

Modeling, Specifying and Forecasting Space Weather

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Utah State University

2020 CEDAR Distinguished Lecture
Presented June 2021



Outline

1. What is Space Weather
2. Physics-Based Ionosphere and Thermosphere Models
3. Upper Atmosphere-Ionosphere Weather Models
4. Multi-Model Ensemble Prediction System (MEPS)
5. Ensemble Averaging of Data Assimilation Models

1. Space Weather

Causes of Space Weather:

Solar Weather

Changes in the Solar Wind

Solar Flares

Coronal Mass Ejections

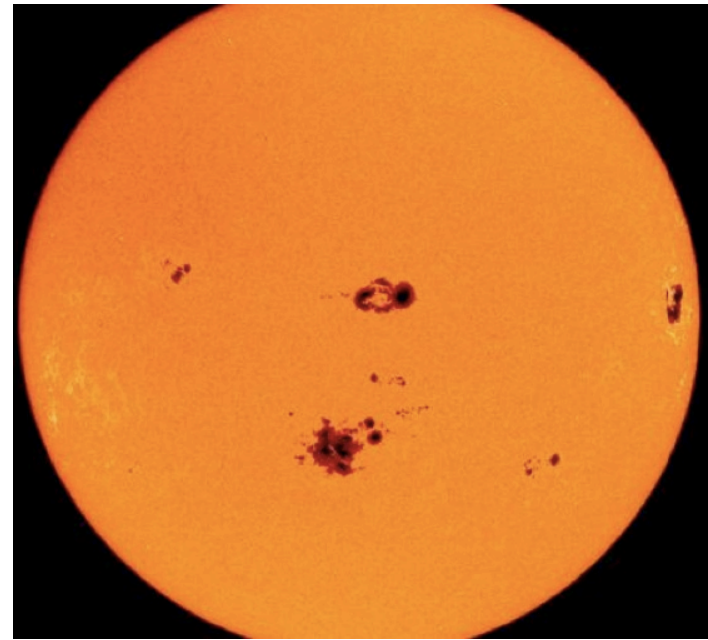
Plasma Instabilities in the Ionosphere

Waves from the Lower Atmosphere

Sunspots



Quiet

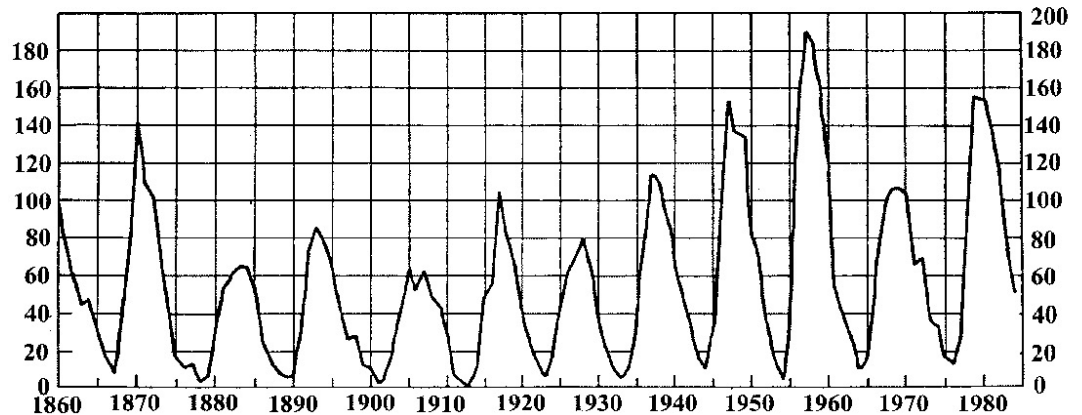
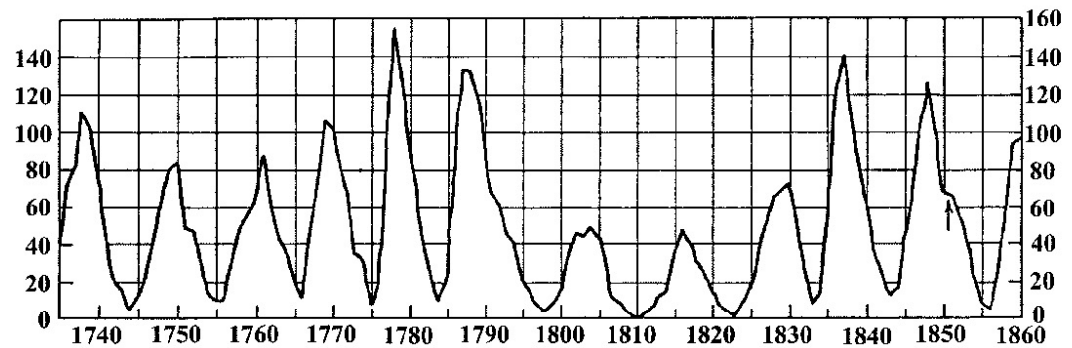
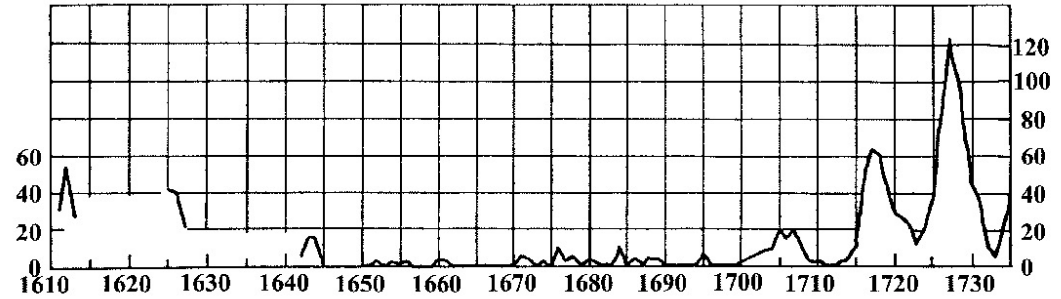


Active

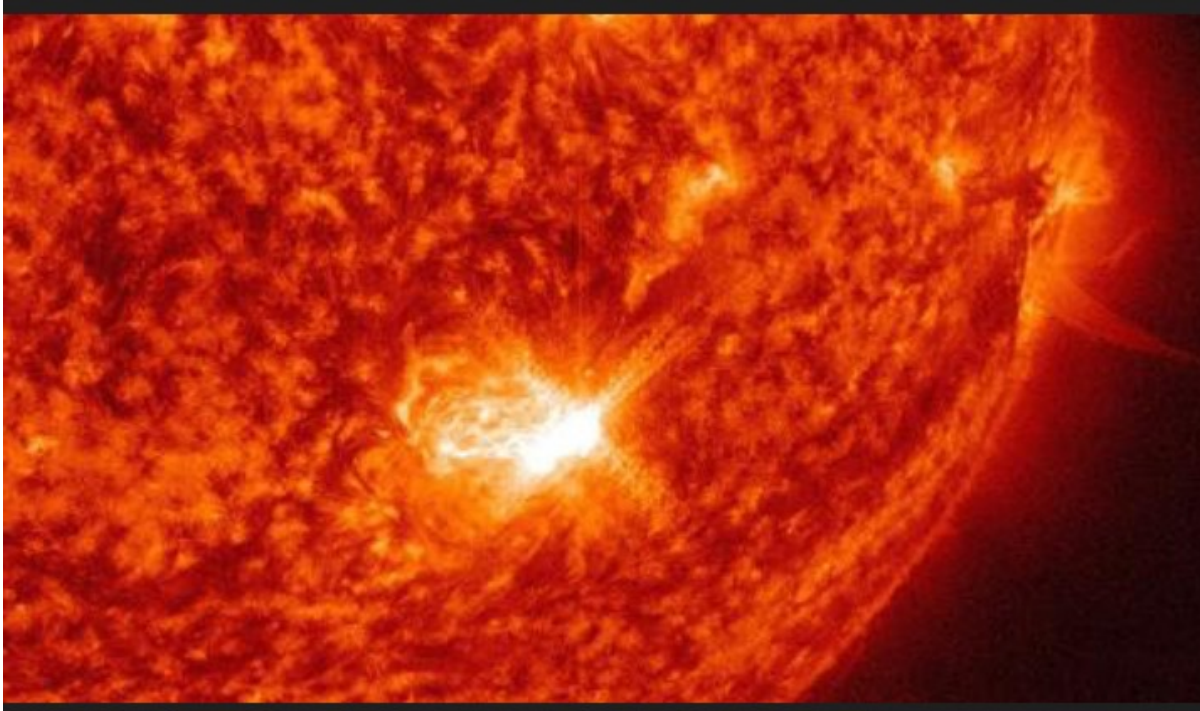
Dark regions in the Sun's lower atmosphere
Cooler than the surrounding region
Associated with Space Weather disturbances

Sunspot Numbers

- 11-Year Solar Cycle
- Annual Averages
- Maunder Minimum
Period 1645-1715



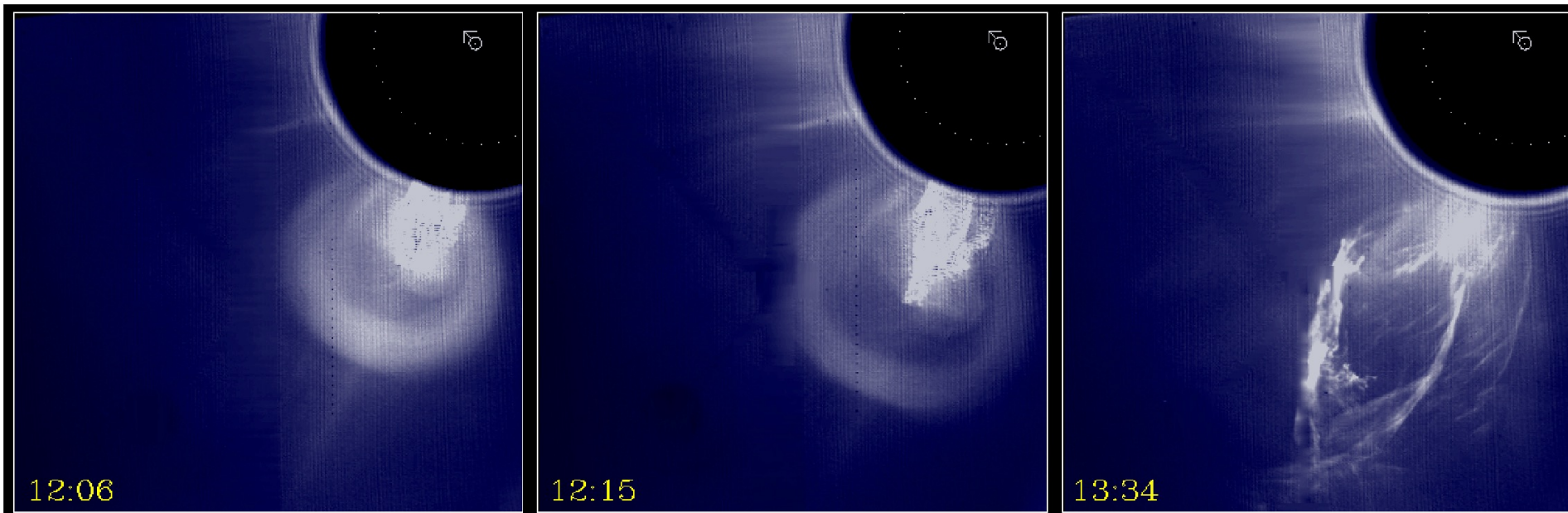
Solar Flare



**A sudden explosion of intense electromagnetic radiation
from the Sun's surface**

Can enhance Ionospheric densities in the D-region

Can affect radio waves and communications



Coronal Mass Ejection

Speed = 1000 km/s (2 million mph)

Mass = 100 billion kg (220 billion lbs)

Explosion = 1 billion hydrogen bombs

Solar Wind



Earth location = 217 Solar Radii from Sun (150 million km)

Light takes 500 sec to reach Earth

Solar Wind takes 2-3 days to reach Earth and at Earth:

