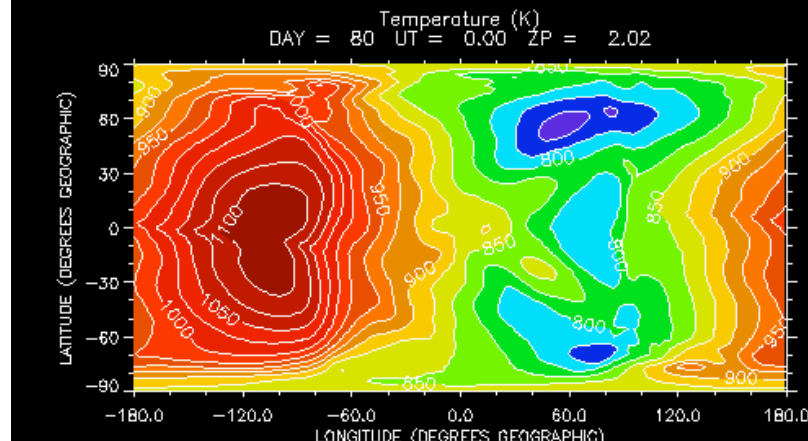
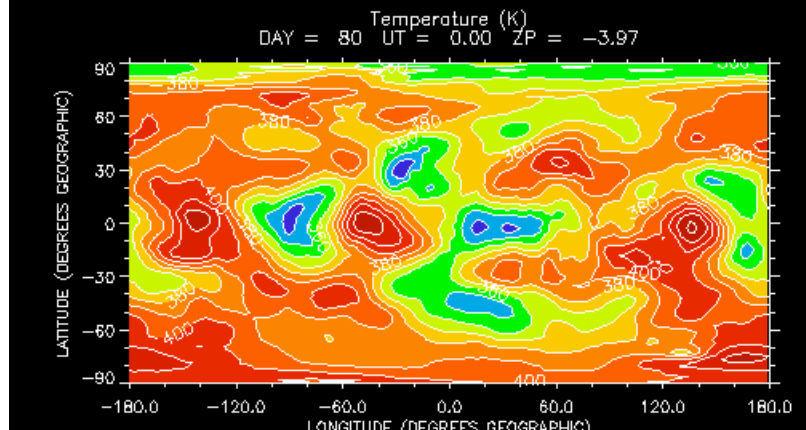
WACCM

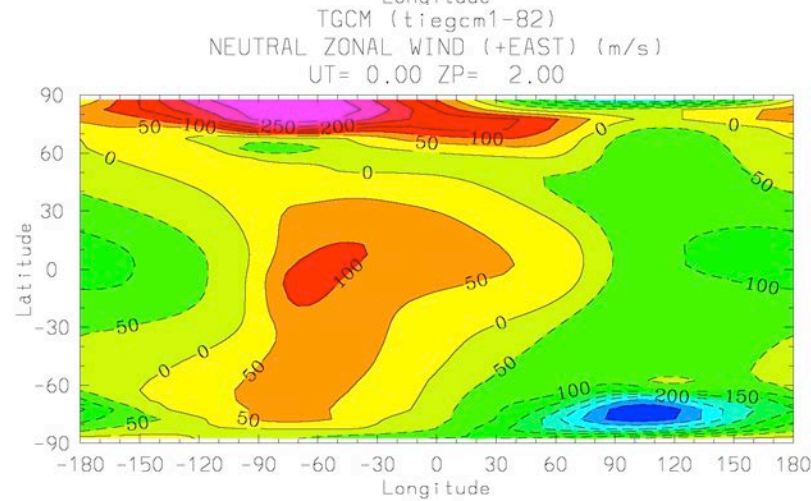
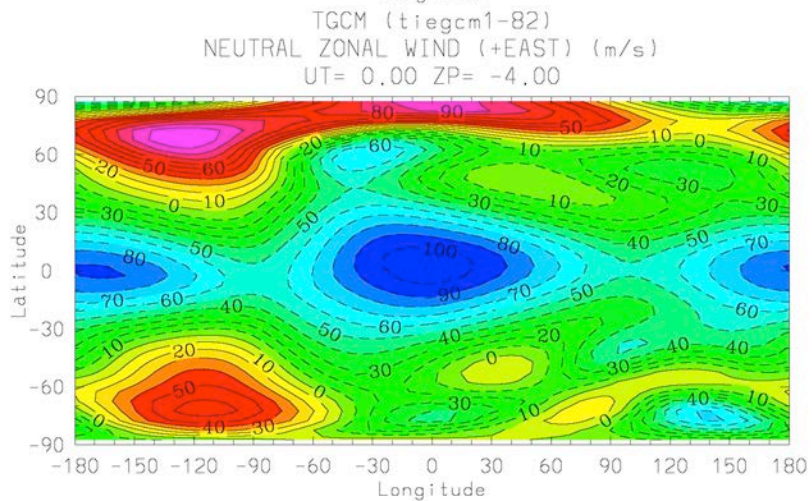
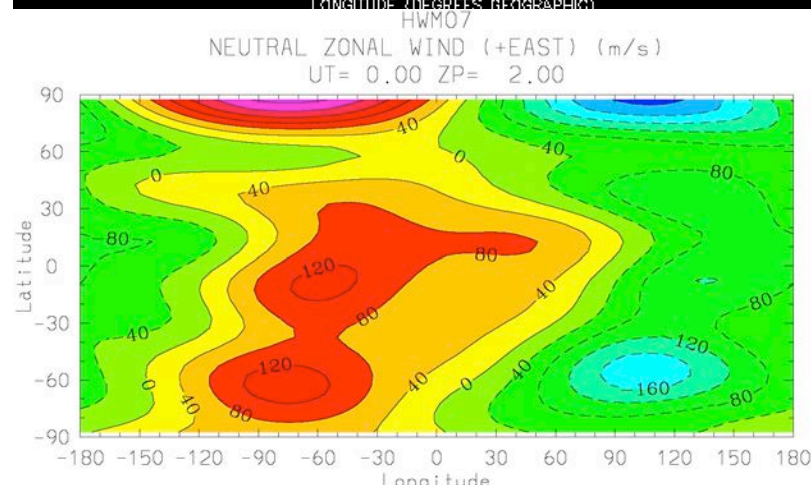
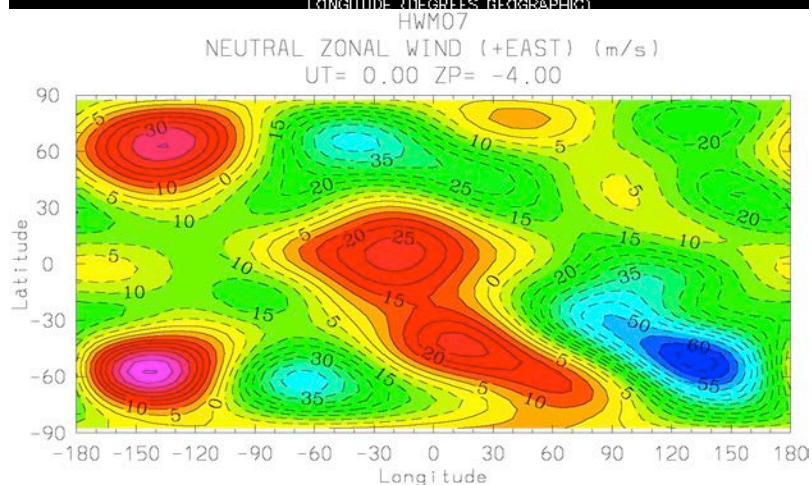
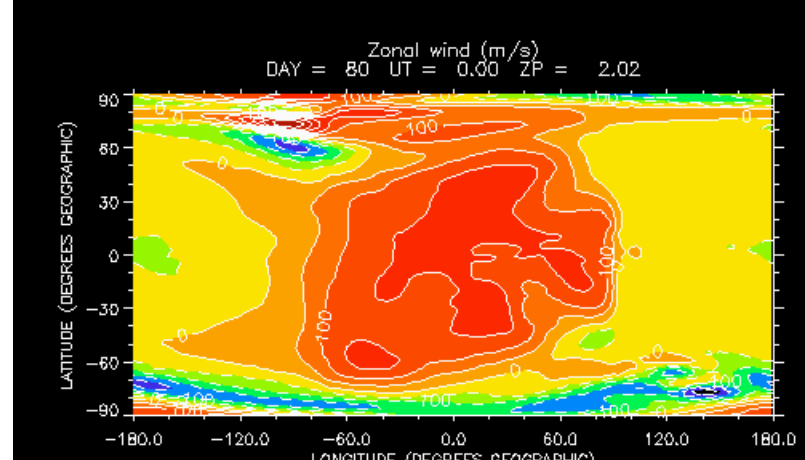
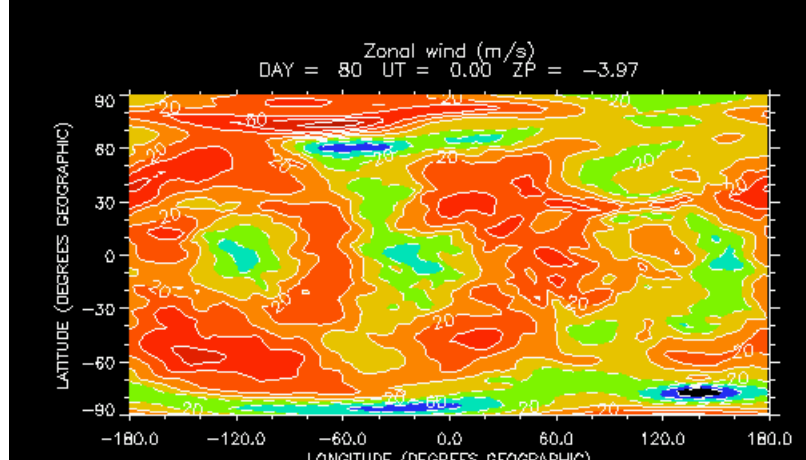
MSIS

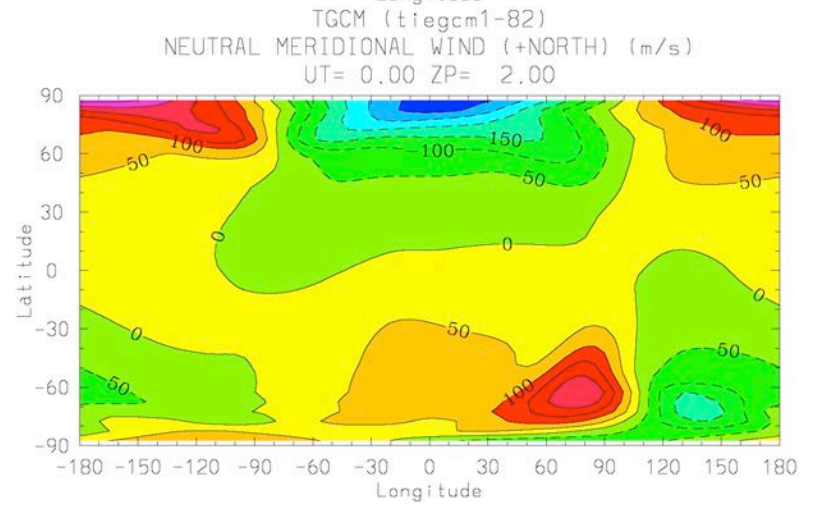
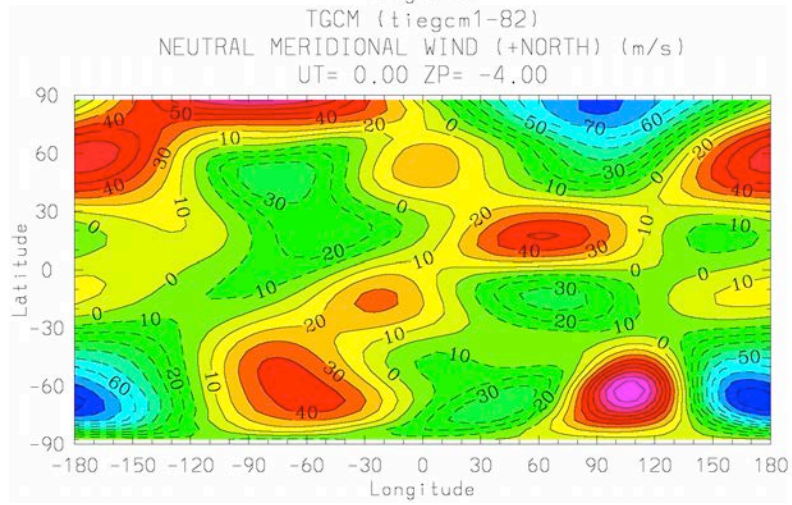
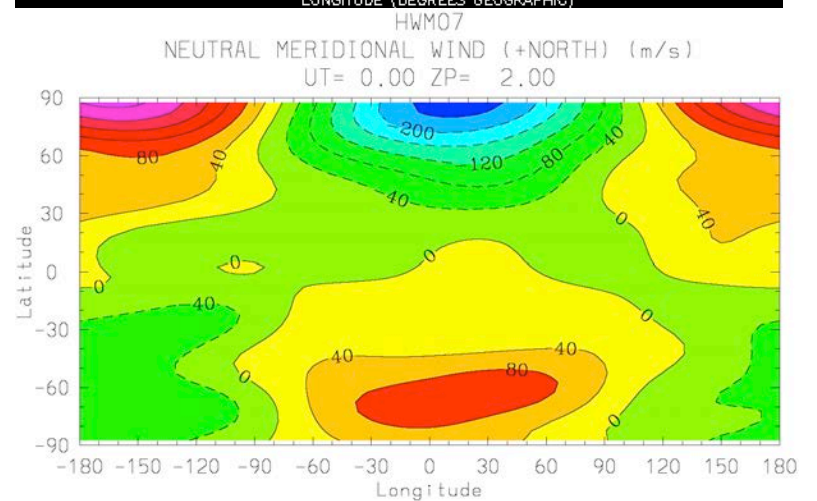
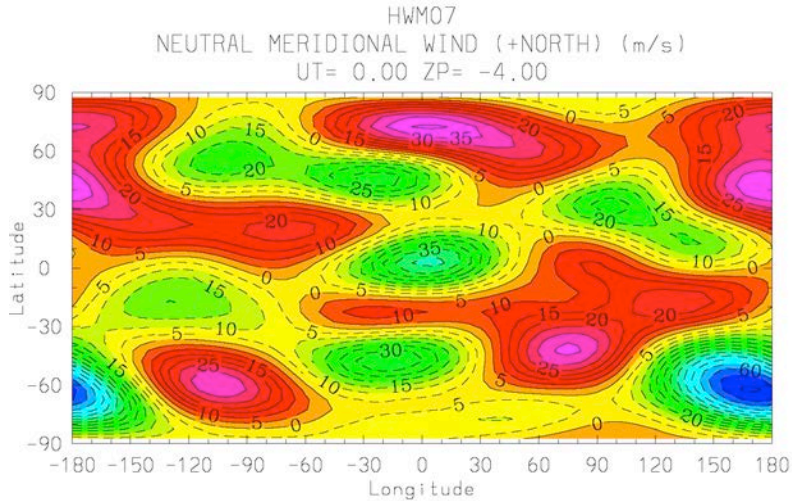
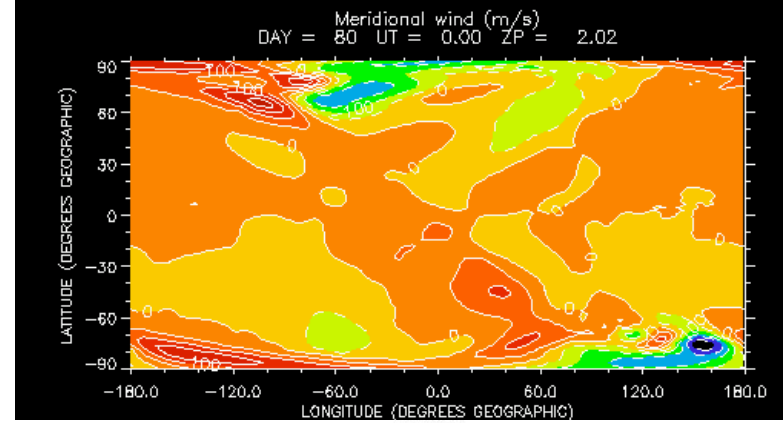
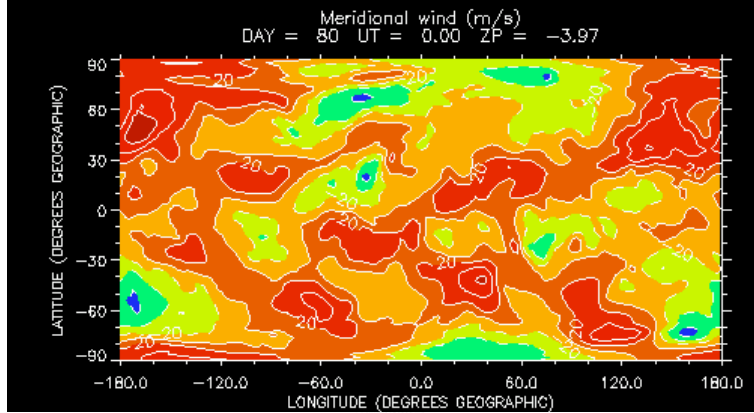


Fuller-Rowell, 1998

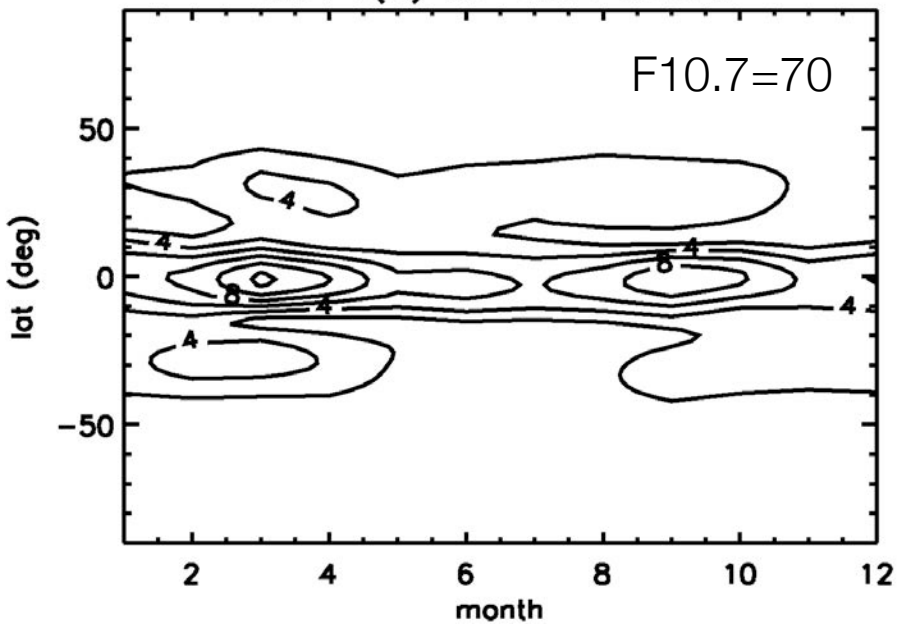
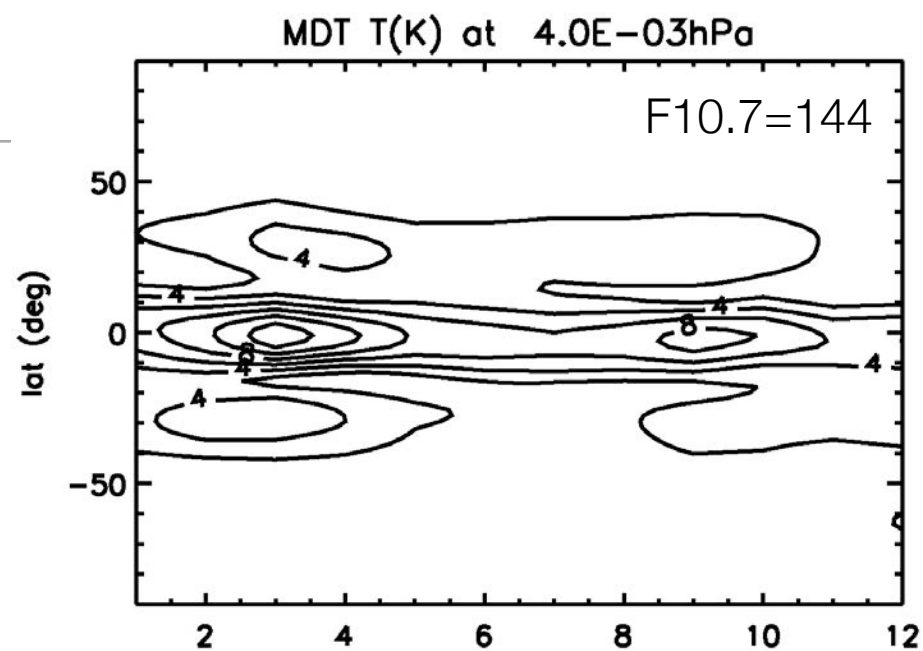
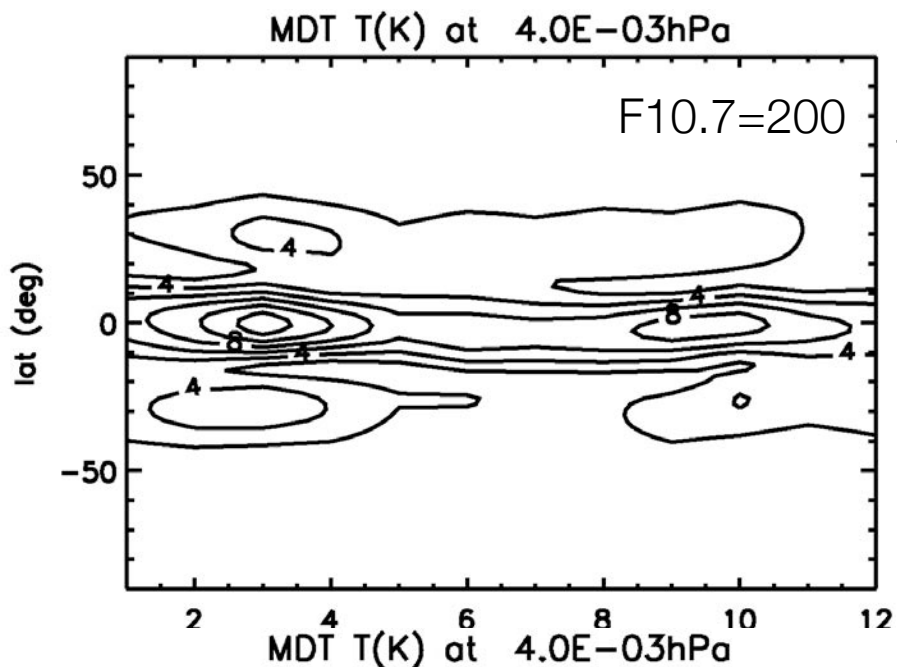
CEDAR, 2009



