

Stormtime Ion Redistribution at Mid-Latitudes: A Coupled Geospace Phenomenon

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Google Earth Imagery
The Space Physics Community

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Massachusetts Institute of Technology



Mesoscale Ionospheric Redistributions

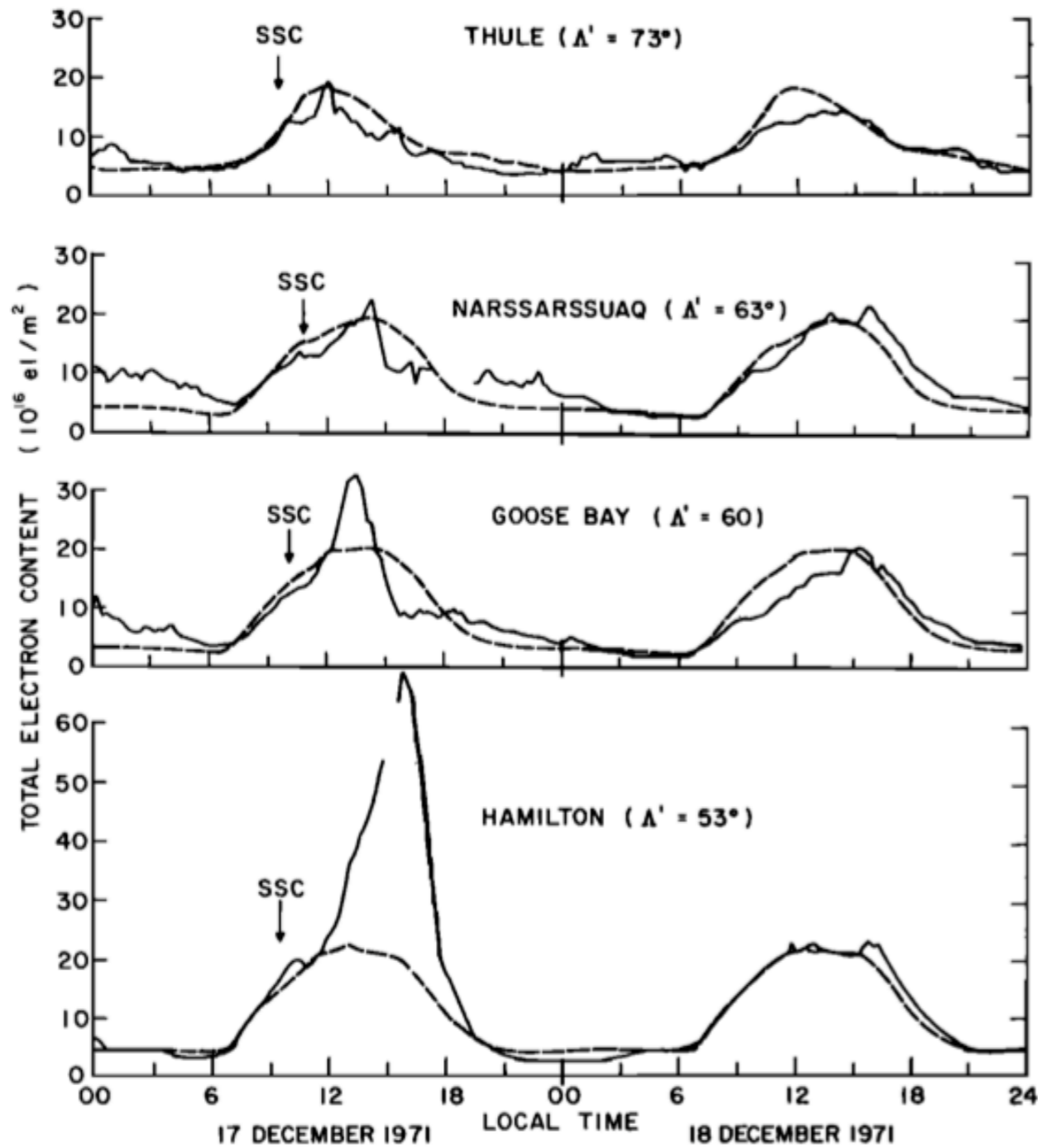


Fig. 5. TEC data from AFRCL facilities at Thule, Narssarssuaq, Goose Bay, and Hamilton for December 17–18, 1971. The dashed curves give the monthly median behavior at each station, and the small arrows mark the local times of the ssc at 1418 UT.

AFRCL network observing
ATS series geosynchronous
VHF beacons

Differential Doppler estimate
of TEC

17 December 1971

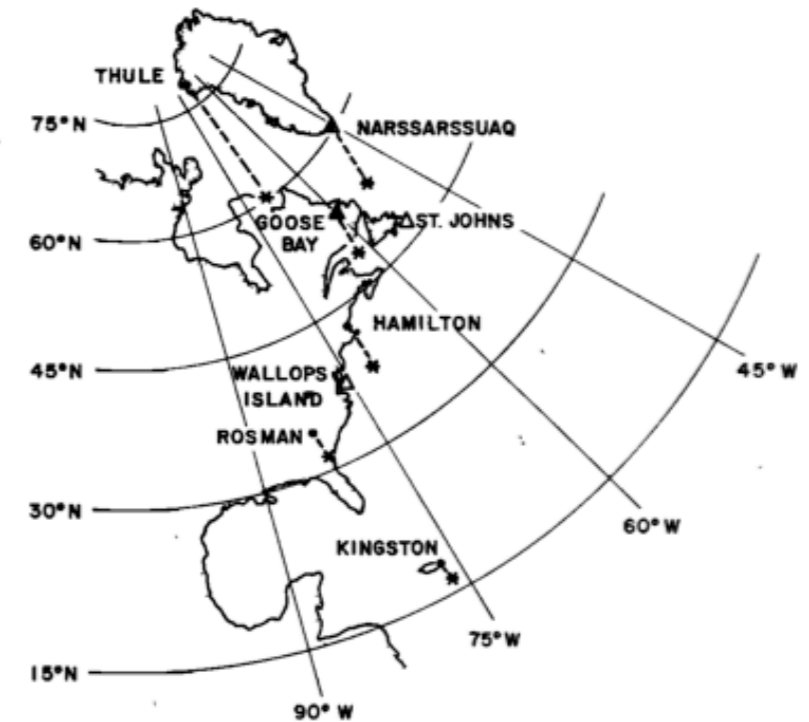
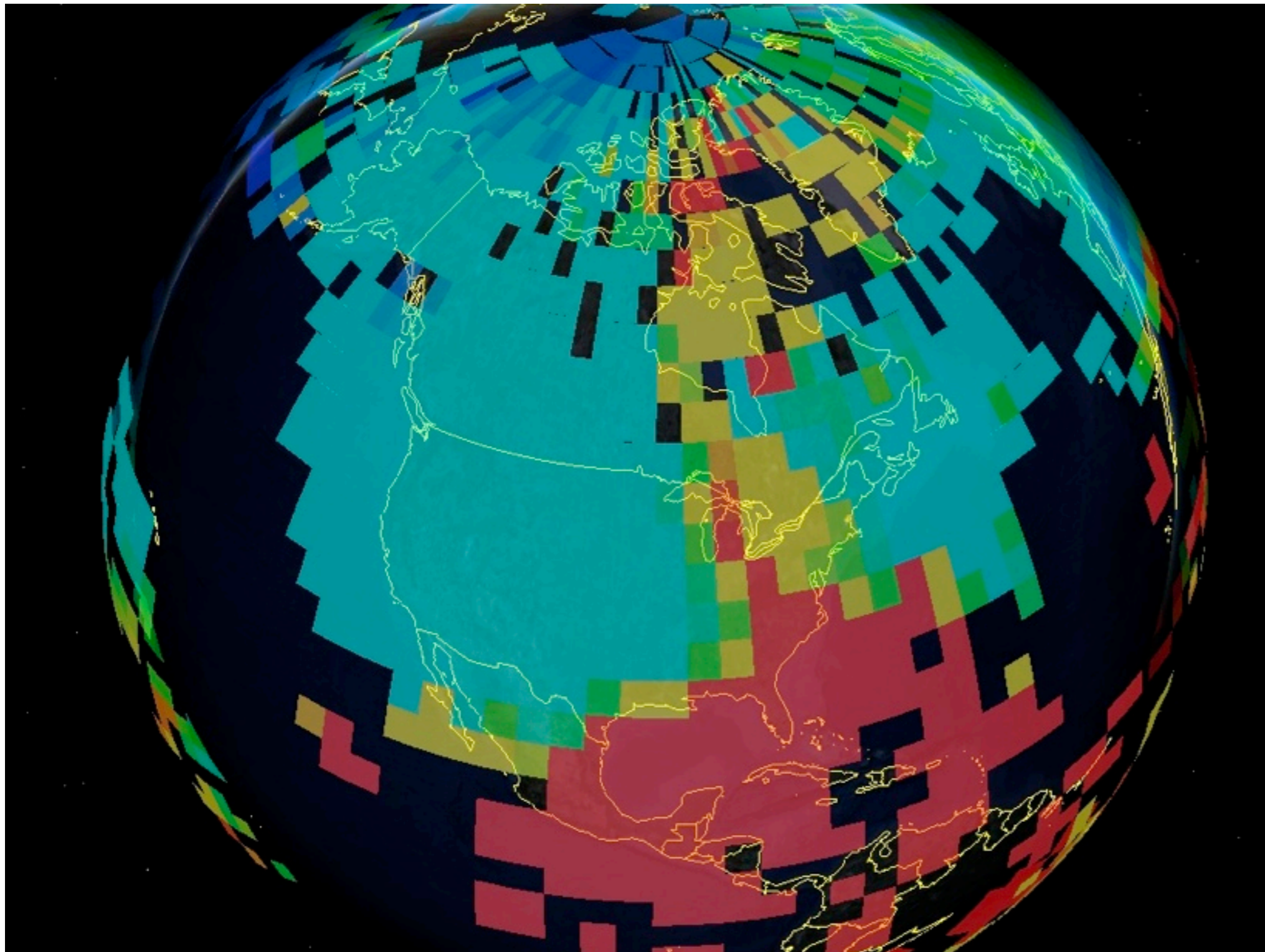


Fig. 1. Locations of TEC observing stations (solid dots), their 420-km subionospheric points (asterisks) for ATS 3 at 70°W, and the nearby ionosonde stations (open triangles).

