

# A Magnetospheric Perspective on Ionospheric Outflow

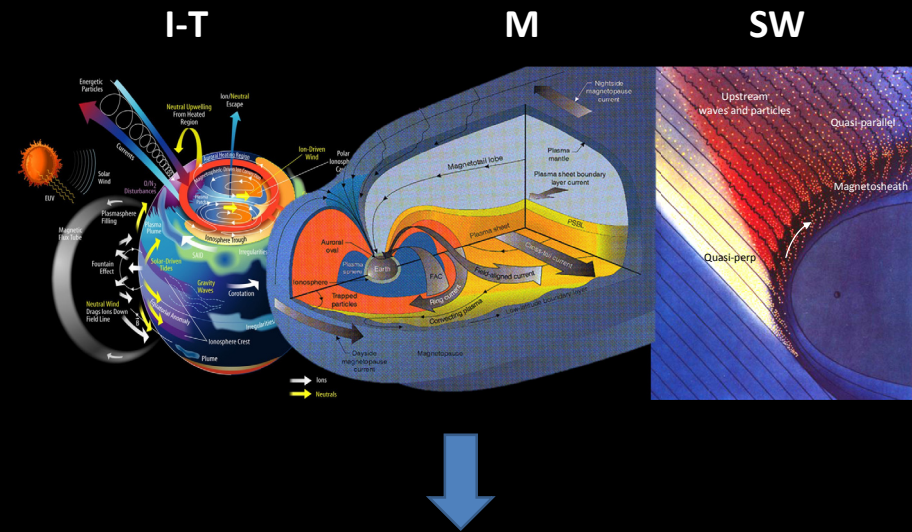
Bill Lotko, Dartmouth College

- Why are outflows important? (*focus on O<sup>+</sup>*)
- What aspects of outflows are important?
- What controls outflows?
- What don't we understand about outflows?

(*Answer: causal relationships*)

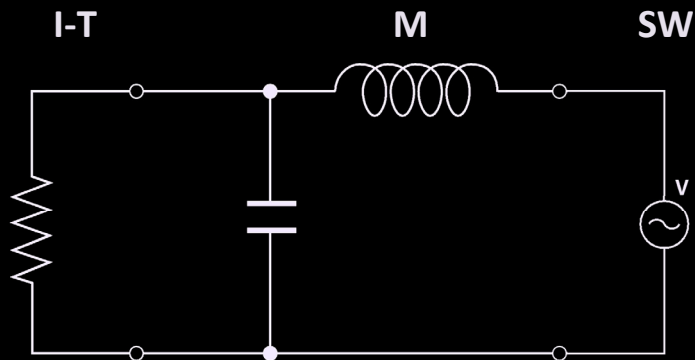
1

## How do we view "other" parts of the geospace system?



2

## How do we view "other" parts of the geospace system?



Mass flow through I-T-M modifies RLC.

How important is the effect in geospace dynamics?

3

## Why are outflows important?

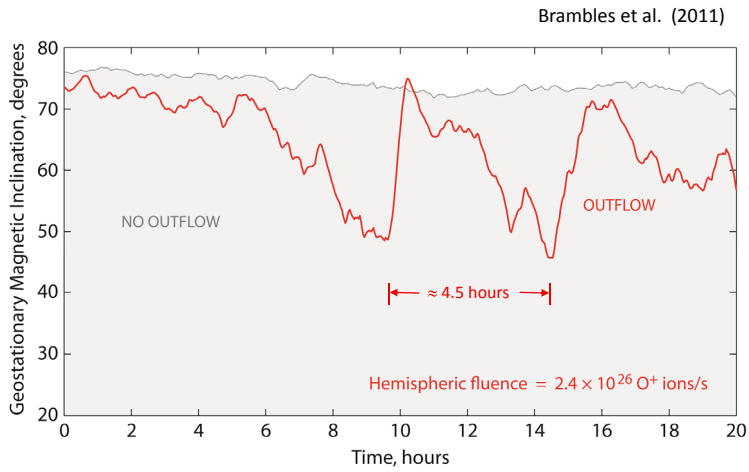
### Impacts on Geospace System (results from models and observations)