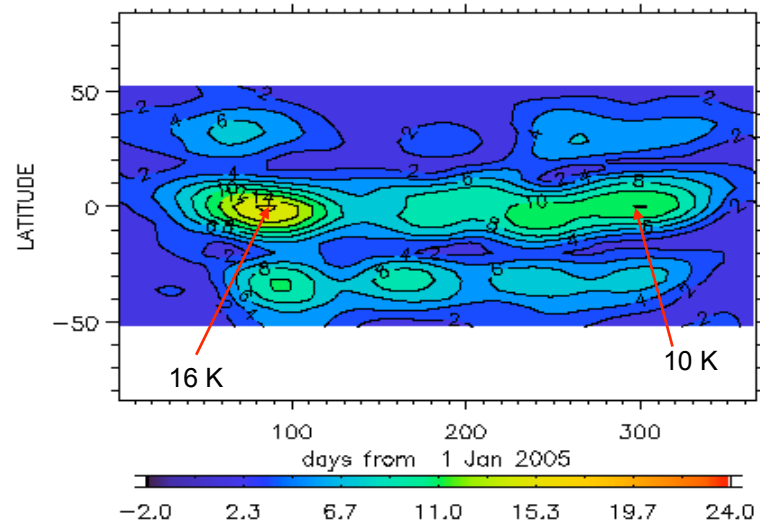




Migrating Diurnal Tide: Temperature

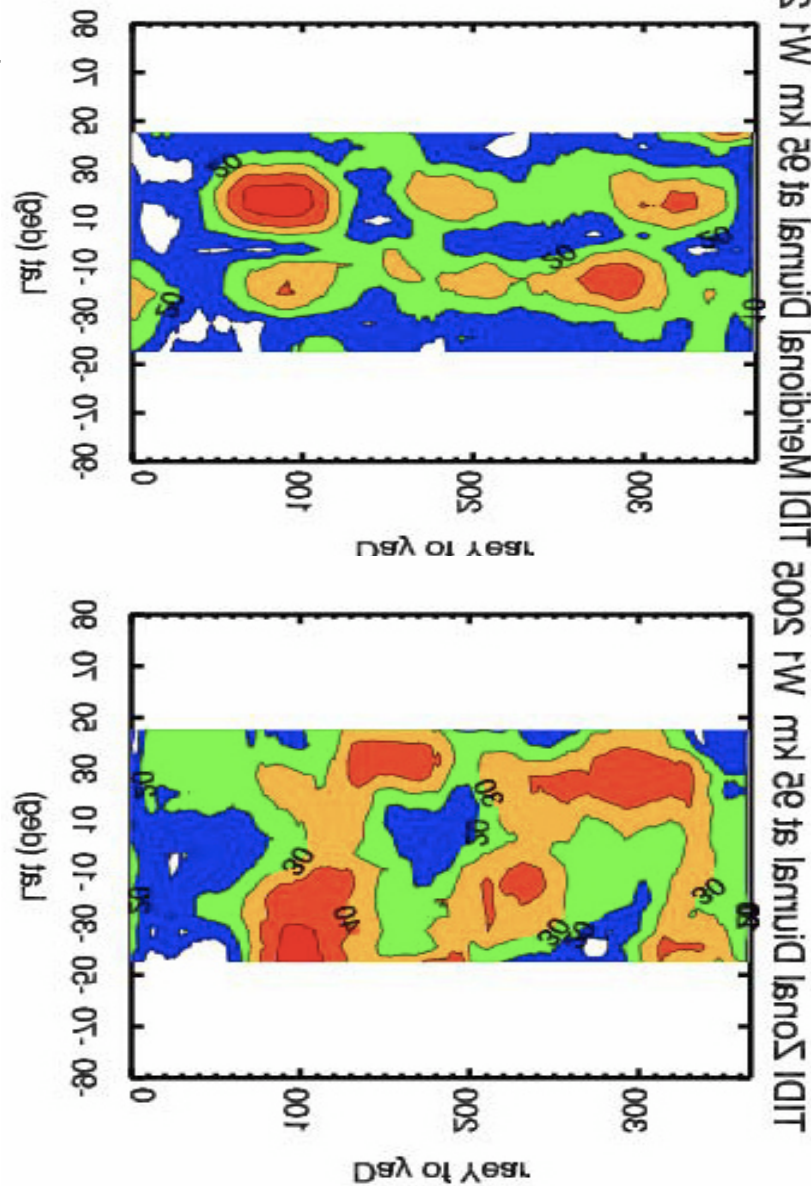
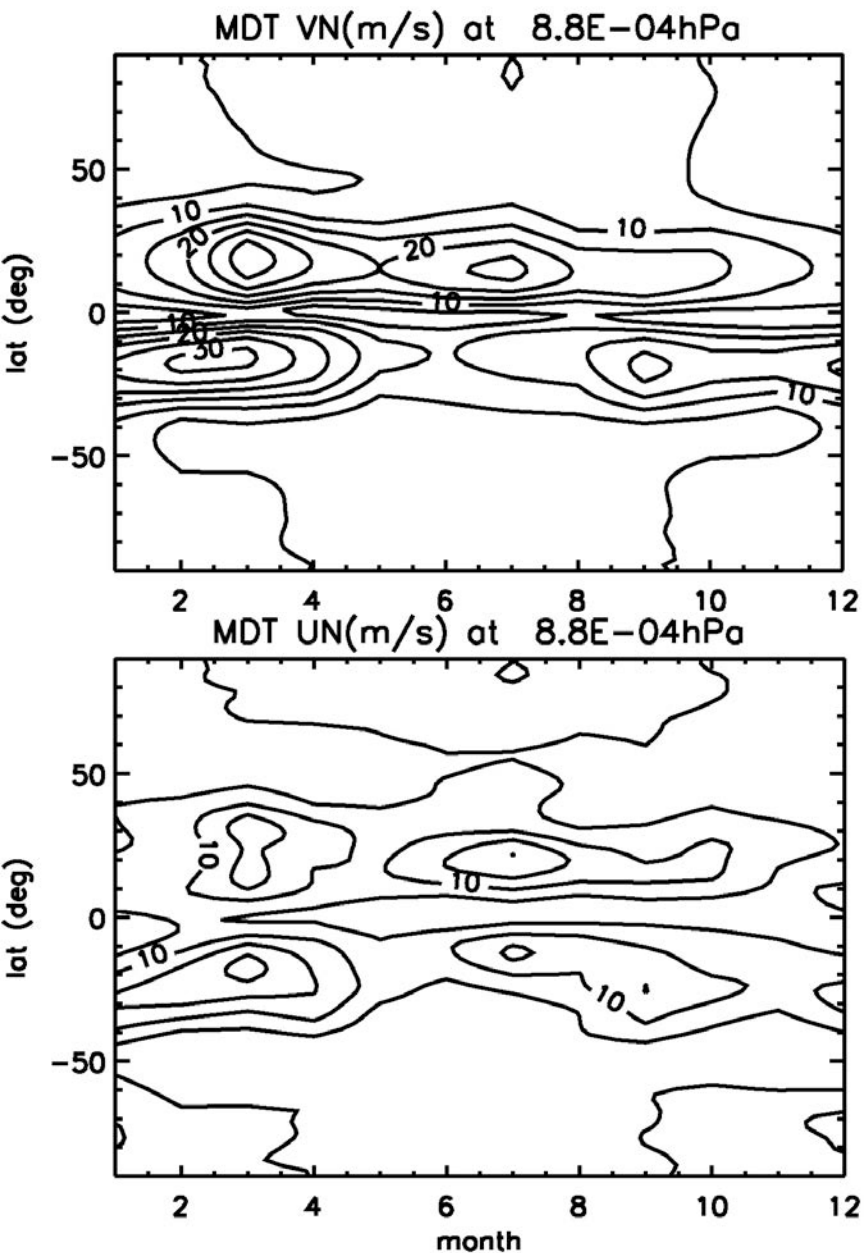


$m=1$, $z=12.5$ sh, $f=(0.986, 1.008)$, $\text{ctr}=2.00$
RMS AMP 1 Jan - 31 Dec 2005



Migrating Diurnal Tide: Horizontal Winds

Wu et al, 2008a

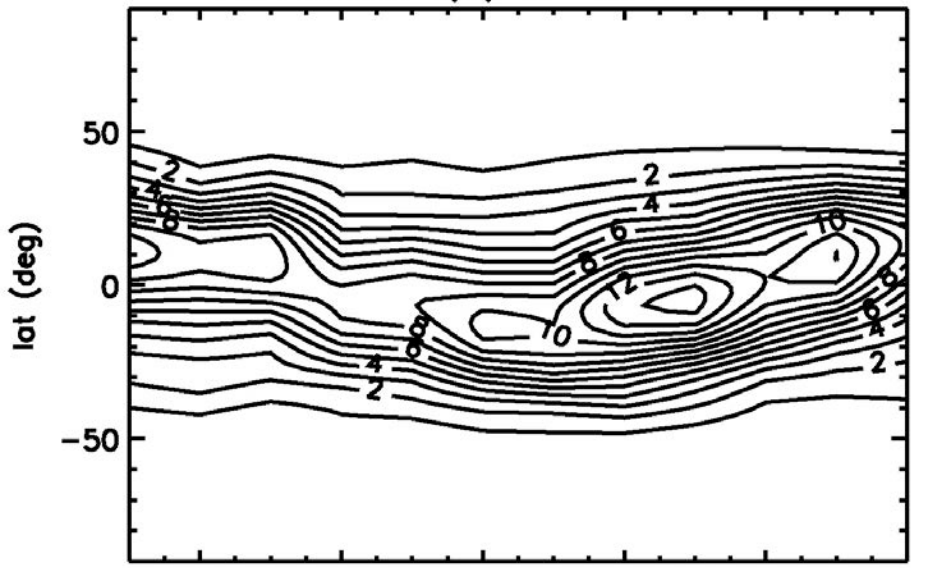


MDT VN(m/s) at $8.8E-04hPa$

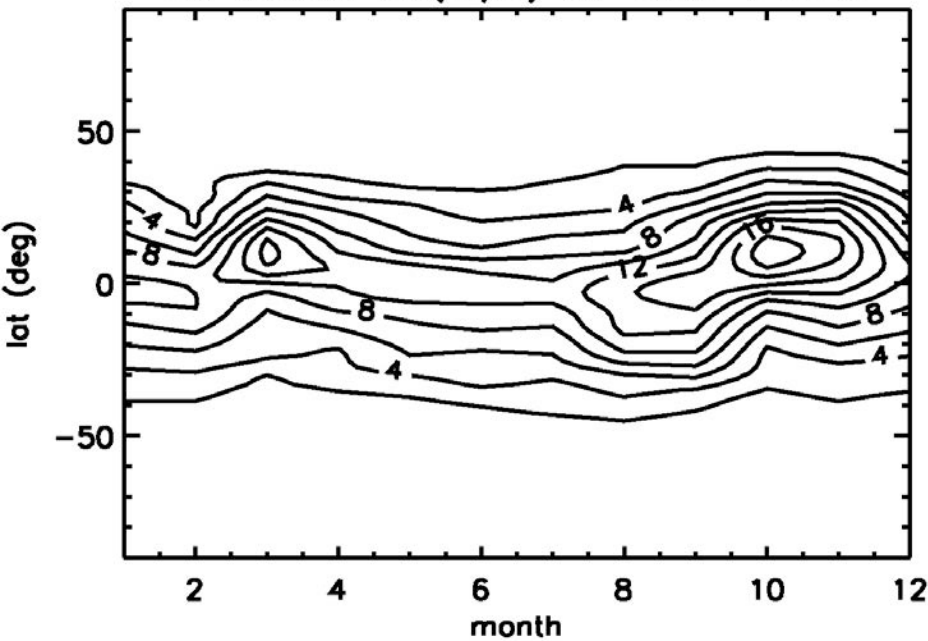
MDT UN(m/s) at $8.8E-04hPa$

Nonmigrating: Diurnal E3

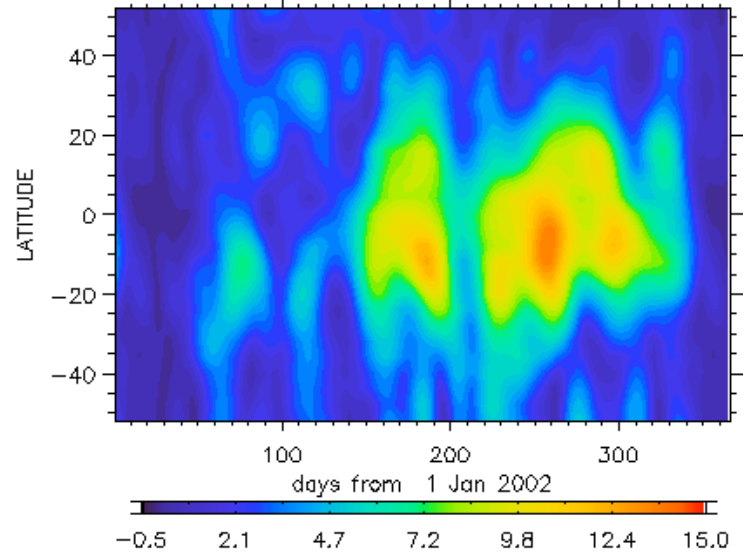
E3 Diurnal T(K) at $1.2E-04\text{hPa}$



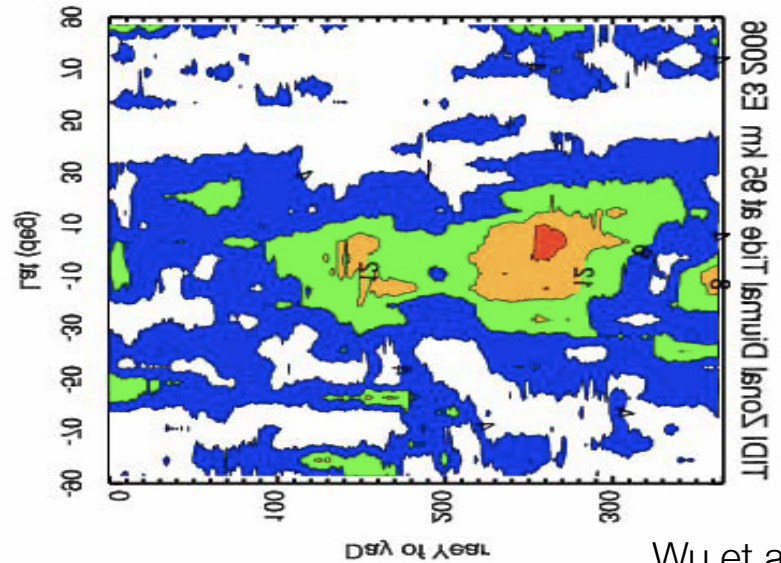
E3 Diurnal UN(m/s) at $8.8E-04\text{hPa}$



$m=3, z=16.1\text{ sh}, f=(-1.030, -0.972), \text{ctr}=0.50$
RMS AMP 1 Jan - 31 Dec 2002

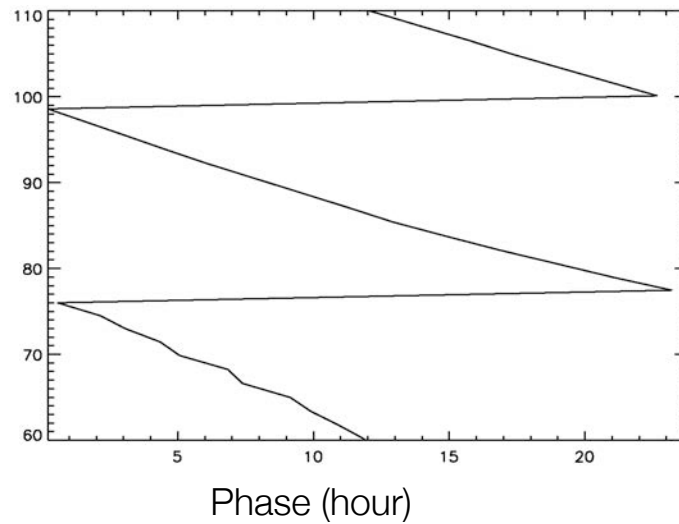
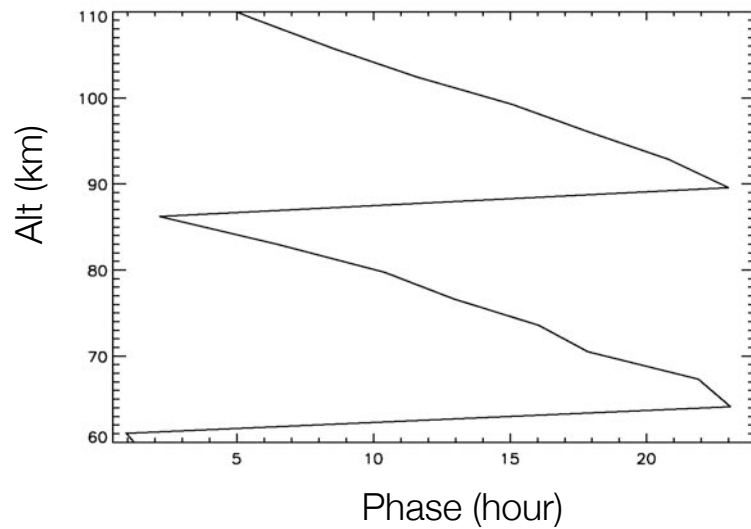
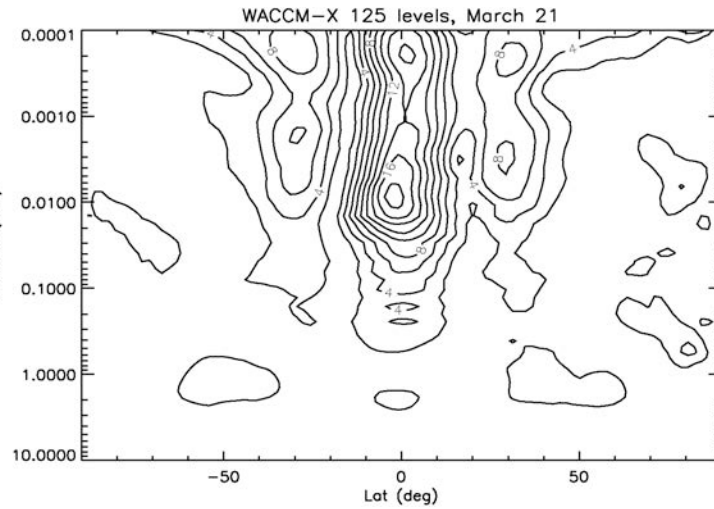
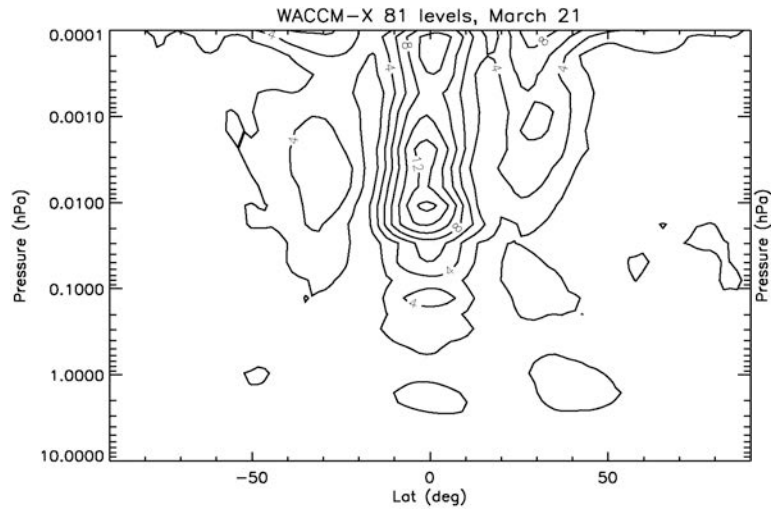