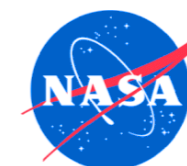
Mendillo and Klobuchar,
1975

Mesoscale Ionospheric Redistributions



GPS TEC
[0, 60] TECu

Nov 20, 2003
1840 - 1900 UTC

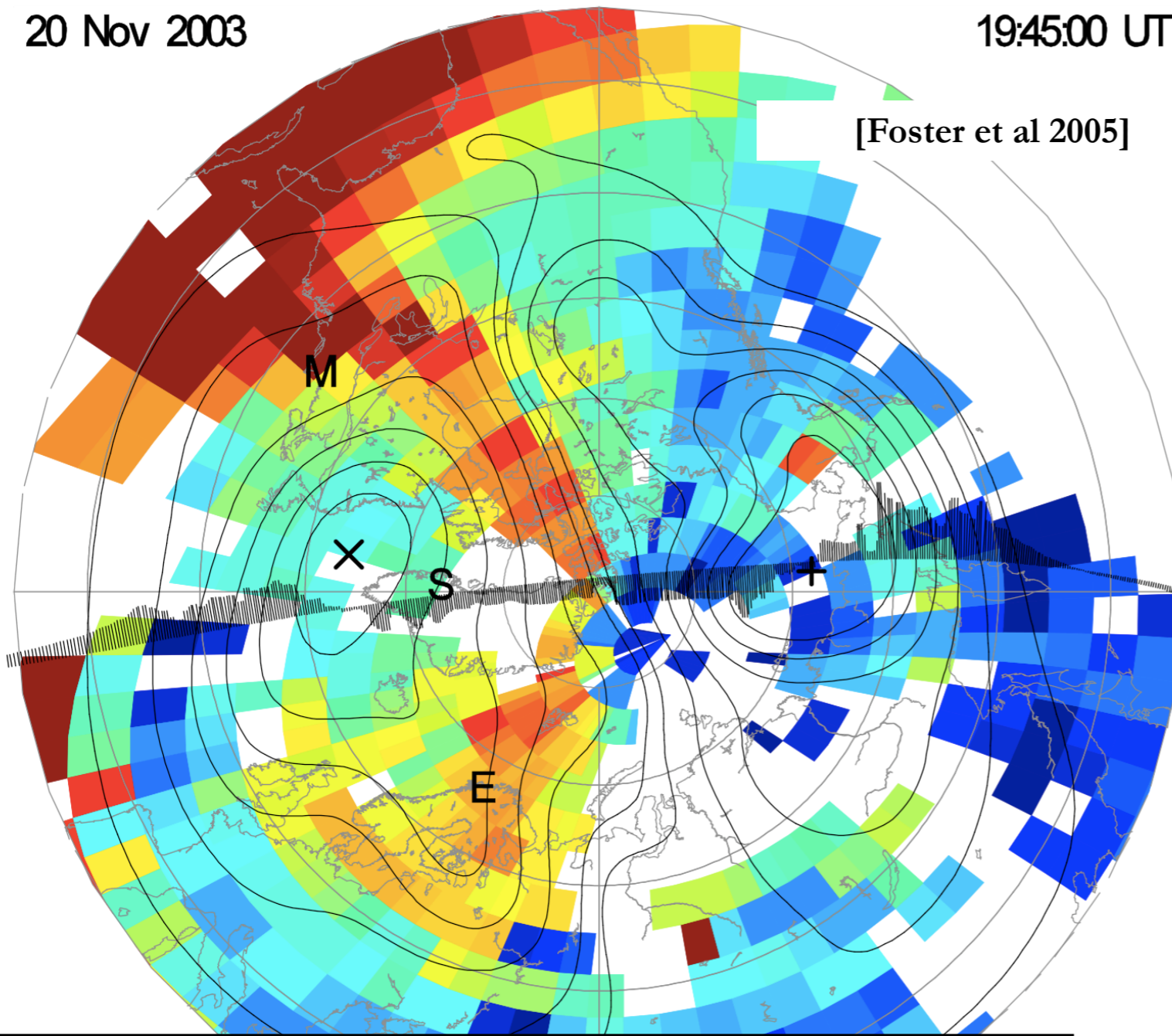


System-Level Redistribution Paths

20 Nov 2003

19:45:00 UT

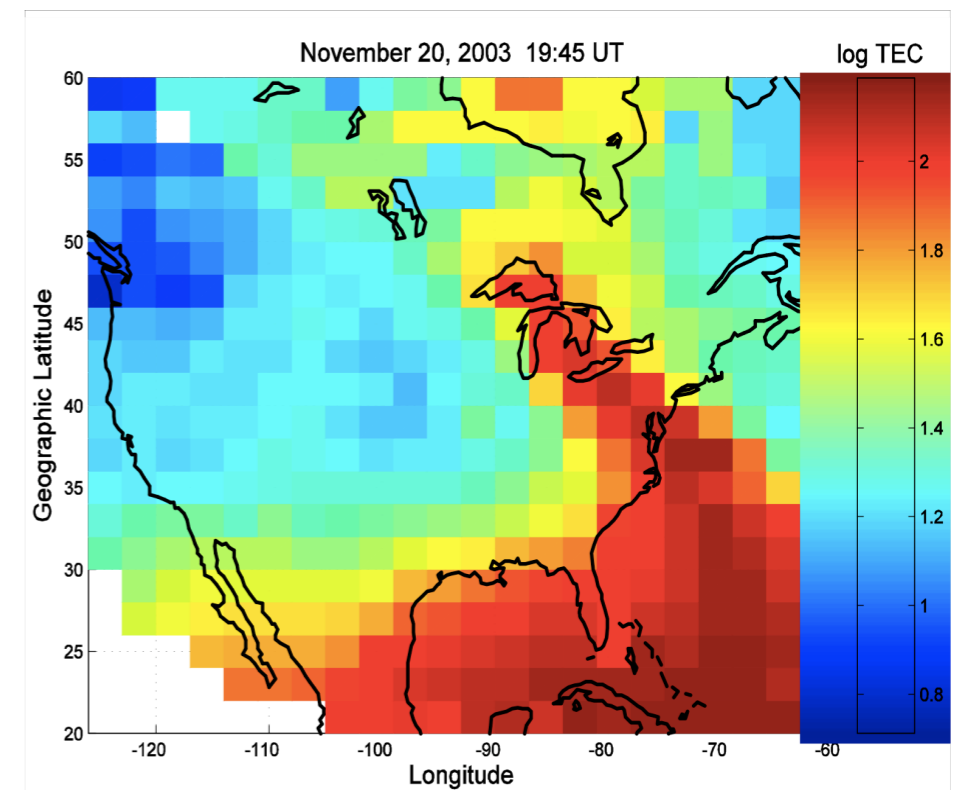
[Foster et al 2005]



Active plasma redistribution couples equatorial and polar latitudes through mid-latitudes

SED, polar tongue of ionization created through actions of region 2 linked SAPS electric fields (M-I coupling)

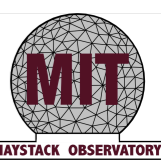
GPS TEC Map



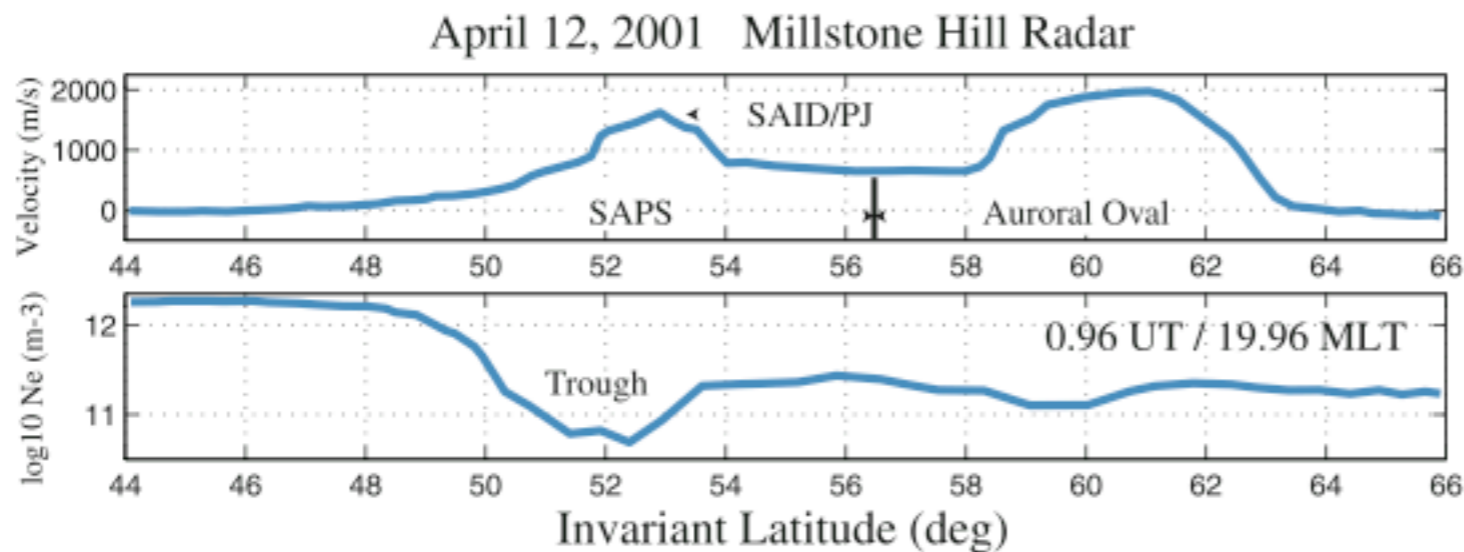
Merged SuperDARN/DMSP Convection
Common projection: maglat/MLT @ 350 km alt



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Sub-Auroral Polarization Stream (SAPS)



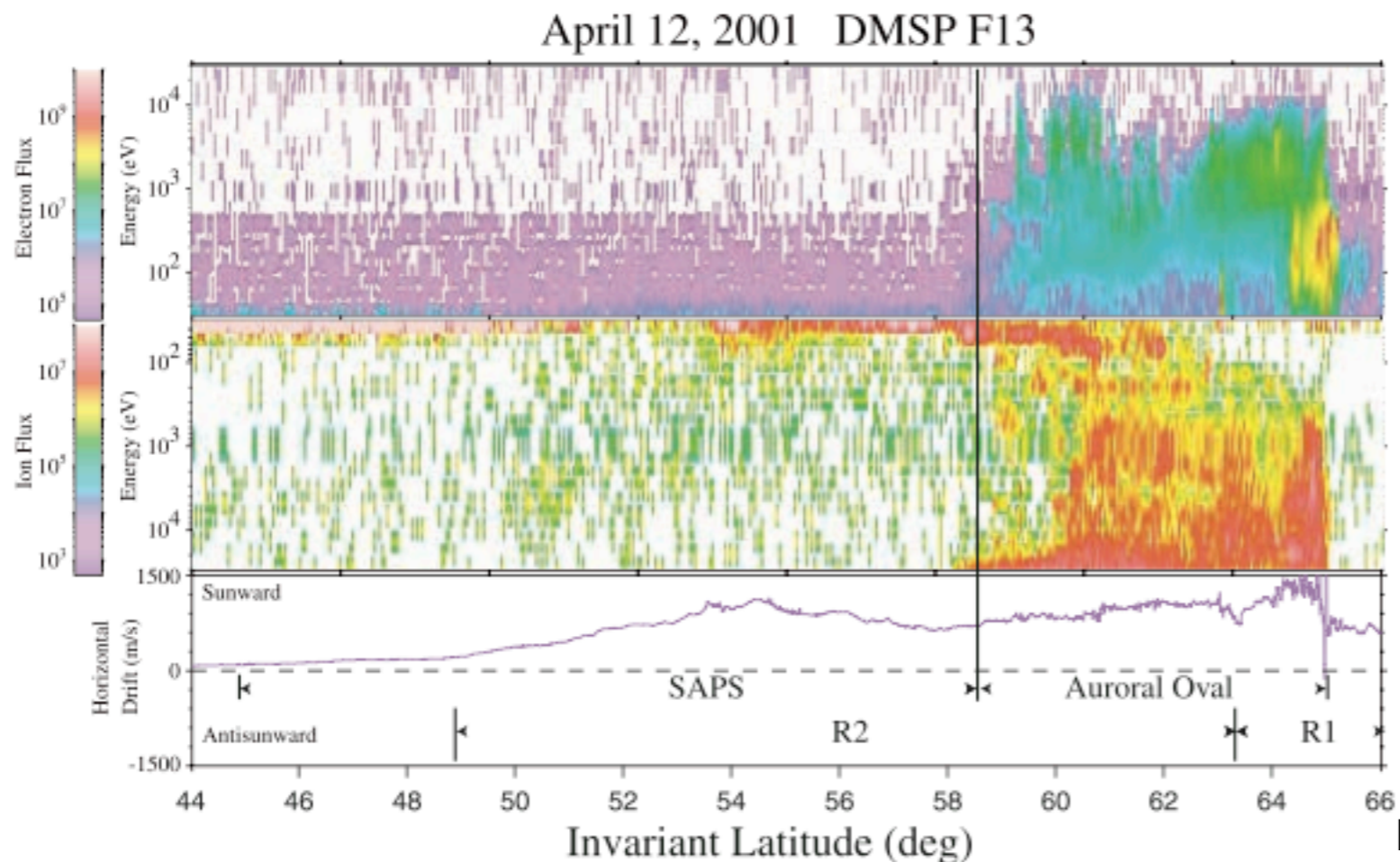
Westward (sunward) subauroral velocity near footprint of region 2 / ring current

2-5 deg wide

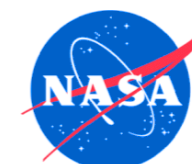
Embedded small and highly variable structures (SAID)

Overlaps edge of storm enhanced density (SED)