Speed = 200-900 km/s

Density = 1-80 cm⁻³

Temperature = 100,000 K

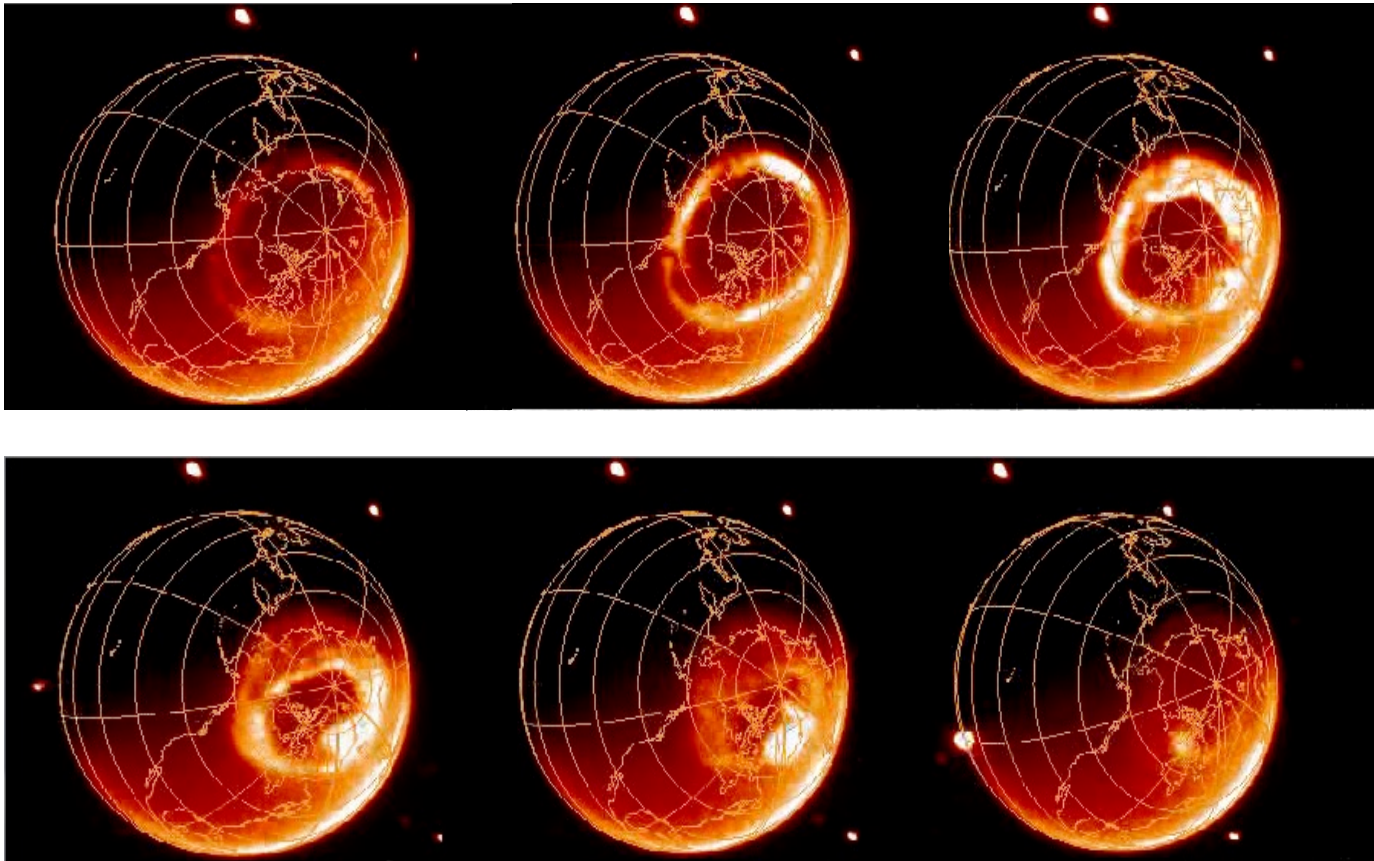
Satellite Images



- Ground Photograph by Jan Curtis.
- FUV Image from the IMAGE Satellite.



Satellite Images



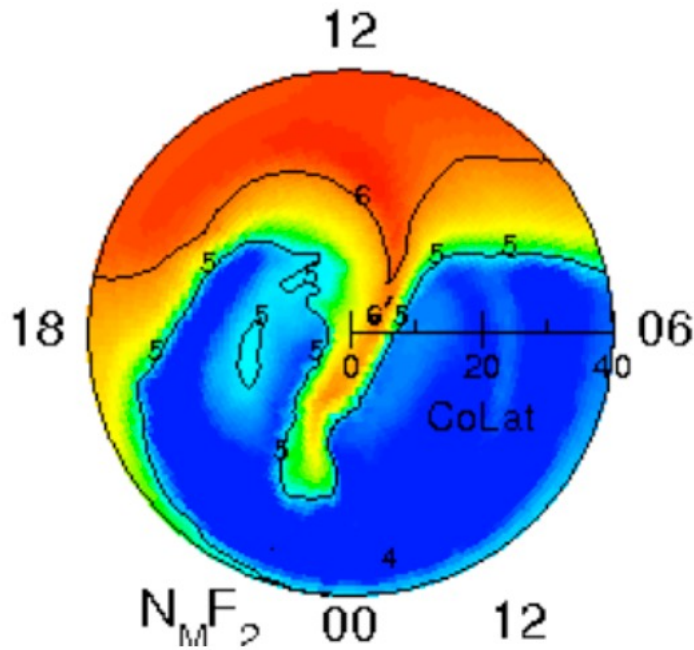
Burch, J. L., *Scientific American*, 284, 72-80, 2001

- **Bastille Day Storm**
- **July 14-15, 2000**
- **Snapshots During a 1-Hour Period**



Space Weather – High Latitudes

Tongue of Ionization



Winter Polar Region

Southward Interplanetary Magnetic Field

Forms in sunlight in the cusp region

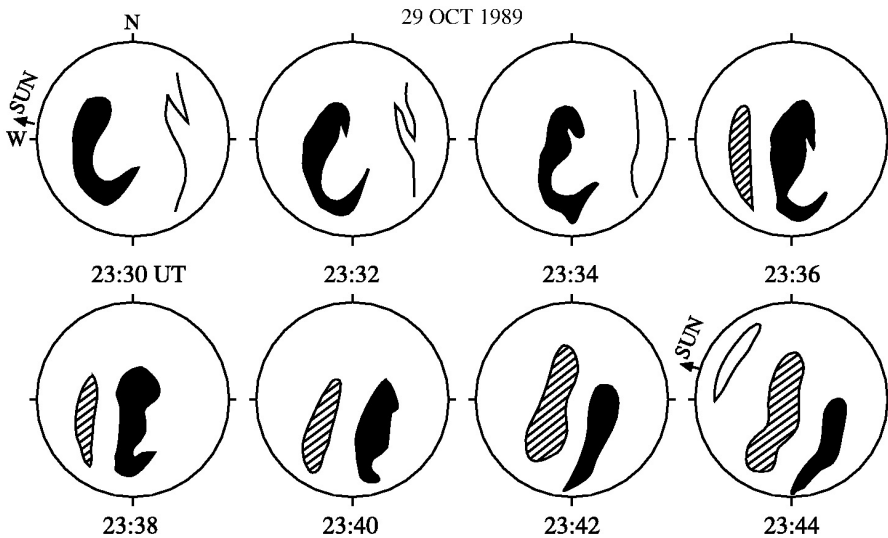
Anti-sunward plasma convection extends the high-density plasma across the dark polar cap, forming the Tongue

Density of the Tongue can be more than a factor of 2 higher than the background plasma density



Space Weather – High Latitudes

Propagating Plasma Patches



Propagating plasma patches observed at Qaanaaq, Greenland, on October 29, 1989. The dials represent a digitization of all-sky images (630-nm) taken at 2-minute intervals.

Winter Polar Region

Southward Interplanetary Magnetic Field

Forms in sunlight equatorward of the auroral zone

Patches can be circular or cigar-shaped

Patch density can be up to a factor of 100 above background plasma density

Plasma patches drift in an anti-sunward direction with the background plasma

Figure shows cigar-shaped plasma patches drifting in an anti-sunward direction across Greenland

Dimensions of the patches are 200 x 1000 km

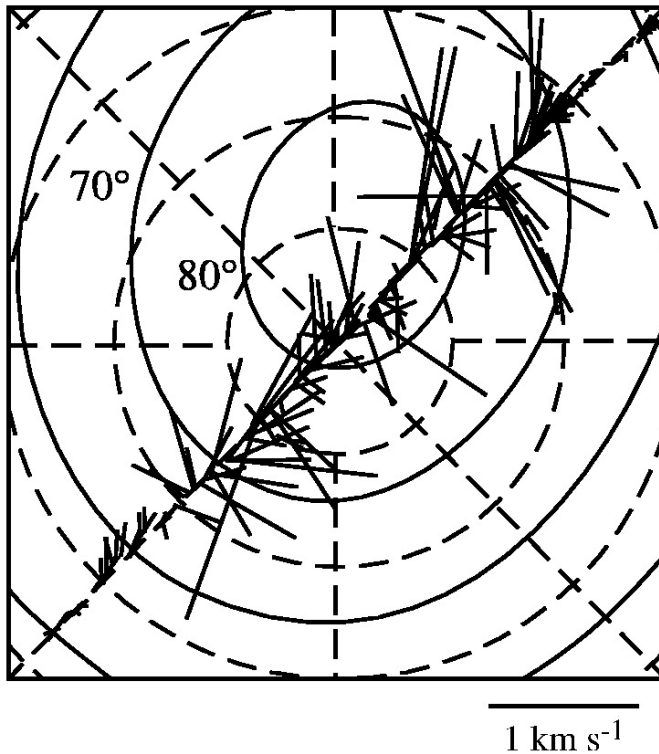
Velocity of the patches is about 730 m/s



Space Weather – High Latitudes

Plasma Drift Velocities

ION DRIFT METER, DE-2
UNIVERSITY OF TEXAS AT DALLAS
OCTOBER 17, 1981
1634-1646 UT



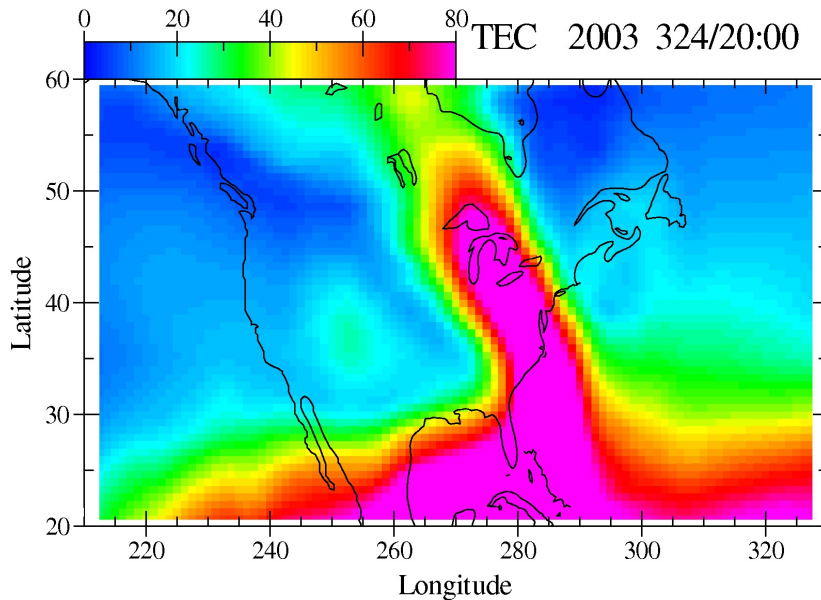
- Northern Polar Region
- Northward Interplanetary Magnetic Field
- Plasma Drift Velocities
- Pattern appears to be turbulent
- Satellite traversal was only 12 minutes
- Probably spatial structure
- Nine reversals in the flow direction



Space Weather – Mid-Latitudes



Snowstorm in the Lower Atmosphere

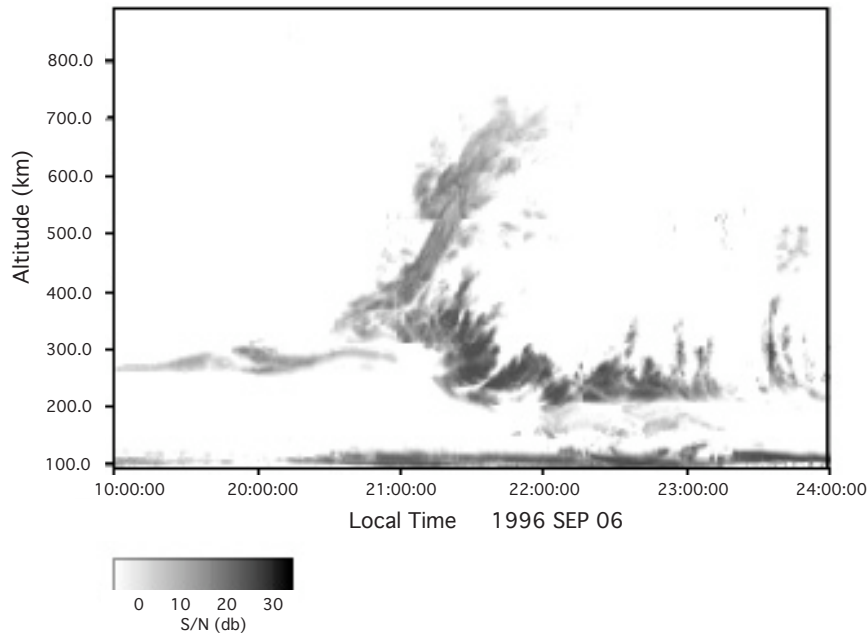


Solar Storm in the Upper Atmosphere-Ionosphere



Spread F/Equatorial Bubbles

JULIA Coherent Scatter Radar



Spread-F event seen by the JULIA coherent scatter radar near Lima, Peru, on September 6, 1996. Plot of coherent backscatter signal-to-noise ratios versus time.