- Enhance stormtime ring current (Nosé et al. 2005; Glocer et al. 2009)
- Regulate magnetospheric composition, EMIC waves, radbelts
- Modify magnetotail, plasmasheet dynamics and structure (Kistler et al. 2005; Nosé et al. 2009; Winglee 2004; Wiltberger et al. 2010)
- M** – Inflate magnetosphere (Garcia et al. 2010; Brambles et al. 2010)
  - ⇒ modifies solar wind – magnetosphere interaction
- Reduce transpolar potential (Winglee et al. 2002; Glocer et al. 2009)
- Change threshold for KH instability at MP (Bouhran et al. 2005)
- Induce periodic (2-5 hr) substorms (Brambles et al. 2011)
- I-T** – *Cause* (?) topside cavitation (Chaston et al. 2006)
- *Correlate* with thermospheric upwelling (Liu et al. 2005)

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What aspects of outflows are important?

**No Outflow vs. Outflow  $\Rightarrow$  "Sawtooth" Substorms**



IMF  $B_z = -10$  nT,  $V_{sw} = 400$  km/s,  $n_{sw} = 5$  cm<sup>-3</sup> // Geo inclination angle  $\equiv \sin^{-1}(B_z/B)$  at  $z = 0.5 R_E$ ,  $r = 6.6 R_E$

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What aspects of outflows are important?

**Case A (conic "like")**

$v_{||} = 50$  km/s,  $T = 100$  eV

Empirical outflow model

$$F_{||O} = 2.1 \times 10^7 S_{dc}^{1.3}$$

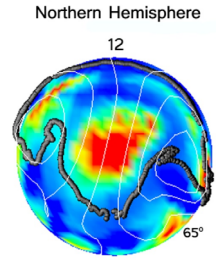
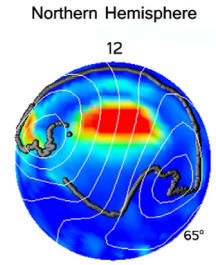
Strangeway et al. 2005

**Case B (polar wind "like")**

$v_{||} = 3$  km/s,  $T = 1$  eV

Projections to  
100 km from  $6 R_E$

$10^9$  O<sup>+</sup> ions/cm<sup>2</sup>-s



Brambles et al. 2010

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What aspects of outflows are important?

"Conic-like" cusp outflow

$v_{||} = 50$  km/s

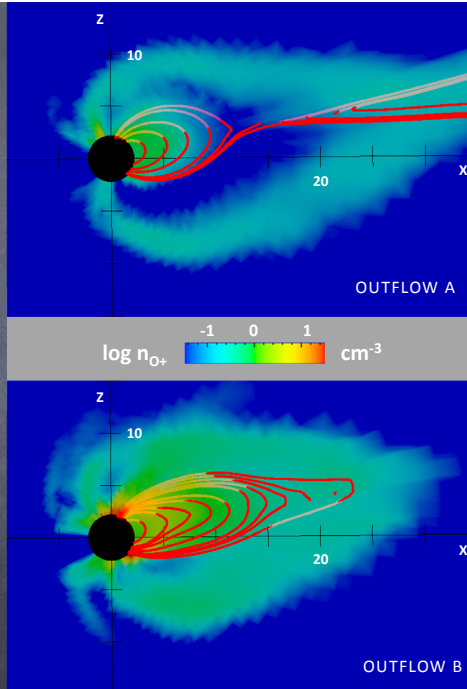
$\rightarrow$  Increases CPCP

O<sup>+</sup> interaction with  
nightside reconnection

"Polar-wind" cusp outflow

$v_{||} = 3$  km/s

$\rightarrow$  Decreases CPCP



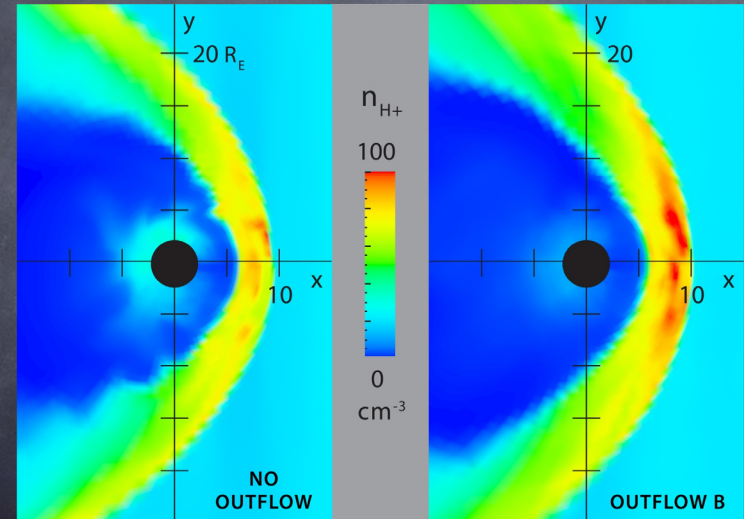
Brambles et al. 2010

OUTFLOW B

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What aspects of outflows are important?