Dusk sector transport of material to noontime cusp

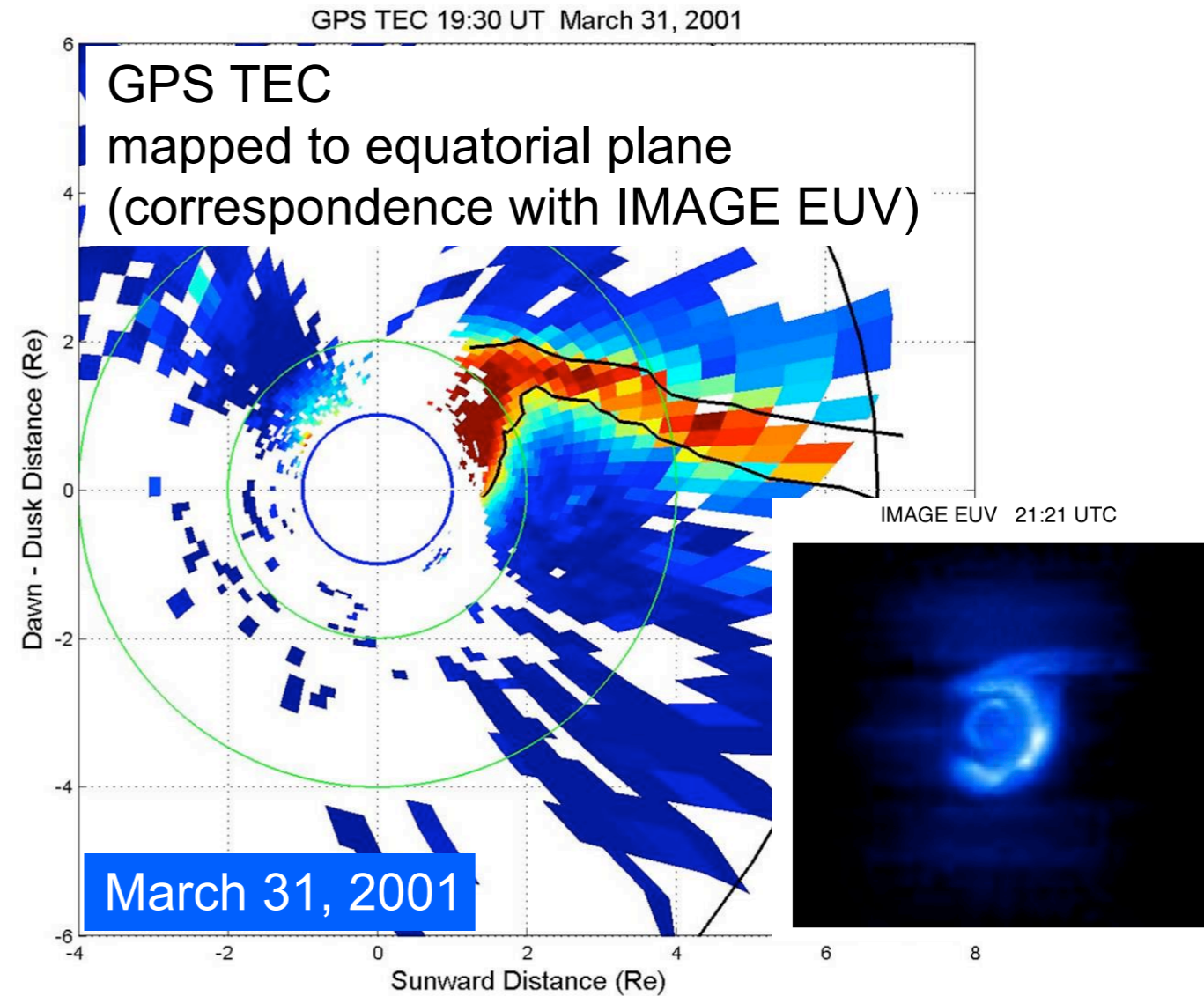
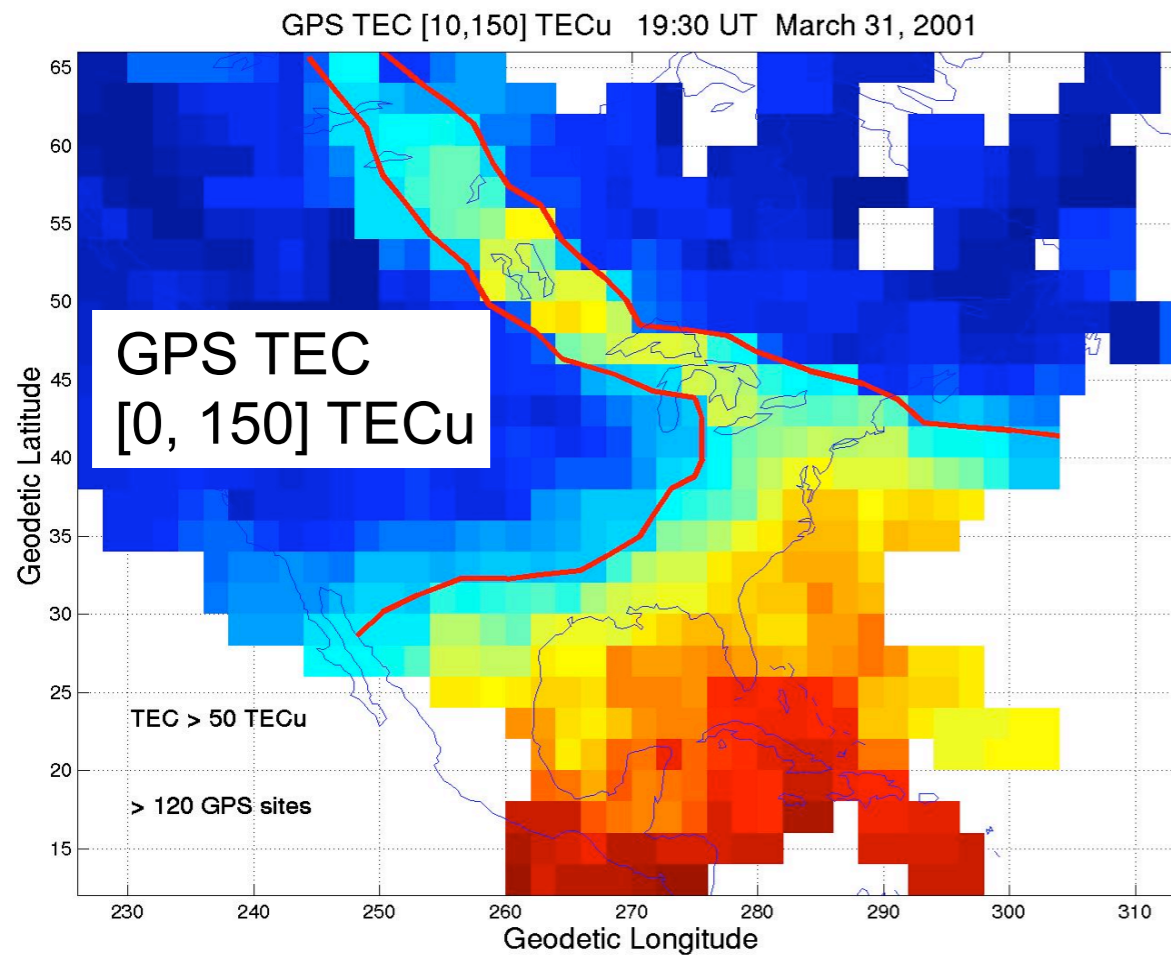


Foster and Vo, 2002



The Coupled Geospace Observational View

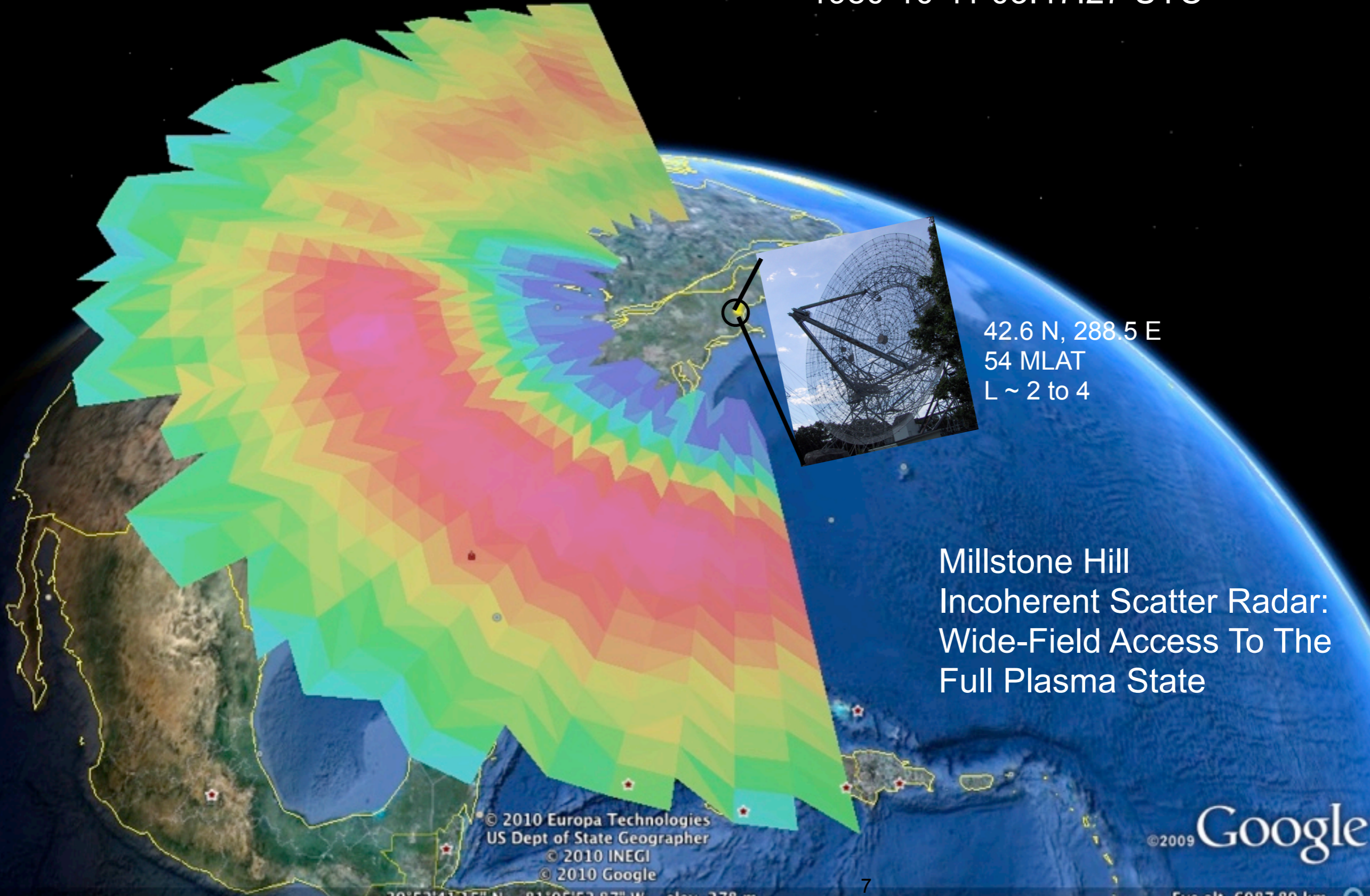
What are the *statistical and system* characteristics of this mesoscale redistribution in the ionosphere?