Plasma instabilities in F-region can lead to Spread-F

In the Equatorial region, spread-F can lead to Plasma Bubbles

Bubble wedge region extends north-south along B-field with an apex altitude as high as 1500 km
East-west extent of the bubble wedge can be several 1000 km



2. Physics-Based Ionosphere and Thermosphere Models

Time Dependent Ionosphere Model (TDIM)

Global Physics-Based Ionosphere Coupled to Global Empirical Thermosphere (MSIS Model)

High Latitudes -- Convection E-Field and Auroral Precipitation Specified

Equatorial E-Field Specified

Ionospheric Forecast Model (IFM)

Same as TDIM but Modified to Run Faster Than Real Time

Ionosphere Plasmasphere Model (IPM)

Global Physics-Based Ionosphere-Plasmasphere Coupled to Global Empirical Thermosphere

IPM-Global

Global Physics-Based Ionosphere-Plasmasphere Extended to Include High Latitudes

Ionosphere at High Latitudes Similar to TDIM but Includes Polar Wind Outflow

Global Thermosphere Model (GTM)

Global Physics-Based Thermosphere can be Coupled to Global Empirical, Physics-Based, or Data Assimilation Ionosphere

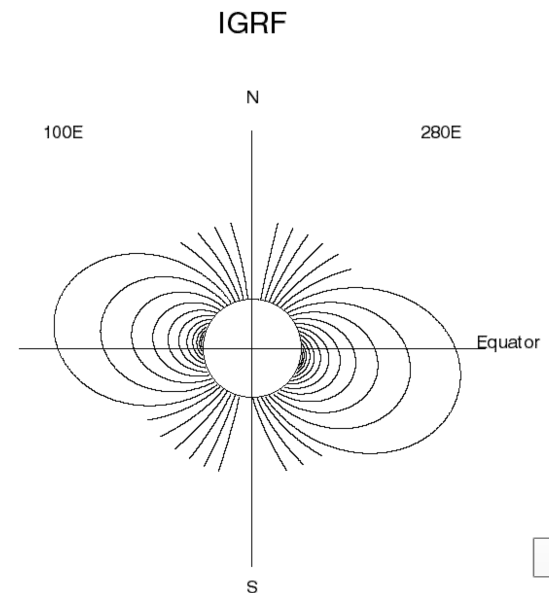
Ionosphere – Plasmasphere Model (IPM)

- 90-30,000 km
- Altitude, Latitude, Longitude Grids Set by User
- Six Ion Species (NO^+ , O_2^+ , N_2^+ , O^+ , H^+ , He^+)
- Realistic Magnetic Field (IGRF)
- Some of the Physical Processes included in IPM:
 - Field-Aligned Diffusion
 - Cross-Field Electrodynamical Drifts
 - Thermospheric Winds
 - Neutral Composition Changes
 - Energy-Dependent Chemical Reactions
 - Ion Production due to:
 - Solar UV/EUV Radiation
 - Auroral Precipitation
 - Star Light

IPM-Global

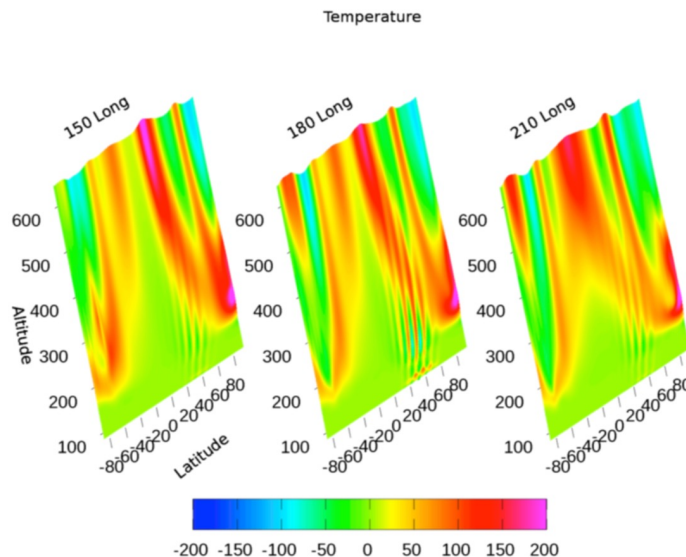
- The Model is Composed of an Ionosphere-Plasmasphere Model (IPM) that Covers Low and Mid-Latitudes and an Ionosphere-Polar Wind Model that Covers High Latitudes.
- 90-30,000 km at Low-Middle Latitudes
- 90-10,000 km at High Latitudes
- Altitude, Latitude, Longitude Grids Set by User
- Output Parameters:
 - $\text{NO}^+, \text{O}_2^+, \text{N}_2^+, \text{O}^+, \text{H}^+, \text{He}^+$
 - T_e, T_i
 - $u_{||}, u_{\perp}$

As the High Latitude Region Expands and Contracts due to Geomagnetic Storms, the B-field Lines Open and Close Accordingly



Global Thermosphere Model (GTM-Ionosphere)

- 40-800 km
- Altitude, Latitude, Longitude Grids Set by User
- Neutrals (N_2 , O_2 , O)
- Non-hydrostatic, Non-linear Flows
- Planetary, Tidal, Gravity, and Sound Waves
- Subsonic, Transonic and Supersonic Winds
- Wave Breaking in the Lower Thermosphere
- Can be Coupled to a Global Empirical, Physics-Based, or Data Assimilation Ionosphere



GTM-Ionosphere simulation of an equatorward propagating Traveling Atmospheric Disturbance interacting with an upward propagating gravity wave from the lower atmosphere (*Gardner and Schunk, 2011b*).



3. Upper Atmosphere-Ionosphere Weather Models

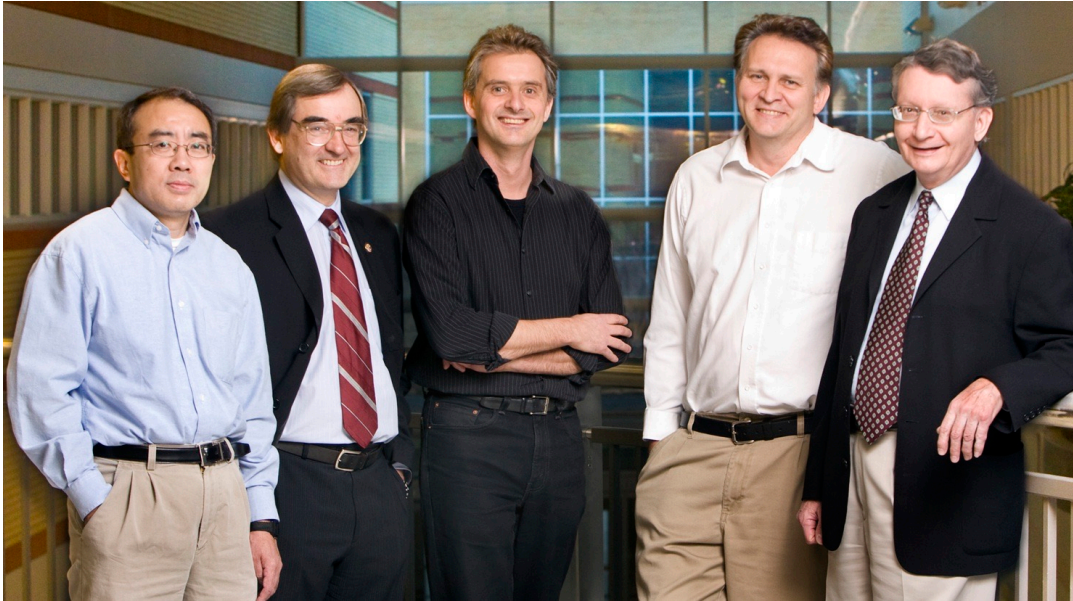
Global Assimilation of Ionospheric Measurements (GAIM)

R. W. Schunk, L. Scherliess, L. C. Gardner,
D. C. Thompson, V. Eccles, J. J. Sojka, and L. Zhu

Similar to Tropospheric Weather Models



GAIM Team



Zhu, Sojka, Scherliess, Thompson, Schunk, Eccles, Gardner



USU Physics-Based Data Assimilation Models

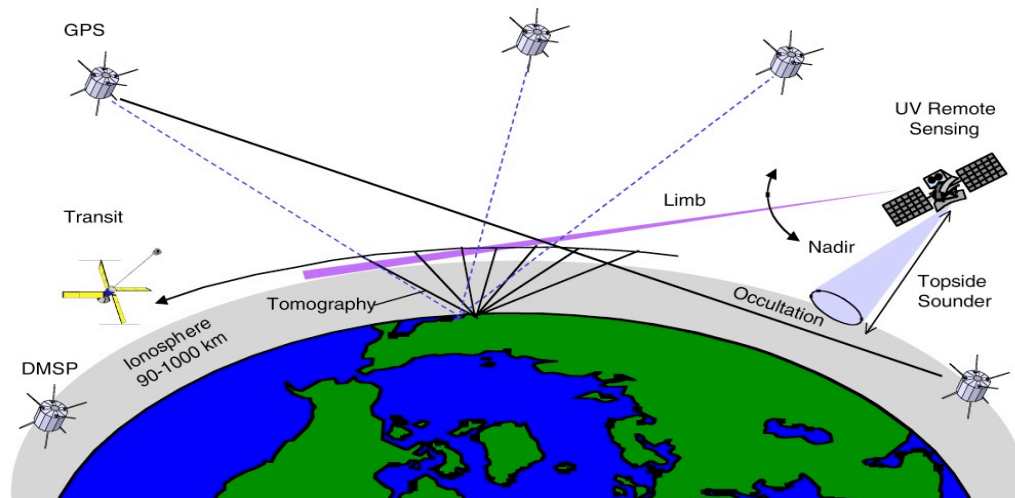
Global Assimilation of Ionospheric Measurements

- **Gauss-Markov Model (GAIM-GM)**
Air Force Operational Model 2006 – 2020
- **Full Physics Model (GAIM-FP)**
Air Force Operational Model 2020

Ensemble Kalman Filter Model for High-Latitude Ionosphere Dynamics & Electrodynamics (IDED-DA)

Currently a Science Model

GAIM-GM and GAIM-FP Assimilate Multiple Data Sources



Data Assimilated Exactly as They Are Measured

- Bottomside N_e Profiles from Digisondes (200)
- Slant TEC from more than 1000 Ground GPS Receivers
- N_e Along Satellite Tracks (4 DMSP satellites)
- Integrated UV Emissions (LORAAS, SSULI, SSUSI, TIP)
- Occultation Data (CHAMP, IOX, SAC-C, COSMIC, C/NOFS)
- Data Assimilated at Low and Mid-Latitudes

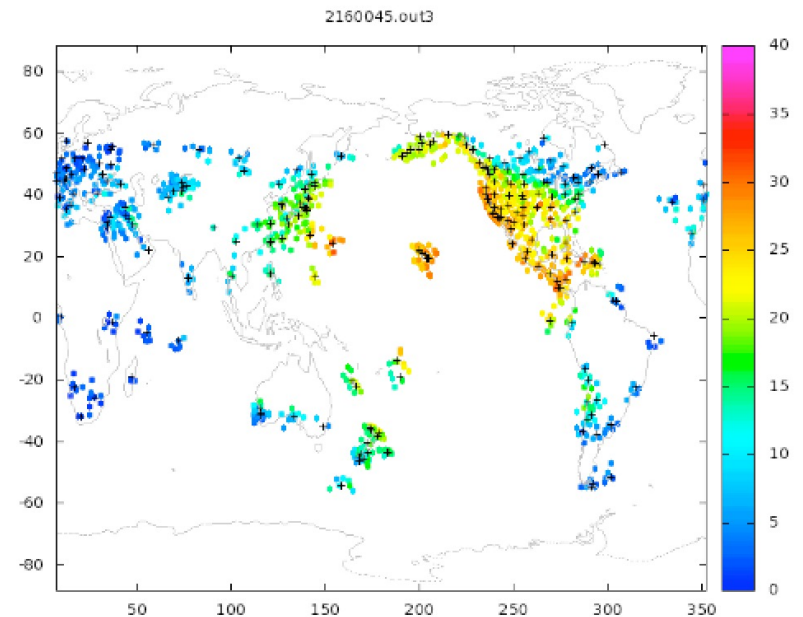
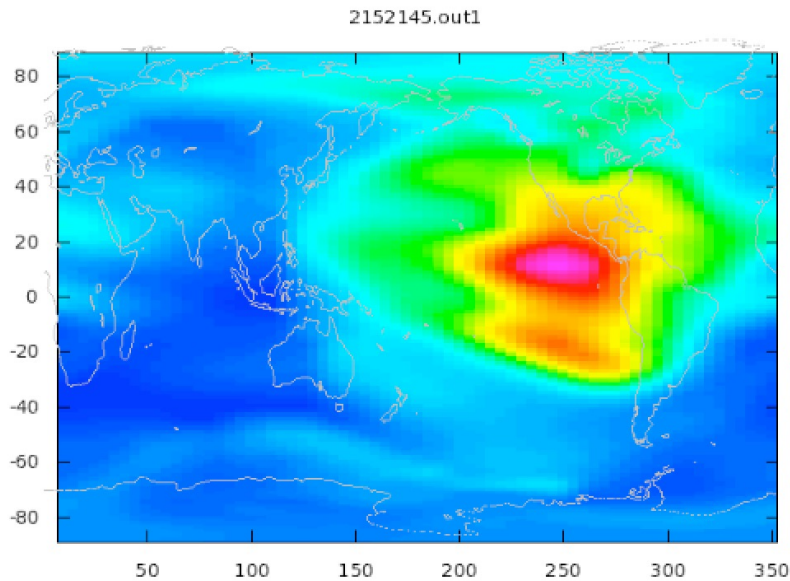
GAIM-GM Model

- Ionosphere Forecast Model (IFM)
- Global physics-based model for background ionosphere
- 90 - 1400 km
- 15 - minute output cadence
- O^+ , H^+ , NO^+ , N_2^+ , O_2^+ , T_e , T_i
 - Only uses N_e
- Kalman Filter solves for deviations from background electron density distribution
- GAIM-GM does not provide information about ionospheric drivers (electric fields, neutral winds, etc.)