Garcia and Lieberman, 2007

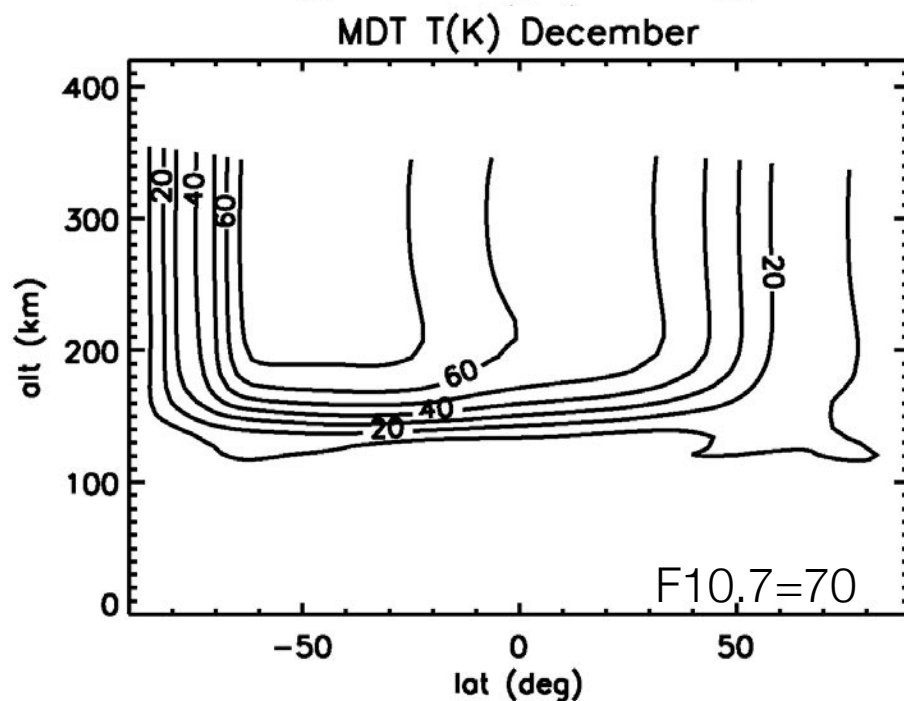
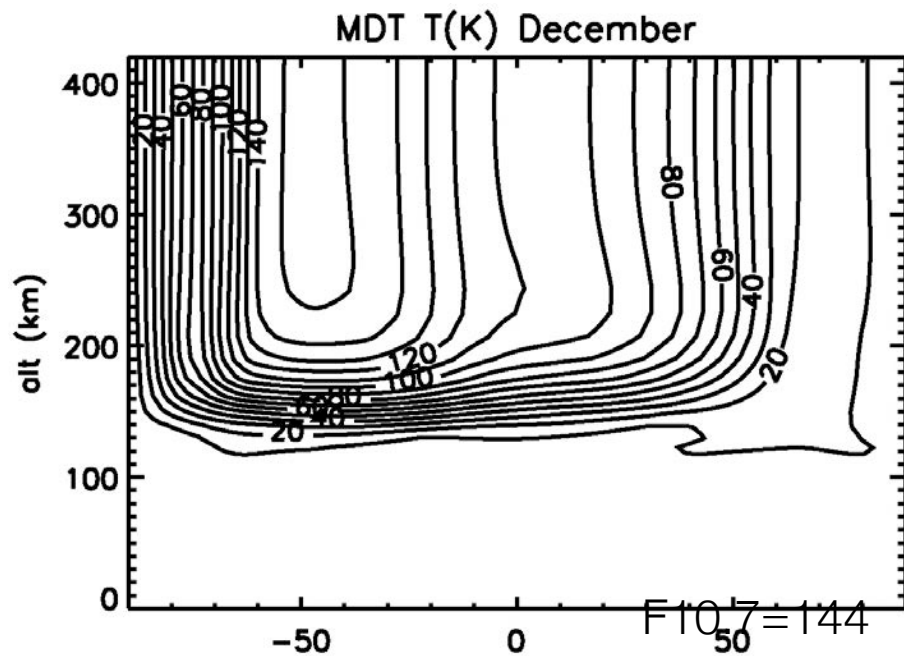
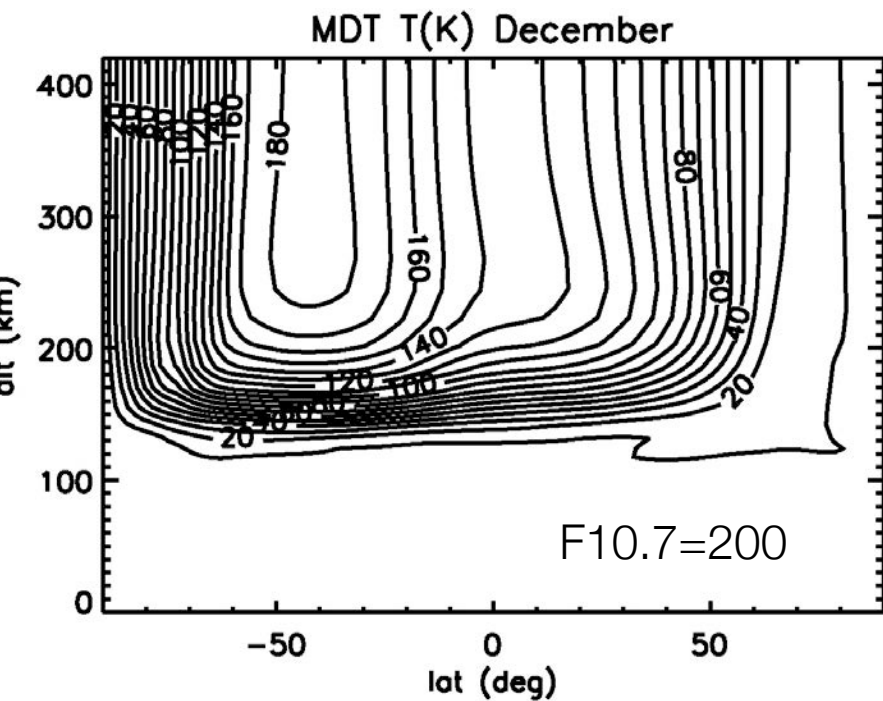


Wu et al, 2008b

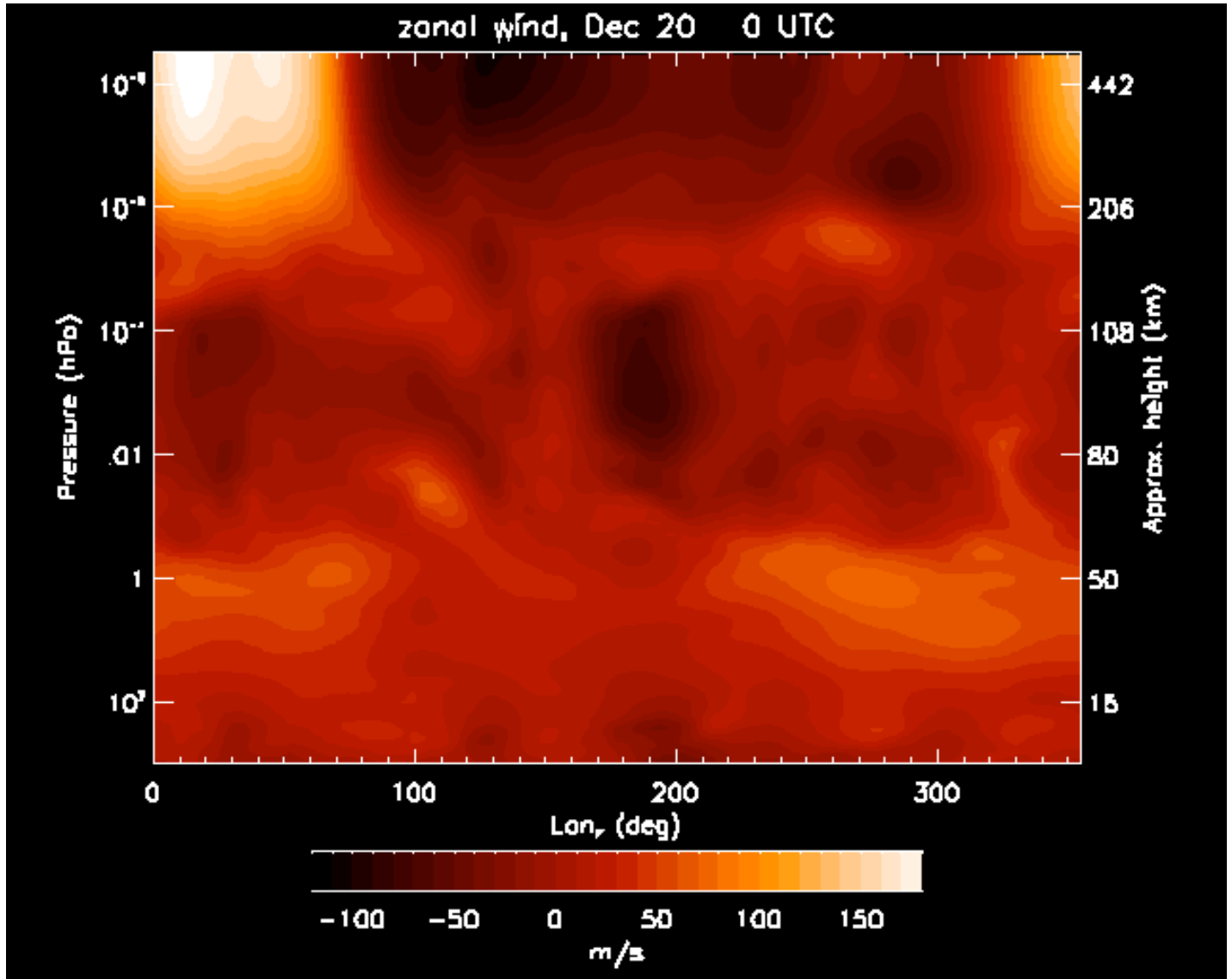
81 vs. 125 levels



Thermospheric migrating diurnal tide



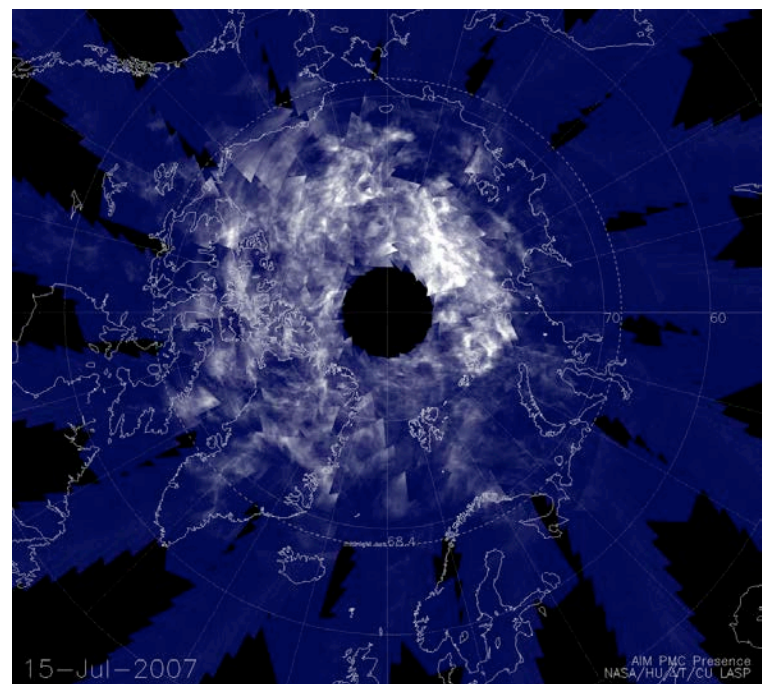
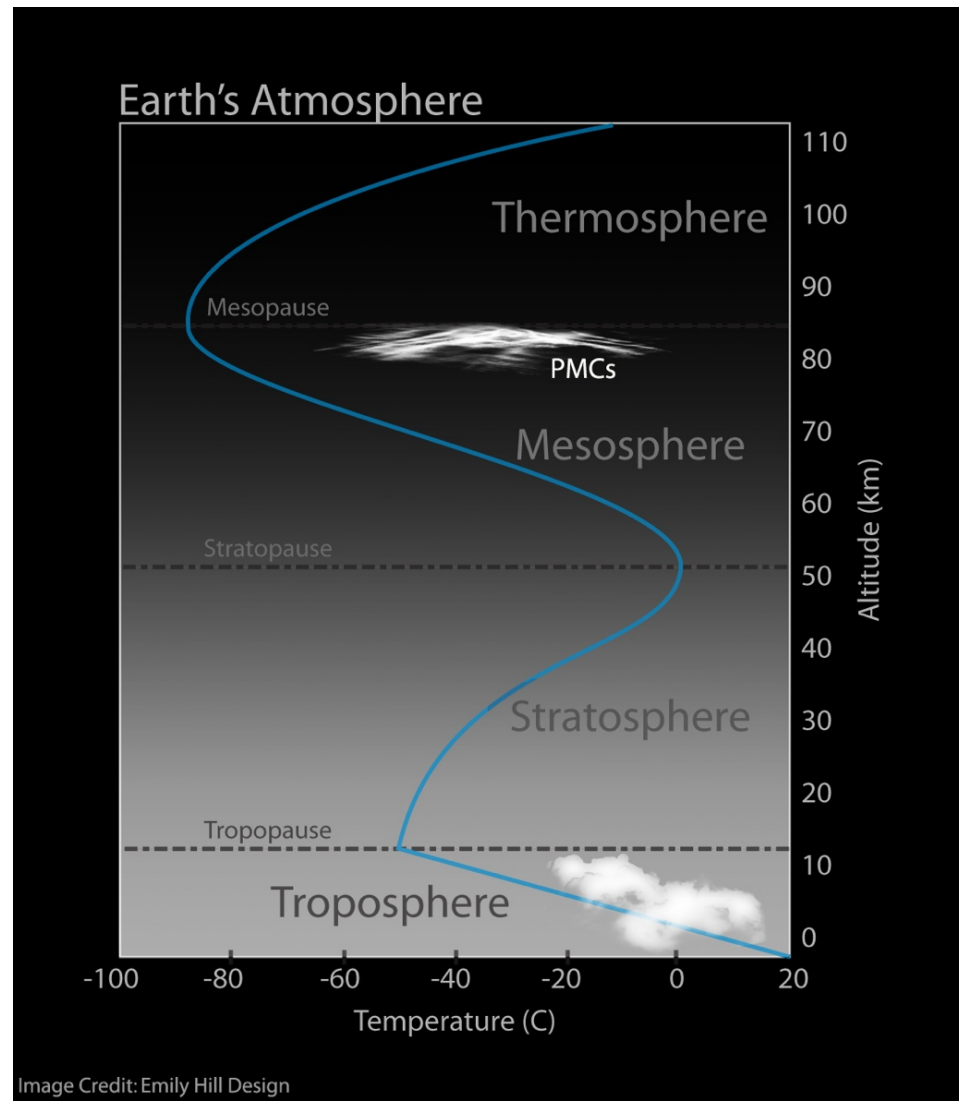
Short-term Variability



Polar mesospheric clouds

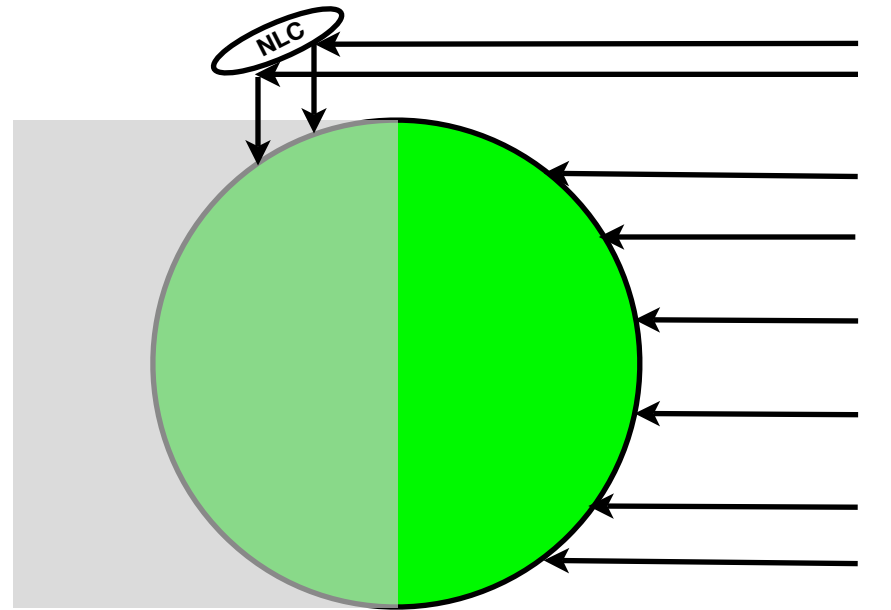
with A. Merkel

PMCs are water ice clouds near the mesopause (~83 km)