(e.g. Foster et al 2004)

Kp = 6 event
F10.7 = 233
DsT -100 nT

Millstone Hill UHF Radar
Azimuth Scan (4 deg EI)
Log Electron Density m⁻³ [10, 12.5]
1980-10-11 03:47:27 UTC



42.6 N, 288.5 E
54 MLAT
L ~ 2 to 4

Millstone Hill
Incoherent Scatter Radar:
Wide-Field Access To The
Full Plasma State

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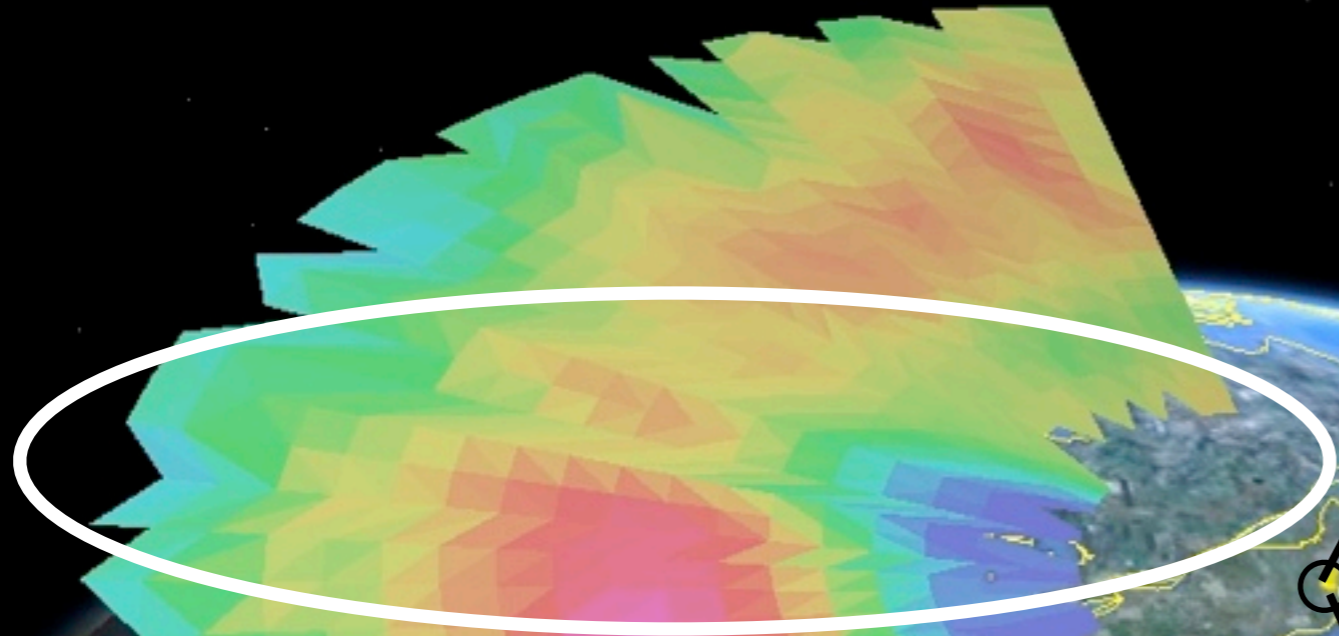
©2009 Google

39°52'41.15" N 81°05'52.87" W elev 278 m

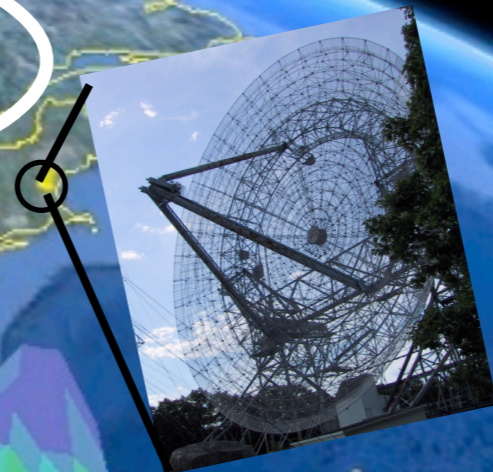
Eye alt 6087.89 km

Kp = 6 event
F10.7 = 233
DsT -100 nT

Millstone Hill UHF Radar
Azimuth Scan (4 deg EI)
Log Electron Density m⁻³ [10, 12.5]
1980-10-11 03:47:27 UTC



Plasmasphere Boundary Layer



42.6 N, 288.5 E
54 MLAT
L ~ 2 to 4

Millstone Hill
Incoherent Scatter Radar:
Wide-Field Access To The
Full Plasma State

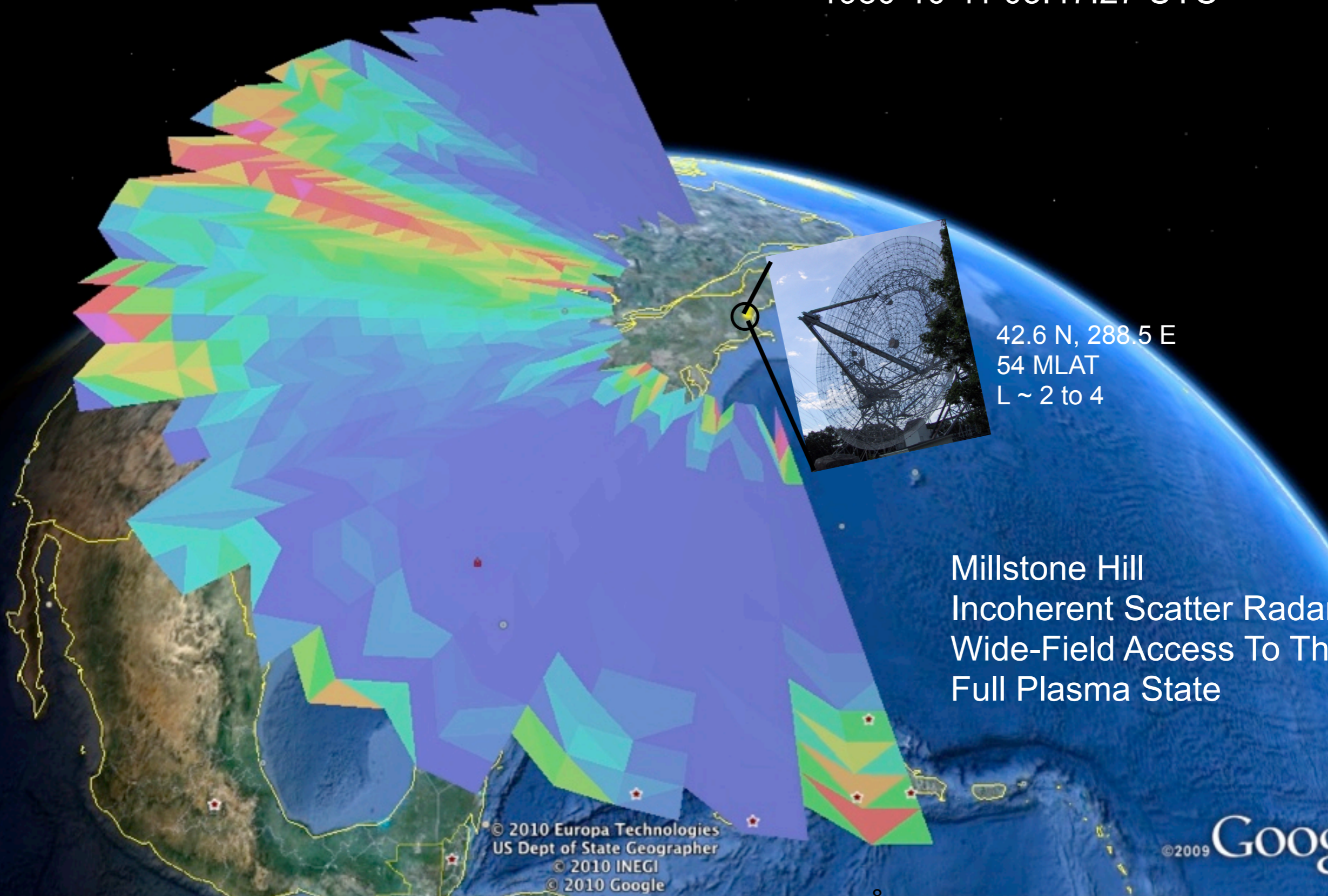
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39°52'41.15" N 81°05'52.87" W elev 278 m

Kp = 6 event
F10.7 = 233
DsT -100 nT

Millstone Hill UHF Radar
Azimuth Scan (4 deg EI)
Line-of-sight Ion Velocity [0,800] m/s
1980-10-11 03:47:27 UTC



42.6 N, 288.5 E
54 MLAT
L ~ 2 to 4

Millstone Hill
Incoherent Scatter Radar:
Wide-Field Access To The
Full Plasma State

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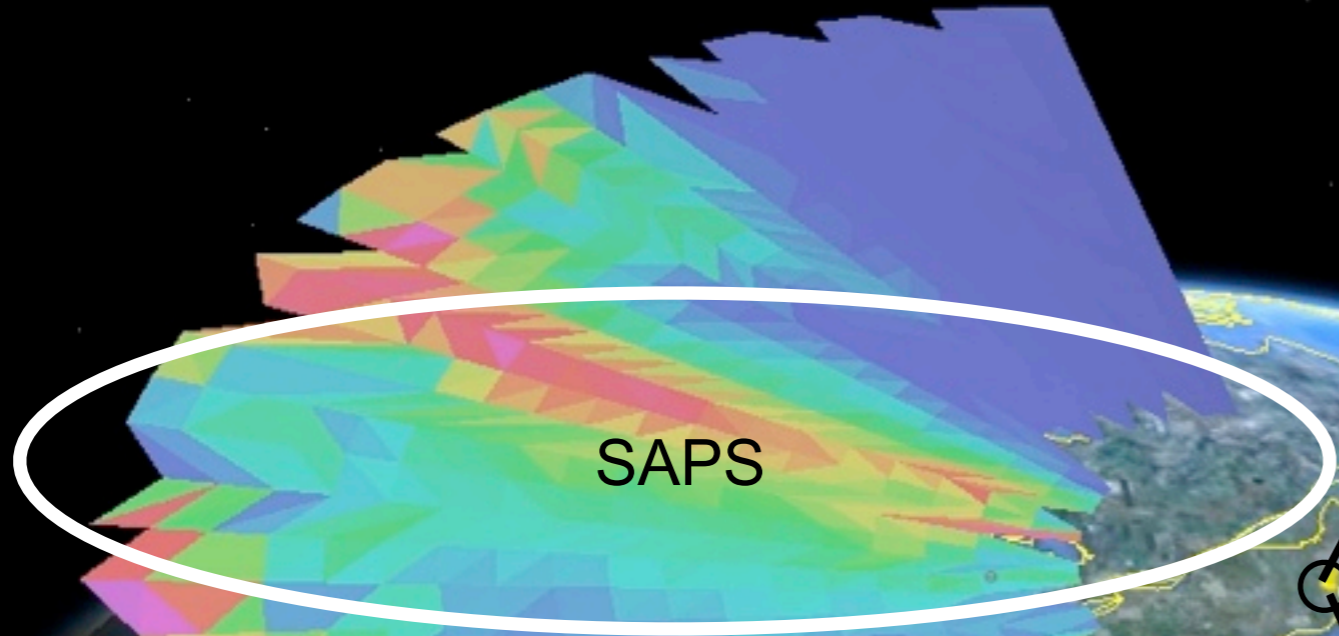
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39°52'41.15" N 81°05'52.87" W elev 278 m

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Millstone Hill UHF Radar
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1980-10-11 03:47:27 UTC



42.6 N, 288.5 E
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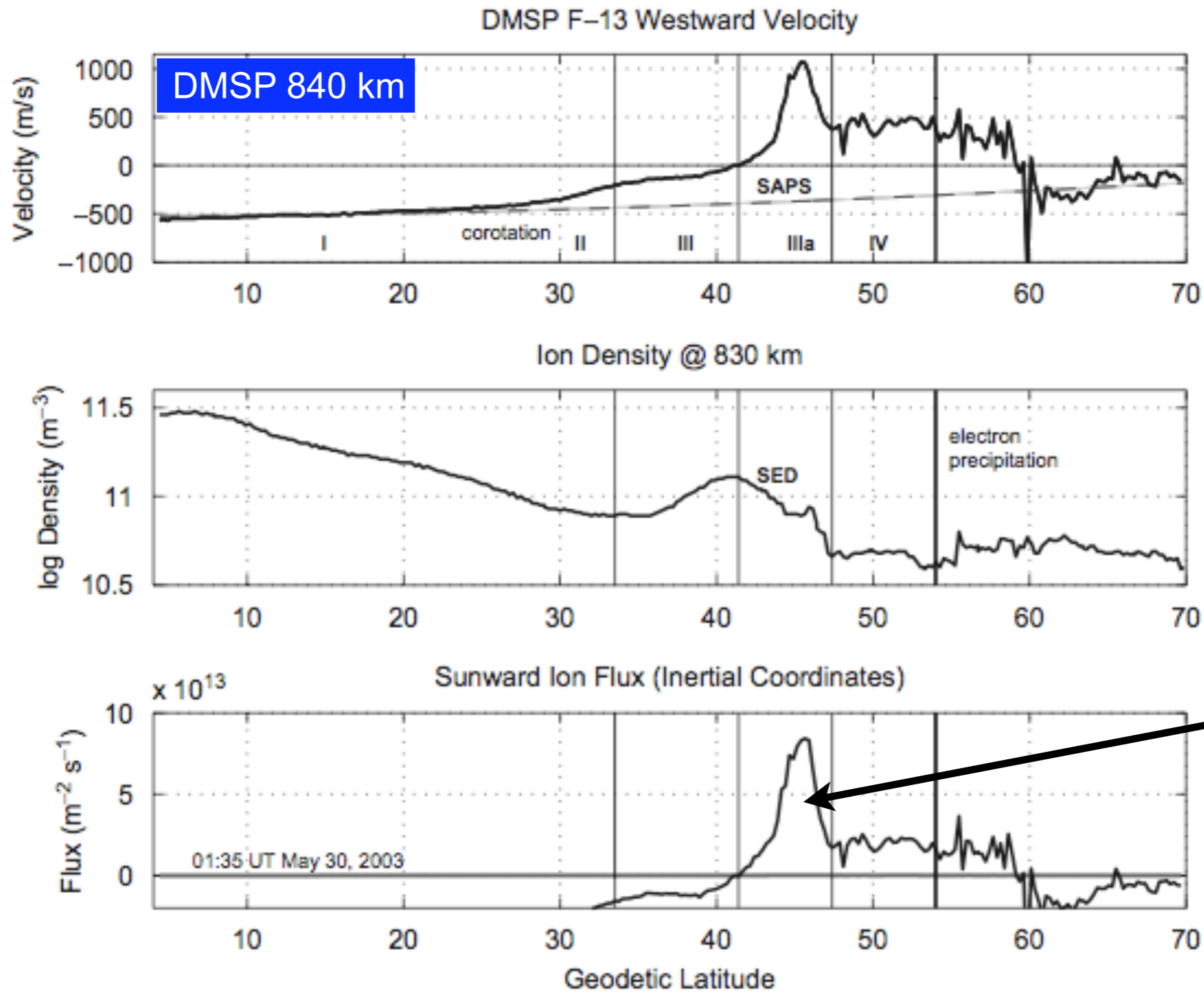
Millstone Hill
Incoherent Scatter Radar:
Wide-Field Access To The
Full Plasma State