GAIM-GM **global** Run

- 357 global TEC stations (IGS network) used in real-time at USU Space Weather Center
- Up to 10,000 measurements assimilated every 15- min



Vertical TEC versus Geographic Latitude (left) and Longitude (right)

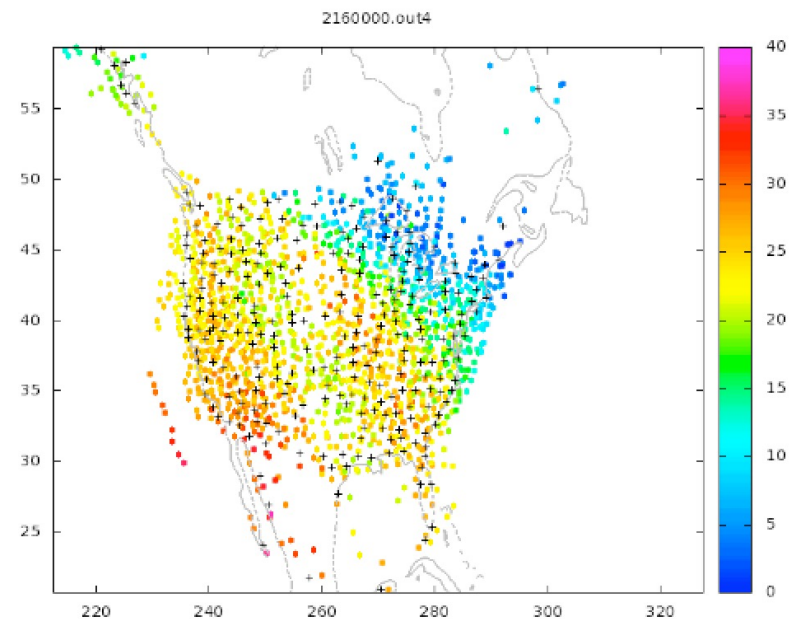
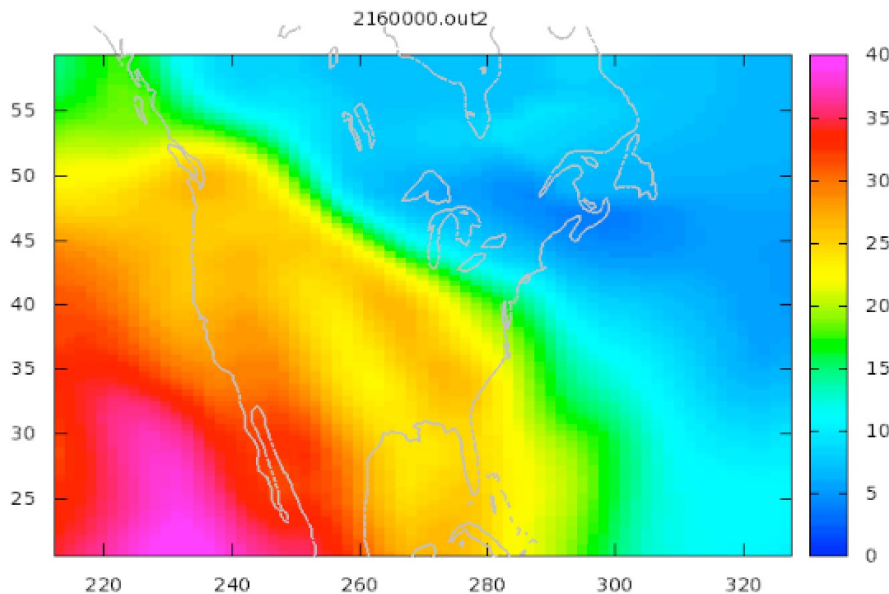


"By joining The Pieces Together"



GAIM-GM regional (High Resolution) Run:

- 424 USTEC stations (CORS network) used in real-time at USU Space Weather Center
- Up to 10,000 measurements assimilated every 15-min



Vertical TEC versus Geographic Latitude (left) and Longitude (right)



"Bring the Pieces Together"



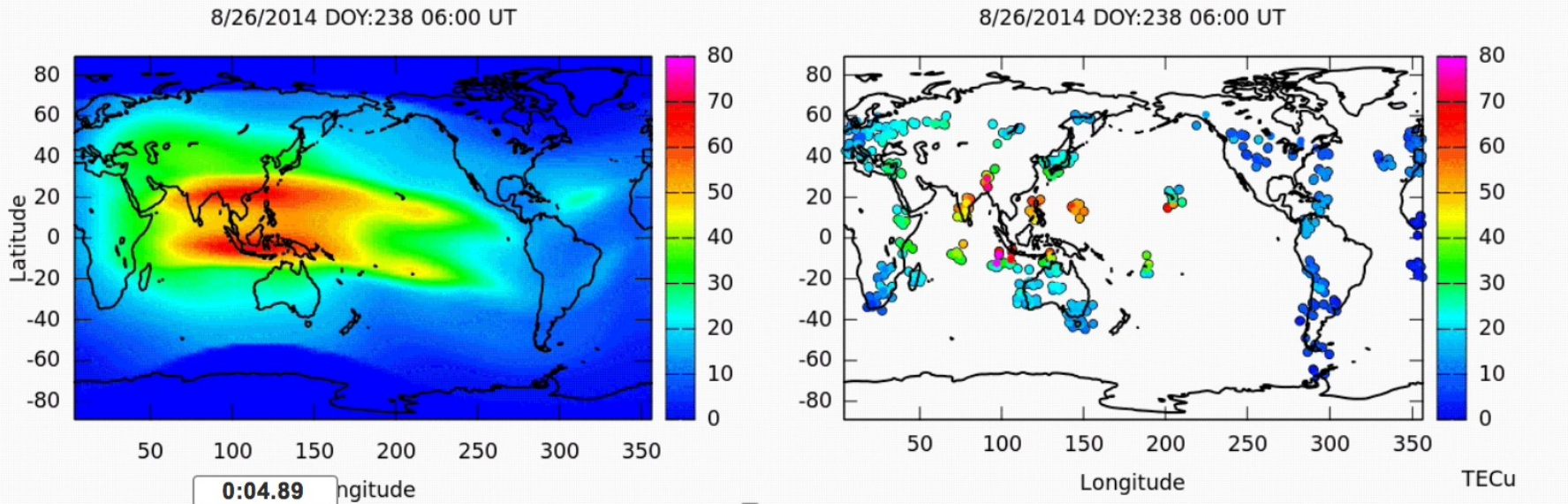
GAIM-FP

- Ensemble Kalman Filter (24-30 CPU/Cores)
- Physics-based Ionosphere-Plasmasphere Model (IPM)
- Incorporates Ionospheric Physics in the Data Assimilation
- Can Assimilate the 5 Data Sources shown in a Previous Slide
- Altitude, Latitude, Longitude Grids Set by User
- Ionospheric Specifications, Forecasts and Drivers
 - Electric Field
 - Neutral Wind
 - Neutral Composition

GAIM-FP Global Run

- 400 global TEC stations (IGS network) used in real-time at USU Space Weather Center
- Up to 10,000 measurements assimilated every 15- min
- 40-50 Ionosondes/Digisondes

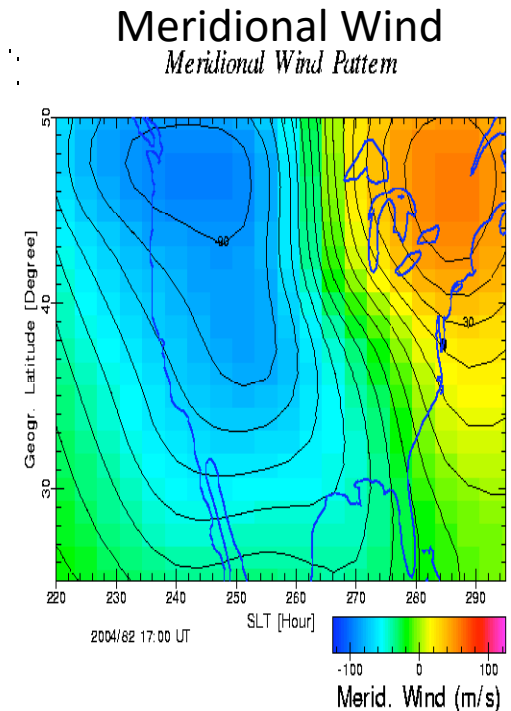
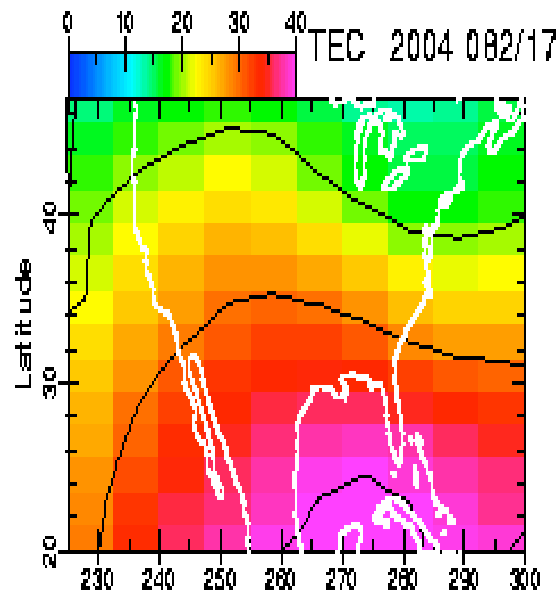
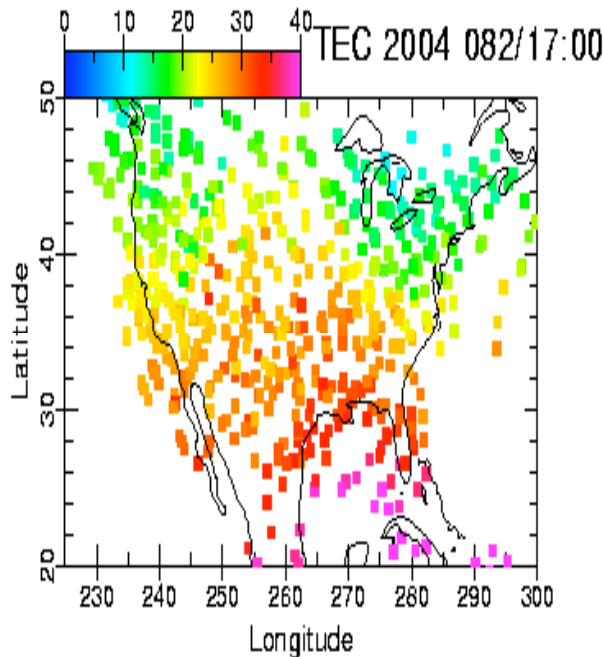
GAIM-FP GLOBAL TEC



"Bringing The Pieces Together"

Reconstructions With Self-Consistent Drivers

GAIM-FP → Regional Run



- Snapshots of TEC measurements (left)
- GAIM-FP reconstruction (middle)
- GAIM-FP neutral wind at 300 km (right)
- 17:00 UT, day 82, 2004

GAIM Data Assimilation Models have been or are currently running at:

- **AFWA (557 Weather Wing)**
- **Northrup Grumman**
- **Air Force Research Laboratory (AFRL)**
- **Naval Research Laboratory (NRL)**
- **USU Space Weather Center (SWC)**
- **Community Coordinated Modeling Center (CCMC)**

Multimodel Ensemble Prediction System (MEPS): Ensemble Modeling with Data Assimilation Models

Utah State University

R. W. Schunk, L. Scherliess, V. Eccles, L. C. Gardner, J. J. Sojka and L. Zhu

Jet Propulsion Laboratory

X. Pi, A. J. Mannucci, and A. Komjathy

University of Southern California

C. Wang and G. Rosen



MEPS Data Assimilation Models

GAIM-BL → Mid & Low Latitudes

GAIM-GM → Mid & Low Latitudes

GAIM-4DVAR → Mid & Low Latitudes, **with Drivers**

GAIM-FP → Mid & Low Latitudes, **with Drivers**

Mid-Low Electro-DA → Ionosphere **with Drivers**

IDED-DA → High Latitudes, **with Drivers**

GTM-DA → Global Thermosphere

TWAM → Thermospheric Wind Assimilation Model

- **Global, Regional & Nested GRID Capabilities**
- **Science, Specifications & Forecasts**