**Magnetospheric Inflation**



Brambles et al. 2010

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What aspects of outflows are important?

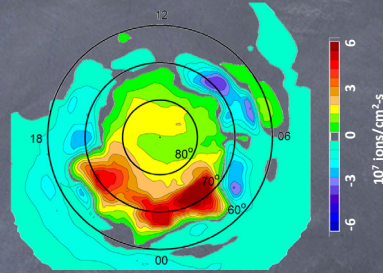
**O<sup>+</sup> Source Location**

Winglee 2009

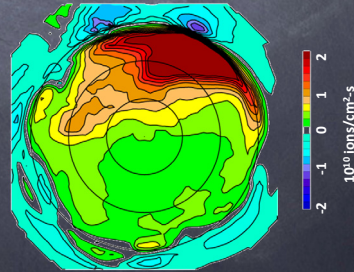
- Substorms: nightside dominant
- Storms: dayside dominant
  - but nightside still significant

- Tail dynamics similar
  - except recurrent activity feeds off preloaded O<sup>+</sup> to give faster responses during later activity

18 Mar 2007 Non-Storm



29 Oct 2003 Storm

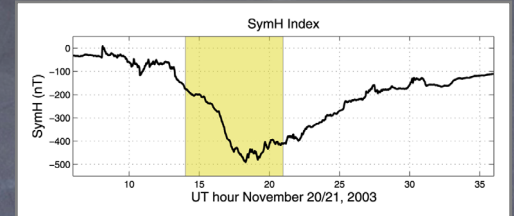


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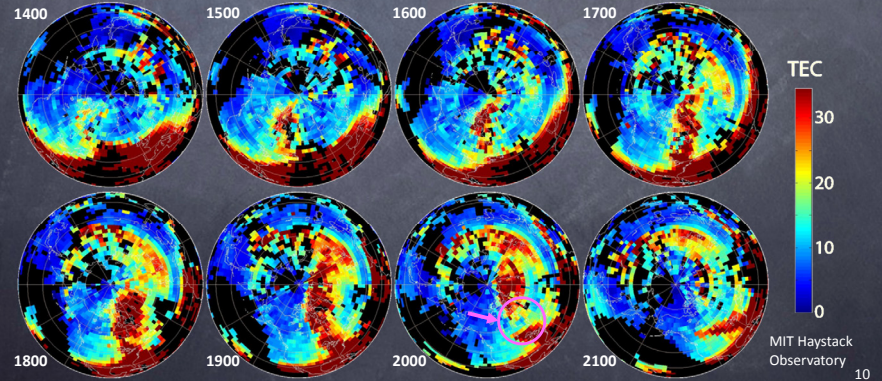
What controls outflows?

**Convective transport to the cusp**

→ 10<sup>10</sup> ions / cm<sup>2</sup>-s



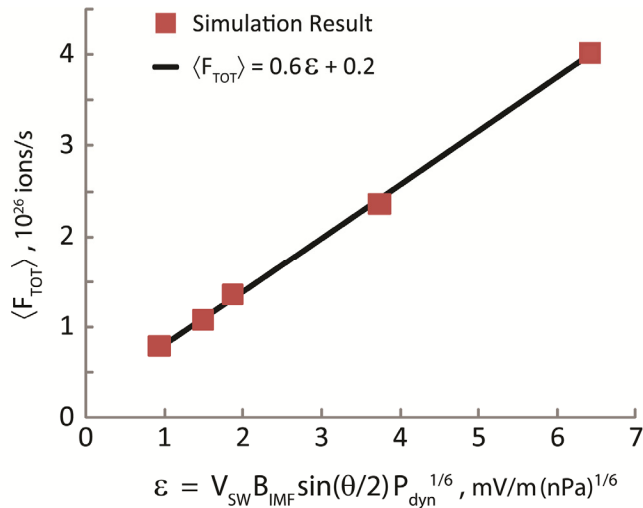
Coster, Erickson, Foster (2009)



MIT Haystack Observatory 10

What controls outflows?