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39°52'41.15" N 81°05'52.87" W elev 278 m

Sunward ion flux driven by SAPS



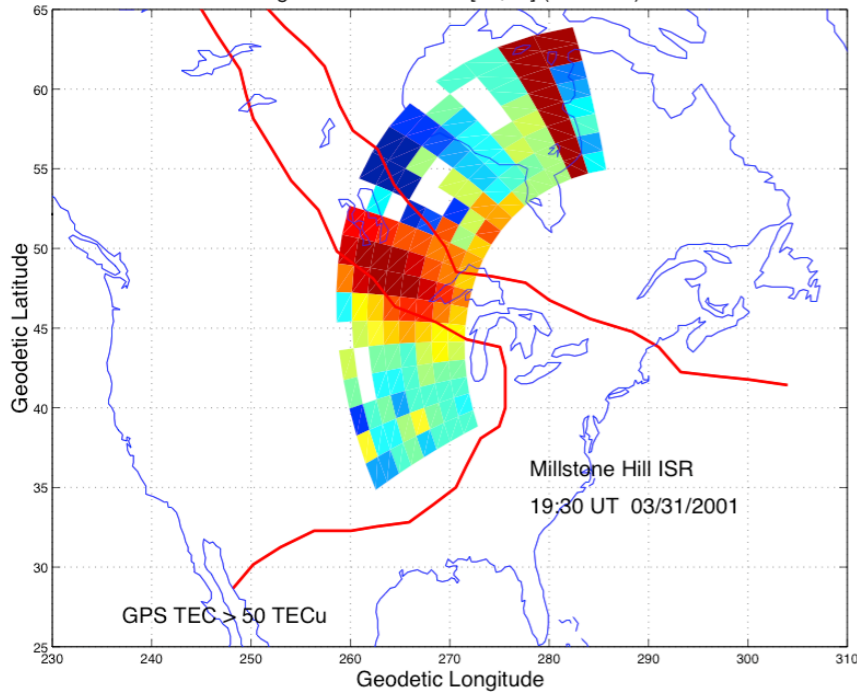
Sunward ion flux caused by SAPS/SED overlap

Foster et al, 2007

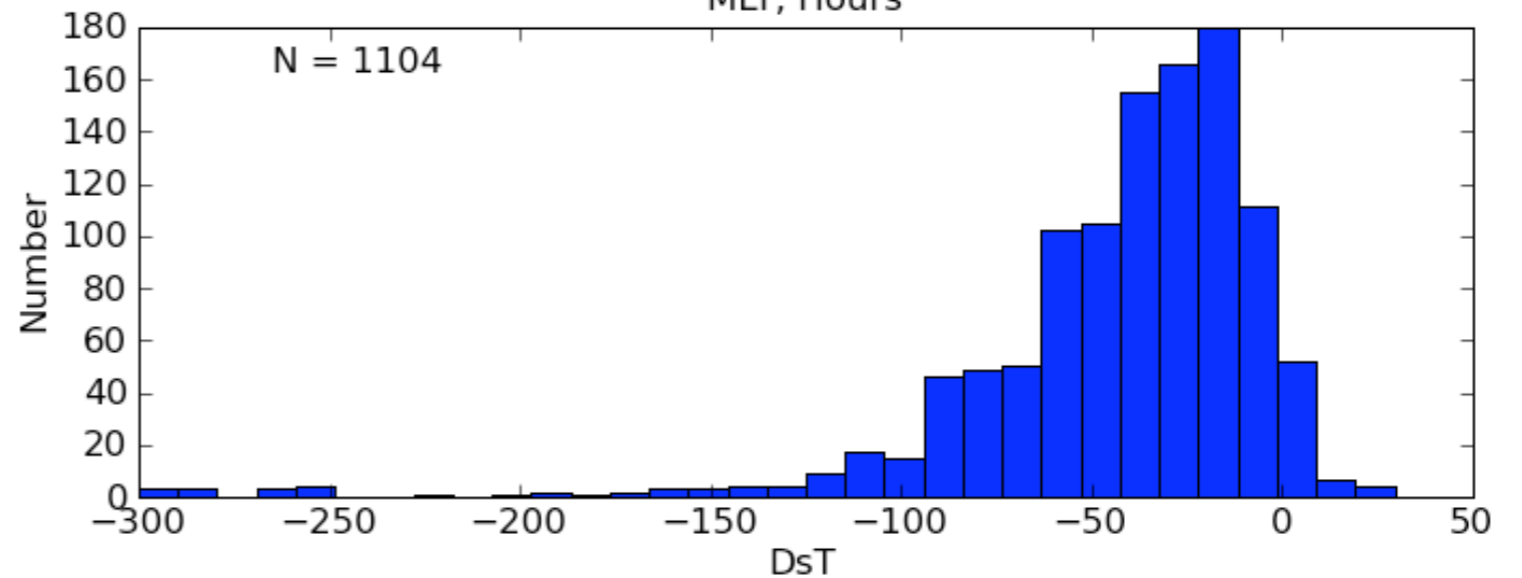
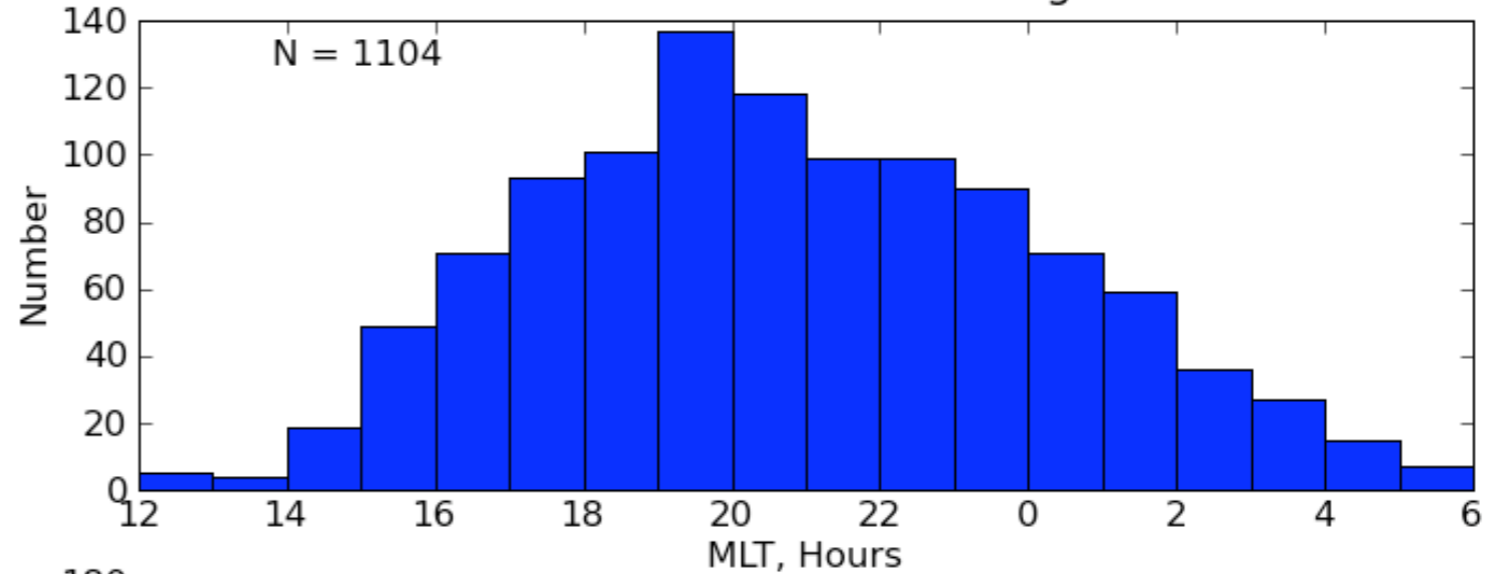
Mid-Latitude Flows: SAPS Statistical Study

Individual events

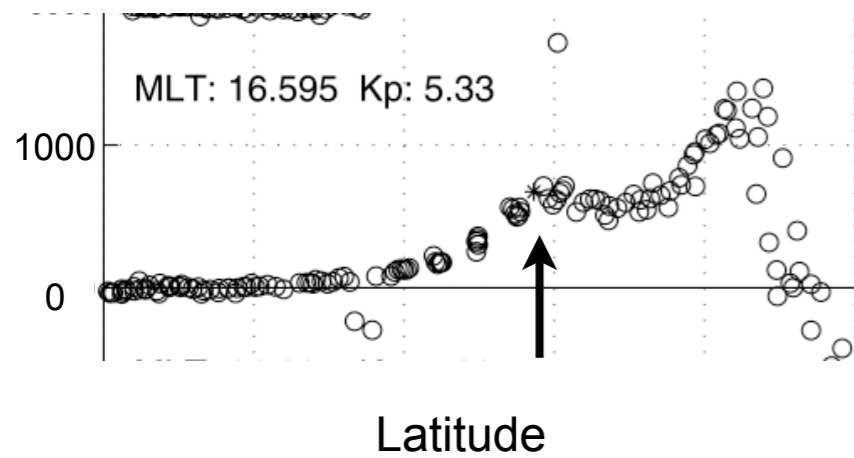
log Sunward Ion Flux [13,15] ($m^{-2} s^{-1}$)