Why Ensemble Modeling



**National Hurricane Center
multi-model ensemble
forecast for hurricane Rita.**

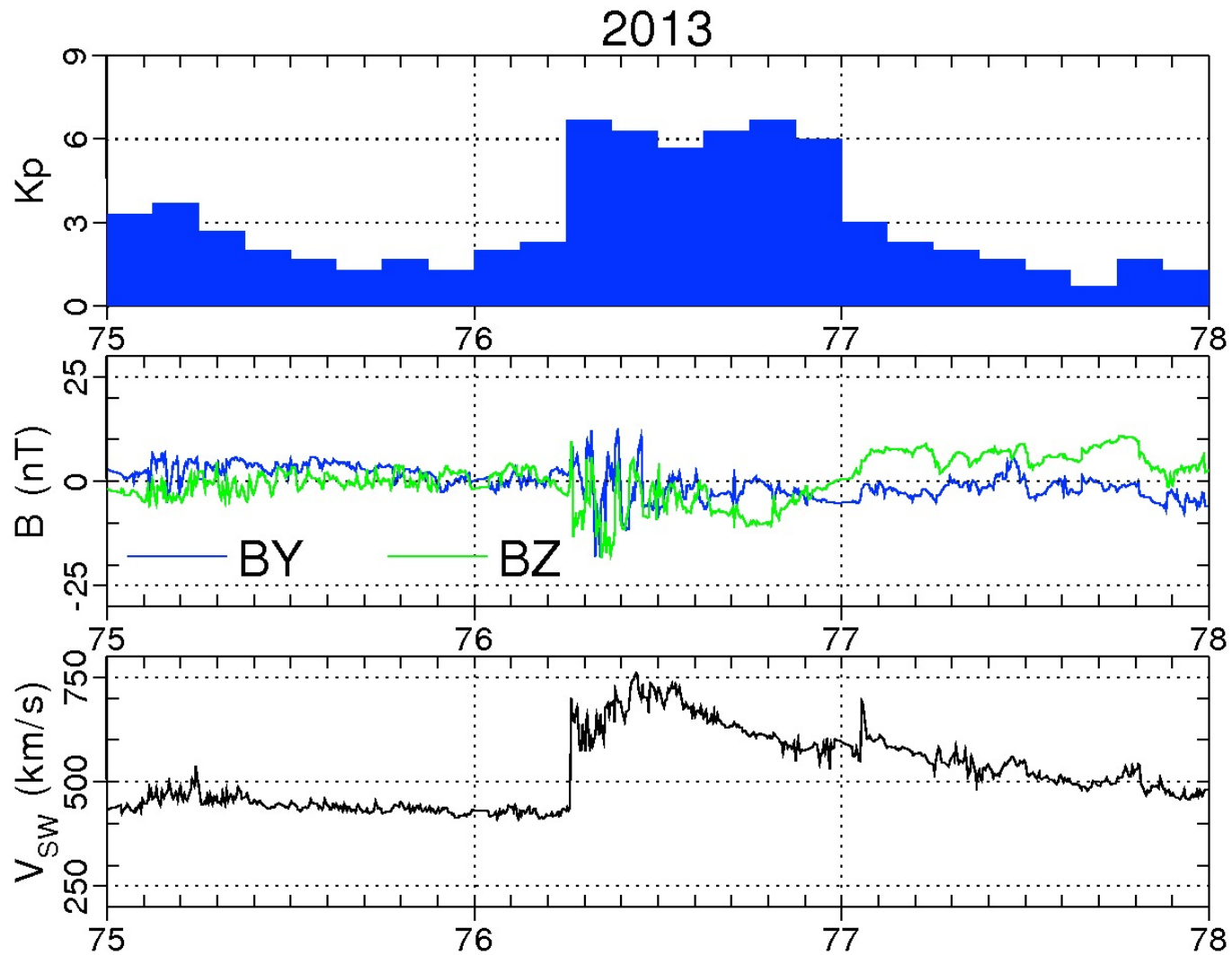


MEPS Initial Tasks

- **Select a Storm Period**
- **Select MEPS Models and Data**
- **Conduct MEPS Ensemble Model Runs**
- **Study the Effect of Different Data Types**
- **Compare MEPS Reconstructions**
- **Study the Usefulness of Ensemble Averaging**



Magnetic Storm Period



Selected MEPS Models and Data

| | | |
|------------|---|--|
| GAIM-BL | → | Mid & Low Latitudes |
| GAIM-GM | → | Mid & Low Latitudes |
| GAIM-4DVAR | → | Mid & Low Latitudes, with Drivers |
| GAIM-FP | → | Mid & Low Latitudes, with Drivers |

BL – Band Limited

GM – Gauss Markov

4DVAR – 4D Adjoint

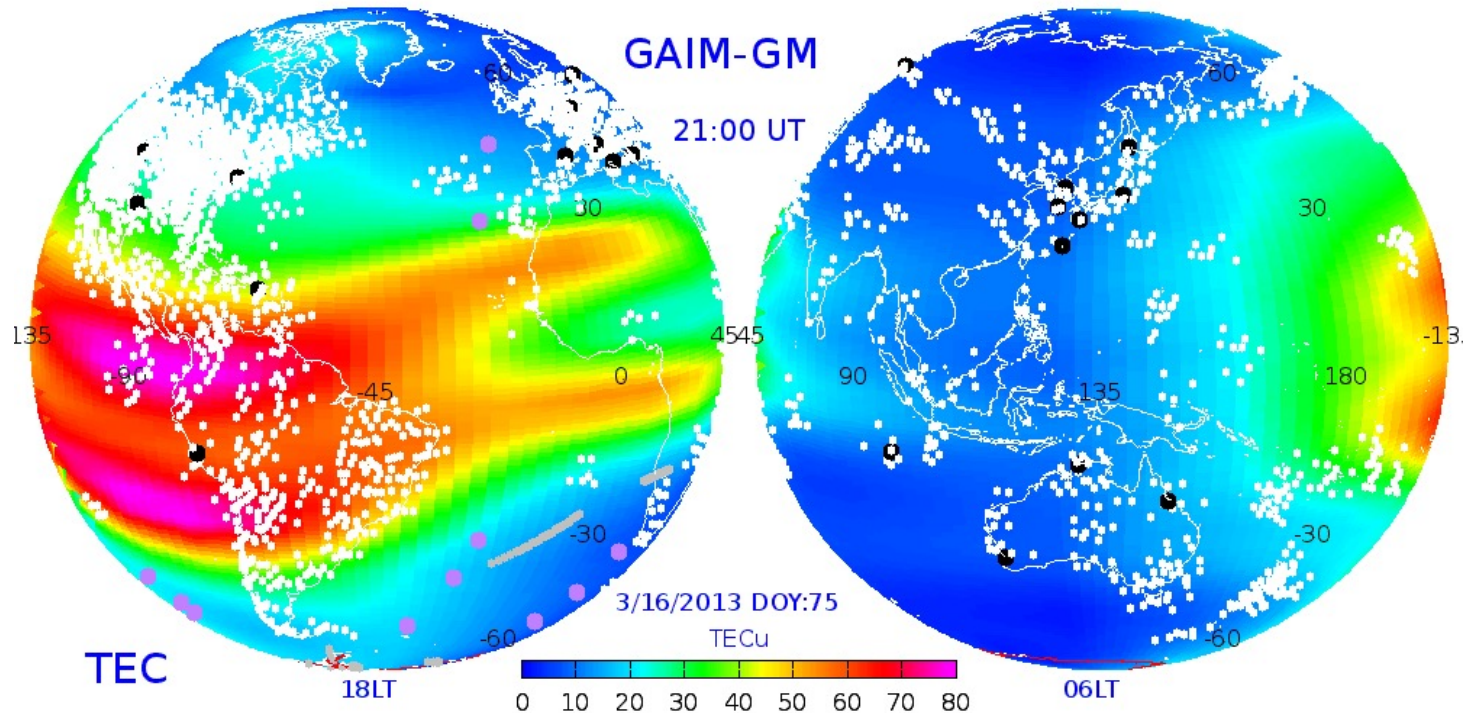
Method

FP – Full Physics

- Ground-Based GPS-TEC
- Satellite to Satellite Occultation
Yields N_e at 800 km
- Ionosonde-Digisonde N_e (sao)
- 911A, 1356A; limb, disk (UV)



Global Data Distribution



White Dots – Locations of 530 ground GPS receivers

Black Dots – Locations of 80 ionosondes/digisondes (sao)

Purple Dots – COSMIC derived electron densities at 800 km

Background – GAIM-GM reconstruction for the quiet day at 21 UT

MEPS – Effect of Different Data Types

- Run with TEC data from 530 ground GPS receivers
- Run with 530 ground GPS receivers & COSMIC occultation data
- Run with 530 ground GPS receivers, occultation data, & 80 digisondes (sao)

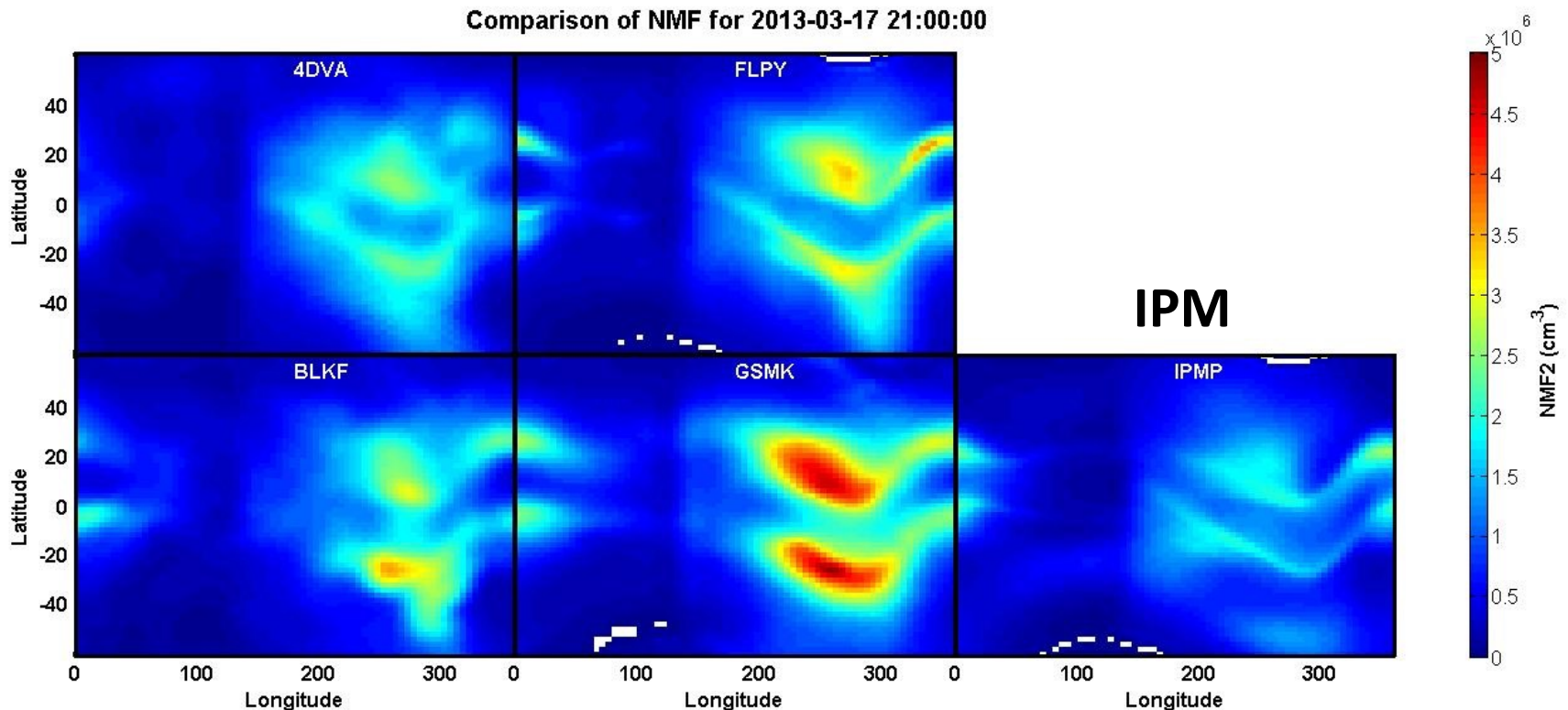
Goal is to see the differences in the model results and to see how the different models handle the same data type



Run the Four Data Assimilation Models with TEC data from 530 ground GPS receivers



NmF2 Comparison for the Storm Day

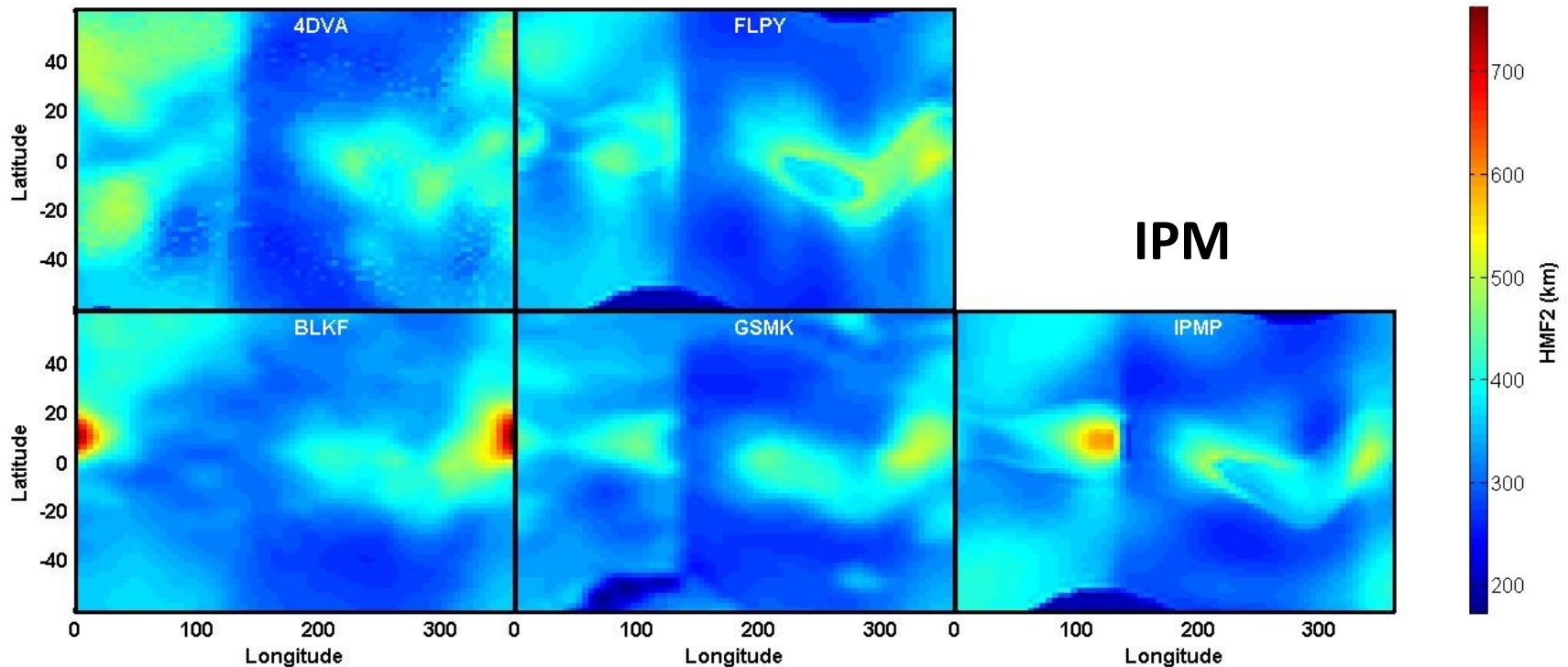


- Differences in magnitude of the equatorial anomaly.
- Some differences in longitude and width of equatorial anomaly
- Four models show enhanced NmF2 in the southern hemisphere beyond 30° latitude
- IPM is background physics-based model for GAIM-FP