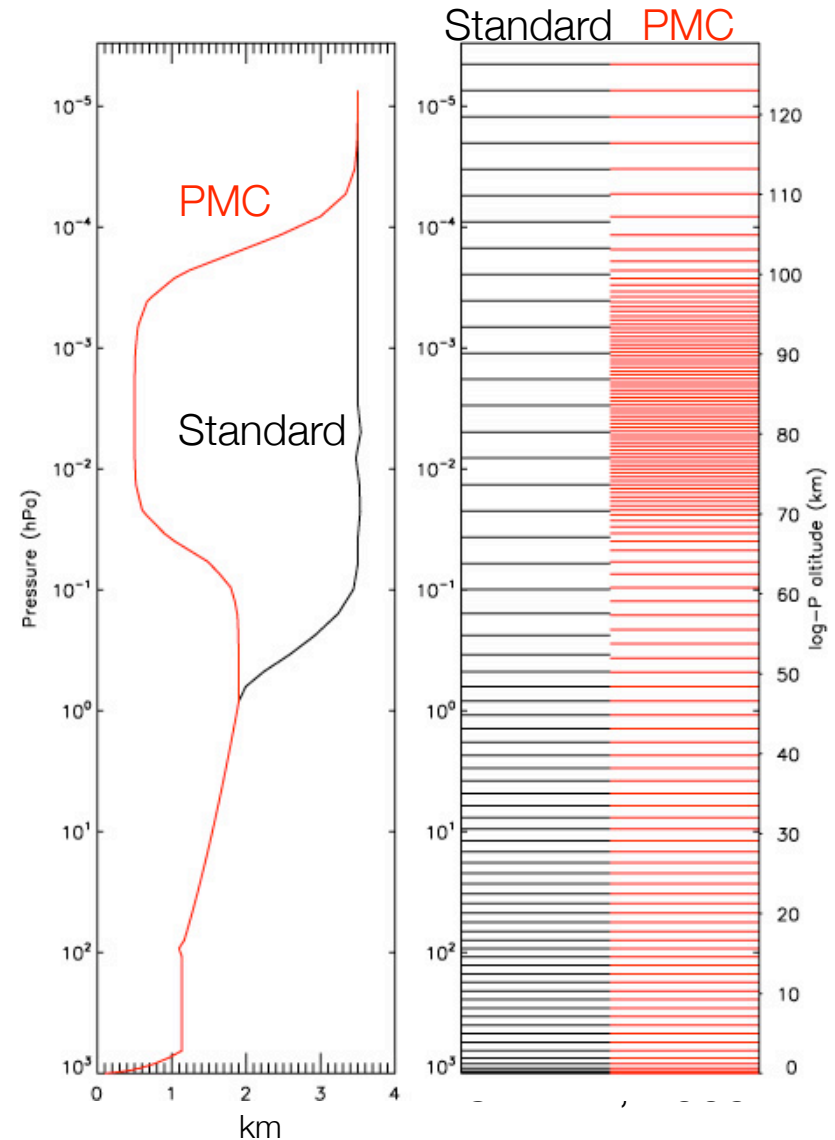


Notilucent clouds (NLC) when viewed from the ground

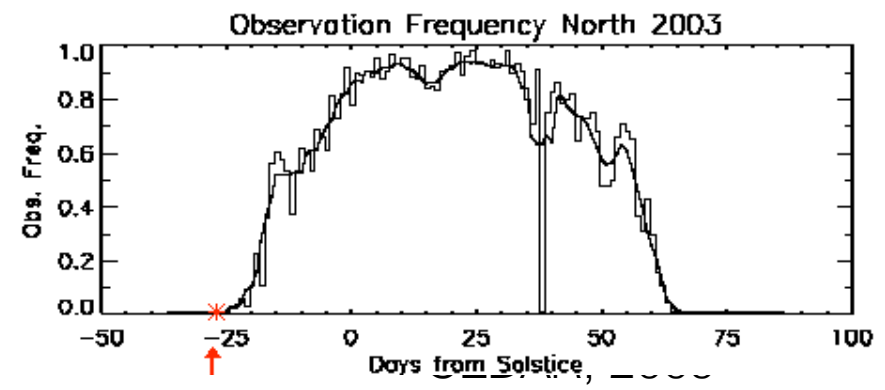
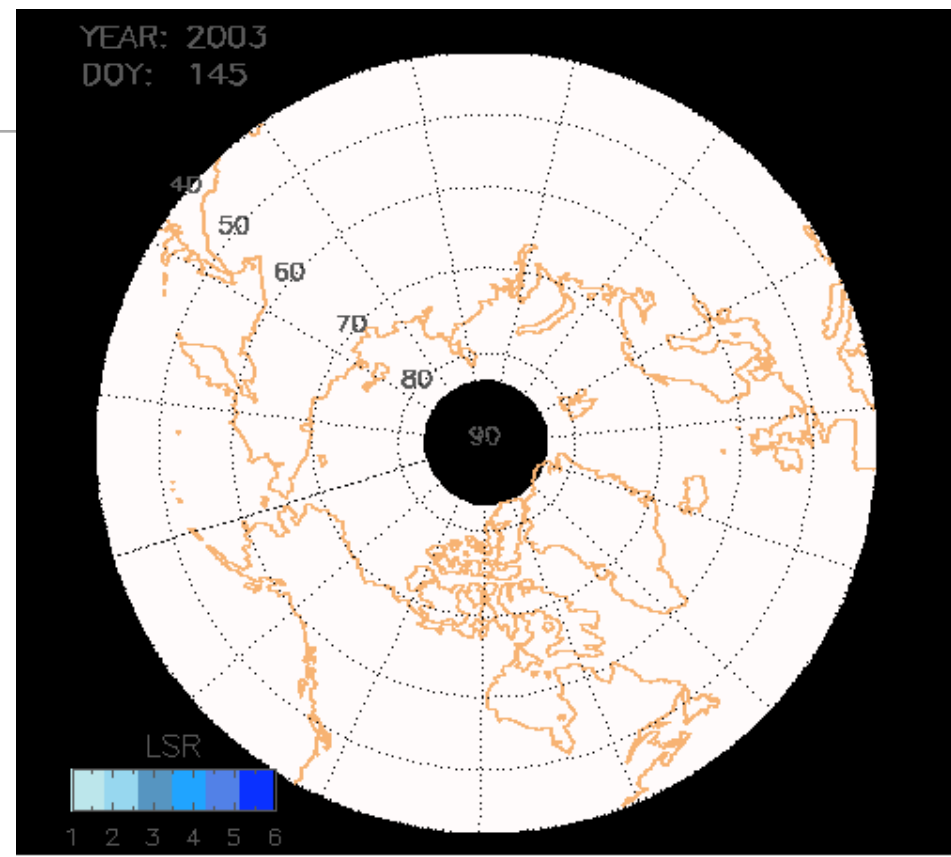
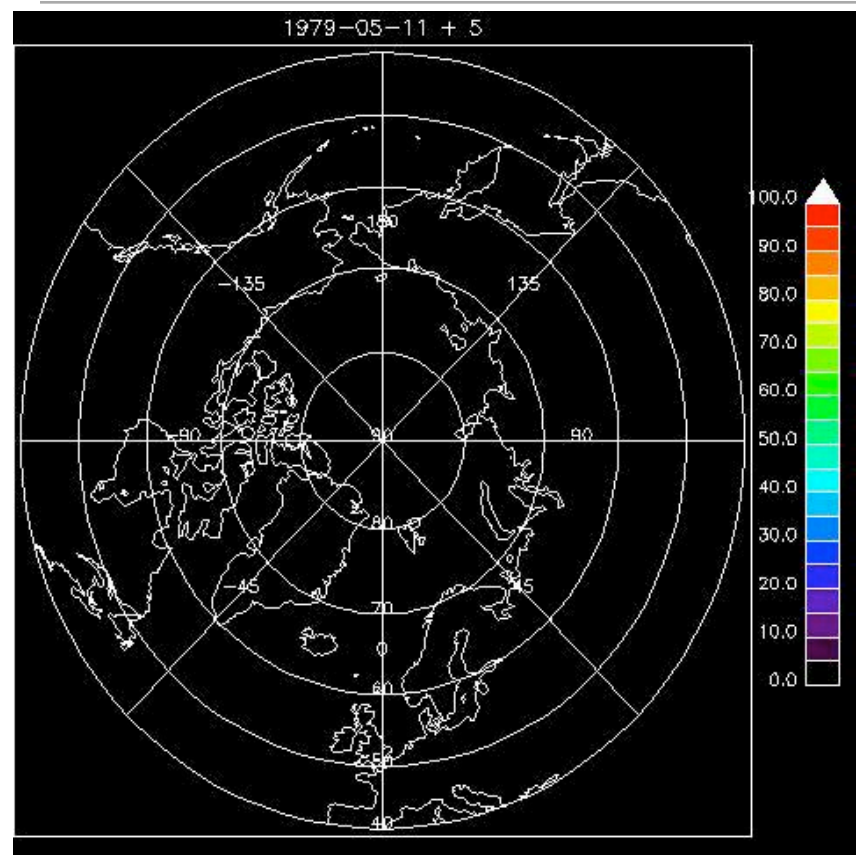


WACCM-3 with PMC parameterization

- 124 levels with 0.5 km resolution near the mesopause
- macroscale PMC parameterization of ice nucleation, growth, and sublimation
- Ice diffused, advected and sedimented
- Intended for long integrations (solar cycle, trends, etc.)
- Depends on local parameters (i.e. not particle-following)
- ~16 wall clock hrs. / year on 192 processors

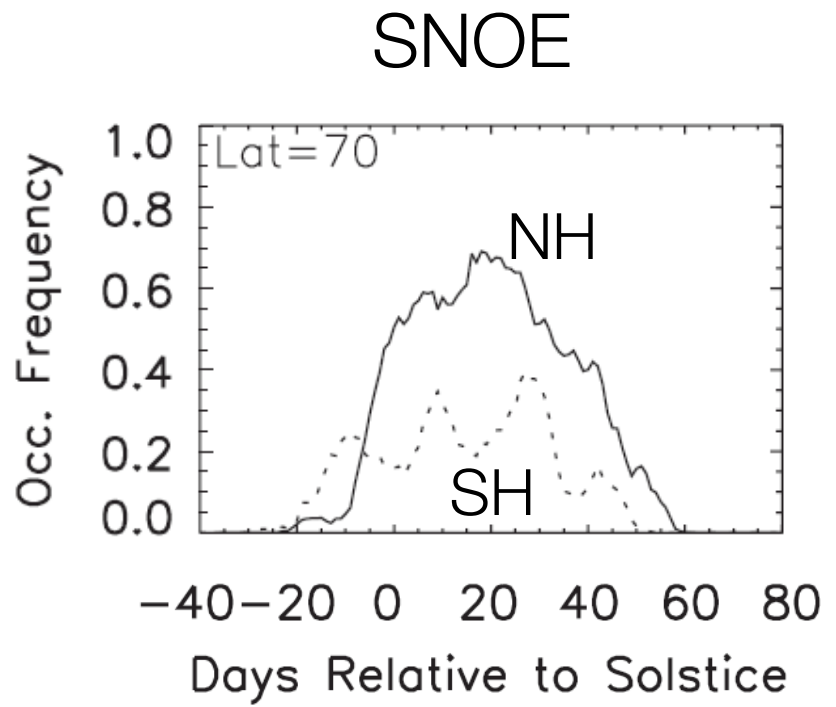
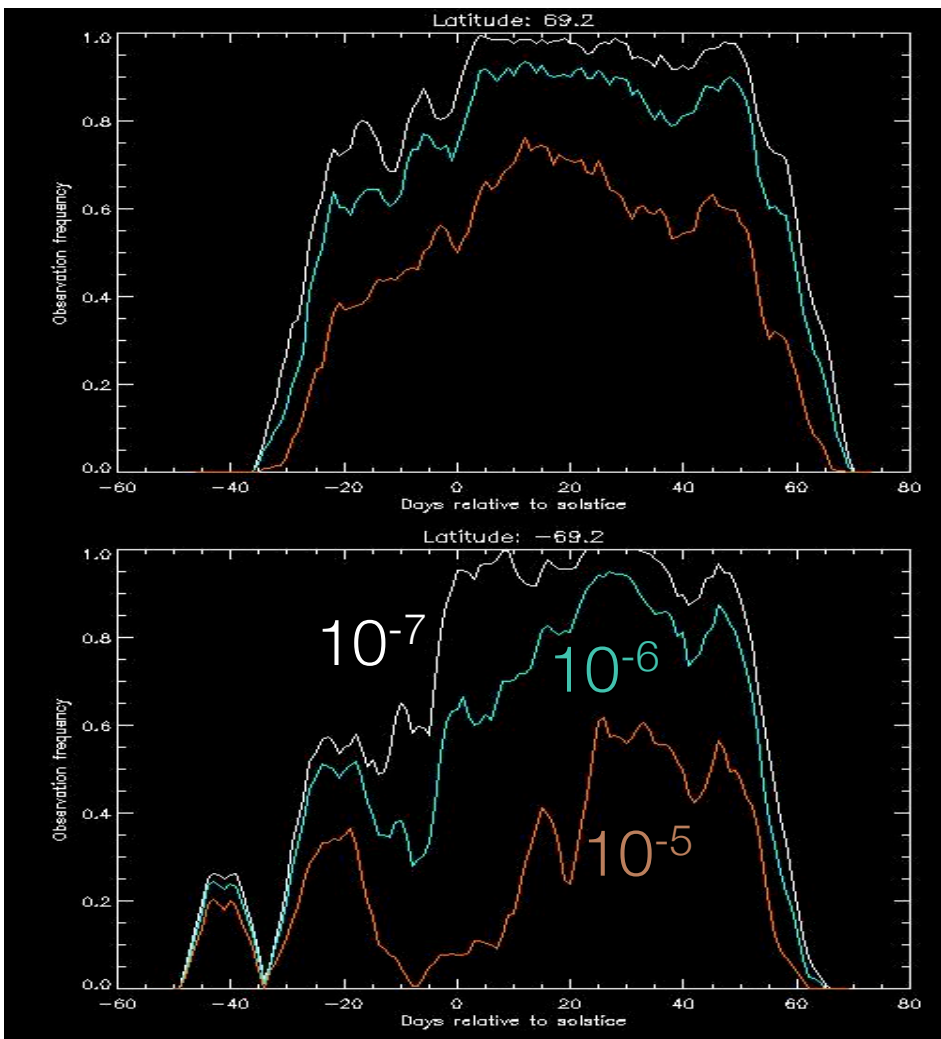


WACCM & SNOE albedo



- Assume Gaussian distrib. width of 13 nm
- Backscatter ratio calculated for 90°

Occurrence frequency at 70°

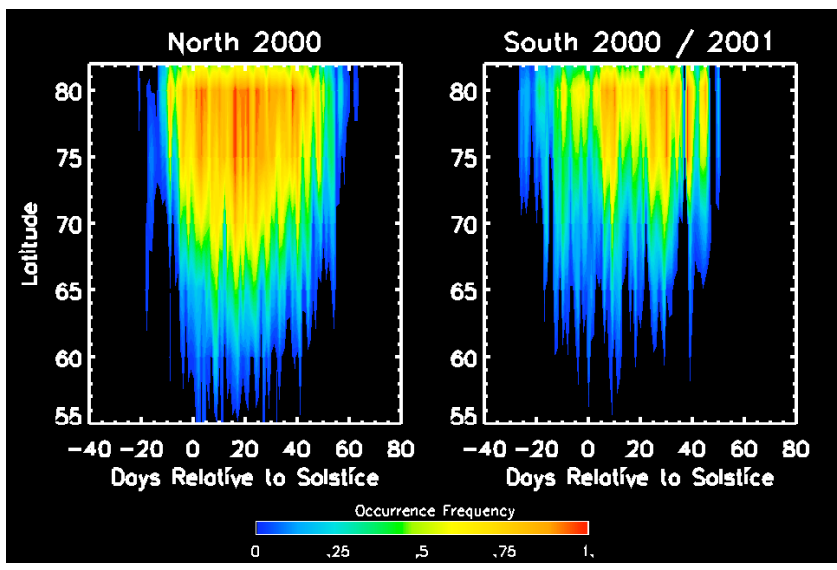


Bailey et al., 2007

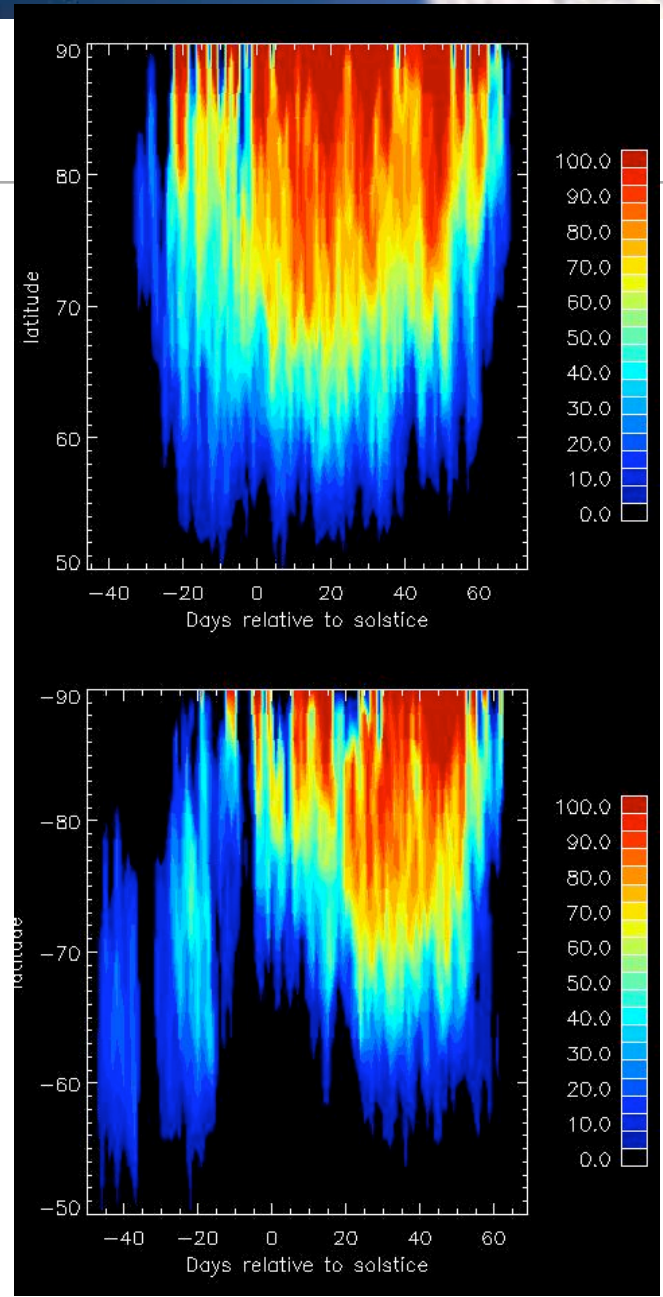
CEDAR, 2009

PMC observation frequency

SNOE



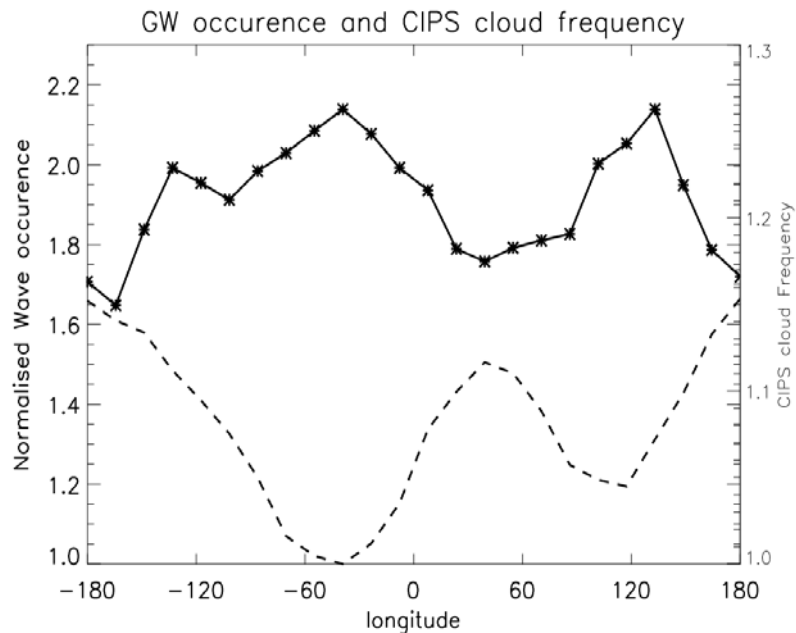
albedo > 10^{-5}



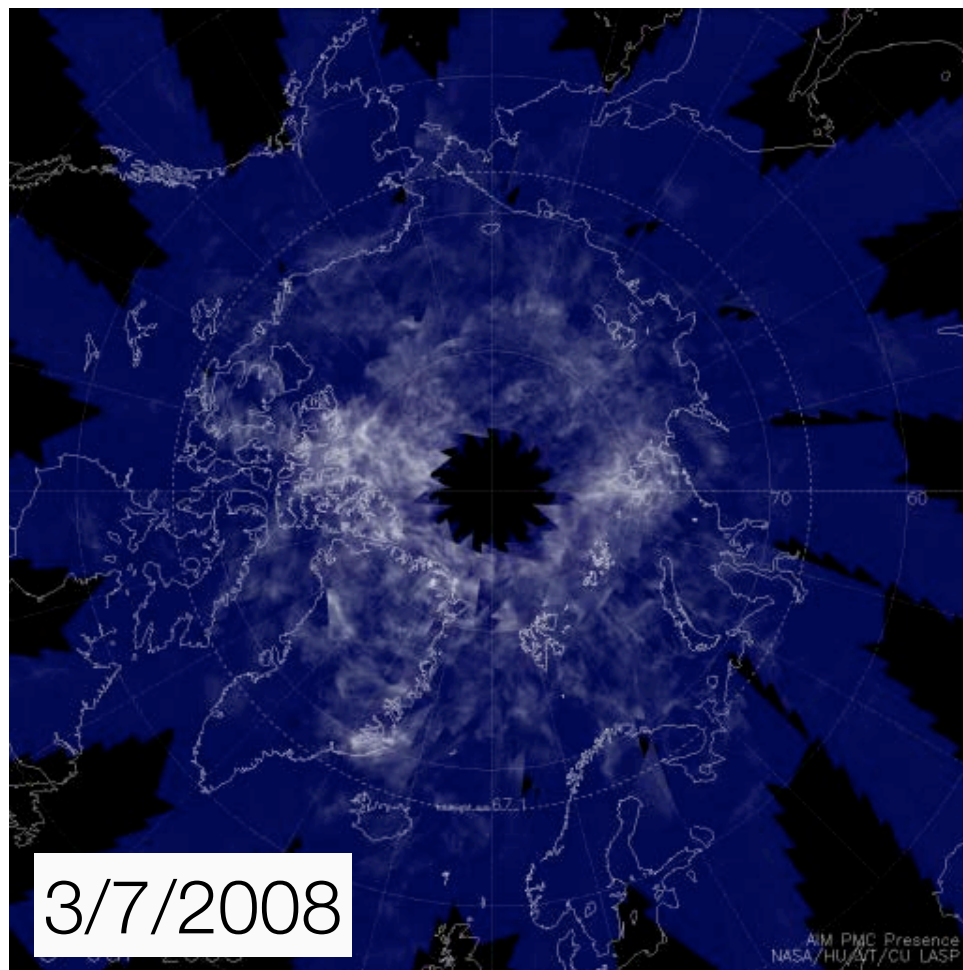
Can solar tides produce zonal asymmetry?