Millstone Hill SAPS Statistical Coverage 1979-2001

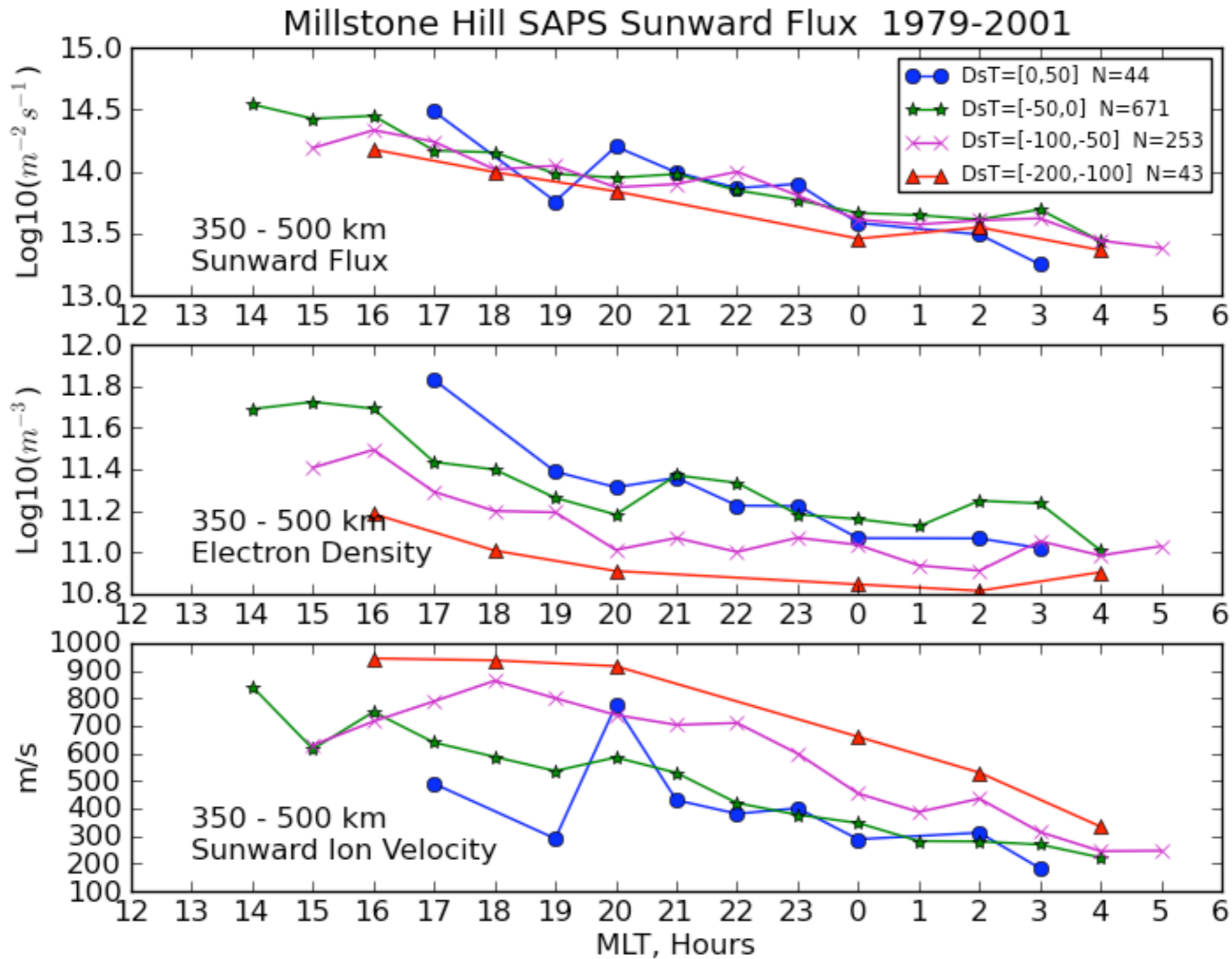


Sunward ion velocity, m/s



Kp = 2 and greater
10,000+ scan database

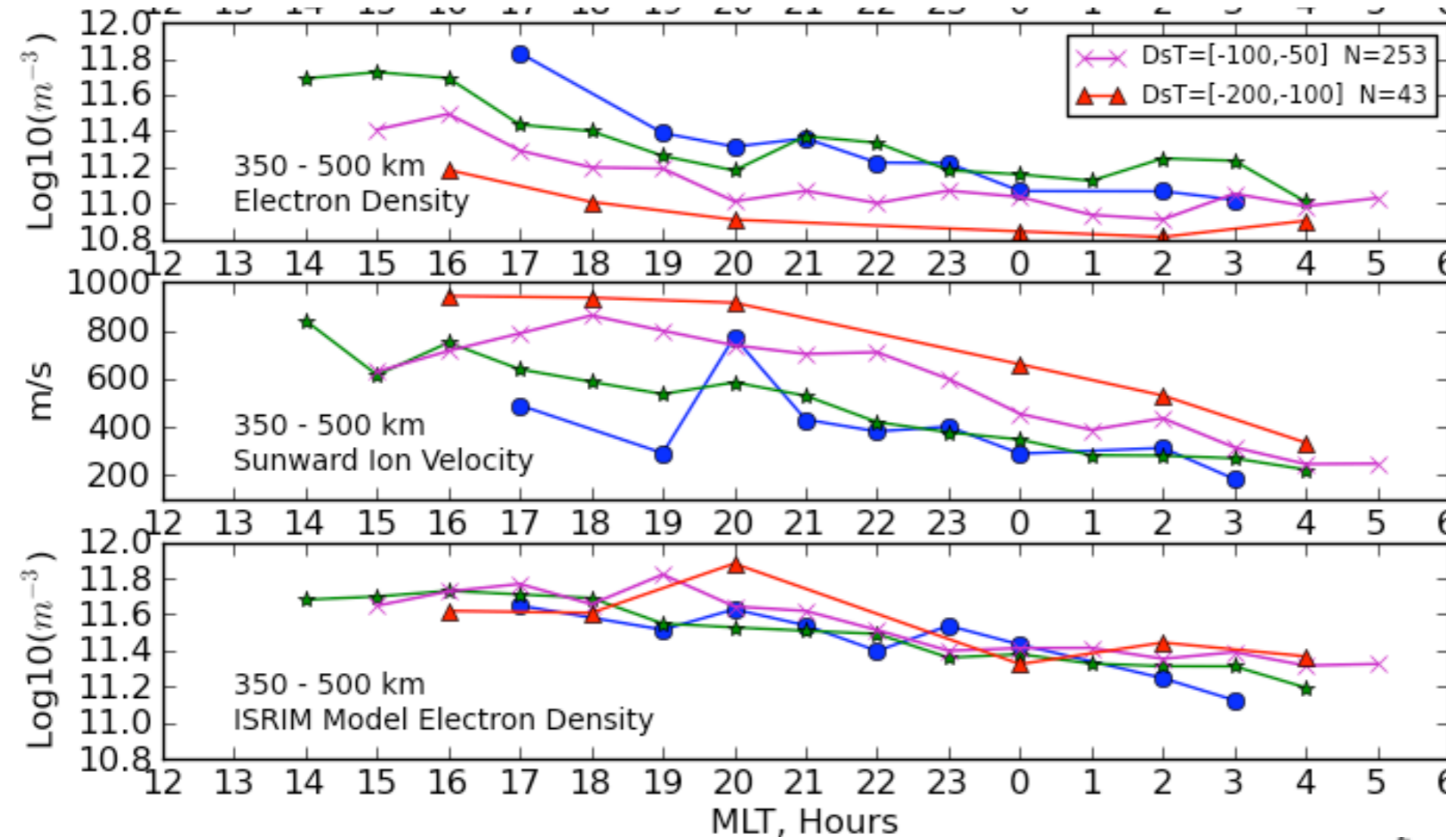
SAPS Flux: Inverse Density/Velocity Relation



Erickson et al
submitted JGR 2010

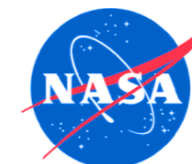
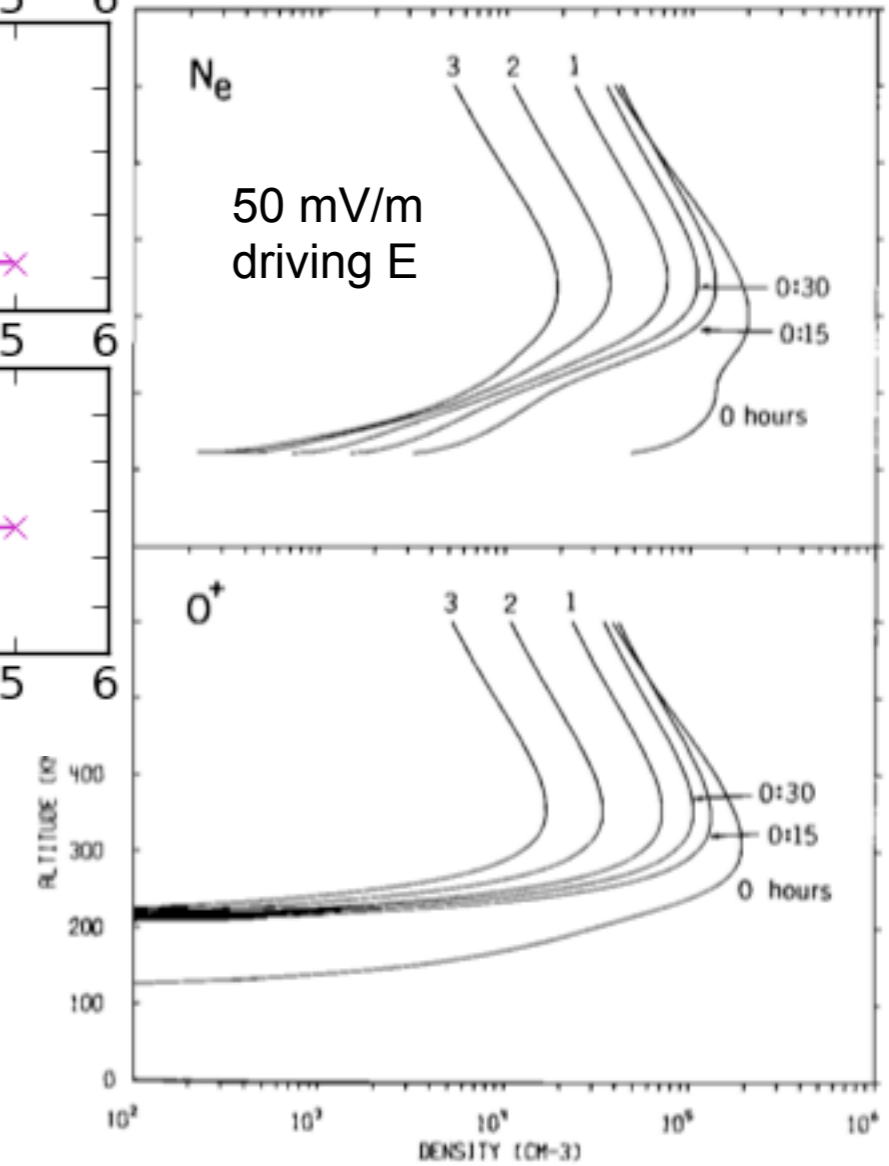