HmF2 Comparison for the Storm Day

Comparison of HMF for 2013-03-17 21:00:00



Differences in

- the equatorial region near 0° and 120° longitude
- middle latitudes in the southern hemisphere
- IPM is background physics-based model for GAIM-FP



**Run the Four Data Assimilation Models with TEC data
from 530 ground GPS receivers and
Satellite to Satellite Occultation data (N_e at 800 km)**



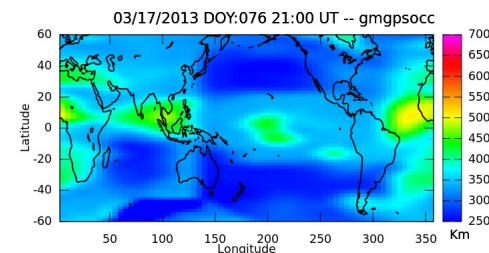
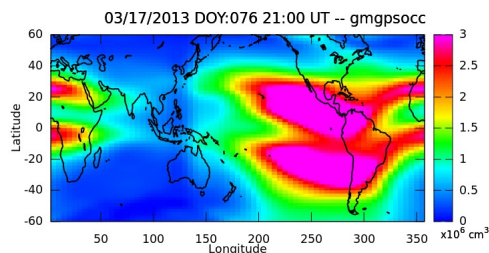
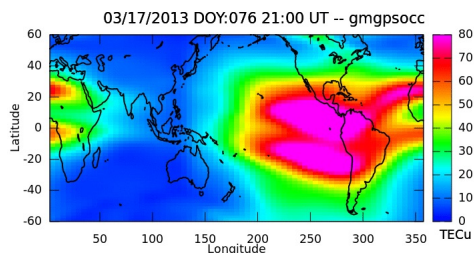
GAIM 2013 Day 76 21:00 UT – GPS + Ne_{COSMIC}(800 km)

TEC

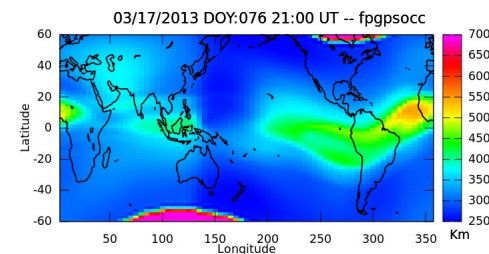
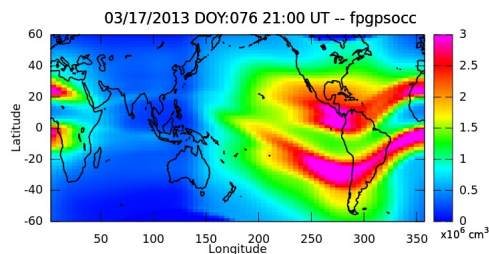
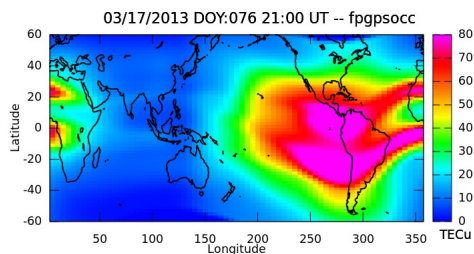
NmF2

hmF2

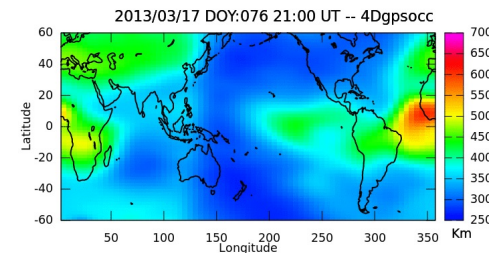
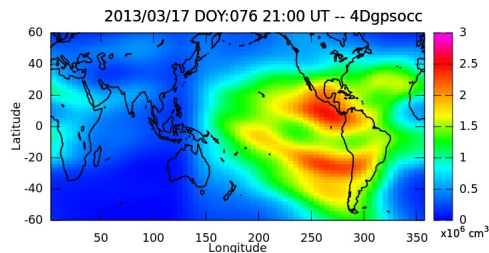
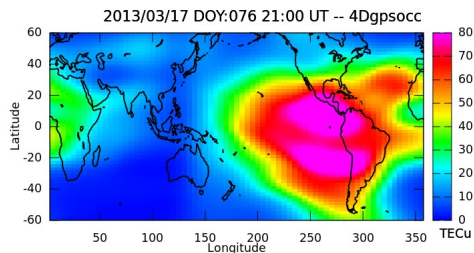
GAIM-GM



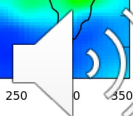
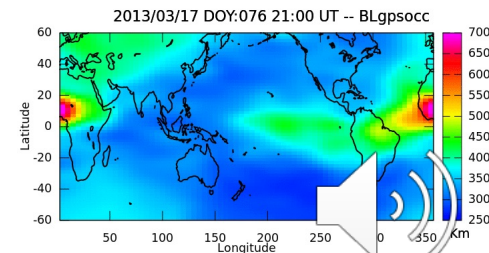
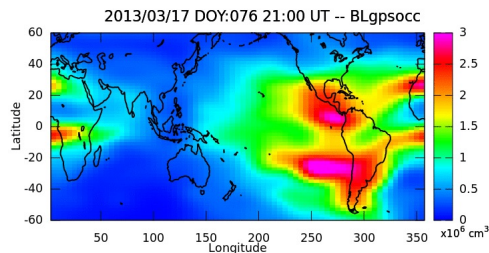
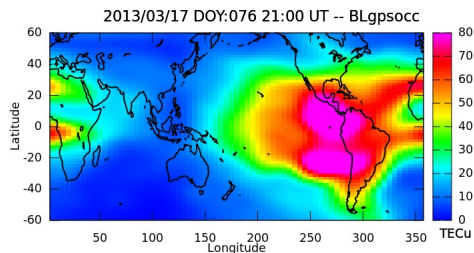
GAIM-FP



GAIM-4D



GAIM-BL



Run the Four Data Assimilation Models with TEC, Occultation, and Digisonde (sao) data



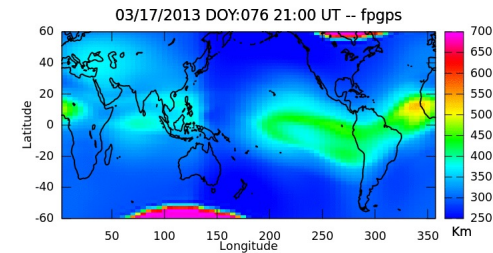
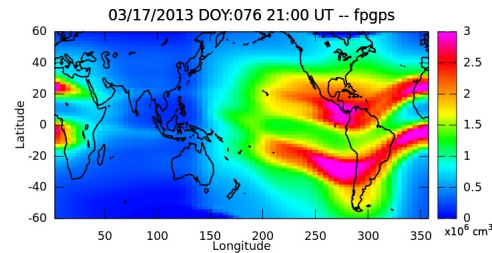
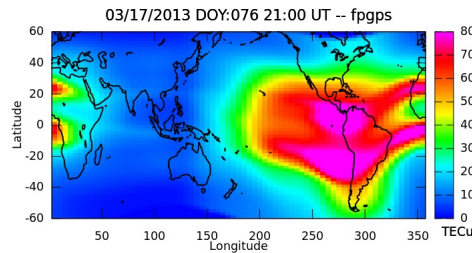
GAIM-FP 2013 Day 76 21:00 UT

TEC

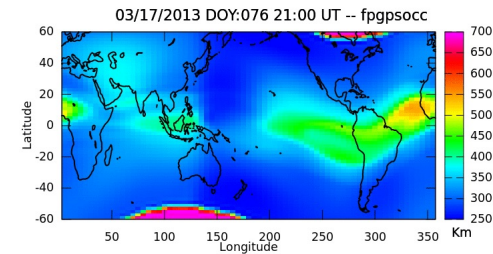
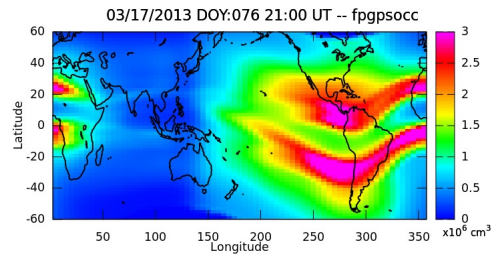
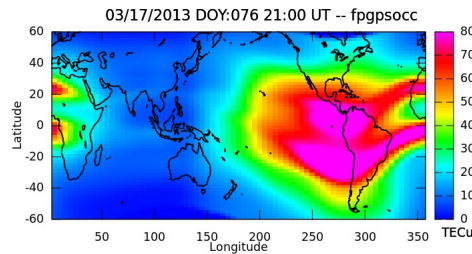
NmF2

hmF2

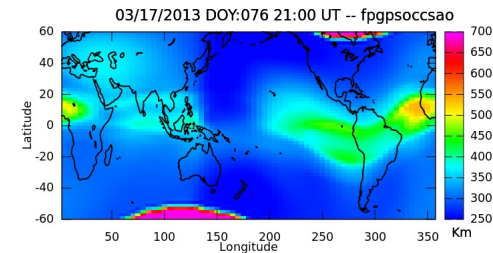
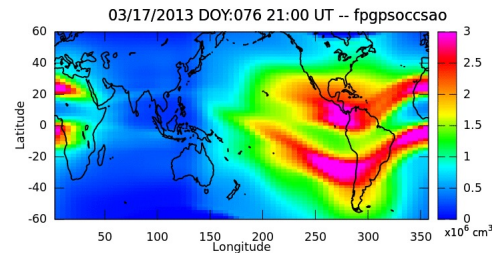
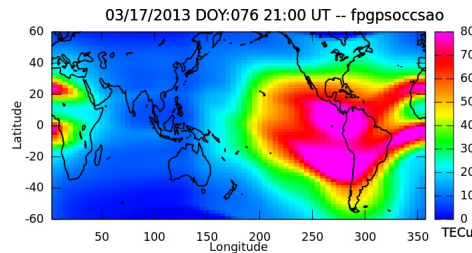
GPS



GPS +
Ne(800 km)



GPS +
Ne(800 km)
+ sa0



Storm Day



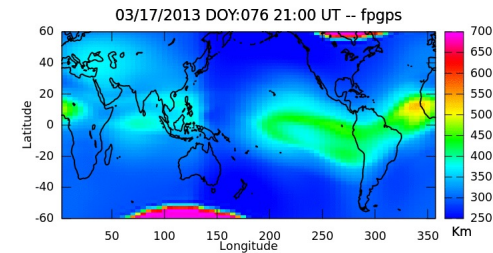
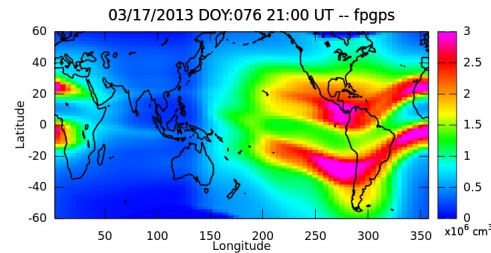
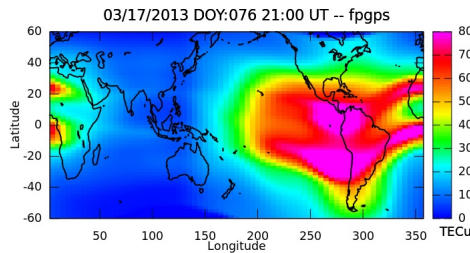
GAIM-FP 2013 Day 76 21:00 UT Diff