CIPS cloud frequency

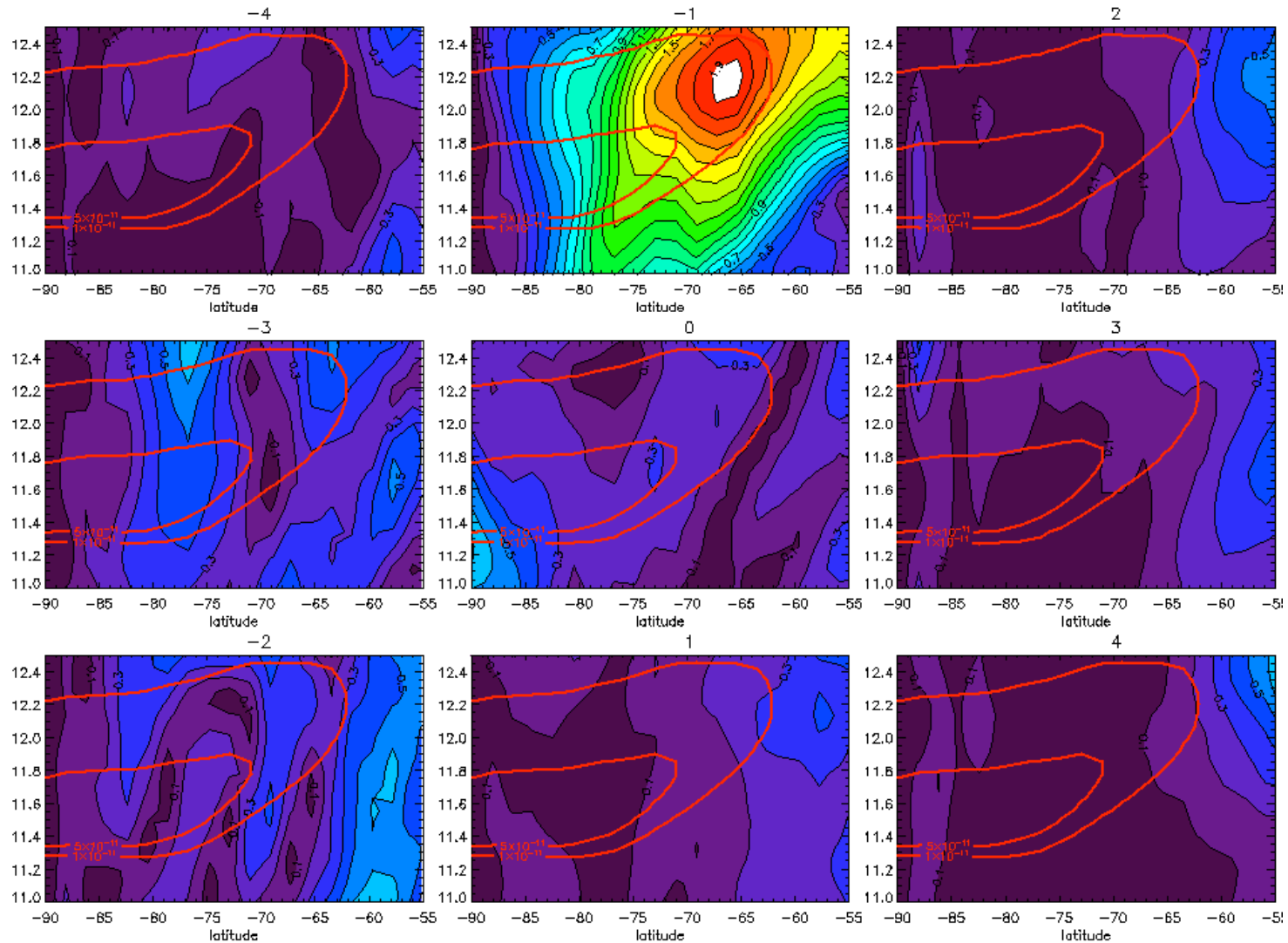
NH Days 172-195, 70-80 N



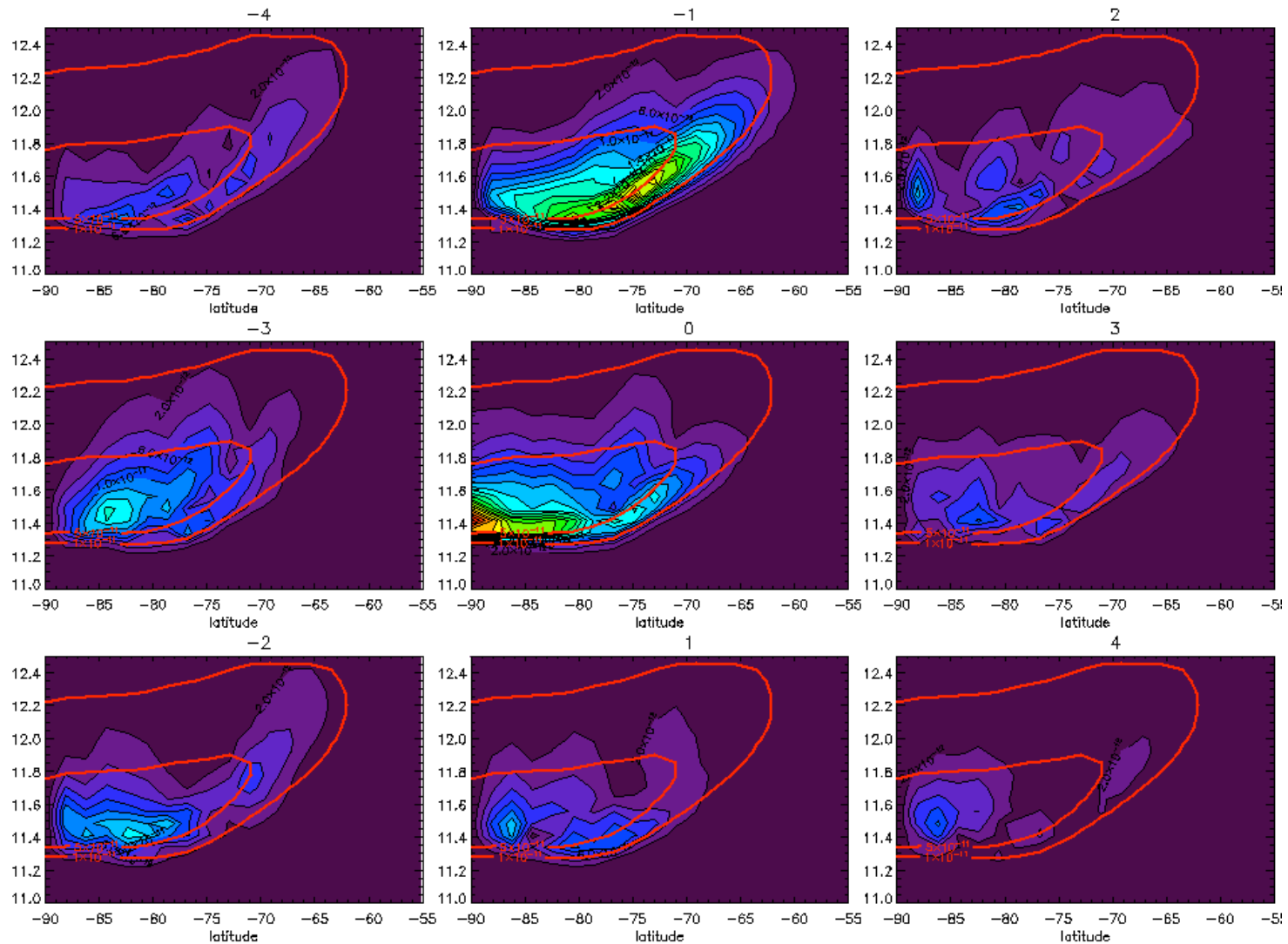
Analysis by Amal Chandran, CU/LASP



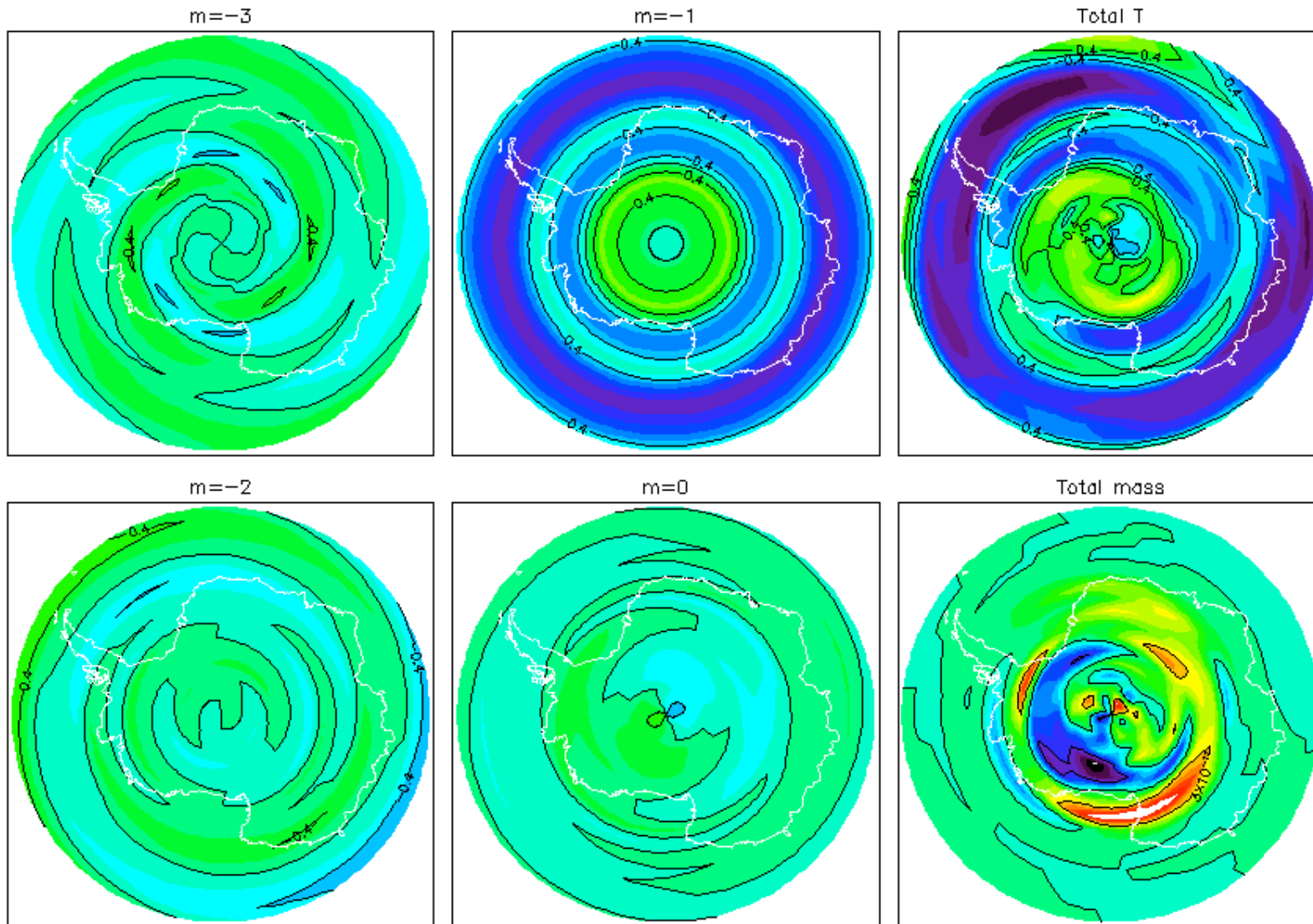
WACCM T tidal amplitudes & mean ice mass



WACCM ice mass tidal amplitudes & mean



Souther hemisphere tidal temperature viewed from sun-synchronous satellite at 12:00 hrs. local time



@11.5
ScaleHts.
~81km

Ice mass



NCAR



WACCM

*Whole Atmosphere
Community Climate Model*

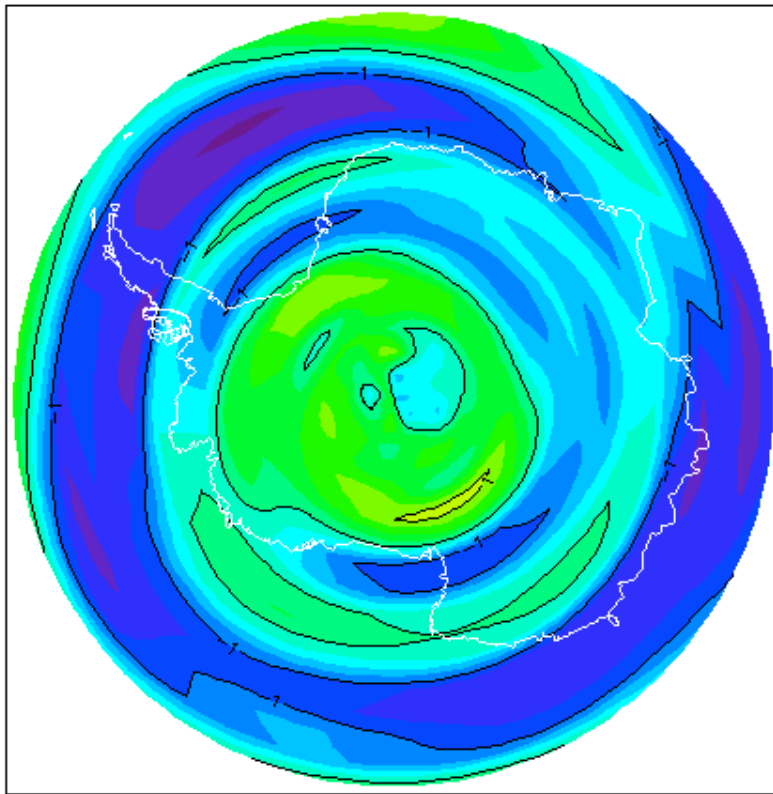


8 x 5-day mean sequence of T and ice mass

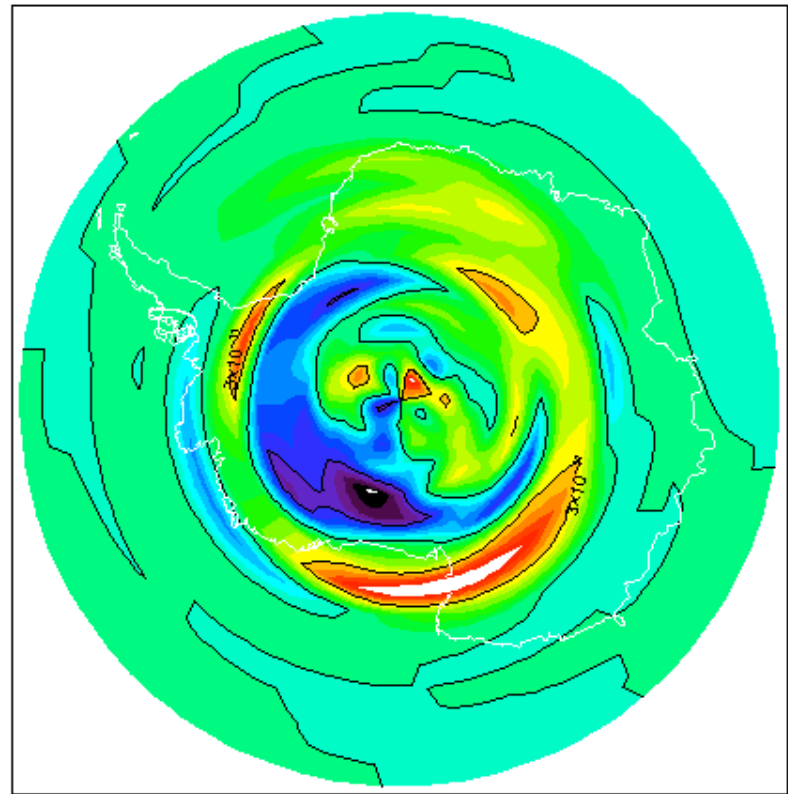
Tidal amplitudes

Southern Hemisphere Jan 11-15 @ 11.5 scale hts.

Temperature



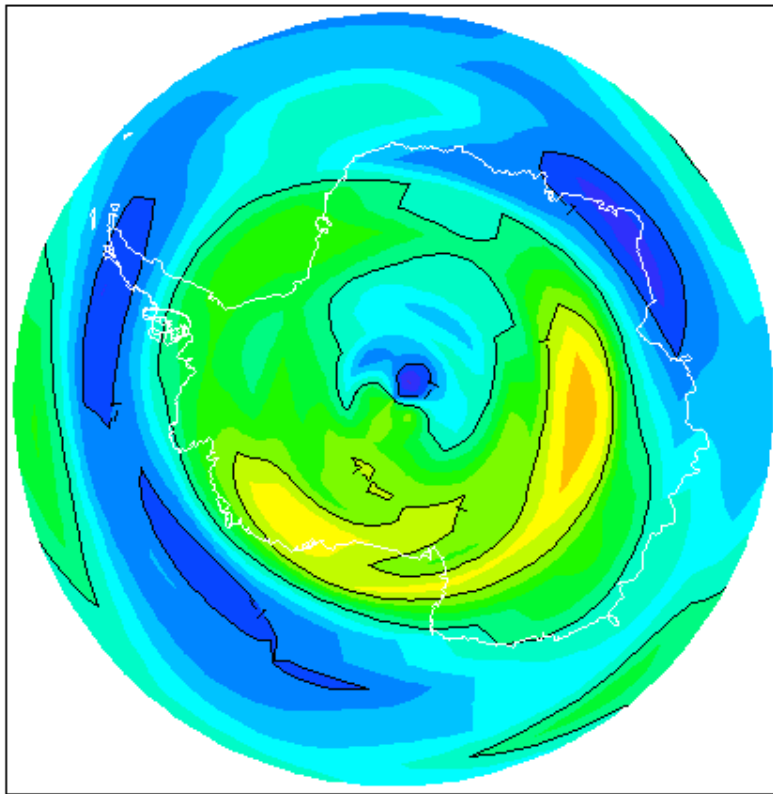
Ice Mass



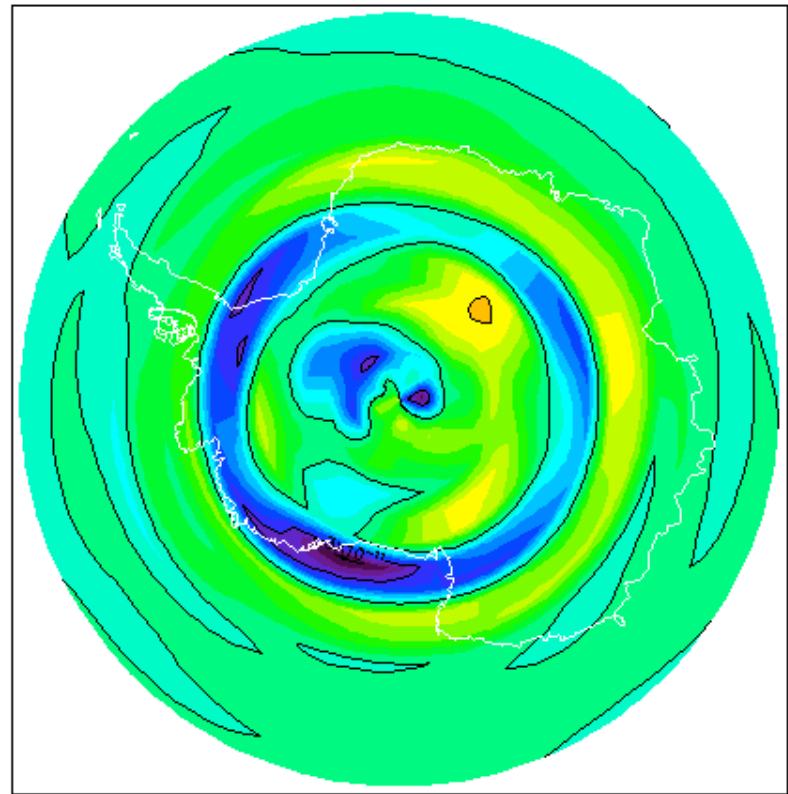
Tidal amplitudes

Southern Hemisphere Jan 16-20 @ 11.5 scale hts.

Temperature



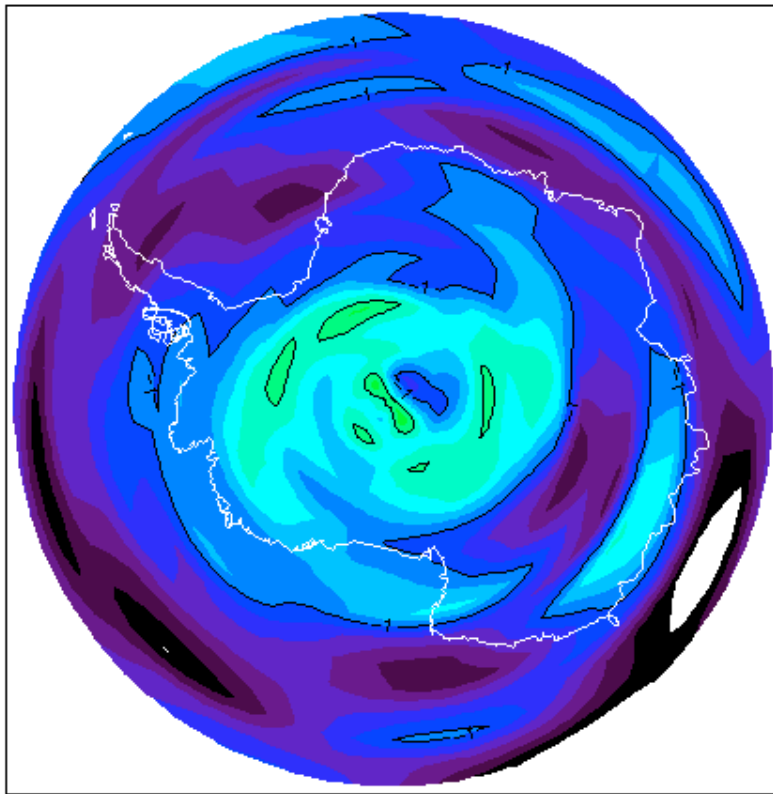
Ice Mass



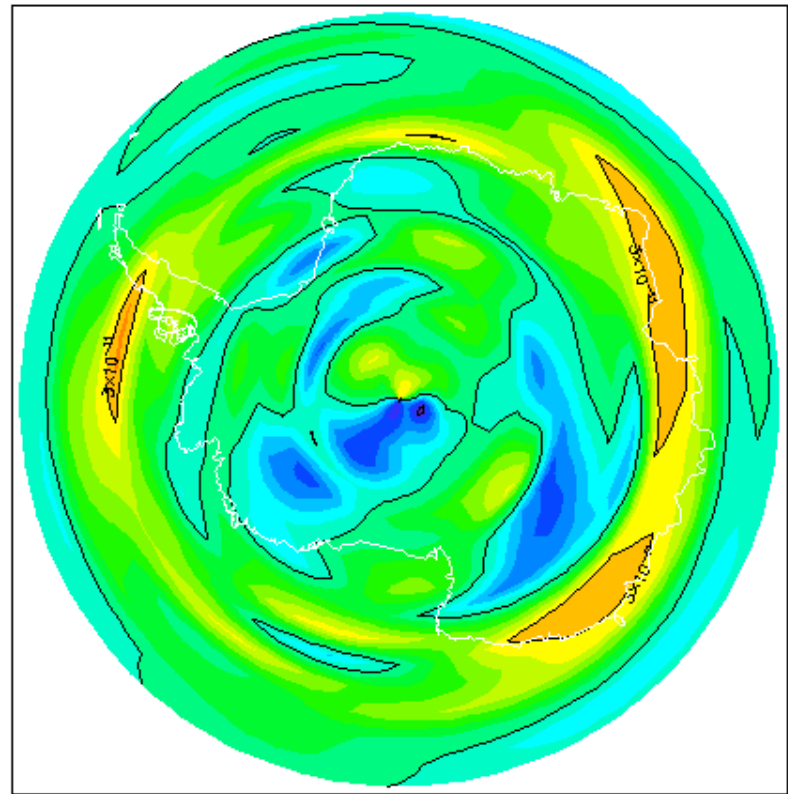
Tidal amplitudes

Southern Hemisphere Jan 21-25 @ 11.5 scale hts.