SAPS Flux: Inverse Density/Velocity Relation

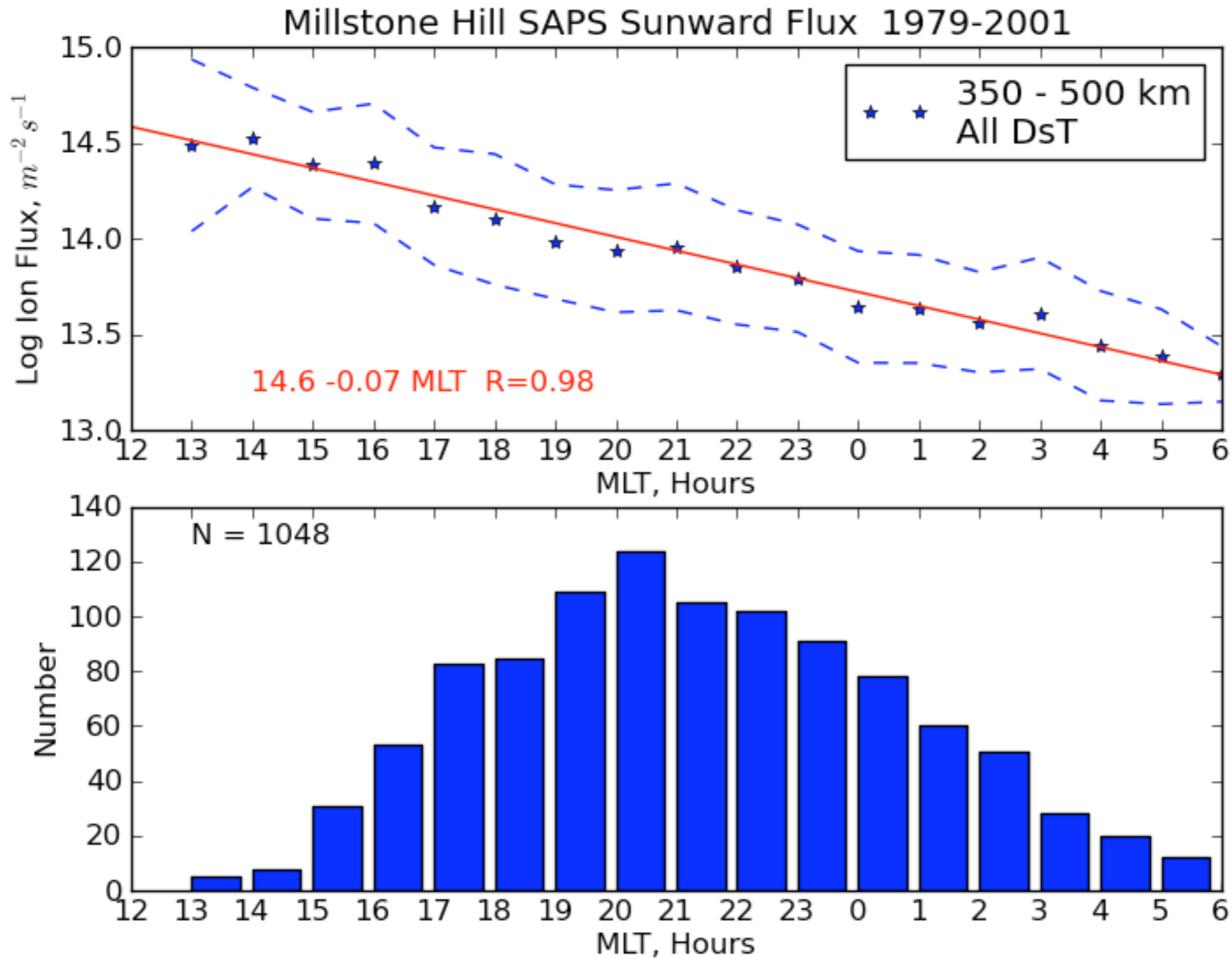


Erickson et al
submitted JGR 2010

Schunk et al 1976

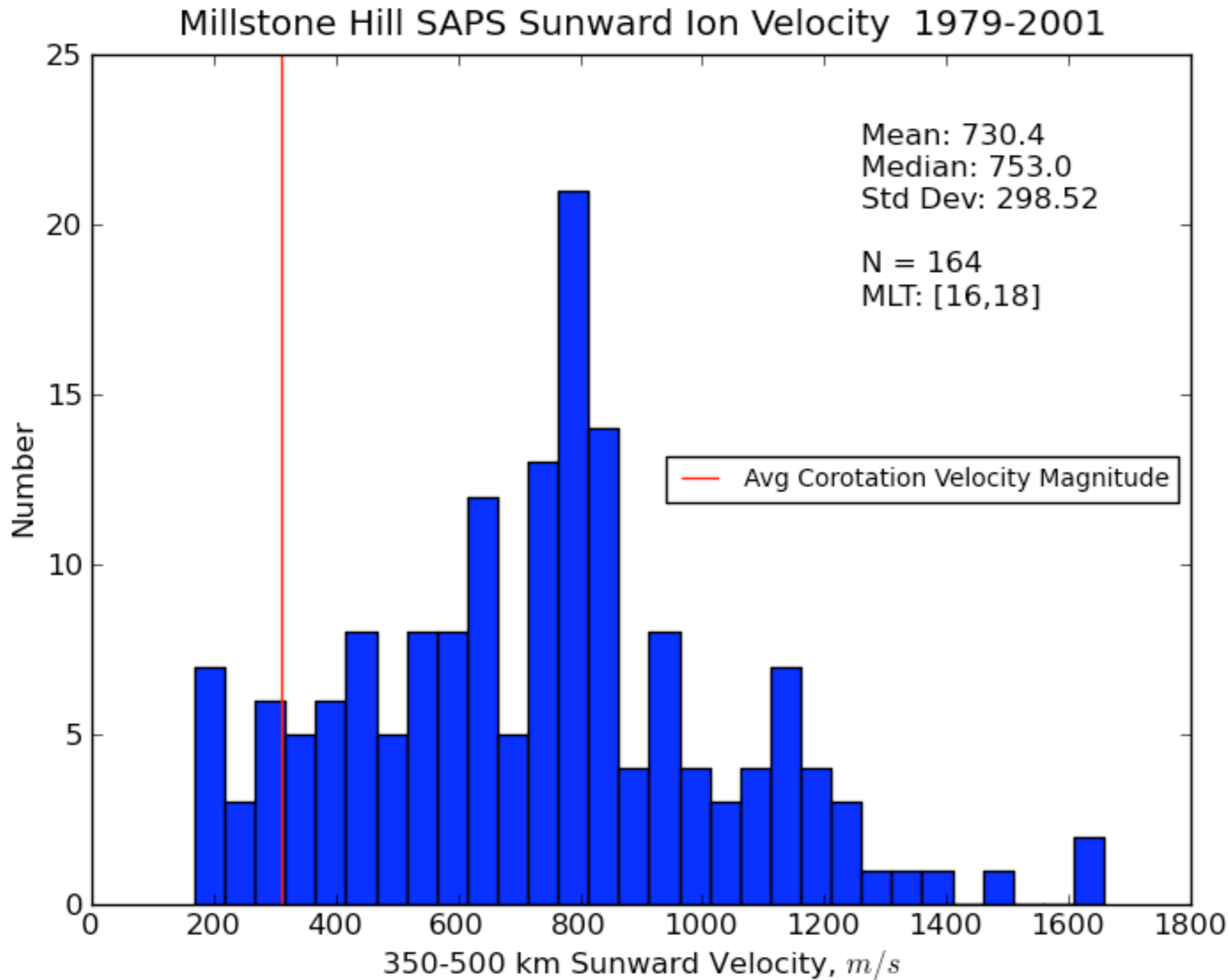


System Regulation: Sunward Flux Invariance



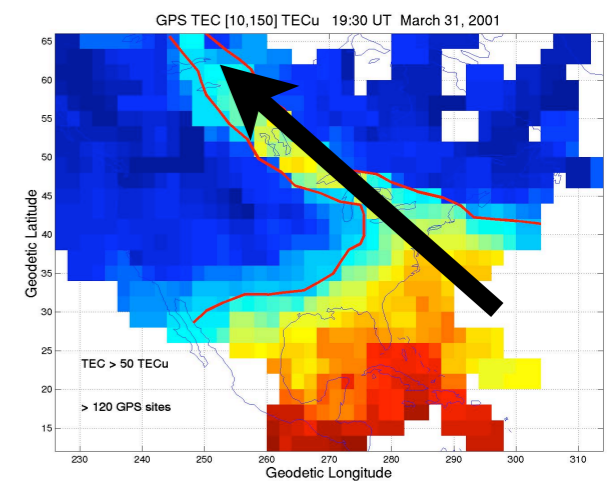
Erickson et al
submitted JGR 2010

System Regulation: Inertially Fixed Flows

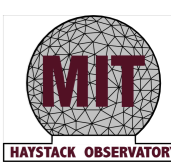


Sunward flow before the terminator moves faster than corotation

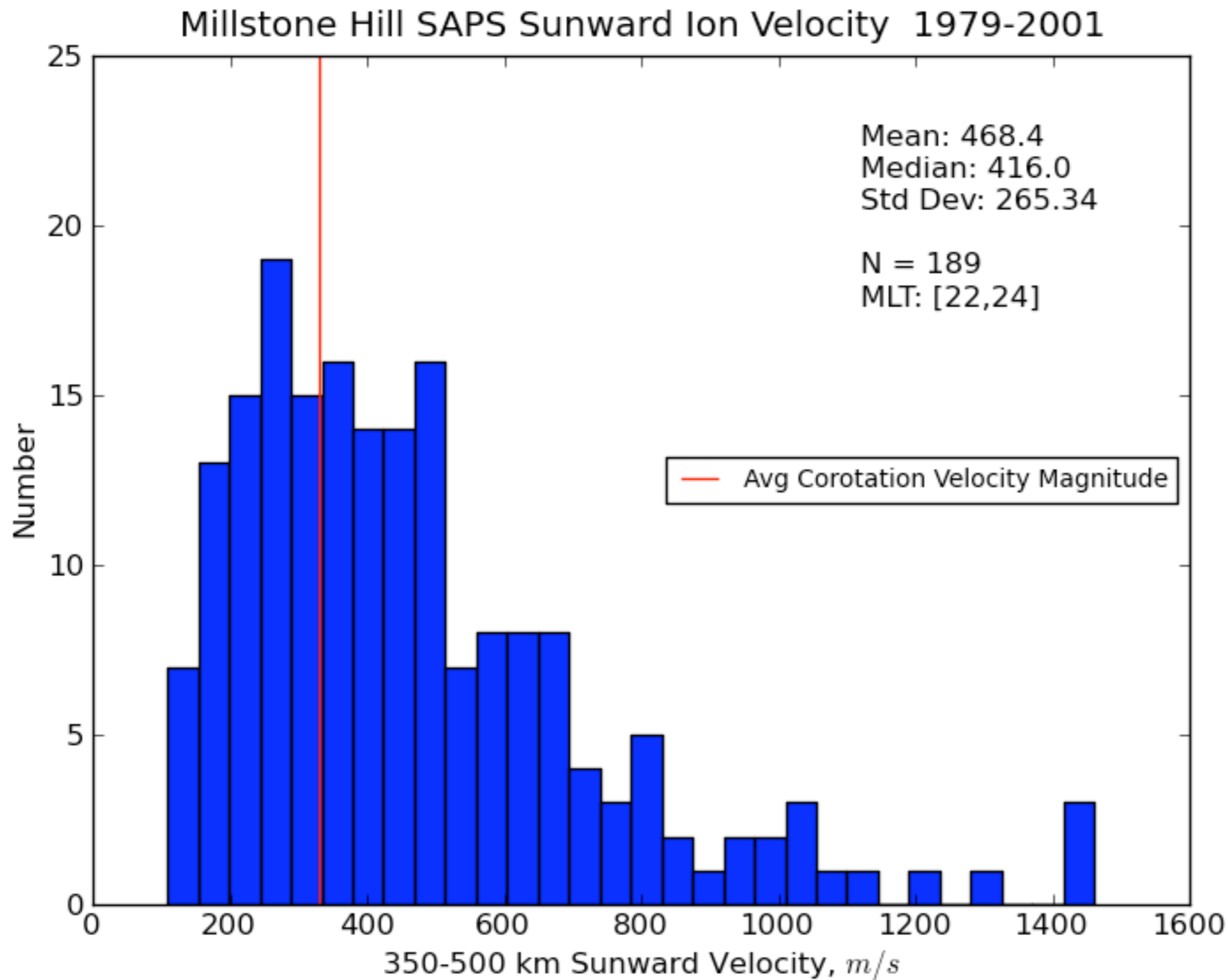
(sweeping material westward to the noon cusp)