TEC

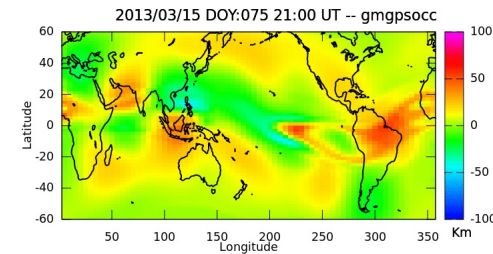
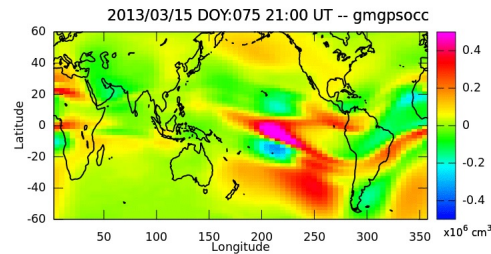
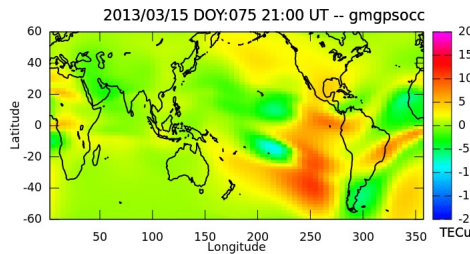
NmF2

hmF2

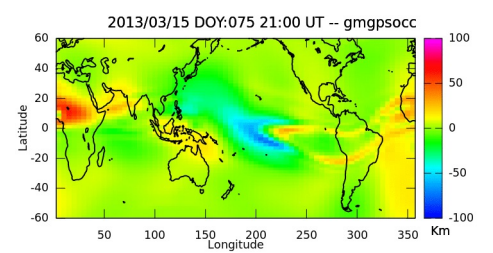
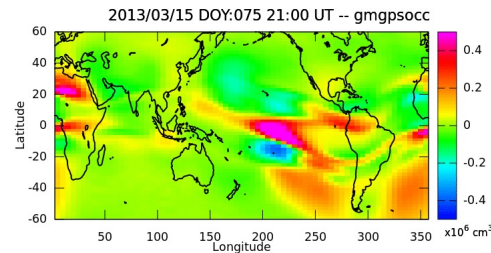
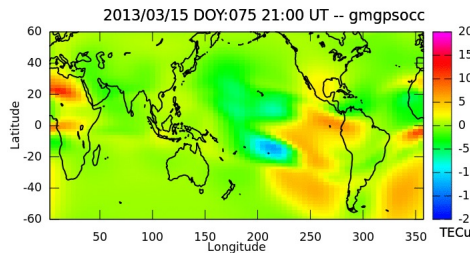
GPS



GPS +
Ne(800 km)



GPS +
Ne(800 km)
+ sa0



Storm Day



Models Reconstructions are Very Similar but there are Differences which are due to:

- **Different Background Physics-Based Models**
- **Different Assimilation Techniques**
- **Different Spatial and Temporal Resolutions**
- **Different Deduced Electrodynamics Drifts**
- **Different Deduced Neutral Winds and O/N₂ Ratios**

Goal is a Systematic Study to Elucidate Causes of Differences



Ensemble Model Averaging Example

March 12-19, 2013

- **5 Data Assimilation & 1 Physics Model**
- **Mid and Low Latitudes**
- **GPS and Occultation Data**
- **Solar Medium, Equinox, Storm**
- **Simple Average of Model Outputs**
 - Sum models, divide by number of models
- **Weighted Average of Model Outputs**
 - Sum models weighted by fit to GPS data, divide by number of models



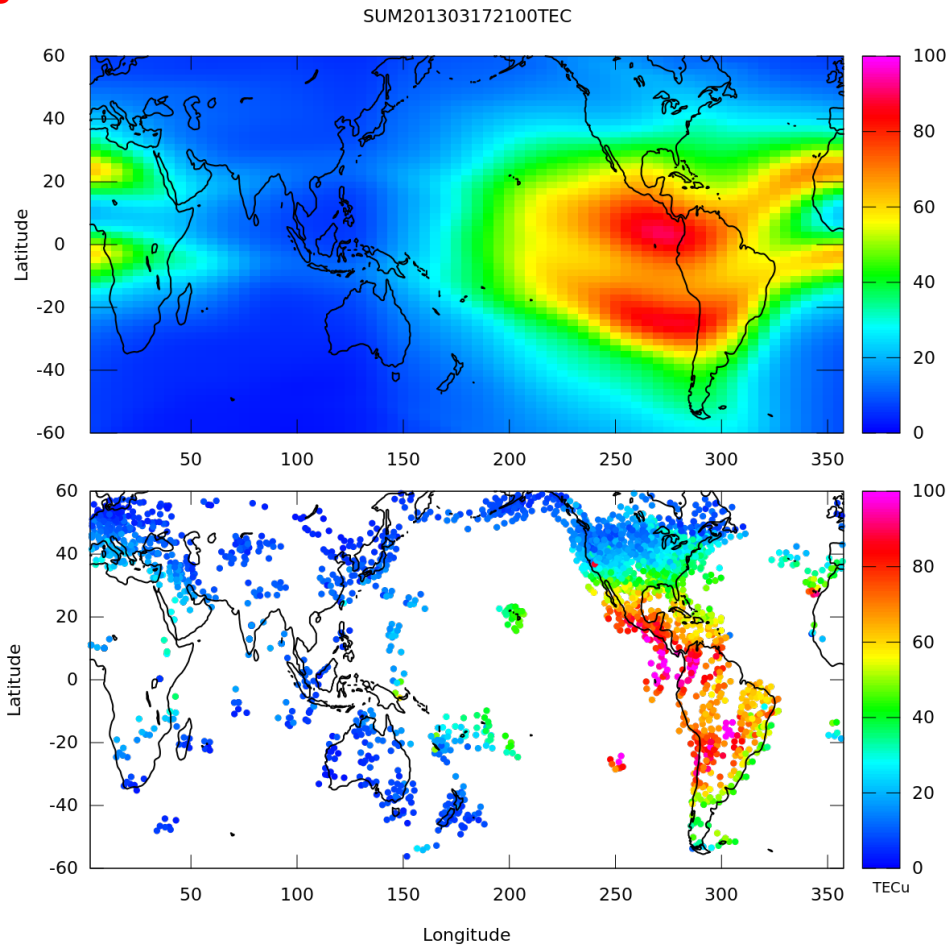
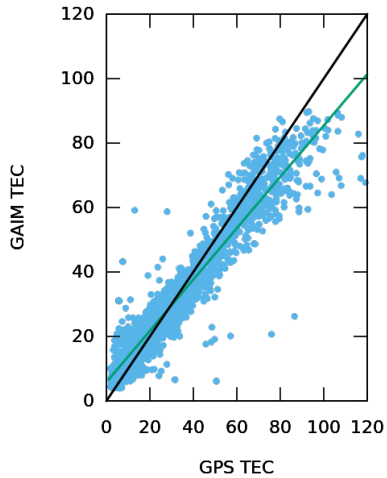
Ensemble Average (6 Models)

- GAIM-BL → Mid & Low Latitudes
- GAIM-GM → Mid & Low Latitudes
- GAIM-4DVAR → Mid & Low Latitudes, **with Drivers**
- GAIM-FP → Mid & Low Latitudes, **with Drivers**
- Mid-Low Electro-DA → Ionosphere **with Drivers**
- IFM Physics-Based Model → **No DA**



Ensemble Average - Simple

Ensemble Mean Vertical TEC Versus TEC Data



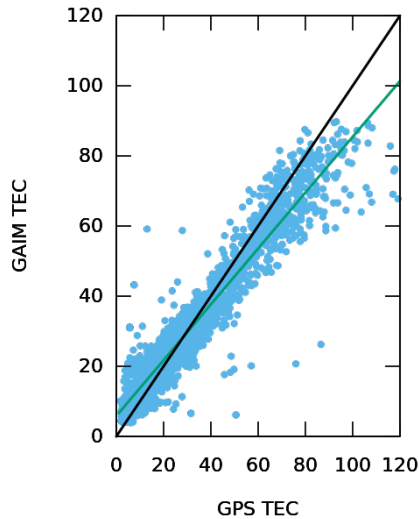
**Ensemble Mean
Vertical TEC
From GPS &
Occultation Run**

**Vertical TEC
Data**

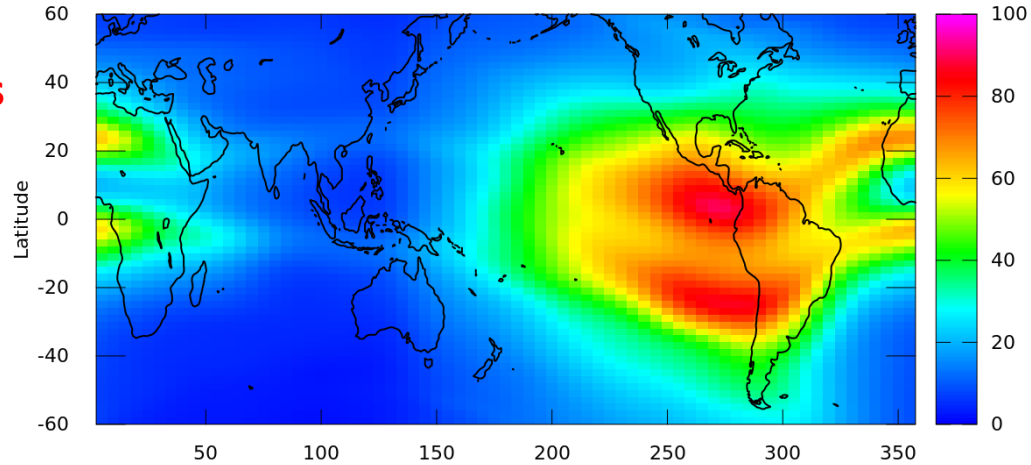


Ensemble Average - Simple

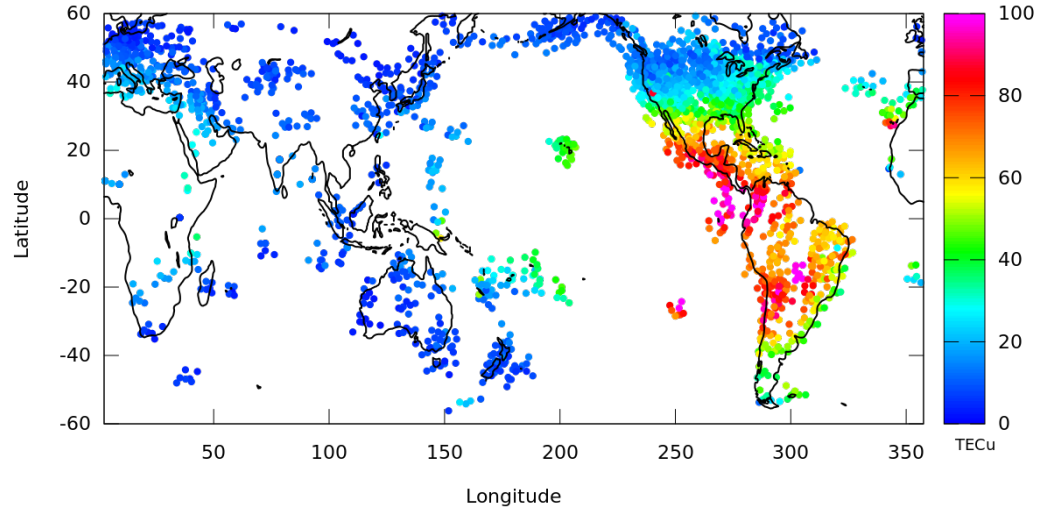
**Ensemble Mean
Vertical TEC Versus
TEC Data**



SUM201303172100TEC



**Ensemble Mean
Vertical TEC
From GPS &
Occultation Run**



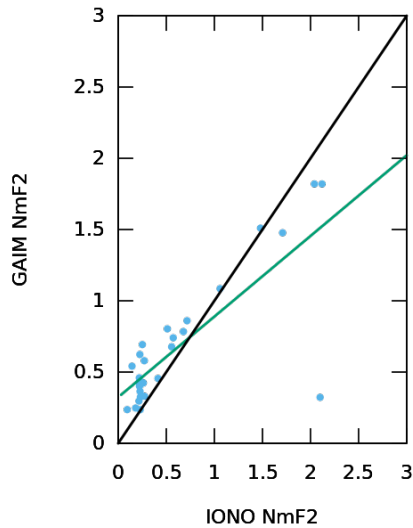
**Vertical TEC
Data**

The Ensemble Mean at all times throughout the 7-day period was better than the Individual Data Assimilation Models