Temperature



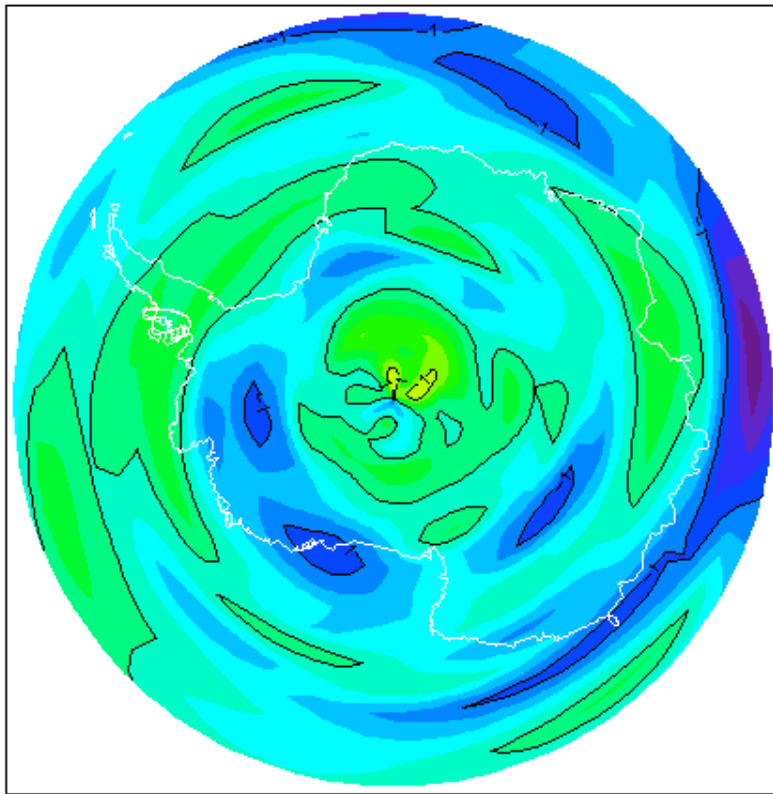
Ice Mass



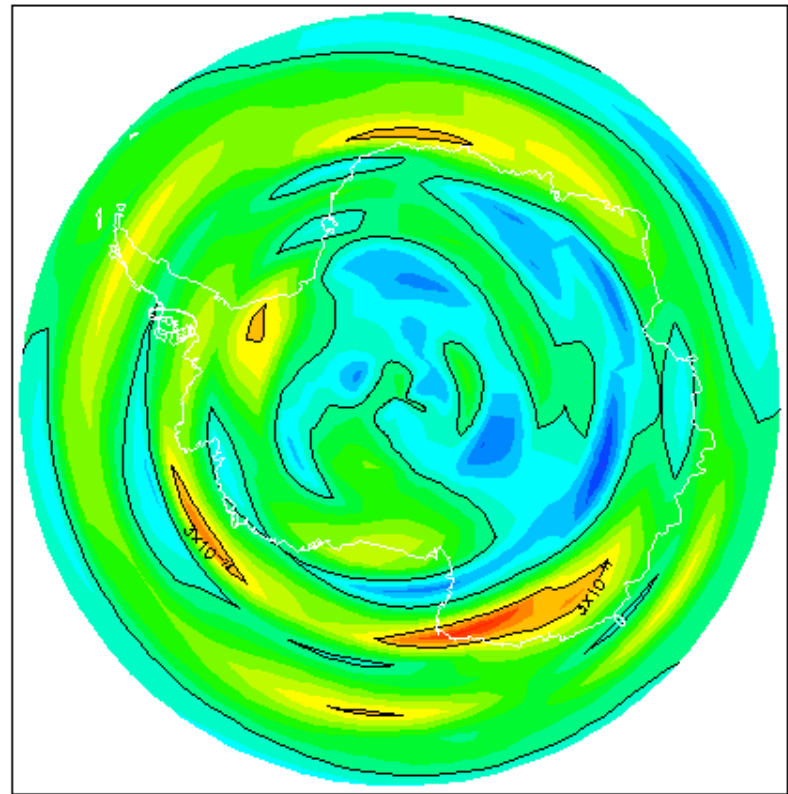
Tidal amplitudes

Southern Hemisphere Jan 26-30 @ 11.5 scale hts.

Temperature



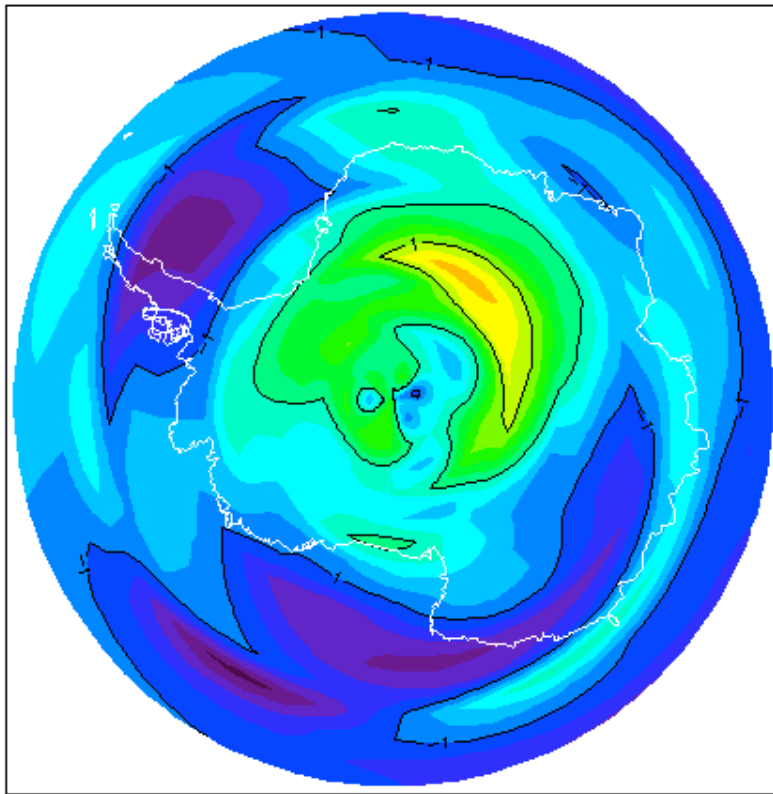
Ice Mass



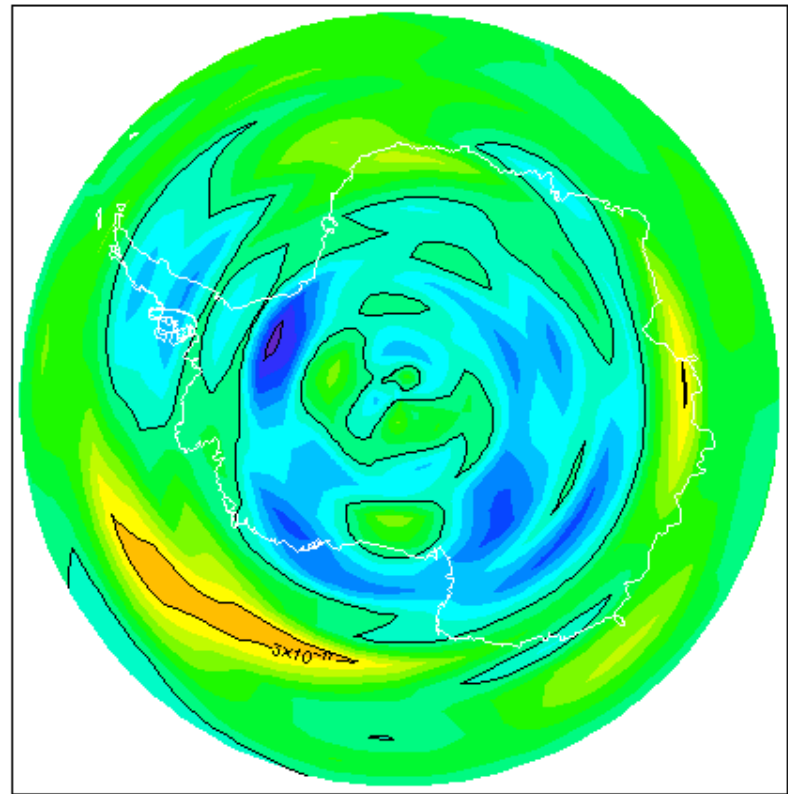
Tidal amplitudes

Southern Hemisphere J31- F04 @ 11.5 scale hts.

Temperature



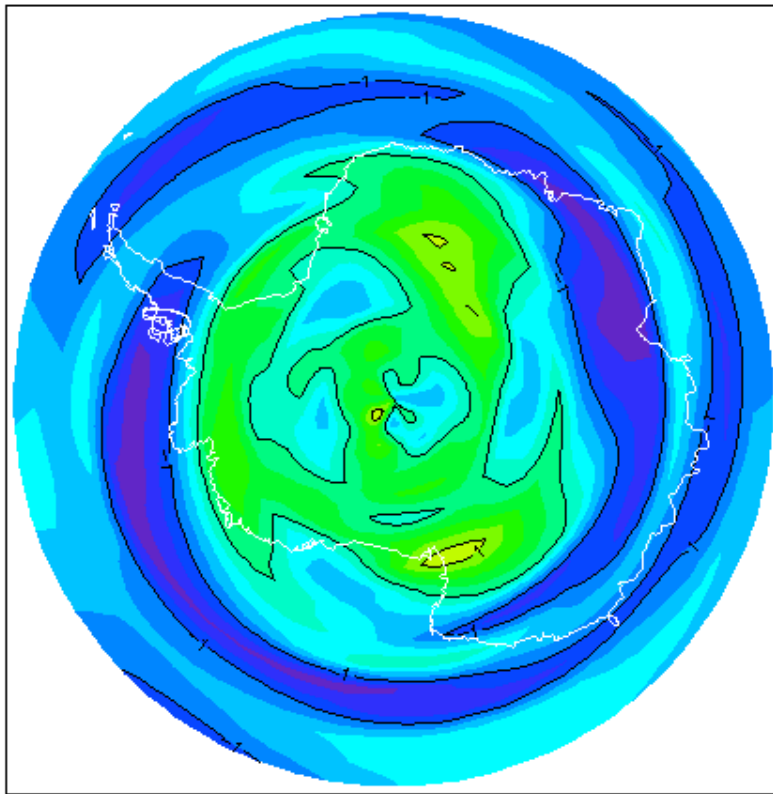
Ice Mass



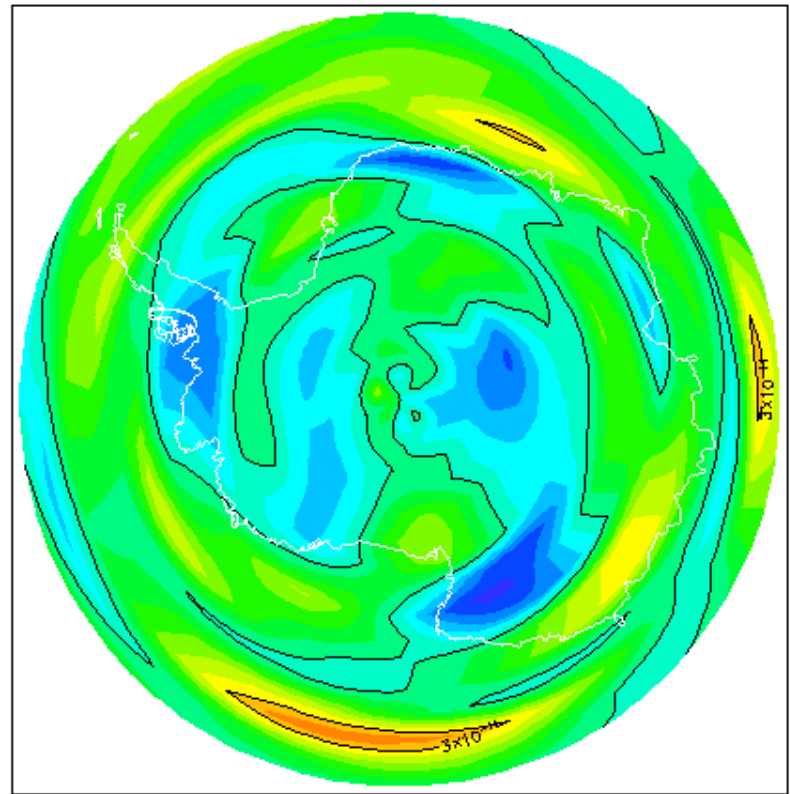
Tidal amplitudes

Southern Hemisphere Feb 05-09 @ 11.5 scale hts.

Temperature



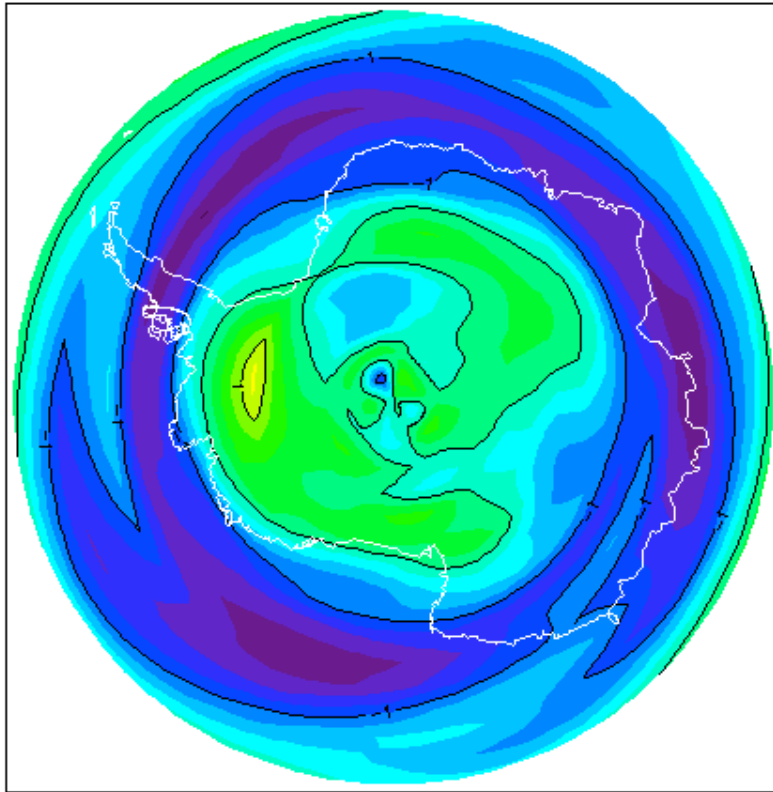
Ice Mass



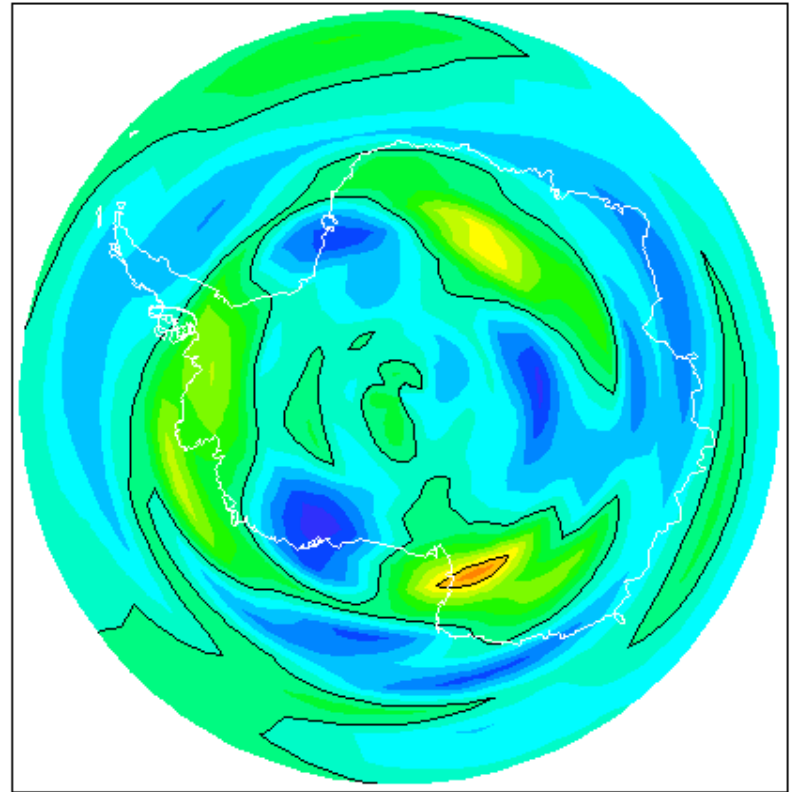
Tidal amplitudes

Southern Hemisphere Feb 10-14 @ 11.5 scale hts.

Temperature



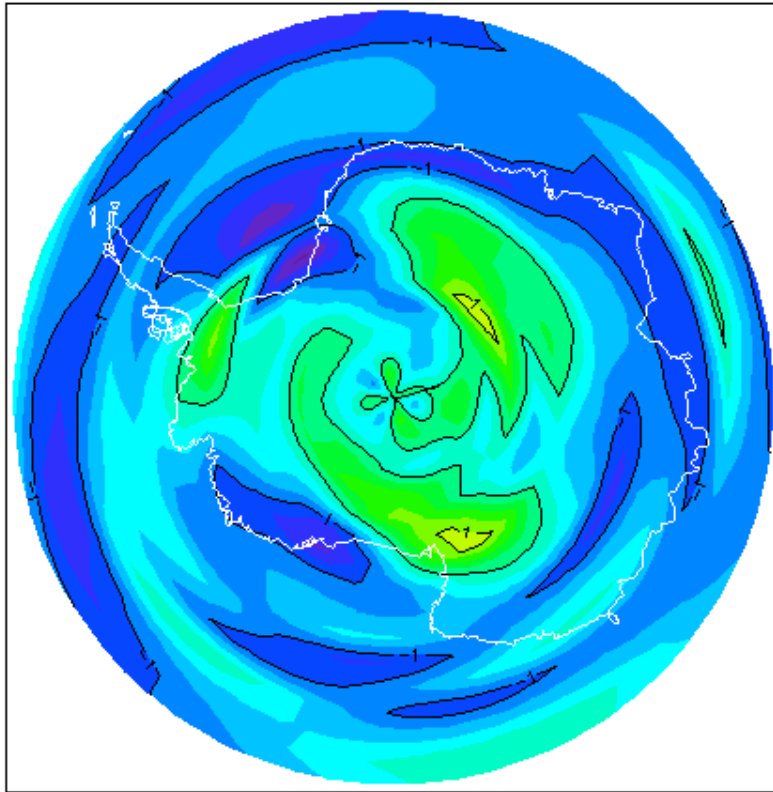
Ice Mass



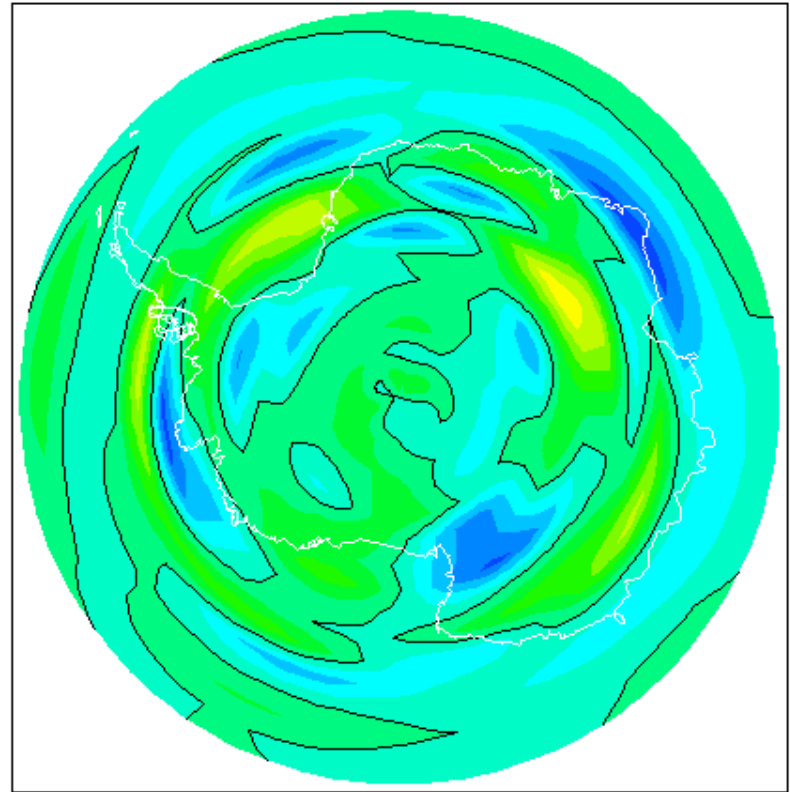
Tidal amplitudes

Southern Hemisphere Feb 15-19 @ 11.5 scale hts.