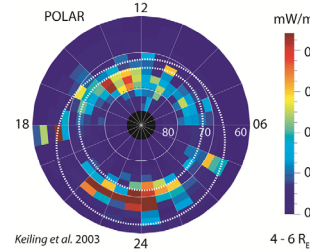
**Outflow Fluence <F<sub>TOT</sub>> vs. SW Driving ε**



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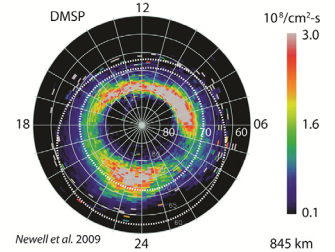
What controls outflows?

**Alfvénic Poynting Flux**



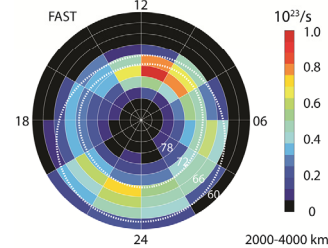
Keiling et al. 2003

**Electron Number Flux**



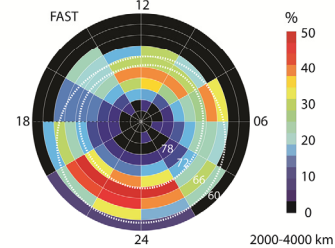
Newell et al. 2009

**Outflow Fluence**



Chaston et al. 2007

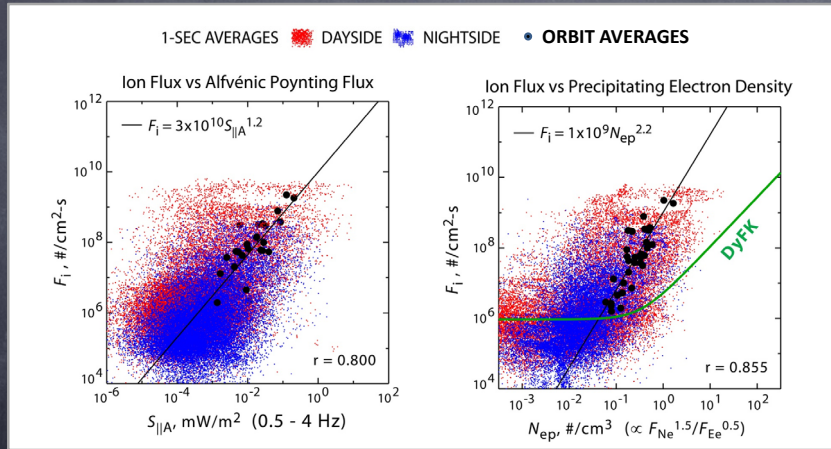
**% Associated With Alfvén Waves**



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What don't we understand about outflows?

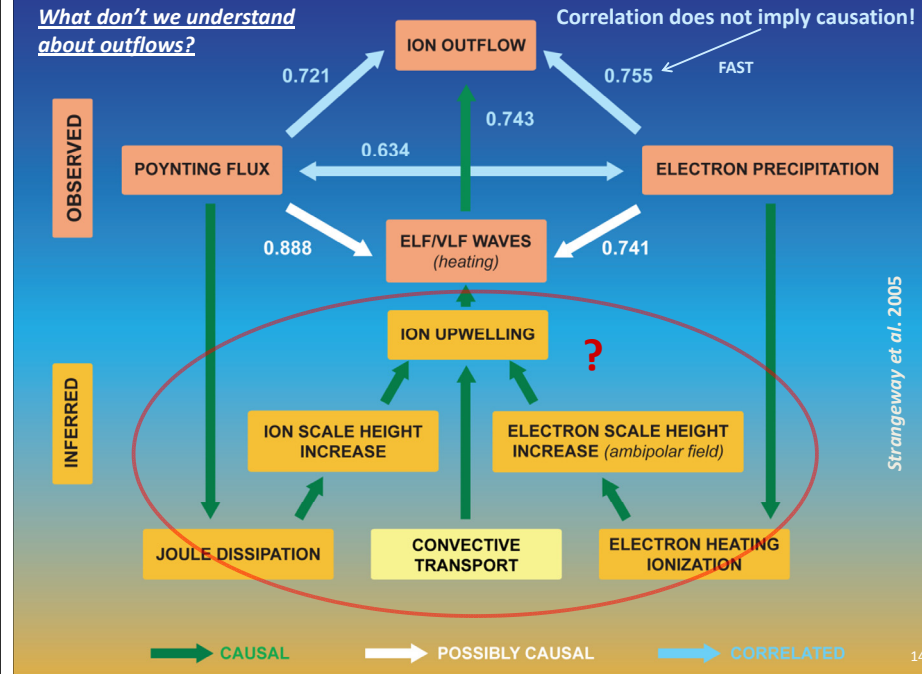
Correlations with outflow flux (FAST)