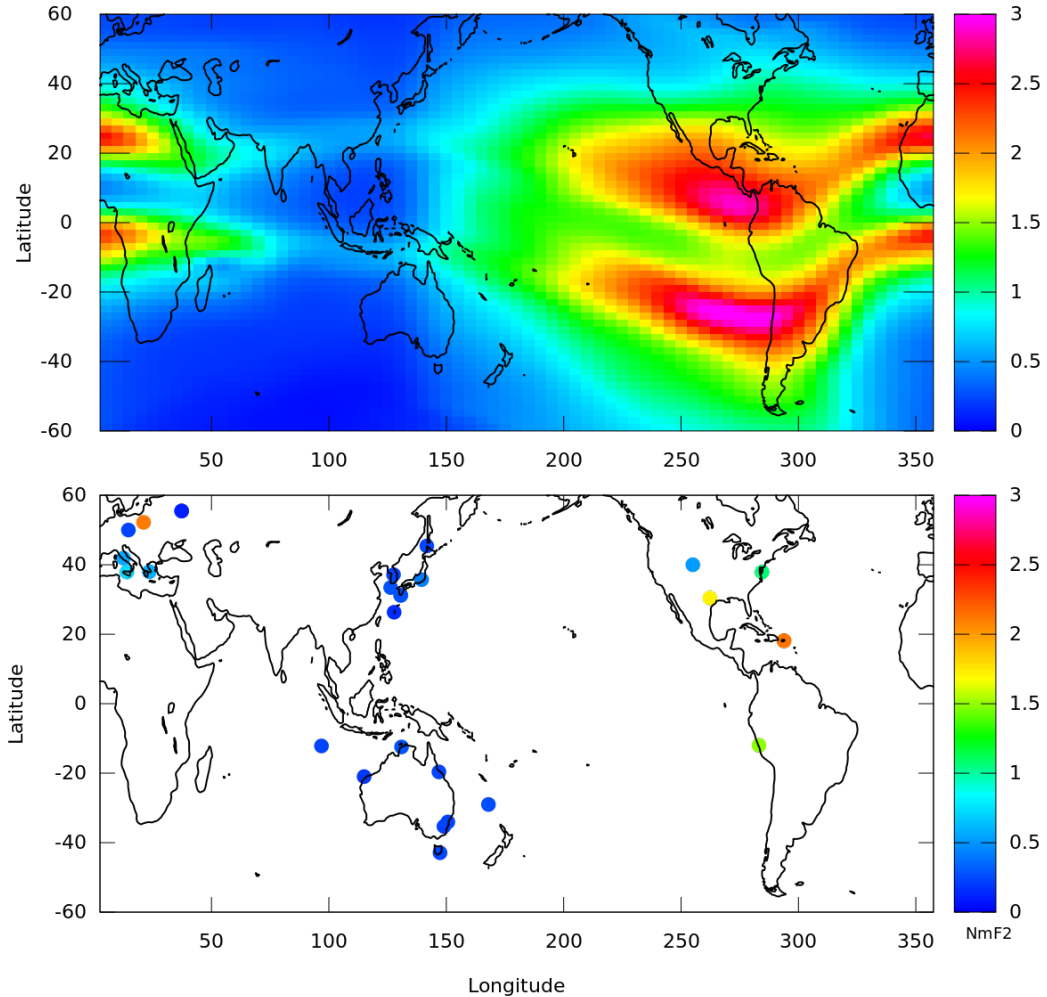


Ensemble Average - Simple

Ensemble Mean $N_m F_2$ Versus $N_m F_2$ Data

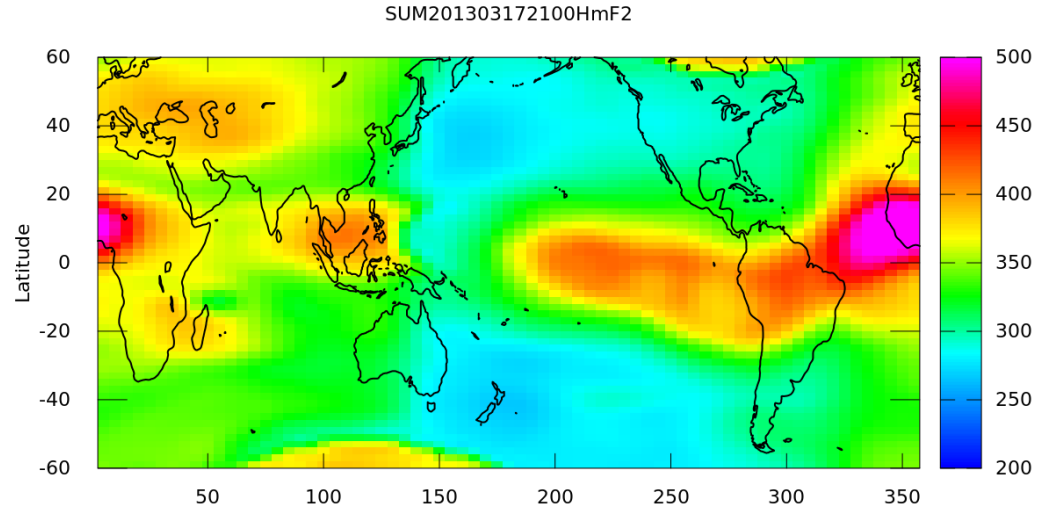
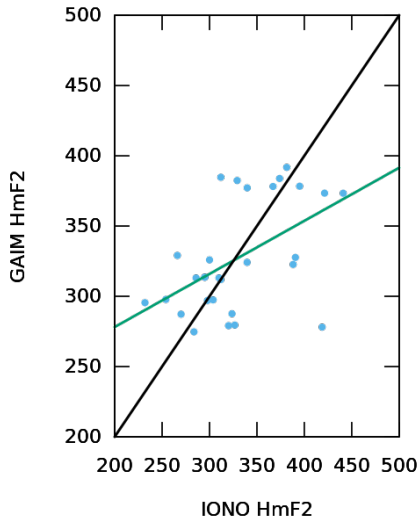


SUM201303172100NmF2

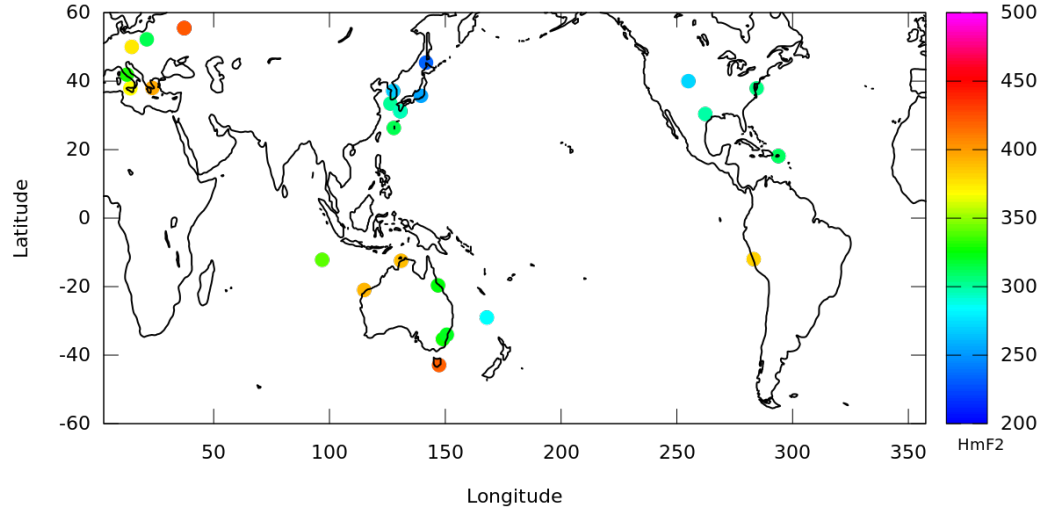


Ensemble Average - Simple

Ensemble Mean $h_m F_2$ Versus $h_m F_2$ Data



Ensemble Mean $h_m F_2$ From GPS & Occultation Run

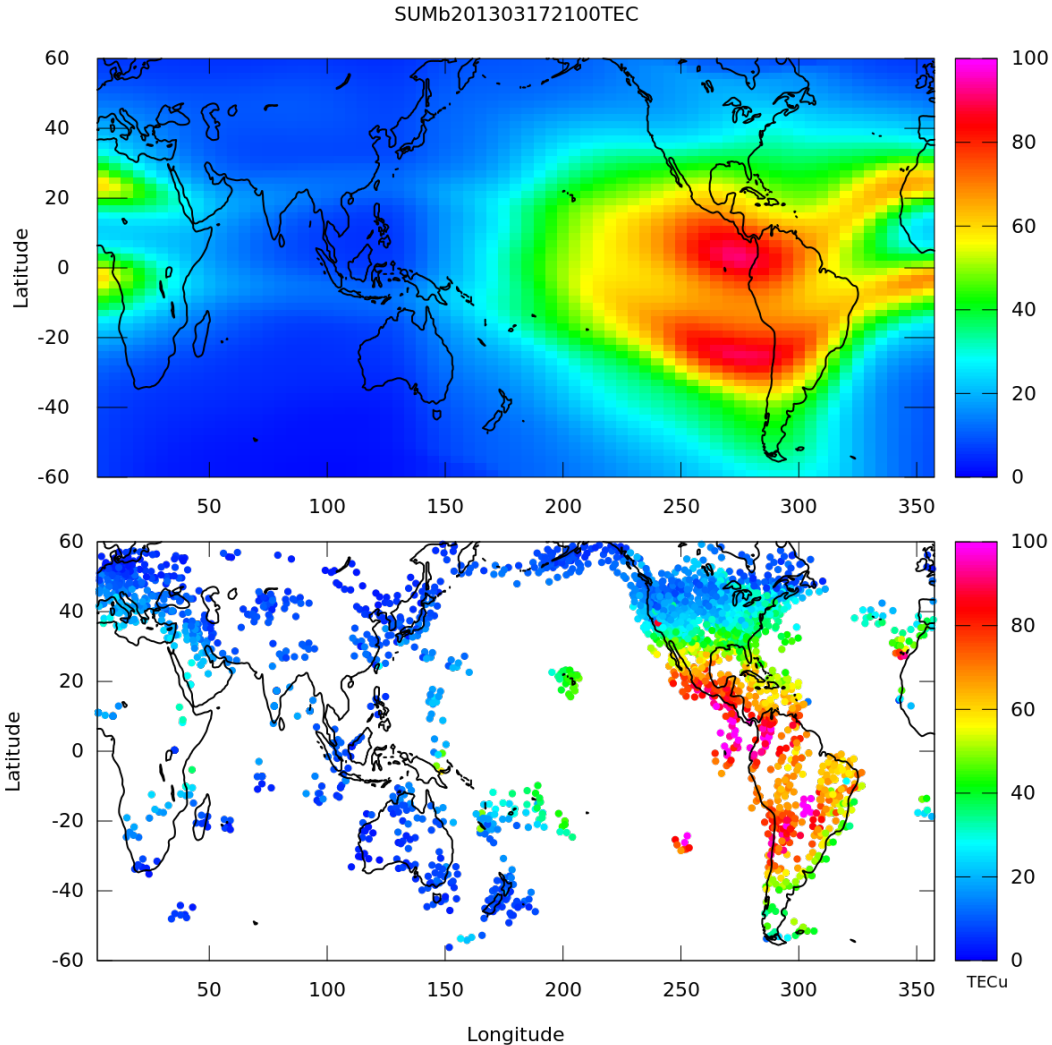
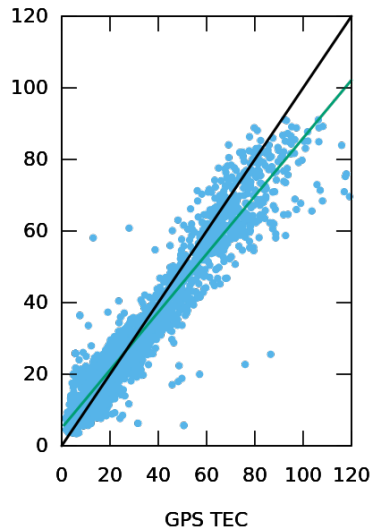


$h_m F_2$ from Ionosondes



Ensemble Average - Weighted

**Ensemble Mean
Vertical TEC
Versus
TEC Data**



**Ensemble Mean
Vertical TEC
From GPS &
Occultation Run**

**Vertical TEC
Data**

**The Weighted Average of the Ensemble of Models
Is slightly better than the Simple Average.**



Forecasting

- We provide a 24-hour Ionosphere Forecast with our GAIM Operational Models based on Persistence of the I-T Drivers
- In general, this Forecast is more reliable during slowly varying magnetic activity
- However, the Forecast is reasonable at mid-latitudes, somewhat reasonable at low latitudes, and unreliable at high latitudes.
- A reliable 24-hour Ionosphere/Thermosphere Forecast requires reasonable Forecasts of the I-T Drivers
 - Convection E-Fields, Auroral Precipitation, Field-Aligned and Horizontal Currents
 - Equatorial E-Fields
 - Upward Propagating Waves from the Lower Atmosphere
- A Significant Challenge for the Next Decade



Summary

- **MEPS → ensemble modeling with different data assimilation models**
- **Data assimilation on multiple spatial & temporal scales**
- **Wide range of ground and space data**
- **An important tool for studying basic physics**
- **Can combine different data sets into a coherent picture**
- **Fills in regions where there are no data**
- **New approach for specifying and forecasting space weather**





Selected Physics-Based I-T Model Publications

Time-Dependent Ionosphere Model (TDIM)

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Schunk, R. W., (1988), A mathematical model of middle and high-latitude ionosphere. *Pure and Applied Geophysics (PAGEOPH)*, 127, 255-303.

Sojka, J. J., (1989), Global scale, physical models of the F region ionosphere. *Reviews of Geophysics* 27, 371-403.

Ionosphere Forecast Model (IFM)

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Zhu L., et al., (2006), Validation study of the Ionosphere Forecast Model (IFM) using the TOPEX total electron content measurements, *Radio Sci.*, 41, RS5S11, doi:10.1029/2005RS003336.

Ionosphere-Plasmasphere Model (IPM)

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Schunk, R. W., et al. (2004), Global Assimilation of Ionospheric Measurements (GAIM). *Radio Science*, 39, RS1S02, doi:10.1029/2002RS002794.

IPM-Global

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Physics-Based I-T Model Publications Continued

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- Eccles J. V. (2004), Assimilation of global-scale and mesoscale electric fields from low-latitude satellites, *Radio Science*, 39, RS1S09, <https://doi.org/10.1029/2002RS002810>.
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USU GAIM and MEPS Publications Continued

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