Temperature

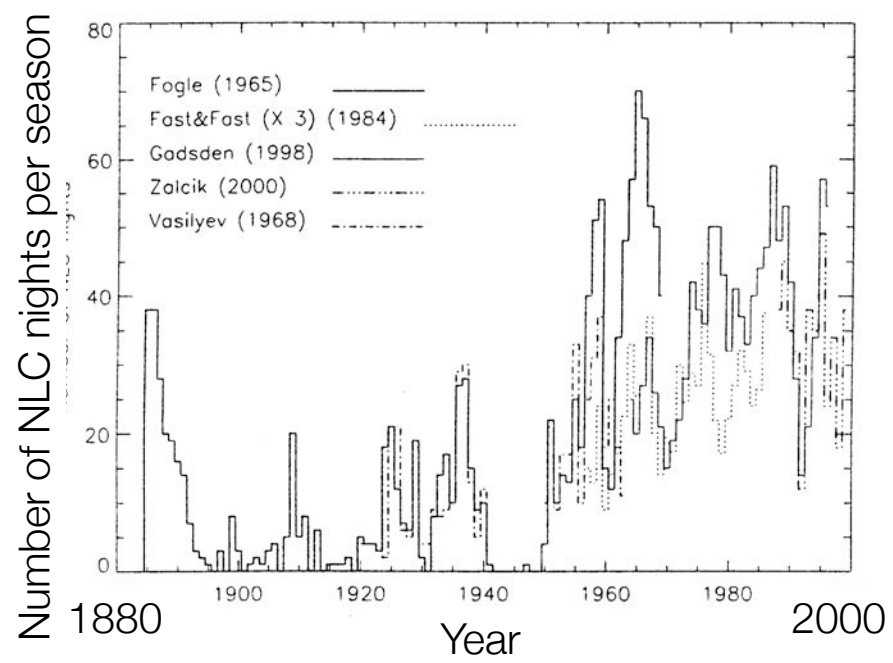


Ice Mass

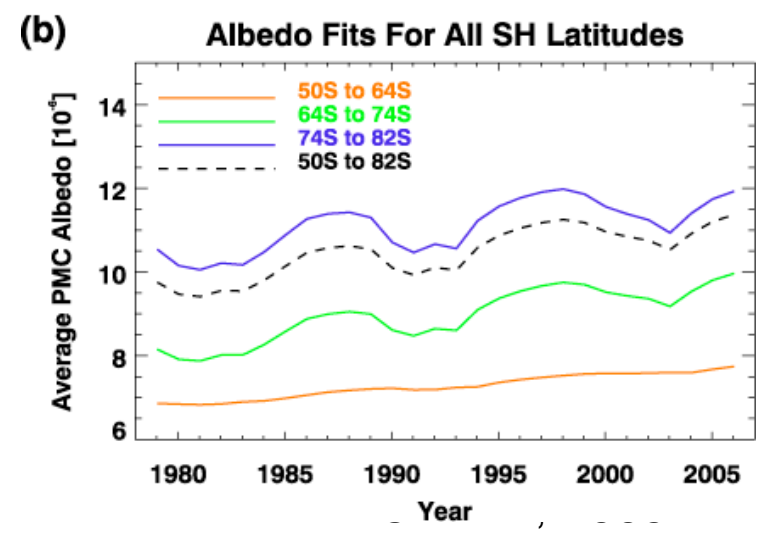
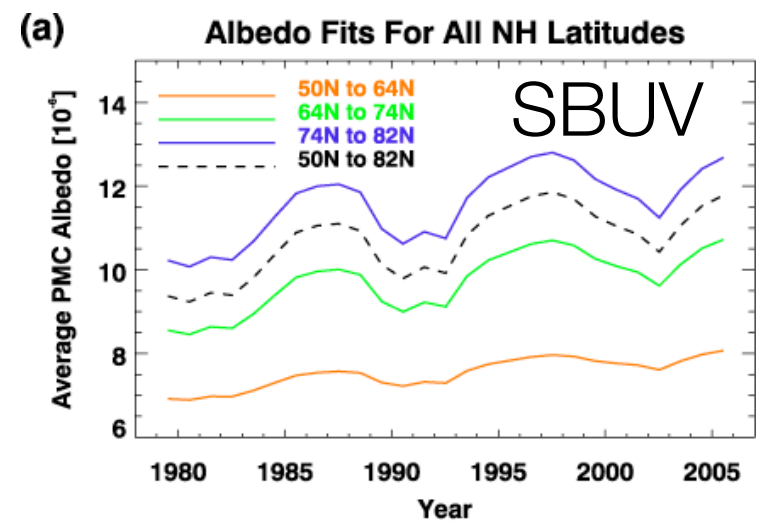


PMC frequency & brightness are increasing

Ground based observations



Thomas and Olivero, 2001

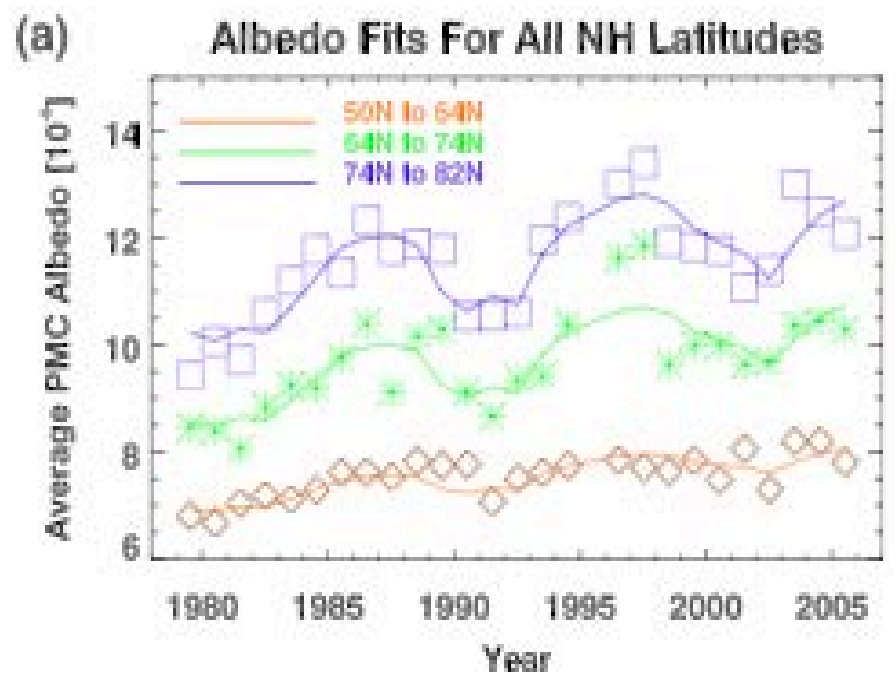


WACCM albedo calculations

- Albedo calculated at 252 nm a 120° scattering angle
- Albedo above detectable threshold calculated over each season
- Season defined as SBUV: -30 to +70 DFS.
- Threshold dependent on latitude (DeLand):
 - 50N to 64N = 5.5G Threshold
 - 64N to 74N = 6.5G threshold
 - 74N to 82N = 7.5G threshold

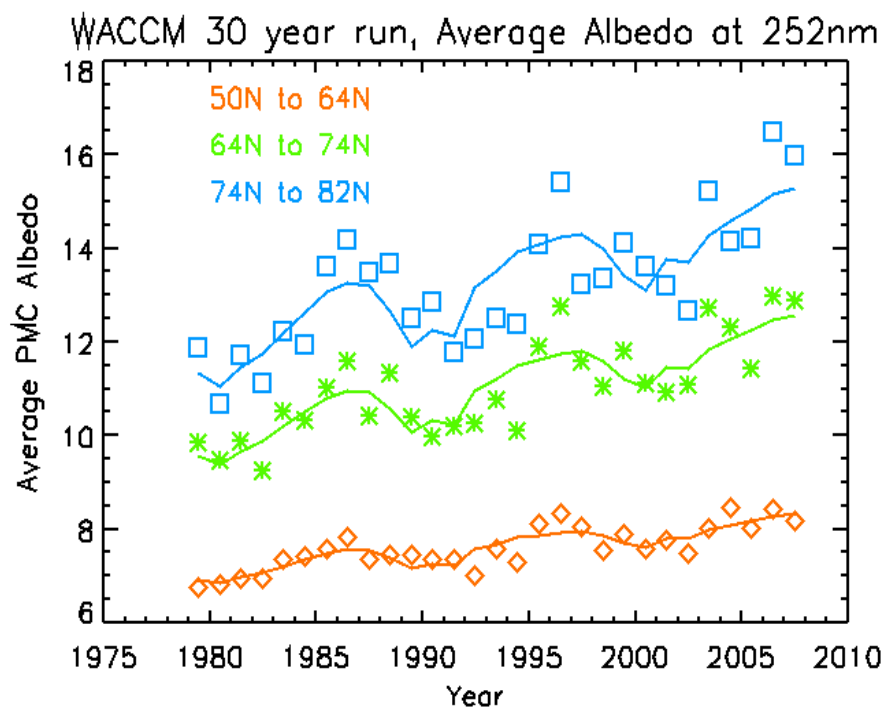
Northern hemisphere

SBUV



DeLand et al. 2007

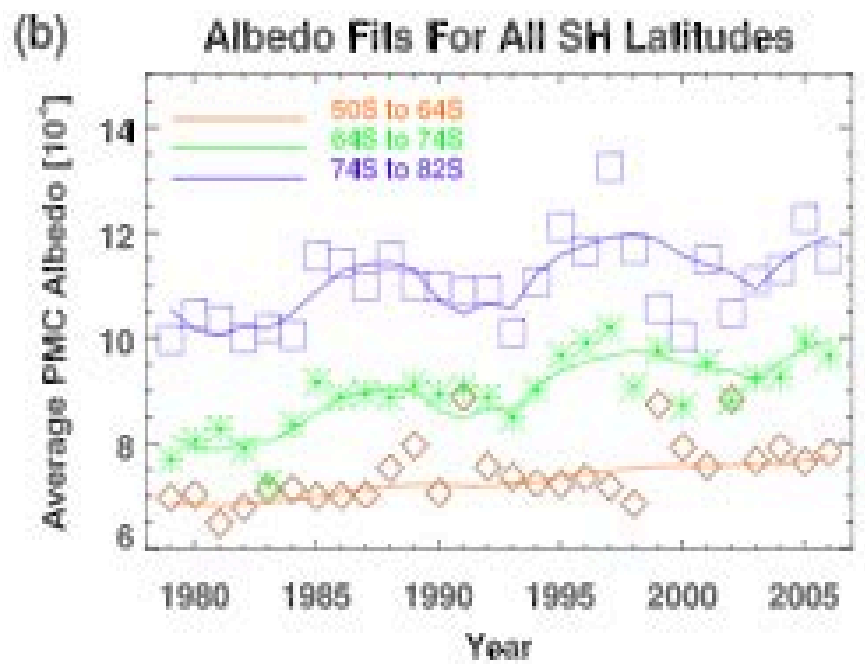
WACCM



CEDAR, 2009

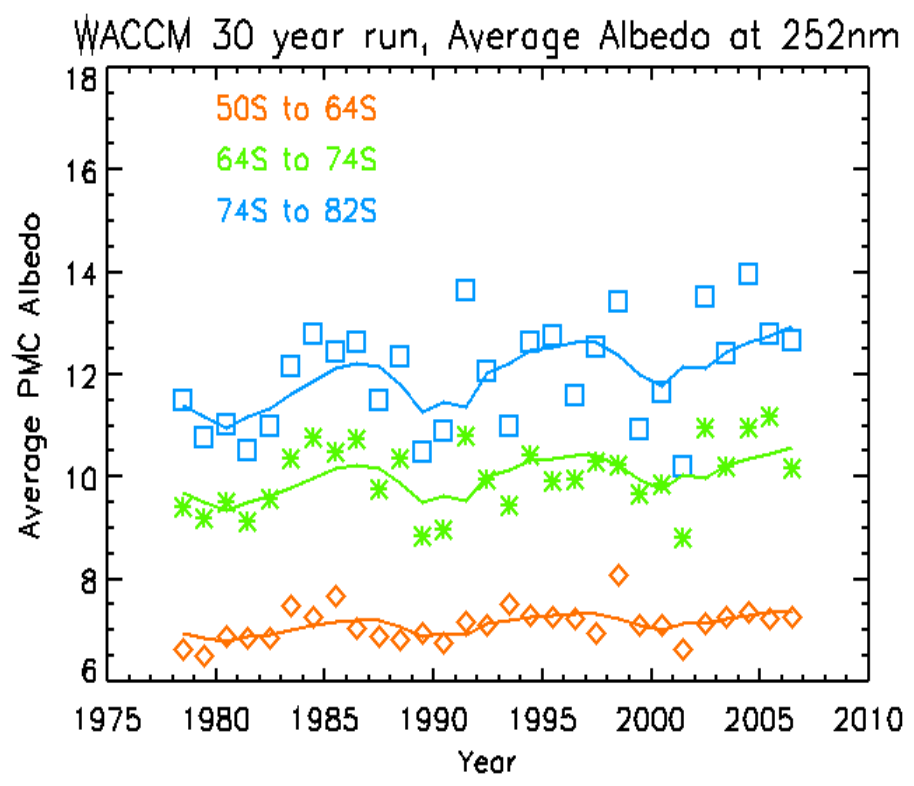
Southern hemisphere

SBUV



DeLand et al. 2007

WACCM



CEDAR, 2009

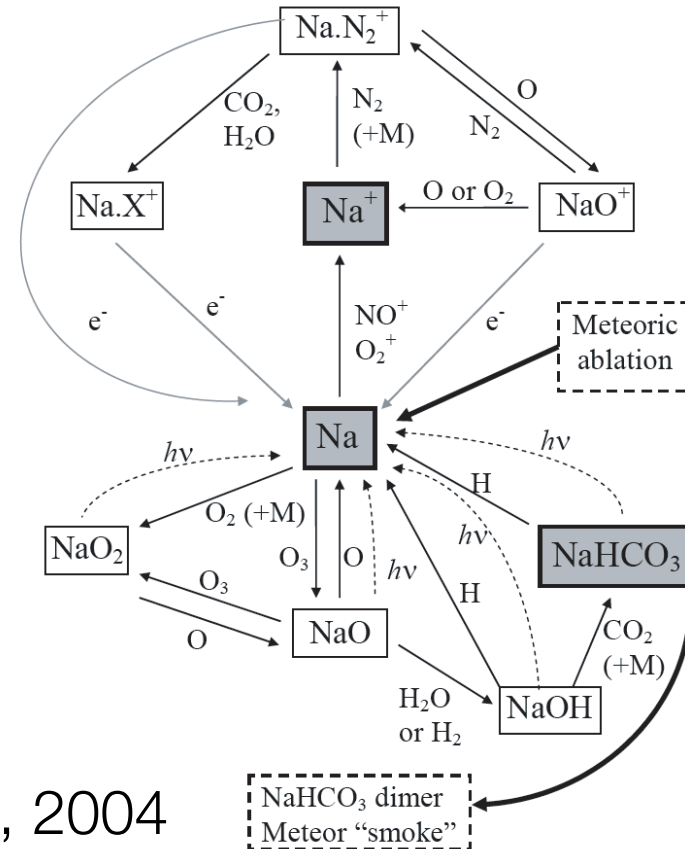
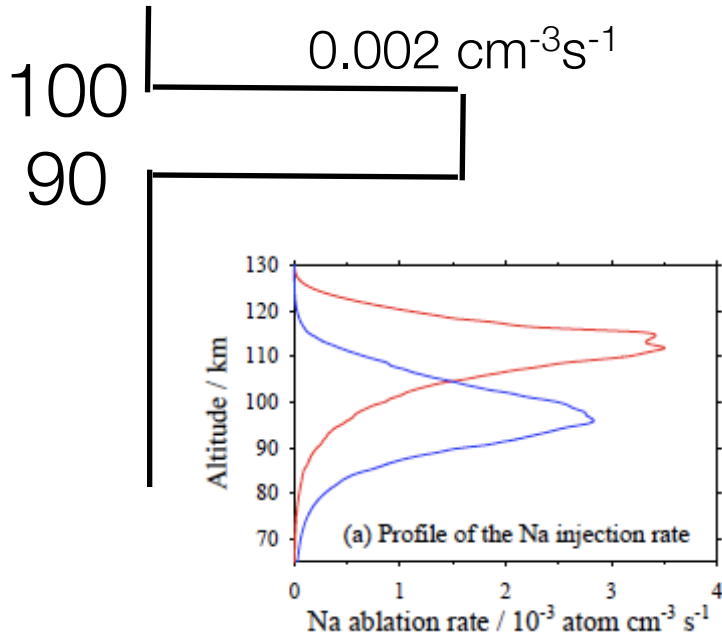


Meteoric sodium

WACCM sodium chemistry model

Chemistry follows Plane [2004]

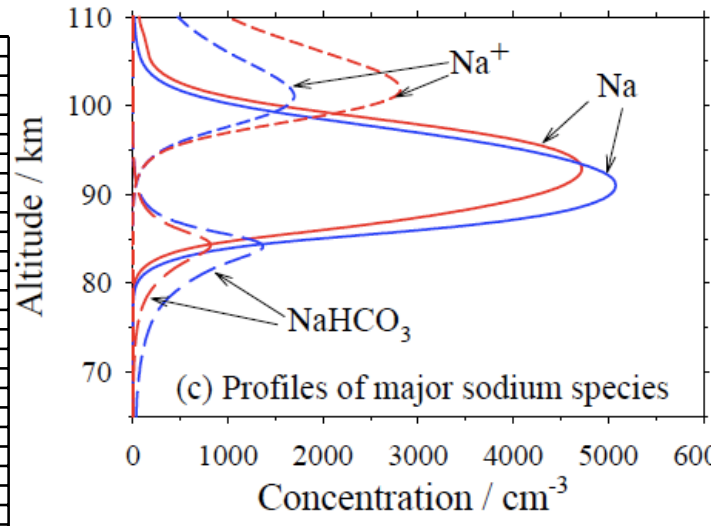
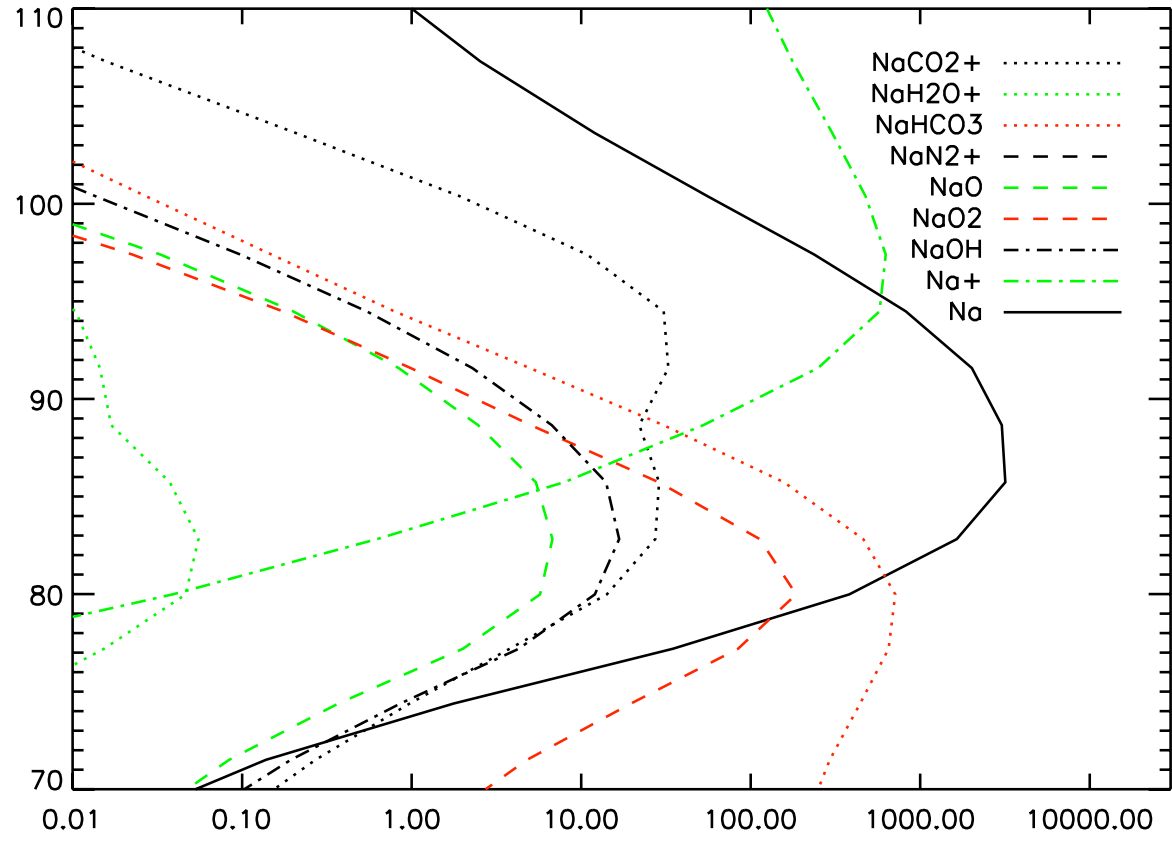
- 26 Na chemical reactions
- 5 photolysis reactions
- *very* simplified constant meteor input function:



Plane, 2004

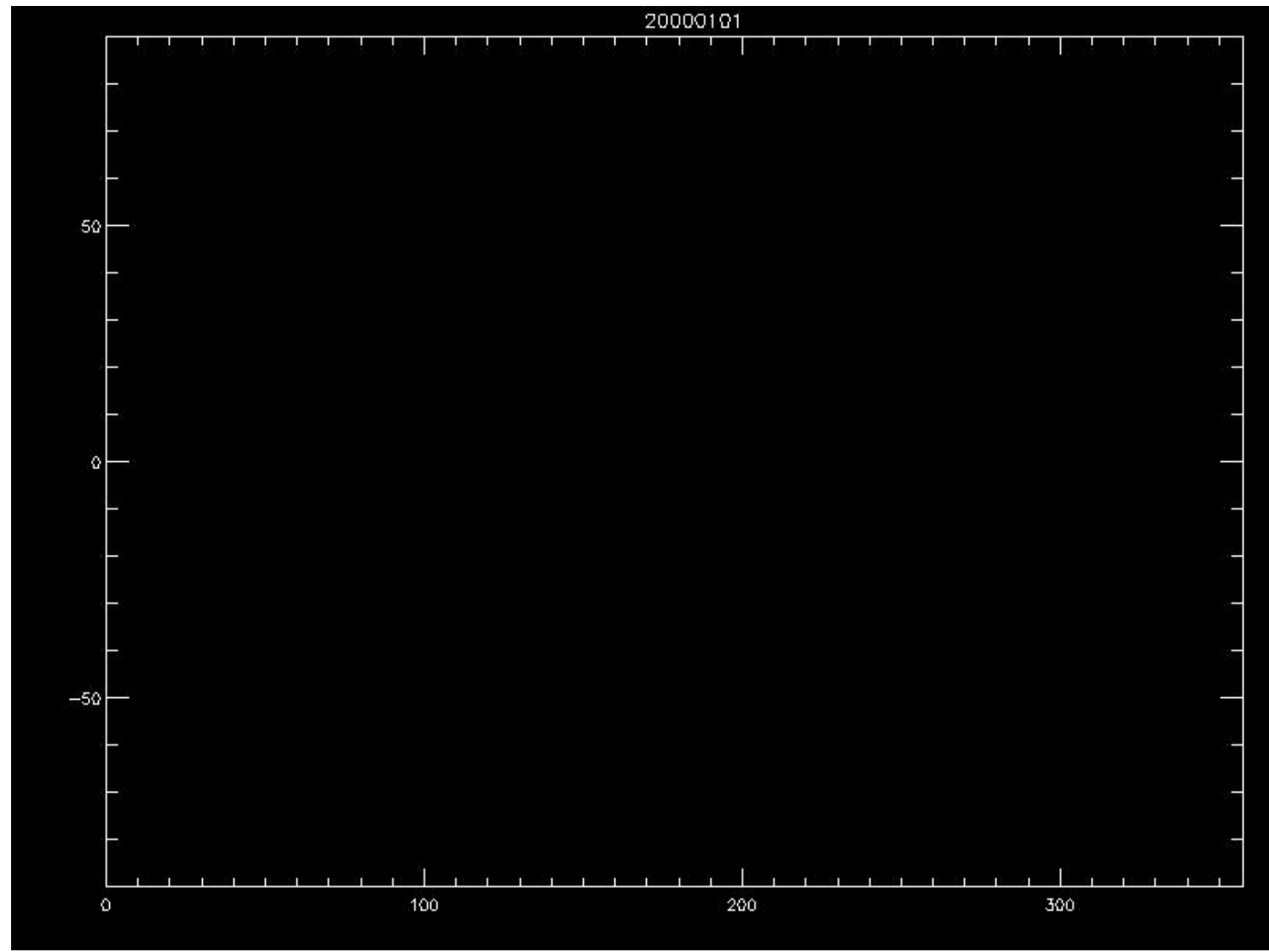
CEDAR, 2009

Global mean constituent profiles



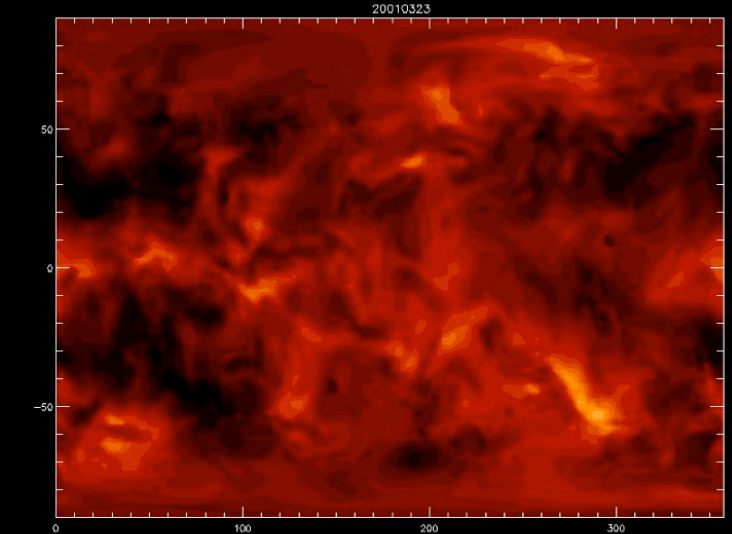
Plane, 2004

Daily 'snapshots' of sodium ~90km / UT00

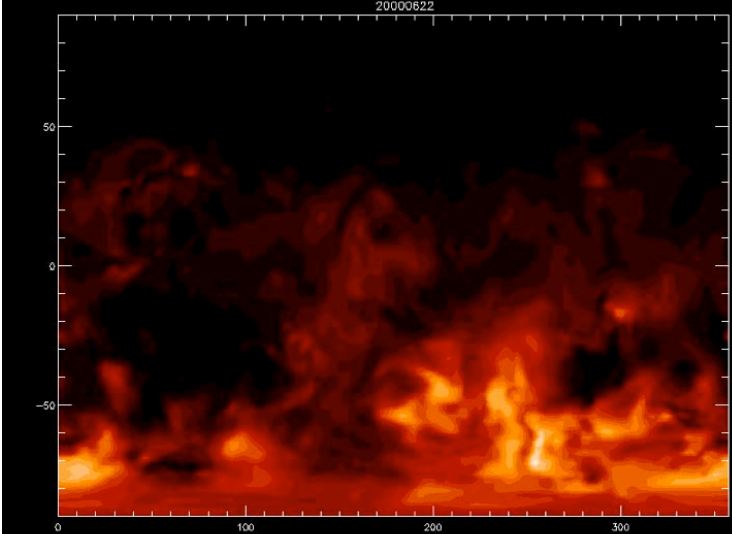


Sodium density at ~90 km

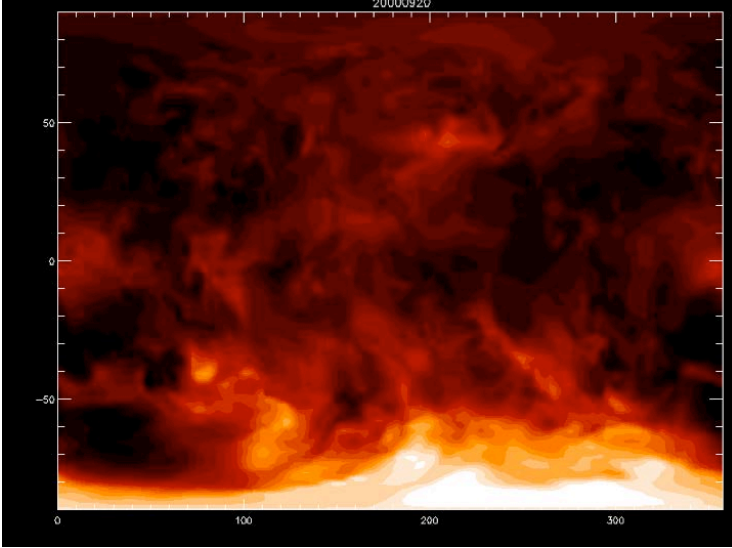
3/23



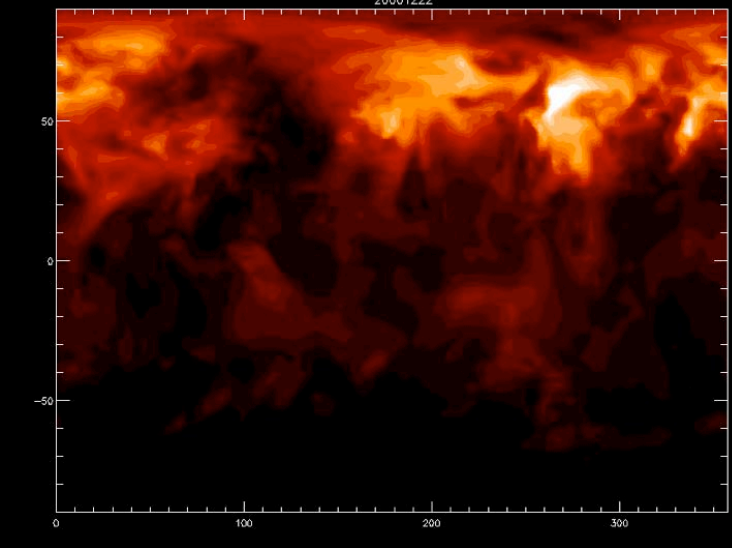
6/22



9/20

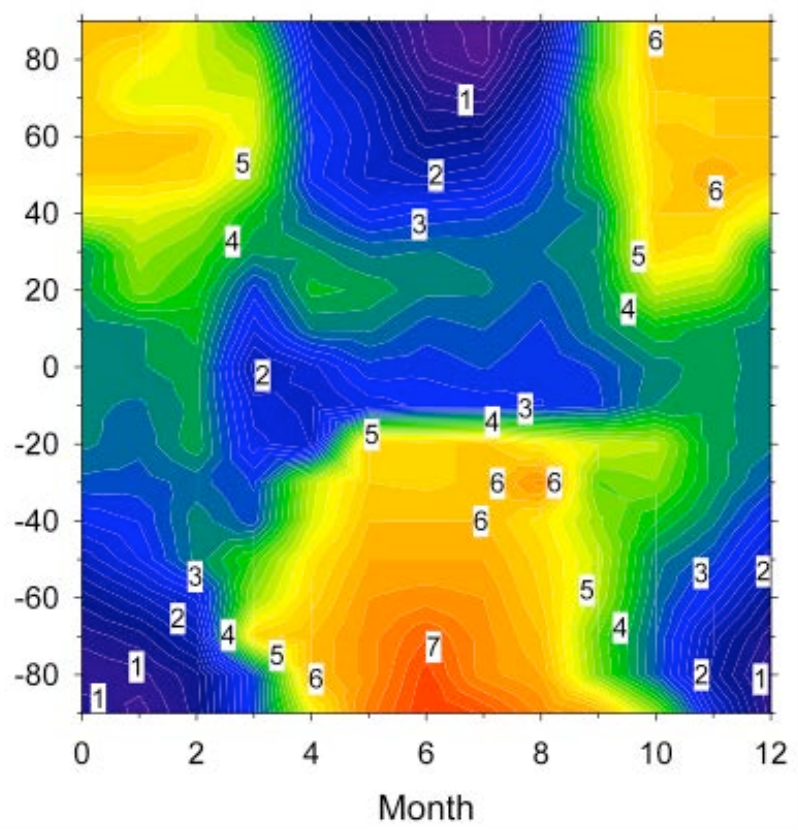


12/22



Seasonal variation of sodium column density

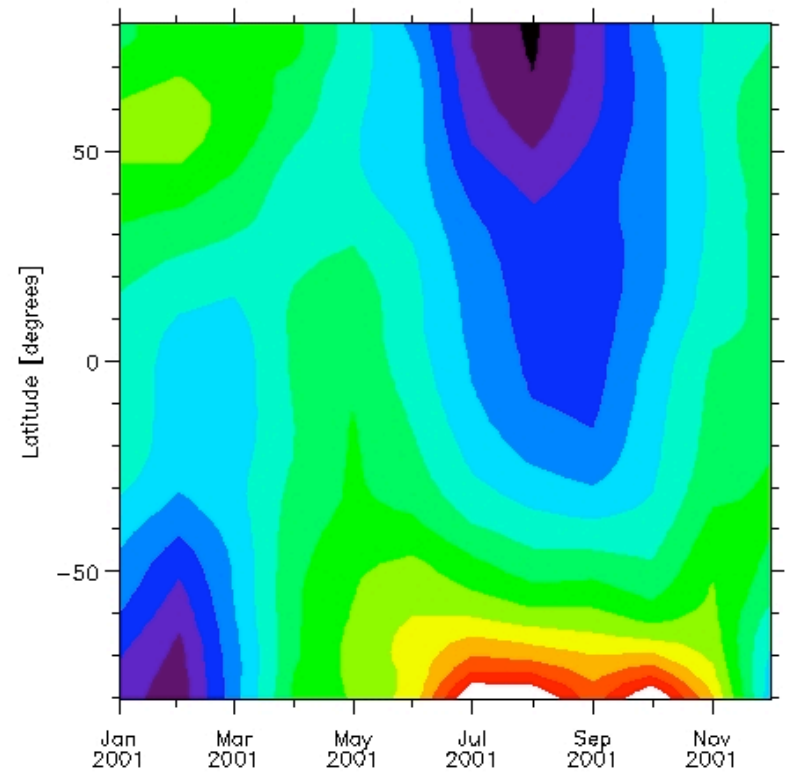
Osiris



Plane, 2009

WACCM

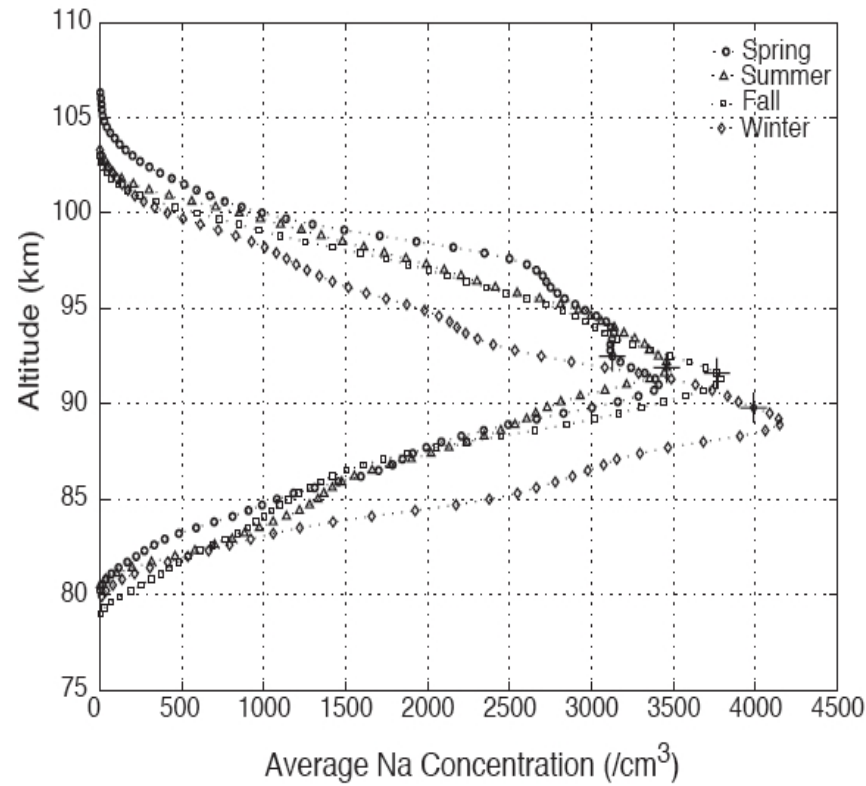
Na Col Dens [cm^{-2}]



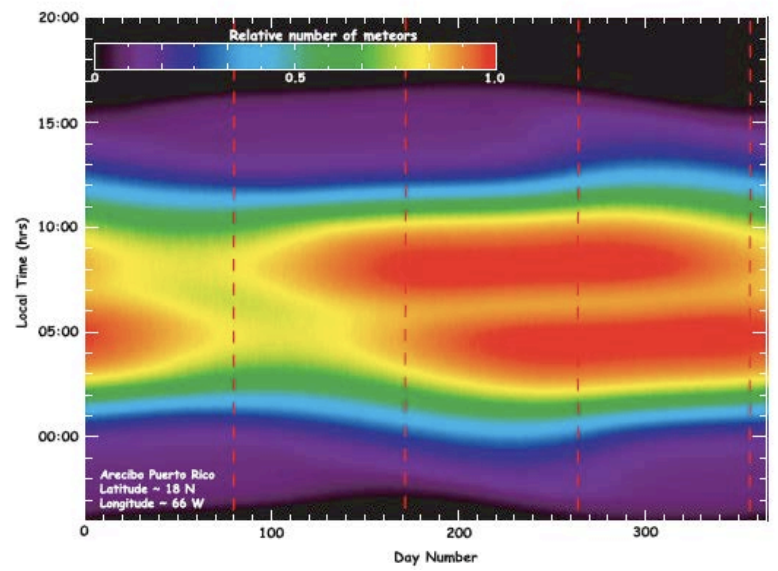
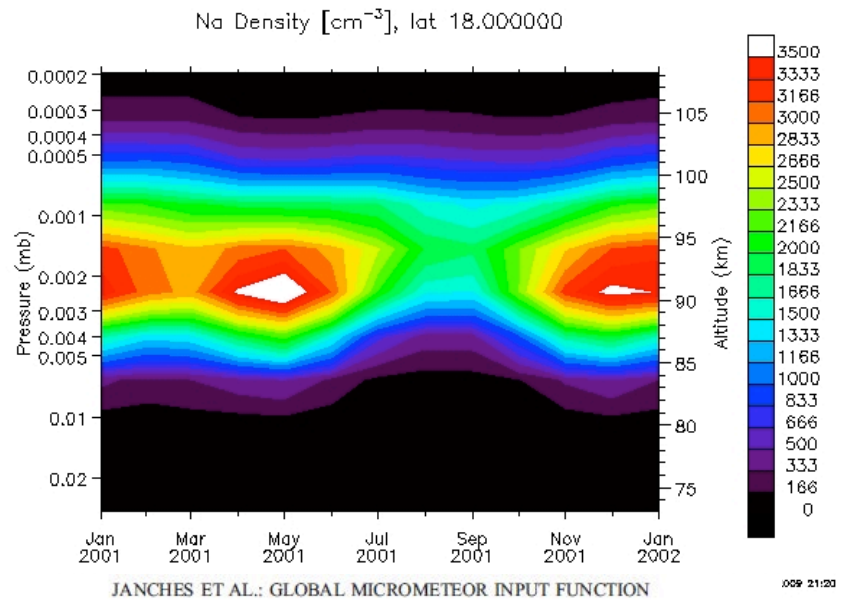
/Users/marsh/Documents/projects/wacm/sodium/wa3548_2c_noam2h02m.na

marsh 28.08.2009 20:45

CEDAR, 2009



Arecibo ISR
Zhou, 2005



CEDAR, 2009



WACCM is available to the scientific community ...

<http://cdp.ucar.edu/>



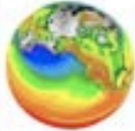
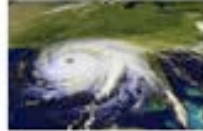








The Community Data Portal (CDP) is a collection of earth science datasets from NCAR, UCAR, UOP, and participating organizations.

Statistics:
8000+ Collections
1,169,041 Files
6.3TB Total Size
4.5TB Downloaded
2107 Registrations

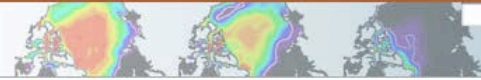
The Community Data Portal is led by NCAR's Computational Science and Information Systems Laboratory (CISL). Support is provided through NCAR's Cyberinfrastructure Strategic Initiative (CSI) and is a collaboration among NCAR, UCAR, UOP, and the National Science Foundation. The technology base for this effort is also a collaborative effort, combining contributions from Unidata, NOAA/PMEL, DLESE, OPeNDAP, OAL, and NASA. [Privacy Policy](#) | [Terms of Use](#)

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Webpage http://www.cesm.ucar.edu/working_groups/

Community Climate System Model


WELCOME TO CCSM

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CCSM WHOLE ATMOSPHERE MODEL WORKING GROUP

The Whole-Atmosphere Community Climate Model (WACCM) is a comprehensive numerical model, spanning the range of altitude from the Earth's surface to the thermosphere. The development of WACCM is an inter-divisional collaboration that unifies certain aspects of the upper atmospheric modeling of HAO, the middle atmosphere modeling of ACD, and the tropospheric modeling of CGD, using the NCAR Community Climate System Model (CCSM) as a common numerical framework.

Upcoming Meetings

WAWG Meeting, NCAR, Mesa Lab, 6 March 2009 [[announcement](#)]

14th Annual CCSM Workshop, The Village at Breckenridge, Breckenridge, CO, 15-18 June 2009.

WAWG Announcements

The WACCM development team is pleased to announce the public release of version 3 of the Whole Atmosphere Community Climate Model (WACCM3). WACCM3 source code, input datasets, run scripts, and documentation can be downloaded from the [Community Data Portal](http://cdp.ucar.edu/) [<http://cdp.ucar.edu/>]. Users are encouraged to participate in meetings of the CCSM WACCM Working Group and to sign up to the mailing list.

WAWG Community Liaison:
TBD

Co-Chair Contact Information:

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CCSM PROJECT

The Community Climate System Model (CCSM) is a fully-coupled, global climate model that provides state-of-the-art computer simulations of the Earth's past, present, and future climate states.

CCSM is sponsored by the National Science Foundation (NSF) and the U.S. Department of Energy (DOE). Administration of the CCSM is maintained by the Climate and Global Dynamics Division (CGD) at the National Center for Atmospheric Research (NCAR).

RELATED

- [CCSM Working Group Co-chairs](#)
- [CCSM Working Group Co-Chairs Terms of Reference](#)
- [WACCM wiki](#)

CONTACT INFORMATION

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Hire
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CEDAR, 2009

WACCM forum on CCSM Bulletin Board

<http://bb.cgd.ucar.edu/>










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Whole Atmosphere Modeling with WACCM

All times are GMT -6. The time now is 06:09 PM.

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CEDAR, 2009

Summary I

- Tides
 - ▶ WACCM produces diurnal/semi-diurnal migrating and non-migrating tides in general agreement with observations
 - ▶ Tidal amplitudes usually lower than observed (~factor 2)
 - ▶ Seasonal and inter-annual variability is well-simulated including QBO modulation
- WACCM-X
 - ▶ Thermospheric extension of WACCM reproduces reasonable composition, temperatures and winds (inc. tides) up to 500km
 - ▶ Ionospheric electrodynamics under development

Summary II

- PMCs
 - ▶ Macroscale parameterization of PMCs incorporated into WACCM reproduces observed seasonal and latitudinal variations ice mass, albedo, frequency
 - ▶ WACCM analysis indicates significant variability induced by tides
 - ▶ Long-term SBUV record well-simulated by WACCM
- WACCM with sodium chemistry
 - ▶ Plane (2004) chemistry implemented with simplified MIF
 - ▶ Significant variability caused by dynamics
 - ▶ Seasonality not as observed - need more realistic MIF

Future activities

- Continued WACCM-X development
- IPCC AR5 simulations using WACCM/CCSM4 with fully-interactive ocean -- investigate “high-top” vs. “low-top” (first with ionosphere?)
- Internally generated QBO
- Couple meteoric metal input function to Na chemistry and compare to LIDAR observations. Extend chemistry to Fe.
- Use metal chemistry model to estimate ‘smoke’ CCN concentrations fed into PMC model



WACCM

*Whole Atmosphere
Community Climate Model*

Thank you