Erickson et al
submitted JGR 2010



System Regulation: Inertially Fixed Flows

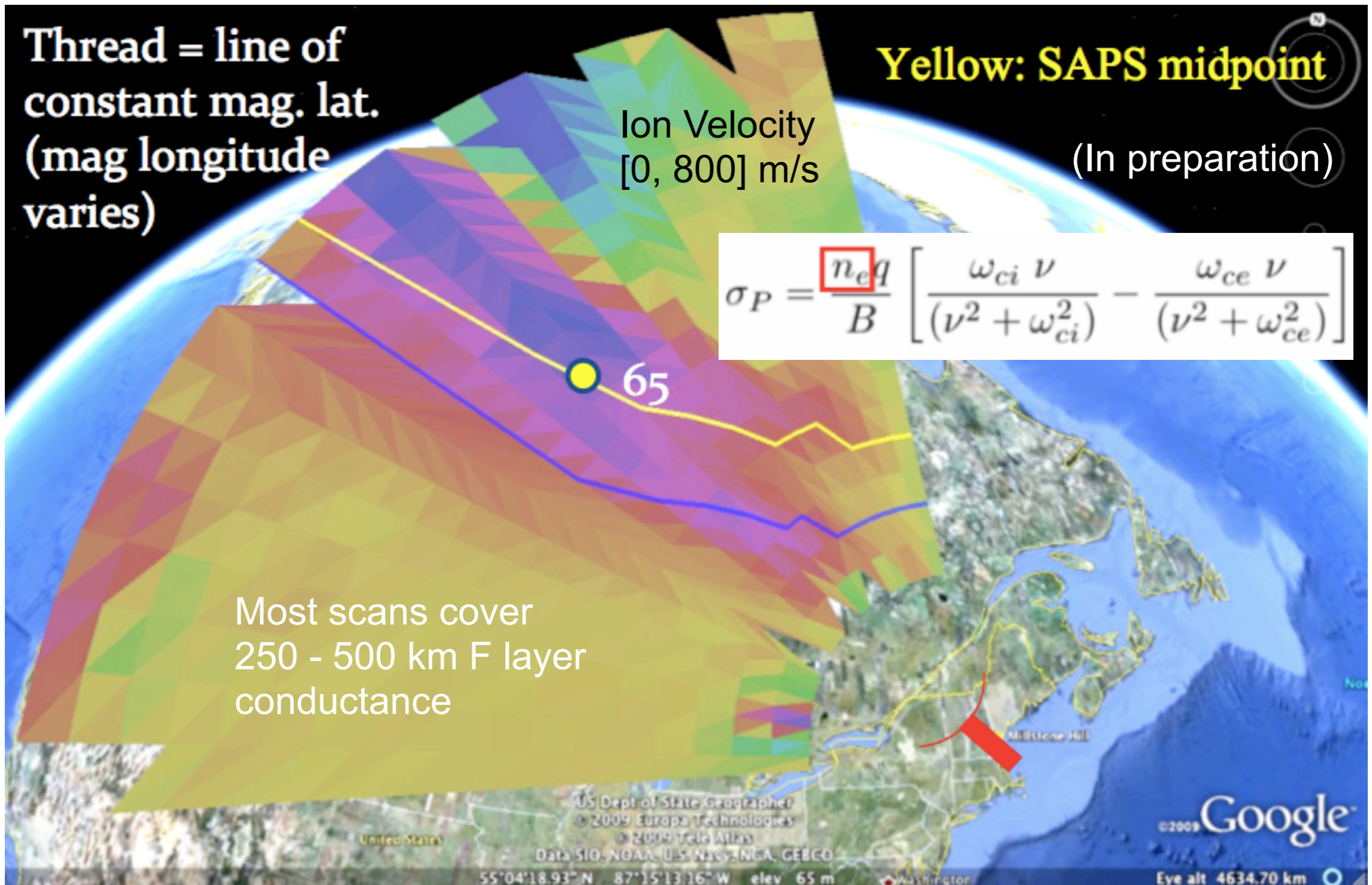


... but after the terminator, flows are nearly opposite corotation

SAPS transport post-sunset is locked inertially at a fixed MLT

Erickson et al
submitted JGR 2010

Field-Aligned Integrated Conductance

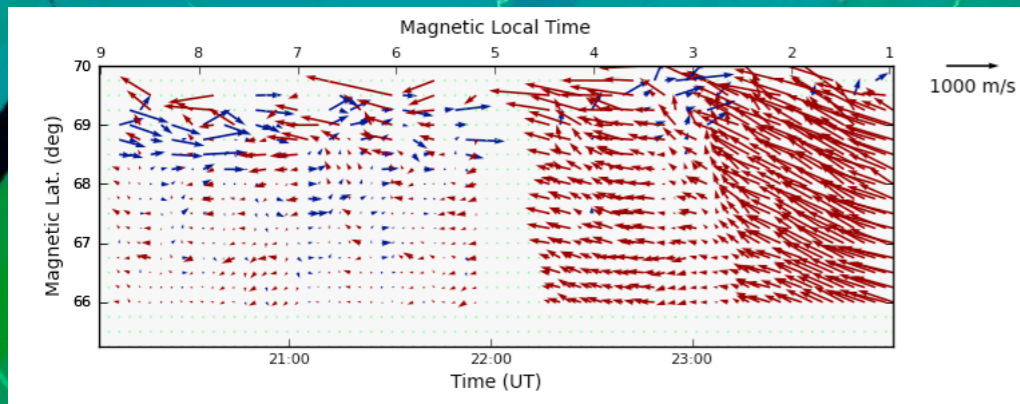


Geospace Observations: DASI

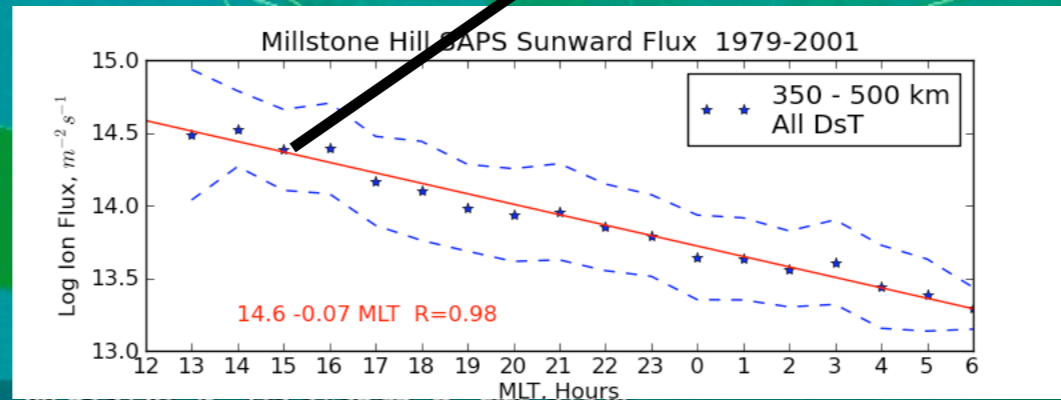
Feb 25, 2010

PFISR
 Veloc Mag
 [0,1000] m/s

Cusp Entry
 Flux > 3E14 m⁻² s⁻¹



(M. Nicolls, SRII)

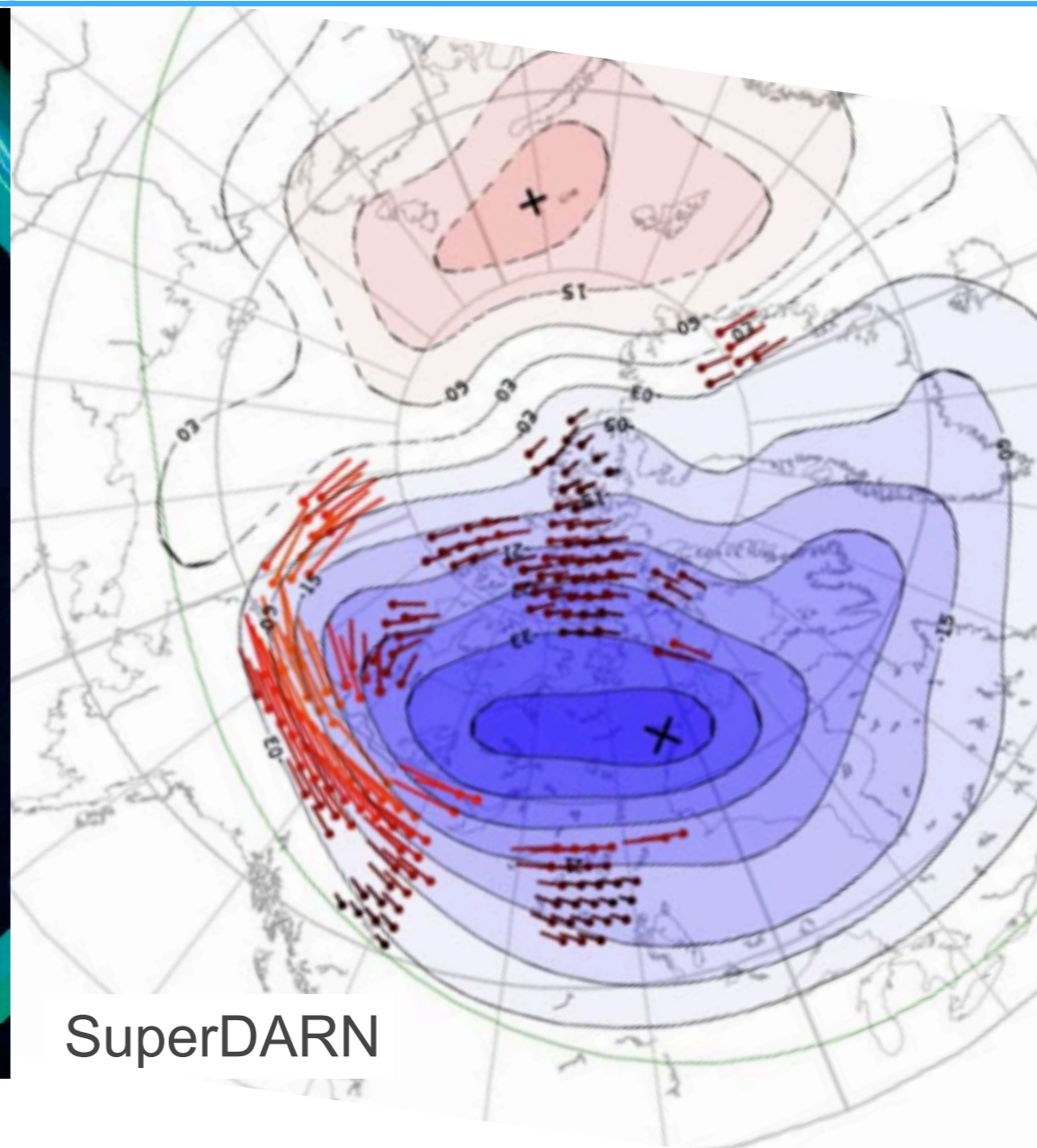
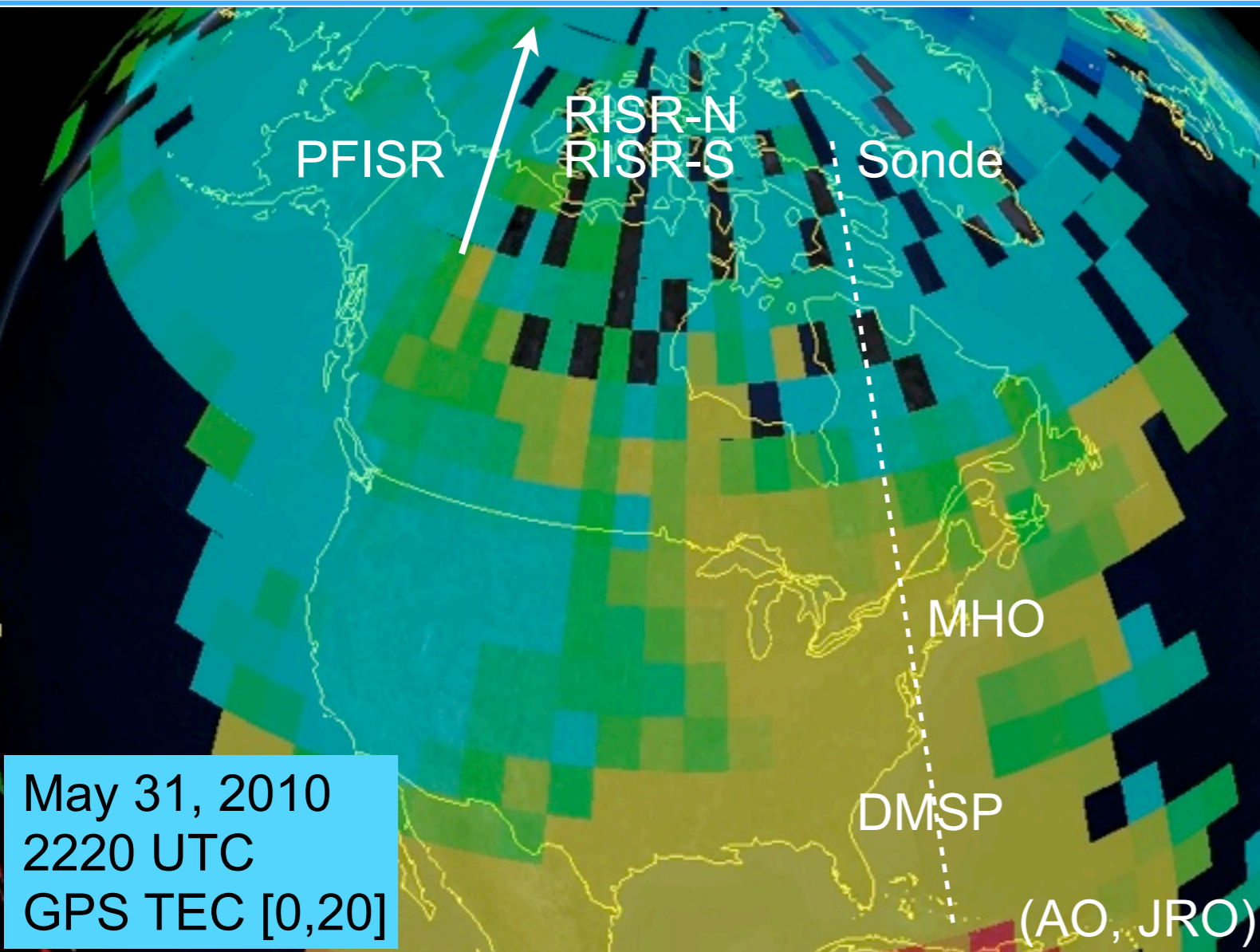


May 31, 2010
 2340 UTC
 GPS TEC [0,20]

©2009 Google

Eye alt 5673.12 km

Geospace Observations: DASI



System Level Responses Require System Level Observations and Science

Summary

- Stormtime ionospheric redistribution is a repeatable phenomenon
- Significant quantities of flux transported horizontally to noontime cusp
- System maintains flux through inverse density/velocity relationship
- System science requires a system and interdisciplinary observational and theoretical perspective

Thanks for your attention!