Strangeway, 2009

Are Alfvén waves direct or indirect drivers?

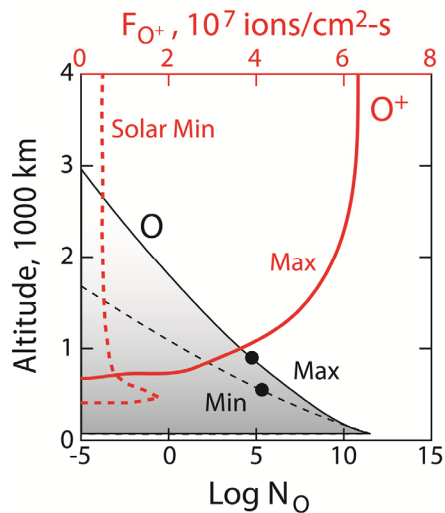
What don't we understand about outflows?



Strangeway et al. 2005

What controls outflows?

Thermospheric Control of Outflows (Cannata and Gombosi, 1989)



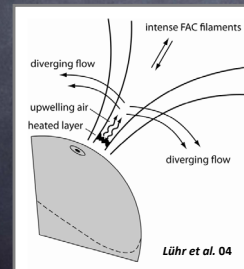
- Accidentally resonant charge exchange chemistry couples H<sup>+</sup>, O<sup>+</sup> flows to the neutral density profiles of atomic H and O.
- Larger O<sup>+</sup> outflows arise at solar max based solely on thermospheric changes
- Higher neutral O scale heights at solar max shifts maximum production level of O<sup>+</sup> to higher altitudes where loss by reaction with O<sub>2</sub> and N<sub>2</sub> is less frequent

What don't we understand about outflows?

Thermospheric Control

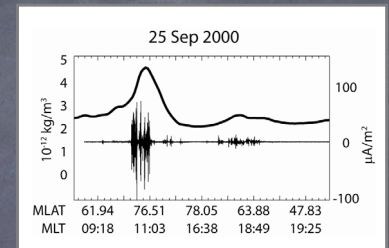
Causes of upwelling

- Large-scale Joule heating
- Ion upflow
- Soft electron precipitation
- Alfvén waves, small-scale FAC

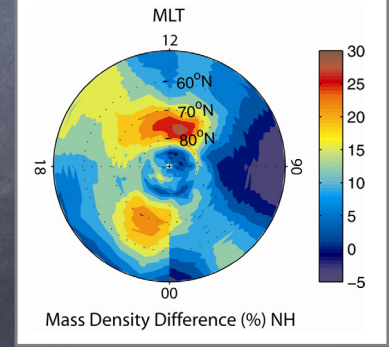


Lühr et al. 04

CHAMP thermospheric density enhancement



Lühr et al. 04



Liu et al. 05

# Open Questions

## • How are ionospheric $O^+$ outflows energized?

- EM energy conversion, particle precipitation, waves/turbulence, ...

## • How do interplanetary and I-T-M conditions control outflows, their distributions and fluxes?

- Outflow distribution, ionosphere-thermosphere upwelling and chemistry, magnetospheric power flows, sub/storms, ...

## • How does the I-T-M system respond to ionospheric outflows?

- Thermosphere-ionosphere upwelling, ionospheric cavitation, episodic dynamics, reconnection potential, transpolar potential, ...

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## References

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